

## The formation of mini-Neptunes

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### Abstract

The results of the Kepler mission show that planets smaller than Neptune ("mini-Neptunes"), presumably composed of non-negligible amounts of hydrogen and helium, are frequent. The formation of this type of objects is in principle challenging for the classical core-accretion paradigm of giant planet formation. We show that mini-Neptunes are a common outcome when including the effect of envelope enrichment by icy planetesimals/pebbles in the formation models. We will discuss as well the planetary composition/structure, and the implications for the interpretation of exoplanet observations.