



Calculation of CME kinematics and propagation directions by constraining STEREO HI-images with in situ signatures at 1 AU

T. Rollett (1), C. Möstl (1,2), M. Temmer (1), A.M. Veronig (1), C.J. Farrugia (3), and H.K. Biernat (4)

(1) University of Graz, Institute of Physics, Austria (tanja.rollett@uni-graz.at), (2) Space Sciences Laboratory, University of California, Berkeley, CA 94720-7450, (3) Space Science Center and Department of Physics, University of New Hampshire, Durham, New Hampshire, USA, (4) Space Research Institute, Austrian Academy of Sciences, Graz, Austria

We present a new approach to combine remote observations and in situ measurements by STEREO/HI and Wind, respectively, to derive the kinematics and propagation directions of interplanetary coronal mass ejections (ICMEs). We use two methods, Fixed-Phi and Harmonic Mean, to convert ICME elongations into distance. The ICME direction is constrained such that the ICME distance-time and speed-time profiles are most consistent with in situ measurements of the arrival time and speed at 1 AU. These methods are applied to two ICME events of 02–06 June 2008 and 13–18 February 2009. Due to the geometrical assumptions HM delivers the propagation direction further away from the observing spacecraft with a mean difference of 25° . This work has received funding from the European Commission FP7 Project COMESEP (263252).