

## Similarities of Tritonian guttae and Martian dark dune spots

H. I. Hargitai  
Eötvös Loránd University, Department of Physical Geography, 1117 Budapest Pázmány P st 1/a Hungary.  
hhargitai@gmail.com

### Abstract

Dark spots with bright halo near the edge of the Triton's photographed south polar cap and Martian Dark Dune Spots (DDS) sometimes with bright halo, found at the Martian circumpolar regions show striking morphological similarities which may imply similar formative processes.

### 1. Introduction

Triton's guttae (drop, lat.), (Fig 1.) originally named 'mushrooms' [1] and now named maculae in the IAU nomenclature, have been photographed by Voyager 2. They are found in groups between 5S and 45S, on or near the edge of the south polar cap [2], on bright smooth terrain (albedo: 0.85) [3]. They are irregular spots with an albedo lower than the surrounding terrain (albedo: 0.7), with brighter (albedo: 0.9) annular aureoles around them whose materials are visually similar to the opaque bright material on the south polar cap [2].

Martian Dark dune spots (Fig. 2.), also called 'Dalmatian spots' and 'fried eggs' [4] are located in the southern and northern circumpolar region, at latitudes higher than about 65° on seasonal-frost covered dark basaltic sand dunes. They are 10s of m across, commonly showing an internal structure with a darker (umbra-like) core and a lighter, ring shaped (penumbra-like) outer area, around which a bright halo (collar) is sometimes also seen [5]. They form groups within the dune covered terrain.

These two features display morphological similarities, although the ones on Triton are much larger. Any analysis based on morphological similarities should take the concept of equifinality (form-convergence) into consideration [6]. In this case there is only a snapshot available from Triton, but putatively corresponding Martian DDS's have been observed

through several seasons revealing their seasonal development sequence. The two features may have similar origin.

### 2. Formative processes

Guttae have been interpreted as residual deposits of a pre-existing, thicker and more extensive polar cap [2]. They are proposed to have been formed as the low albedo interior is sublimating N<sub>2</sub> at high rate, which rises local N<sub>2</sub> pressure that radiates outwards, where under the locally increased partial pressure, condensation temperature is higher and consequently N<sub>2</sub> re-condensates as frost (or sublimates at slower rate), resulting in the bright annulus [3]. They have been compared to outliers of the permanent Martian cap shaped by ablation [2] or CO<sub>2</sub> frost halos on the south polar residual cap, where frost accumulation occurs due to higher sublimation rates on sun-facing slopes [3].

In this paper I propose that guttae have similar origin to seasonally formed inliers on the Martian polar seasonal frost cover that form over dark dunes. DDS's have been proposed to form by defrosting, sublimation of CO<sub>2</sub> frost, exposing the underlying dark surface. Defrosting occurs as low albedo polar sand heats beneath an optically thin layer of frost, causing the frost to evaporate. This vapor is re-precipitated in the colder outer zones forming a bright halo. As the core evolves, the bright halo narrows and disappears, but the spot develops a dark interior and a grey exterior zone. [7].

On Triton only a dark interior and a bright outer zone have been observed, both with sharp boundaries, composed of N<sub>2</sub> frost. On Mars, DDS's have a prominent dark core with a temporary grey outer zone with a diffuse bright halo, composed of CO<sub>2</sub> frost. The Tritonian example may be a temporal state in seasonal guttae development.

### 3. Implications

If guttae have similar origin to that of DDS's, a seasonal change on their distribution and morphology should be supposed for the Tritonian features, similar to that shown by dark dunes spots on Mars.

If guttae and DDS's formed similarly, it has implications on the Dark Dune Spots Mars Surface Organism (DDSMSO) model [8]: it favors a model that can be applied to both Triton's and Mars' near polar cap environments.



Fig. 1. Maculae on Triton, PIA00056, Voyager 2

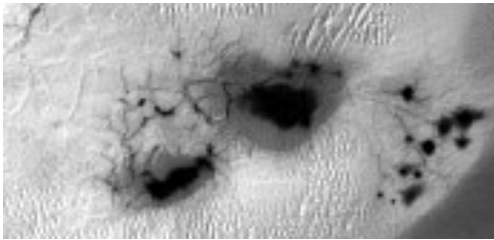


Fig. 2. DDS on Mars, Richardson Crater, HiRISE PSP\_003175\_1080

### References

- [1] Smith BA, Soderblom LA, Banfield D, Bartnet C et al.: Voyager 2 at Neptune: Imaging Science Results Science 246; 4936; 1422-1449, 1989.
- [2] Croft SK: Geomorphology of Triton's Polar Materials. LPSC XXIV 345-346, 1993.
- [3] Becerna P, Byrne S: Mars' South Polar Halos and Triton's Aureoles: Sublimation- Driven Models of Formation. New Horizons Workshop on Icy Surface Processes, 2011.

[4] Kieffer HH: Behavior of solid CO<sub>2</sub> on mars: still a zoo. Third Mars Polar Science Conference #8083, 2003.

[5] Kereszturi A, Vincendon M, Schmidt F: Water ice patches in Richardson crater. Planetary and Space Science 59, 26-42., 2010.

[6] von Bertalanffy, L.: General System Theory. George Braziller Inc., New York, 1968.

[7] Malin MC, Edgett KS: Frosting and Defrosting of Martian Polar Dunes. LPS XXXI #1056, 2000.

[8] Gánti T, Pócs T, Bérczi Sz, Horváth A, Kereszturi A, Sik A, Szathmáry E: Ideal microhabitats on Mars: the astrobiological potential of polar dunes. 40th Lunar and Planetary Science Conference #1618, 2009.