



Venus Express/VMC observations of the Venus O₂ visible nightglow

Antonio García Muñoz (1,2), Ricardo Hueso (2), Agustin Sanchez-Lavega (2), Wojciech Markiewicz (3), Andrea Opitz (1), Olivier Witasse (1), and Dmitrij Titov (1)

(1) ESA-ESTEC, SRE-SM, Netherlands (tonhingm@gmail.com), (2) Grupo de Ciencias Planetarias, UPV/EHU, Bilbao, Spain, (3) Max Planck Institute for Solar System Research, Katlenburg-Lindau, Germany

We are analyzing the images of the Venus night side obtained with the Venus Monitoring Camera (VMC) aboard Venus Express at visible wavelengths (passband of 502-568 nm at 1/4 maximum transmission). The images show a faint but distinct emitting layer at about 100 km altitude attributed to the O₂ visible nightglow discovered by the Venera 9/10 missions [1]. The visible filter is most sensitive to the $v''=9, 10$ bands of the $c(0)-X(v'')$ progression, that occur at 513 and 551 nm, respectively. The VMC images allow us to investigate day-to-day variations in the nightglow intensity, that typically ranges from 200 to 400 kiloRayleighs in limb viewing, over the Venus disk, thus expanding on past studies from either space-borne or ground-based telescopes. In the presentation, we will discuss the status of our analysis of nearly five years of O₂ visible nightglow data with VMC. Ref.: [1] Krasnospolsky et al. (1977), Cosmic Res., 14, 687.