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Solar wind driving of asymmetries in the magnetosheath - magnetosphere system

Andrew Dimmock (1), Tuija Pulkkinen (1), Adnane Osmane (1), and Katariina Nykyri (2) (1) School of Electrical Engineering, Aalto University, Espoo, Finland, (2) Embry-Riddle Aeronautical University, Daytona Beach, Florida, USA

Over the decades of in-situ measurements of the terrestrial magnetosphere it has been suggested and experimentally shown that various parameter dawn-dusk asymmetries arise. What is also apparent is that such asymmetries are delicately coupled to the properties of the solar wind. The IMF configuration has a considerable impact since its orientation dictates the shock geometry, thus driving different dawn-dusk plasma properties downstream. Magnetosheath asymmetries are notably important since the magnetosheath effectively modifies and reconfigures plasma before it enters the inner magnetosphere and therefore may play a role in driving asymmetries in the inner magnetosphere. We apply our existing statistical mapping tool which uses over 7 years of THEMIS and OMNI data to create statistical maps of plasma properties in the global magnetospheric system. We look at asymmetries of both steady state properties (e.g. B, V, n), and also transient/kinetic features such as mirror mode activity. We focus specifically on 1. solar wind dependence and 2. the co-dependence between the magnetosheath and magnetospheric regions.