



Revised age for Iapetus

D. L. Matson (1), J. C. Castillo-Rogez (1), N. Turner (1), M. H. Lee (2), T. V. Johnson (1), and J. I. Lunine (3)
(1) JPL/Caltech, 4800 Oak Grove Drive, Pasadena, CA 91109(dmatson@jpl.nasa.gov), (2) Department of Earth Sciences, University of Hong Kong, Hong Kong, (3) Lunar and Planetary Lab, University of Arizona, Tucson, AZ, USA

Decay energies used for computing the heat produced by the decay of ^{26}Al in geophysical models have recently been reviewed [1,2]. The ^{26}Al decay scheme is complex and the values used range by a factor ~ 3 . This is a major issue because ^{26}Al is a primary heat source for planetary objects formed in the early solar system. They [1,2] recommend a heat production value of 3.12 MeV per decay and a half-life of 0.717 My. This heat value is a factor of ~ 2.4 higher than used for Iapetus by Castillo-Rogez et al. [3]. The new value does not change their conclusions but does shift their time of formation by about 1 My, (from $\sim 2.5 - 5.0$) to between ~ 3.4 and 5.5 My after the formation of the Ca-Al inclusions (CAIs). Since Saturn had to be in place before Iapetus could form, Saturn formed less than ~ 5.5 My after CAIs. This constraint is fully consistent with the growing number of observed protoplanetary disks that have cleared lanes, indicating giant planet formation in less than 8 My [4]. This work was performed at the Jet Propulsion Laboratory-California Institute of Technology, under contract with NASA. [1] Castillo-Rogez et al. (2009) *Icarus*, submitted. [2] Matson et al. (2009) *Lunar and Planet. Sci. Conf.* 40, submitted. [3] Castillo-Rogez J. et al. (2007) *Icarus*, 190, 179-202. [4] Cieza et al. (2007) *ApJ* 667, 308-328.