



Assessment of the Impact of Middle-Atmosphere Solar Tides on Gravity Waves in a WKB Gravity-Wave Model Based on Wave-Action Phase-Space Density

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abstract

Gravity waves (*GWs*) and solar tides (*STs*) are main constituents of the dynamical coupling between troposphere and mesosphere-lower-thermosphere (*MLT*). Via momentum deposition, *GWs* control to a large extent the mesospheric mean circulation. *STs* are large scale waves, mostly due to tropospheric and stratospheric diurnal heating processes, that modulate all dynamical fields in the mesosphere. *GWs* and *STs* also interact strongly with each other.

Conventional *GW* parameterizations used to describe this interaction (e.g. [1]) neglect the time-dependence and horizontal gradients of the background flow, with fatal effects (e.g. [2]). We study here the propagation of *GWs* in a time-dependent middle-atmosphere background flow, using a new (caustics free) *WKB GW* model (ray tracer). The background flow is composed by a climatological mean and tidal fields extracted from a general circulation model (*HAMMONIA*, see [3]). In order to avoid caustics, inevitable in classic ray-tracer implementations, we implemented a new wave-action phase-space density conservation scheme [4, 5]. The scheme attaches to each ray a finite volume in the location & wavenumber phase-space. The location-wavenumber volume is conserved during the propagation, responding in shape to the local stretching and squeezing in wave-number space. From the propagation of *GWs* we evaluate the deposition of momentum and buoyancy. Rayleigh-friction and temperature-relaxation coefficients are also evaluated.

In this extension of the study by [2] it is shown, with an amplitude scheme more stable against numerical instabilities, due to the avoidance of caustics, that *STs* (and so the time dependence of the background flow) modulate the propagation of *GWs*. Via Rayleigh-friction and temperature-relaxation coefficients, we also quantify how the pseudo-momentum-, momentum-, and enthalpy-deposition of *GWs* can influence the amplitude and the phase structure of *STs*. Finally, we compare momentum and buoyancy fluxes from the propagation of *GWs* with results from a simple scale analysis of the problem. These explain the amplitudes obtained by the scheme quite well.

Key words: Middle-Atmosphere dynamics, Solar Tides, Gravity Waves, WKB model

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