

Observation of Ammonia and Methanol in comet 67P with MIRO onboard Rosetta

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Abstract

The Microwave Instrument on the Rosetta Orbiter (MIRO) [5] makes submillimeter- and millimeter-wavelength observations of the nucleus and coma of the target of the Rosetta mission: comet 67P/Churyumov-Gerasimenko. MIRO is probing both the continuum emission of the nucleus at 0.5 and 1.6 mm wavelength and the gaseous emission via high-resolution spectroscopic measurements of 8 molecular lines in the submillimeter range ($\rm H_2O$, $\rm H_2^{17}O$, $\rm H_2^{18}O$, $\rm CO$, $\rm NH_3$, and three lines of $\rm CH_3OH$).

Since Rosetta arrived in the vicinity of comet 67P/Churyumov-Gerasimenko in July 2014, MIRO has been observing the coma almost continuously. Water emission at 556.9 GHz has been detected since early June 2014 [6] and the water outgassing of the comet and its spatial distribution has been regularly monitored and mapped [4, 6, 7].

MIRO is also observing the fundamental rotational transition J_K =(10 - 00) of ammonia at 572.5 GHz and three lines of methanol: J_K =(3-2 - 2-1E) at 568.6 GHz, (8+1 - 70E) at 553.1 GHz and (12-1 - 11-1E) at 579.2 GHz. The ammonia line and the lowest energy (3-2 - 2-1E) methanol line have been detected in absorption against the nucleus of the comet (Fig.1,2) and in emission in the coma since August 2014, when Rosetta was at distances between \approx 10 to \geq 100 km from the nucleus.

We will present the results of one year of monitoring of the ammonia and methanol emissions. The measured NH₃ and CH₃OH abundances relative to water in comet 67P/Churyumov-Gerasimenko will be compared to remote measurements made in several other comets [1, 2, 3].

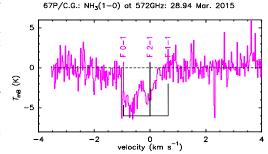


Figure 1: Ammonia detection in absorption against the nucleus of comet 67P on 28 March 2015 with MIRO.

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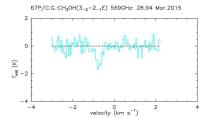


Figure 2: Methanol detection in absorption against the nucleus of comet 67P on 28 March 2015 with MIRO.

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