



## **CME-CME interaction during the 2010 August 1 events**

M. Temmer (1), B. Vrsnak (2), T. Rollett (1), B. Bein (1), C.A. deKoning (3), Y. Liu (4), E. Bosman (5), J.A. Davies (6), C. Möstl (4,1), T. Zic (2), A.M. Veronig (1), V. Bothmer (5), R. Harrison (6), N. Nitta (7), M. Bisi (8,9), O. Flor (1), J. Eastwood (10), D. Odstrcil (11), and R. Forsyth (10)

(1) University of Graz, Institute of Physics, IGAM-Kanzelhöhe Observatory, Graz, Austria (manuela.temmer@uni-graz.at), (2) Hvar Observatory, Faculty of Geodesy, University of Zagreb, Kaciceva 26, HR-10000 Zagreb, Croatia, (3) NOAA Space Weather Prediction Center, Boulder Colorado, CO 80305, USA, (4) Space Sciences Laboratory, University of California, Berkeley, CA 94720, USA, (5) Institut für Astrophysik, Göttingen University, Friedrich-Hund Platz 1, D-37077 Göttingen, Germany, (6) RAL Space, Rutherford Appleton Laboratory, Harwell Oxford, Didcot OX11 0QX, UK, (7) Solar and Astrophysics Laboratory, Lockheed Martin Advanced Technology Centre, Palo Alto, CA 94304-1191, USA, (8) Institute of Mathematics and Physics, Aberystwyth University, Ceredigion SY23 3BZ, Wales, UK, (9) Center for Astrophysics and Space Sciences, University of California, San Diego, La Jolla, CA 92093-0424, USA, (10) The Blackett Laboratory, Imperial College London, London SW7 2AZ, UK and, (11) George Mason University/NASA Goddard Space Flight Center, Mail code 674, Greenbelt, MD 20771, USA

We study a CME-CME interaction that occurred during the 2010 August 1 events using STEREO/SECCHI data (COR and HI). The CMEs were Earth directed where clear signatures of magnetic flux ropes could be measured from in situ Wind data. To give evidence of the actual interaction we derive the direction of motion for both CMEs applying several independent methods. From this we obtain that both CMEs head into similar directions enabling us to actually observe the merging in the HI1 field-of-view (and rule out the possibility that this is just a line of sight effect). The full de-projected kinematics of the faster CME from Sun to Earth is derived when combining data points from remote observations with in situ parameters of the ICME measured at 1 AU. We study the evolution of the kinematical profile of the faster CME by applying a drag based model.