



## Phase relationships between orbital forcing and the composition of air trapped in Antarctic ice cores

Lucie Bazin (1), Amaelle Landais (1), Valérie Masson-Delmotte (1), Catherine Ritz (2), Ghislain Picard (2), Emilie Capron (3), Jean Jouzel (1), Marie Dumont (4), Markus Leuenberger (5), and Frédéric Prié (1)

(1) Laboratoire des Sciences du Climat et de l'Environnement, UMR8212, CEA-CNRS-UVSQ, Orme des Merisiers, Gif sur Yvette, France, (2) Laboratoire de Glaciologie et Géophysique de l'Environnement, UMR 5183, Univ. Grenoble Alpes-CNRS, Grenoble, France, (3) British Antarctic Survey, NERC, Cambridge, UK, (4) Météo-France-CNRS, CNRM-GAME UMR 3589, CEN, Grenoble, France, (5) Climate and Environmental Physics, Physics Institute and Oeschger Center for Climate Change Research, University of Bern, Bern, Switzerland

Orbital tuning is central for ice core chronologies beyond annual layer counting, available back to 60 ka (i.e. thousands of years before 1950) for Greenland ice cores. While several complementary orbital tuning tools have recently been developed using  $\delta^{18}\text{O}_{atm}$ ,  $\delta\text{O}_2/\text{N}_2$  and air content with different orbital targets, quantifying their uncertainties remains a challenge. Indeed, the exact processes linking variations of these parameters, measured in the air trapped in ice, to their orbital targets are not yet fully understood. Here, we provide new series of  $\delta\text{O}_2/\text{N}_2$  and  $\delta^{18}\text{O}_{atm}$  data encompassing Marine Isotopic Stage (MIS) 5 (between 100-160 ka) and the oldest part (380-800 ka) of the East Antarctic EPICA Dome C (EDC) ice core. For the first time, the measurements over MIS 5 allow an inter-comparison of  $\delta\text{O}_2/\text{N}_2$  and  $\delta^{18}\text{O}_{atm}$  records from three East Antarctic ice core sites (EDC, Vostok and Dome F). This comparison highlights a site-specific relationship between  $\delta\text{O}_2/\text{N}_2$  and its local summer solstice insolation that increases the uncertainty associated with the use of  $\delta\text{O}_2/\text{N}_2$  as a tool for orbital tuning. Combining records of  $\delta^{18}\text{O}_{atm}$  and  $\delta\text{O}_2/\text{N}_2$  from Vostok and EDC, we evidence a loss of orbital signature for these two parameters during periods of minimum eccentricity ( $\sim 400$  ka,  $\sim 720$ -800 ka). Our dataset reveals a time-varying lag between  $\delta\text{O}_2/\text{N}_2$  and  $\delta^{18}\text{O}_{atm}$  over the last 800 ka that we interpret as variations of the lag between  $\delta^{18}\text{O}_{atm}$  and precession. Large lags of  $\sim 5$  ka are identified during Terminations I and II, associated with strong Heinrich events. On the opposite, minimal lags ( $\sim 1$ -2 ka) are identified during four periods characterized by high eccentricity, intermediate ice volume and no Heinrich events (MIS 6-7, the end of MIS 9, MIS 15 and MIS 17). We therefore suggest that the occurrence of Heinrich events influences the response of  $\delta^{18}\text{O}_{atm}$  to precession.