Geophysical Research Abstracts Vol. 15, EGU2013-5121-1, 2013 EGU General Assembly 2013 © Author(s) 2013. CC Attribution 3.0 License.



Periodic modulation of ion velocities within the magnetodisk of Saturn

Zoltan Nemeth (1), Karoly Szego (1), Lajos Foldy (1), Margaret Kivelson (2,3), Xianzhe Jia (2), Katherine Ramer (3), Stanley W. H. Cowley (4), and Gabrielle Provan (4)

(1) MTA Wigner Research Center RMI, Budapest, Hungary (nemeth@rmki.kfki.hu), (2) Atmospheric, Oceanic, and Space Sciences, University of Michigan, Ann Arbor, MI, USA, (3) Earth and Space Sciences, University of California, Los Angeles, Los Angeles, CA, USA, (4) Department of Physics and Astronomy, University of Leicester, Leicester, UK

The results of recent numerical simulations [Jia and Kivelson, 2012] show that in the magnetodisk of Saturn fluctuating magnetic field perturbations are accompanied by other oscillatory phenomena. They investigated the magnetotail response for a dual periodicity driver in the case when the solar wind flow was perpendicular to the rotation axis. As demonstrated in Fig. 7 of their paper the components of the flow velocity extracted at different radial distances plotted versus time exhibit periodic modulation.

Andrews et al. [2012] and Provan et al. [2012] investigated magnetic field modulations in Saturn's magnetosphere. They have shown that in the high latitude regions single period modulations can be observed, but near the current sheet dual periodicities are characteristic.

In this study we investigate periodicities in the azimuthal flow velocities using the numerical ion moments derived from the measurements of Cassini Plasma Spectrometer. Ramer et al. [2012] investigated these periodicities in the inner magnetosphere near the equatorial plane. We extend our study to include higher latitude passes in the outer magnetosphere to the orbit of Titan and beyond. To have a close match with the model assumptions of Jia and Kivelson, we investigated the behaviour of the ion velocities in the time range DOY 092–285, 2009, around Saturnian equinox, along different passes all containing a Titan flyby, and crossing the magnetodisk at different angles.

We have found that the azimuthal velocities show oscillatory behavior. The amplitude of the oscillation is comparable with the corrotation speed. We have also compared the structure of the velocity variations to the variations of ion densities along the different passes, and compared those to the model as well. The results of the current study is reported in this presentation.

References

Andrews, D. J., S. W. H. Cowley, M. K. Dougherty, L. Lamy, G. Provan, and D. J. Southwood (2012), Planetary period oscillations in Saturn's magnetosphere: Evolution of magnetic oscillation properties from southern summer to postequinox, J. Geophys. Res., 117, A04224, doi:10.1029/2011JA017444.

Jia and Kivelson (2012), Driving Saturn's magnetospheric periodicities from the upper atmosphere/ionosphere: Magnetotail response to dual sources, J. Geophys. Res., 117, A11219, doi:10.1029/2012JA018183.

Ramer, K. M., M. G. Kivelson1, K. K. Khurana, N. Sergis, R. J. Walke1, X. Jia (2012), Forces and Phases: An Investigation of Azimuthal Plasma and Field Periodicities in Saturn's Inner Magnetosphere, AGU Fall Meeting 2012, SM51B-2303

Provan, G., D. J. Andrews, C. S. Arridge, A. J. Coates, S. W. H. Cowley, G. Cox, M. K. Dougherty, and C. M. Jackman (2012), Dual periodicities in planetary-period magnetic field oscillations in Saturn's tail, J. Geophys. Res., 117, A01209, doi:10.1029/2011JA017104.