



MESSENGER Observations of Magnetotail Dynamics at Mercury

J. A. Slavin (1), S. M. Imber (2), G. A. DiBraccio (1,3), T. Sundberg (3), S. A. Boardsen (3,4), B. J. Anderson (5), H. Korth (5), D. N. Baker (6), R. L. McNutt (5), and S. C. Solomon (7)

(1) Department of Atmospheric, Oceanic and Space Sciences, University of Michigan, Ann Arbor, USA
(jaslavin@umich.edu), (2) Department of Physics and Astronomy, University of Leicester, Leicester, United Kingdom, (3) Heliophysics Science Division, NASA Goddard Space Flight Center, Greenbelt, USA, (4) Goddard Earth Science and Technology Center, University of Maryland, Baltimore County, USA, (5) The Johns Hopkins University Applied Physics Laboratory, Laurel, USA, (6) Laboratory for Atmospheric and Space Physics, University of Colorado, Boulder, USA, (7) Department of Terrestrial Magnetism, Carnegie Institution of Washington, Washington, D.C., USA

MESSENGER magnetic field observations taken in Mercury's magnetotail at distances of ~ 2 to $4 R_M$ (where R_M is Mercury's radius or 2440 km) frequently include tail loading and unloading events, traveling compression regions (TCRs) in the lobes, and plasmoid ejection in the plasma sheet. During these loading–unloading events, the magnitude of the lobe magnetic field can increase and then decrease by up to a factor of ~ 2 , as observed during MESSENGER's third flyby. The duration of these events is usually ~ 1 –3 min, but in some case they can last for as long as ~ 10 min. The transition from loading to unloading is often marked by one or more TCRs signaling the onset of magnetic reconnection in the cross-tail current sheet and the ejection of plasmoids. When MESSENGER crosses the plasma sheet at $\sim 3 R_M$ downstream of Mercury during the orbital phase of the mission, plasmoids are frequently present in the magnetic field measurements, especially when the interplanetary magnetic field has been southward. These plasmoids can occur singly, in groups of several events, or in quasi-periodic series with durations as long at ~ 10 min and with individual plasmoids separated by ~ 5 –10 s. These phenomena at Mercury are qualitatively similar to dynamic events observed during terrestrial magnetospheric substorms and continuous dissipation events, albeit on time scales of seconds to minutes compared with tens of minutes to hours at Earth.