



***k*-filtering applied to *Cluster* density measurements in the Solar Wind: Early findings**

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Studies of solar wind turbulence indicate that a large proportion of the energy is Alfvénic (incompressible) at inertial scales. The properties of the turbulence found in the dissipation range are still under debate ~ while it is widely believed that kinetic Alfvén waves form the dominant component, the constituents of the remaining compressible turbulence are disputed. Using *k*-filtering, the power can be measured without assuming the validity of Taylor's hypothesis, and its distribution in (ω, \mathbf{k}) -space can be determined to assist the identification of weak turbulence components. This technique is applied to *Cluster* electron density measurements and compared to the power in $|\mathbf{B}(t)|$. As the direct electron density measurements from the WHISPER instrument have a low cadency of only 2.2s, proxy data derived from the spacecraft potential, measured every 0.2s by the EFW instrument, are used to extend this study to ion scales.