



## **A 3D Two-Temperature Solar Wind Model with Alfvén Wave Heating**

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The Space Weather Modeling Framework (SWMF), developed at the University of Michigan, is a high performance simulation tool to model a wide variety of physics domains ranging from solar corona and heliosphere to the magnetosphere, ionosphere and thermosphere of the Earth and can run these models in a coupled fashion.

We have developed a new physically consistent, three-dimensional solar wind model within the SWMF, that account for the different electron and ion temperatures. We use field-aligned heat conduction for both the electron and ions. The collisions between the electrons and ions are taken into account: Close to the sun the electrons and ions are tightly coupled, but completely decouple beyond approximately two solar radii. In our model, we use Alfvén waves to accelerate the wind. The ions are heated by the Kolmogorov dissipation of the Alfvén waves. The velocities at 1 AU obtained from the semi-empirical Wang-Sheeley-Argé model in combination with conservation of total energy along the field lines determines the Alfvén wave amplitude at the photospheric level. We compare this two-temperature model with a one-temperature model for Carrington rotation 2077.