

Sulfuric acid vapor in the atmosphere of Venus as observed by the Venus Express Radio Science experiment VeRa

Janusz Oschlisniok (1), Martin Pätzold (1), Bernd Häusler (2), Silvia Tellmann (1), Mike Bird (1,3), Thomas Andert (2), and Stefan Remus (4)

(1) Rheinisches Institut für Umweltforschung, Planetenforschung, Cologne, Germany (joschlis@uni-koeln.de), (2) Institut für Raumfahrttechnik, Universität der Bundeswehr München, Neubiberg, (3) Argelander – Institut für Astronomie, Bonn, (4) European Space Astronomy Centre (ESAC), Villanueva, Spanien

The cloud deck within Venus' atmosphere, which covers the entire planet between approx. 50 and 70 km altitude, consists mostly of liquid and gaseous sulfuric acid. The gaseous part increases strongly just below the main clouds and builds an approx. 15 km thick haze layer of H2SO4. This region is responsible for a strong absorption of radio waves as seen in VeRa radio science observations. The absorption of the radio signals during occultations is used to derive the abundance of gaseous sulfuric acid. VeRa probes the atmosphere of Venus since 2006 with radio signals at 13 cm (s-band) and 3.6 cm (x-band) wavelengths. The collection of nine years of radio science data provides a picture of the global distribution of the sulfuric acid vapor distribution within Venus' atmosphere. We present H2SO4 profiles retrieved with VeRa and compare those with H2SO4 profiles observed by previous missions.