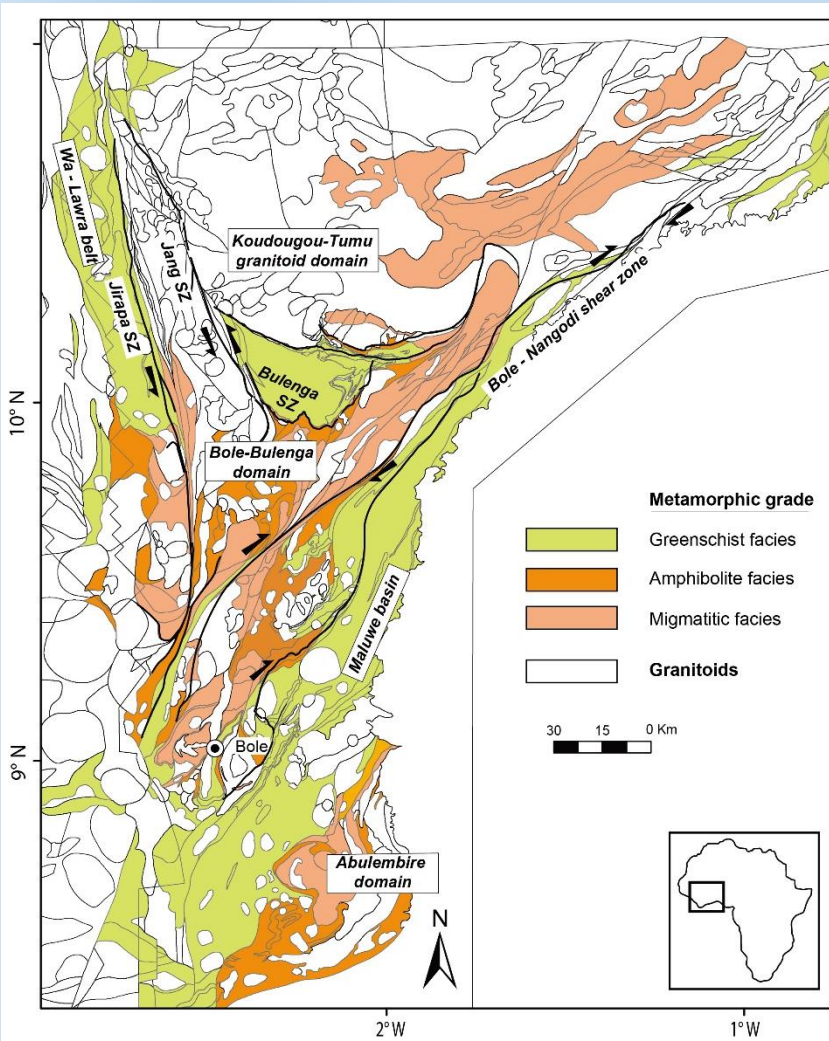


3-D numerical modelling of the influence of pre-existing faults and boundary conditions on the distribution of deformation: example of North-Western Ghana

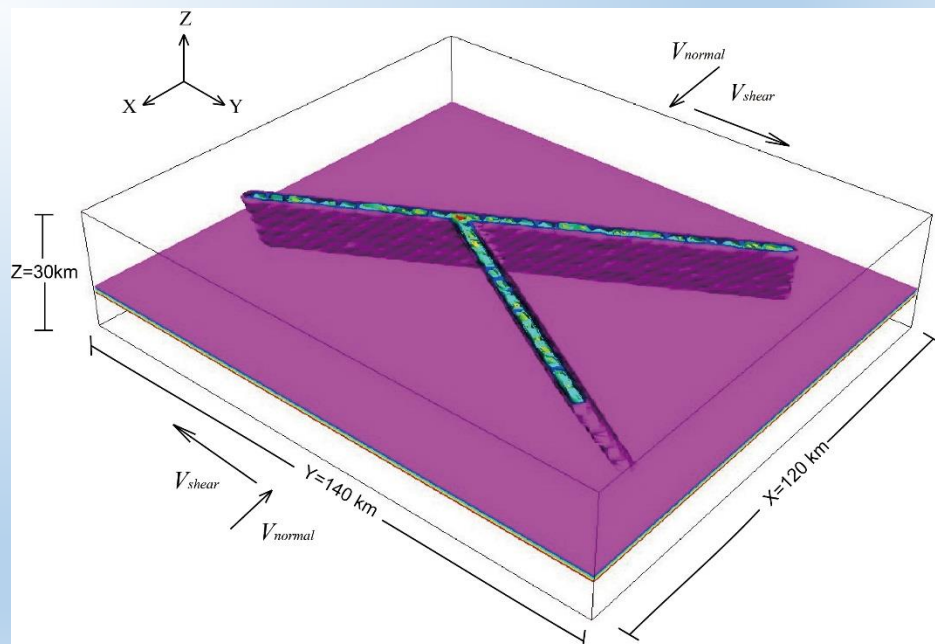
By X.Feng, M.Gerbault, R.Martin, J.Ganne, M.W.Jessell

QUESTION ?

- ? How pre-existing fault orientations influence the distribution of high-strain zones?
- ? Tensile Vs. Compressional areas?

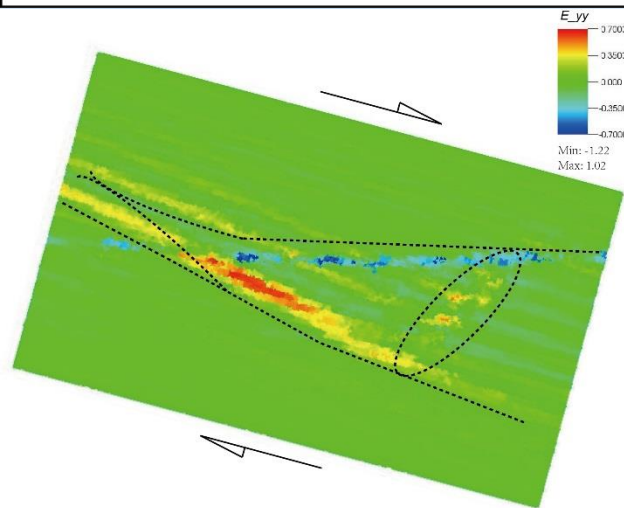
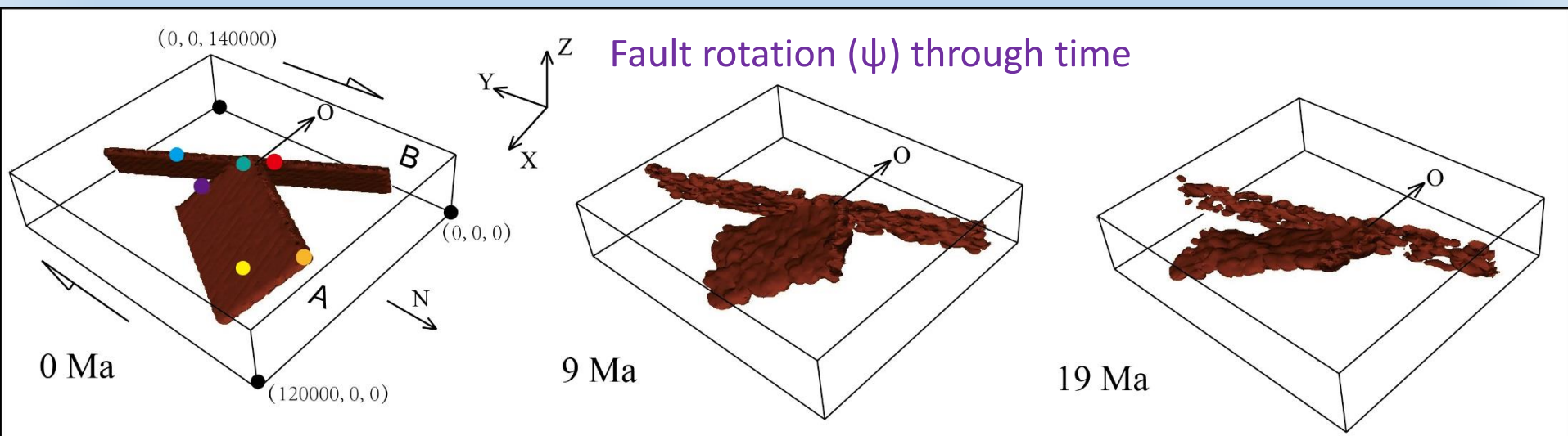


Geological setting NW Ghana



Numerical Model (with Underworld)

NUMERICAL ANSWER



Sketch showing the evolution of faults geometries

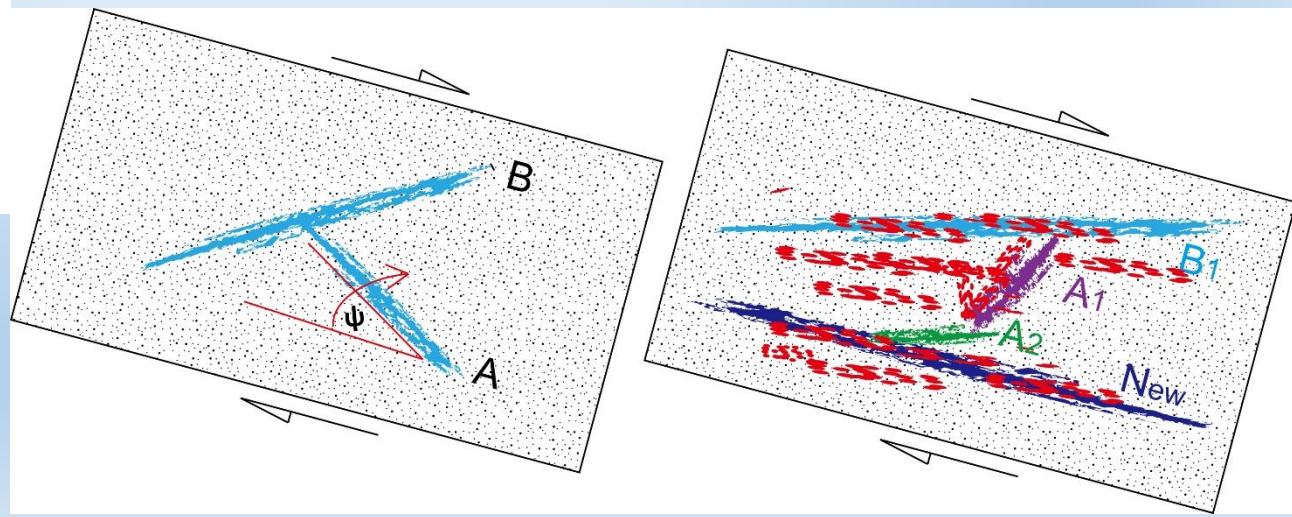
Tensile: A1 + A2 + New

Shear: New + B1

Finite strain zones

Tensile in Red (A1 + A2 + New)

Compressional in Blue



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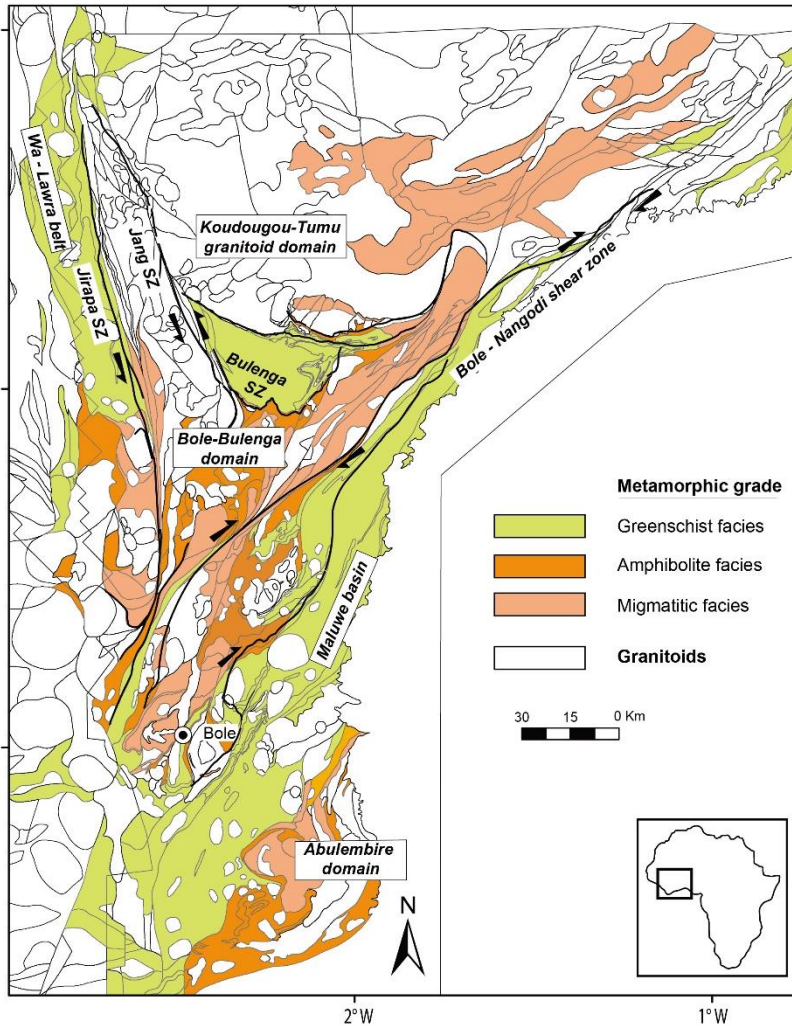
WAXI - West African Exploration Initiative

IXOA - L'Initiative d'Exploration Ouest Africaine

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1. Geological setting NW Ghana



- ❖ The study area formed during the Eburnean orogeny (2.15–2.00 Ga)
- ❖ High-strain zones (Y shape fault system) bound and separate the high-grade terranes from low-grade greenstone belts
- ❖ High-grade assemblages are mainly observed between the Jirapa and Bole-Nangodi shear zones and partly around the town Bole.
- ❖ Bounded by granitoid domains, two main pulses of intrusions, ranging between 2213 Ma and 2080 Ma.

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2. Numerical modelling with Underworld

The governing equations of momentum (1), mass (2) and energy (3) conservation are expressed as follows:

$$\frac{\partial \sigma_{ij}}{\partial x_j} = \rho(T) g_i \quad (1)$$

$$\nabla \cdot u = 0 \quad (2)$$

$$\rho C_p \frac{DT}{Dt} + (u \cdot \nabla) T = \frac{\partial}{\partial x_i} \left(\frac{k_i \partial T}{\partial x_i} \right) + \rho H \quad (3)$$

The depth dependent Drucker - Prager yield criterion (4) and a temperature-dependent non - Newtonian power law (5) viscosity were used:

$$\tau_{yield} = P \cdot \tan \varphi + \sigma_p \quad (4)$$

$$\eta_m = 0.25 \cdot (0.75A)^{(-1/n)} \cdot \dot{\epsilon}^{((1/n)-1)} \cdot \exp(Q/nRT) \quad (5)$$

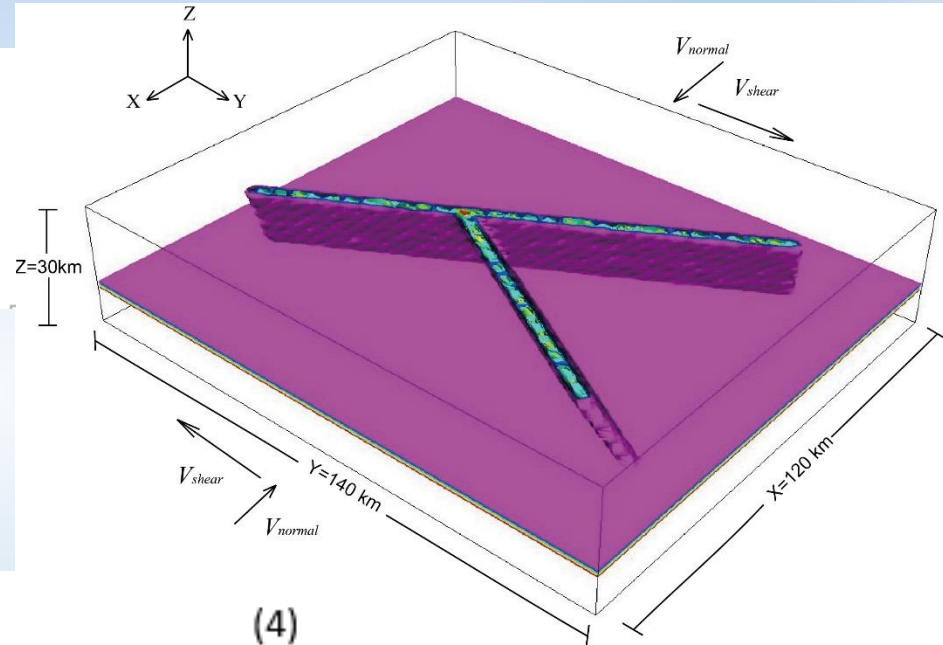


Table. 1. Rheological properties of crustal layers. The non-Newtonian power law is used for the upper and lower crusts, whereas we employ a constant viscosity for fault zones.

Layer	A MPa ⁻ⁿ / s	n	Q KJ/mol
Upper crust	2.0 x 10 ⁻⁴	3.4	260
Lower crust	1.3 x 10 ⁻³	2.4	219
Fault	1 x 10 ²⁰ Pas		

2. Numerical modelling with Underworld

Table. 2. Model parameters common to all experiments. $\ast\rho_{fault}$, $\ast\Delta T$ and $\ast\varphi$ indicate the values that change according to different models.

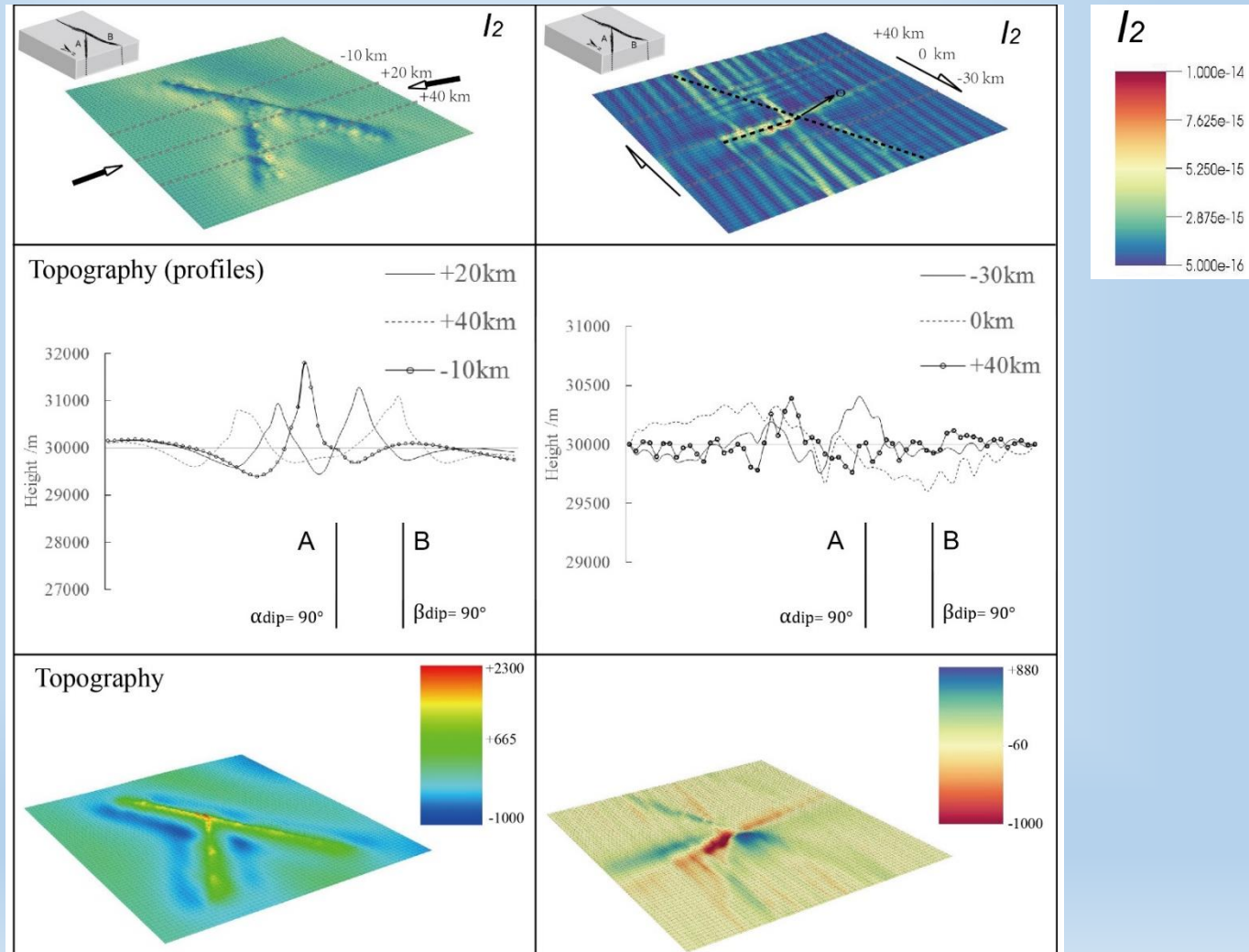
Parameter	Symbol	Value-units
Thickness	Hupper	20 km
	Hlower	10 km
Density	$\ast\rho_{upper}$	2800 kg/m ³
	ρ_{lower}	3000 kg/m ³
	$\ast\rho_{fault}$	2600 kg/m ³
Dislocation creep power law pre-exponential factor	A	MPa ⁻ⁿ / s
Dislocation creep power law exponent	n	-
Power law creep activation energy	Q	KJ/mol
Temperature gradient	$\ast\Delta T$	°C/Km
Gas constant	R	8.314 J mol ⁻¹ K ⁻¹
Heat capacity	C _p	1000 J/kg·K
Thermal diffusivity	γ	m ² /s
Thermal expansion	λ	K ⁻¹
Gravitational acceleration	g	9.81 m/s ²
Friction angle	$\ast\phi$	2.5~36.5 °
Cohesion for Druker-Prager	σ _c	1-20 MPa
Viscosity limit	η	1e+20<η<1e+23 Pa s

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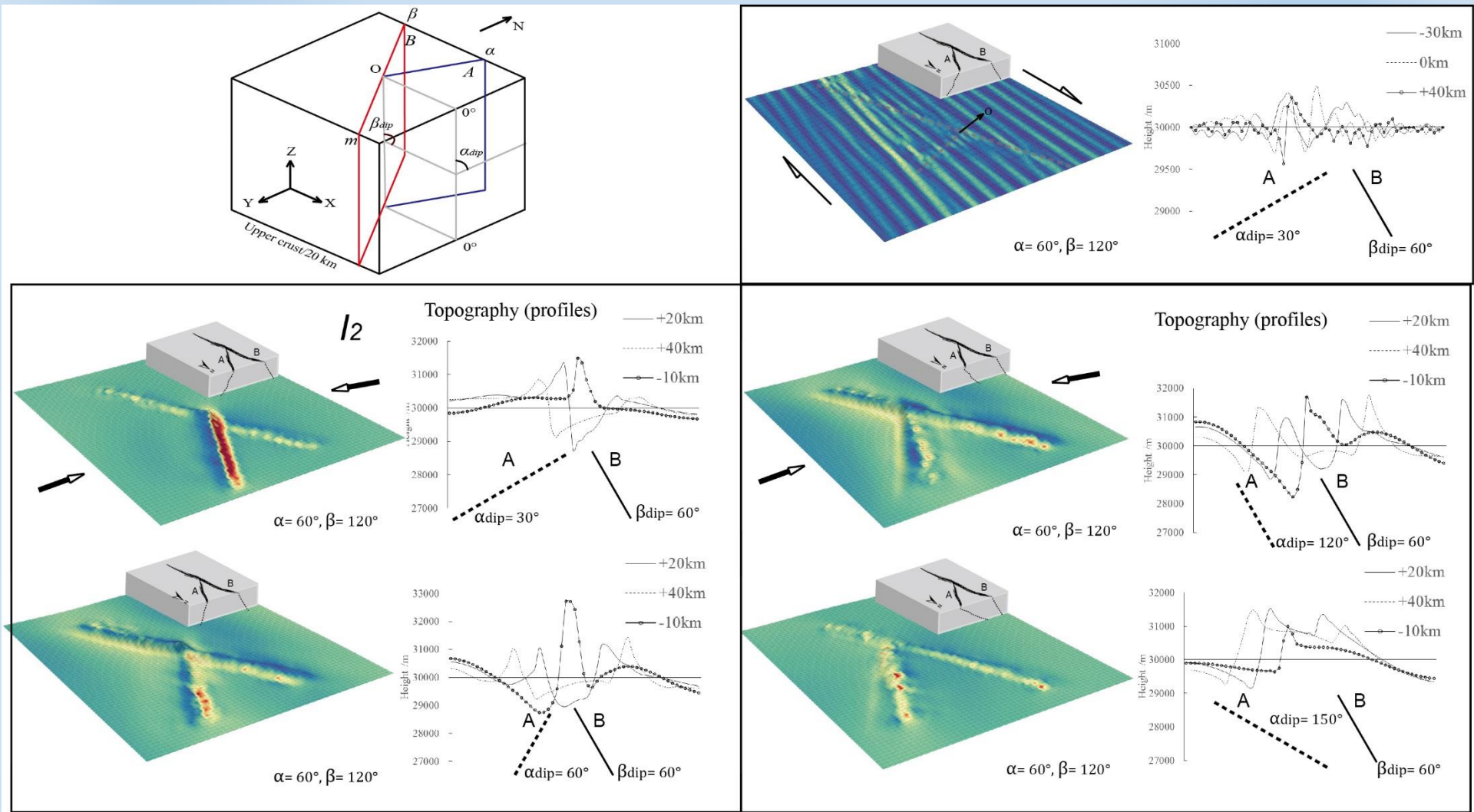
3. Result

Two vertical faults under compressional (**left**) and simple shear (**right**) boundary conditions.



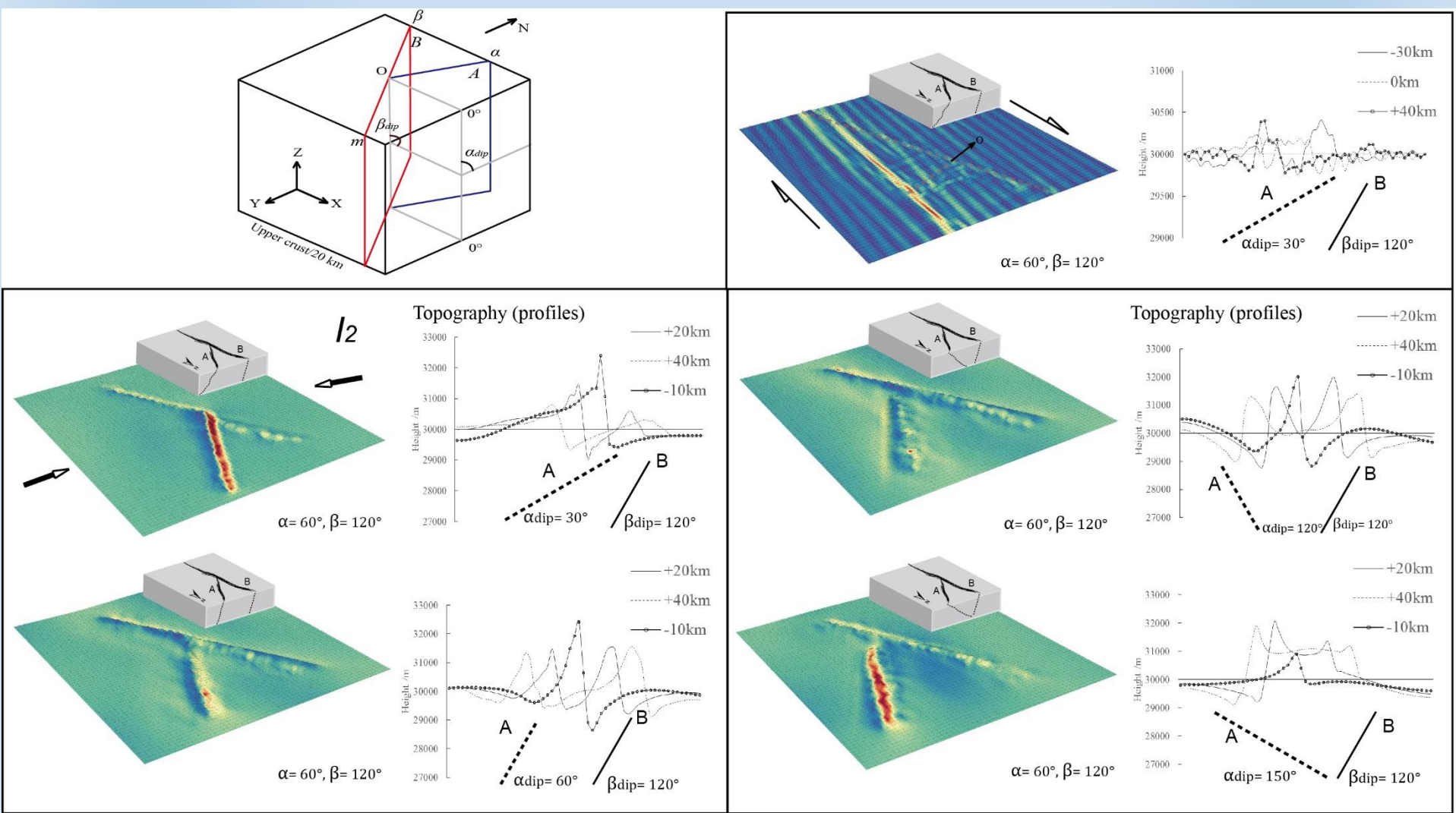
3. Result

Distribution of deformation and Topo Vs. fault orientations



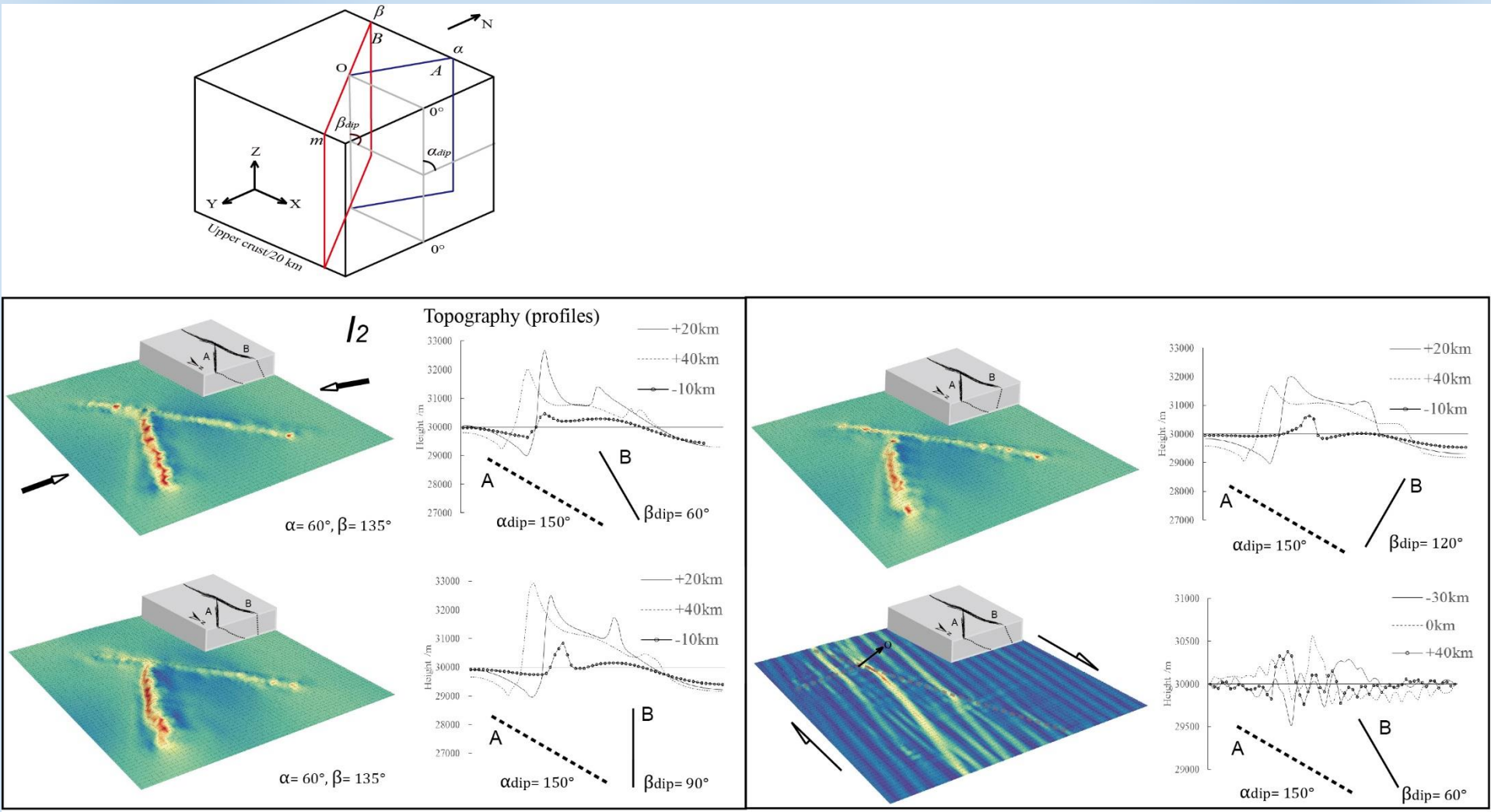
3. Result

Distribution of deformation and Topo Vs. fault orientations



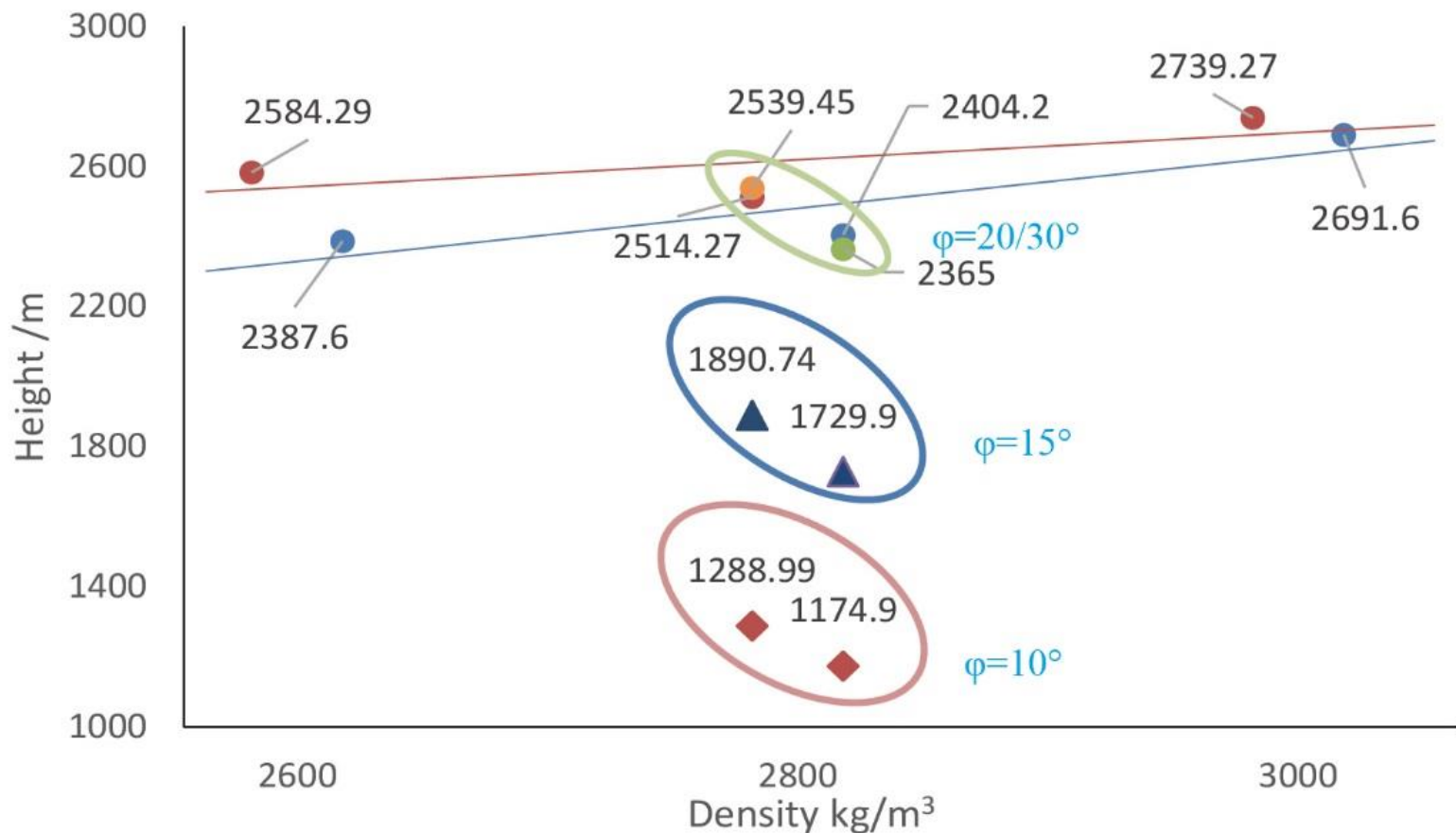
3. Result

Distribution of deformation and Topo Vs. fault orientations



3. Result

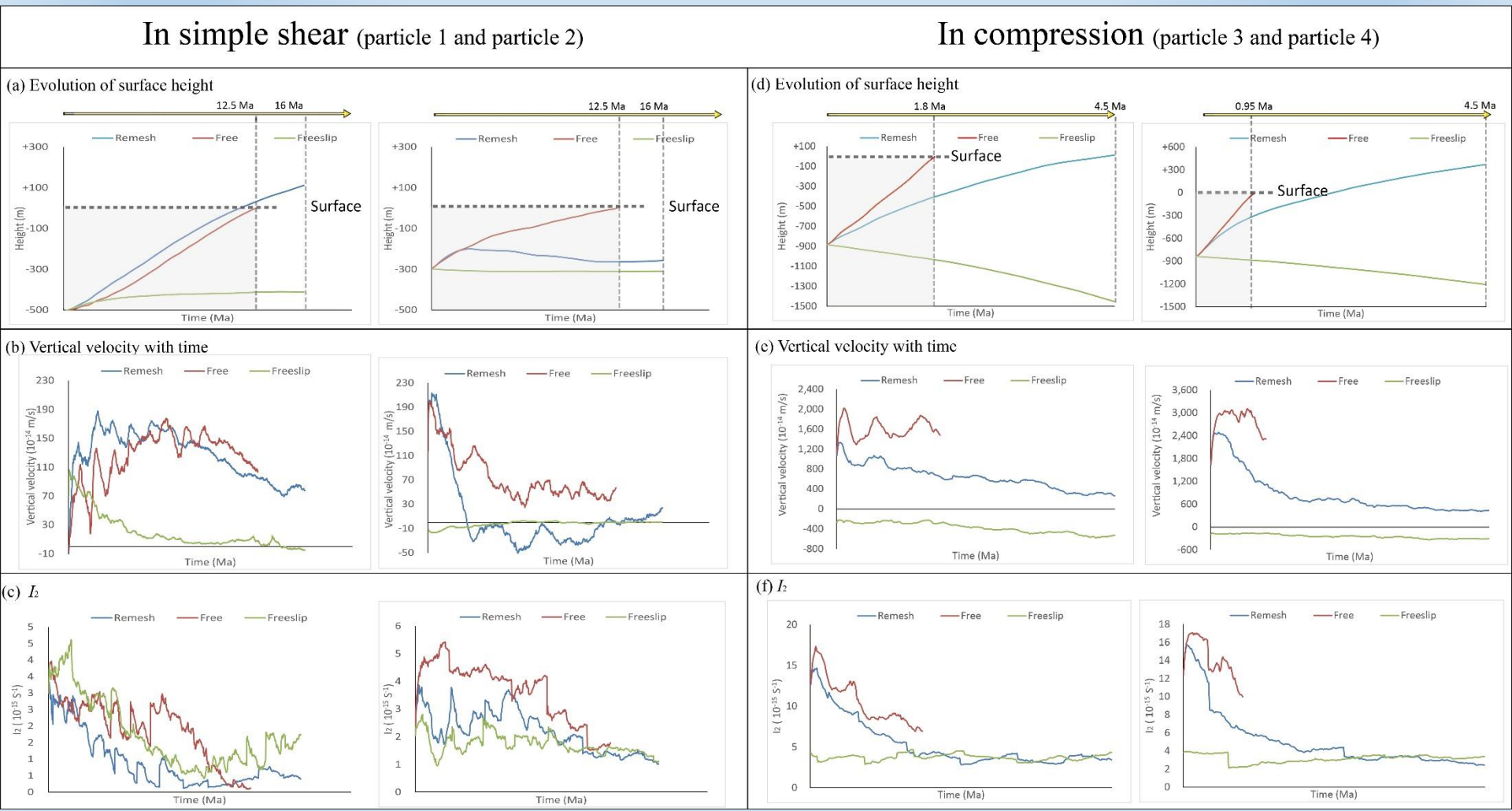
Density variations Vs. friction angle in building Topo



3. Result

Effect of the top surface boundary conditions

1. Freeslip, 2. Free, 3. Remesh

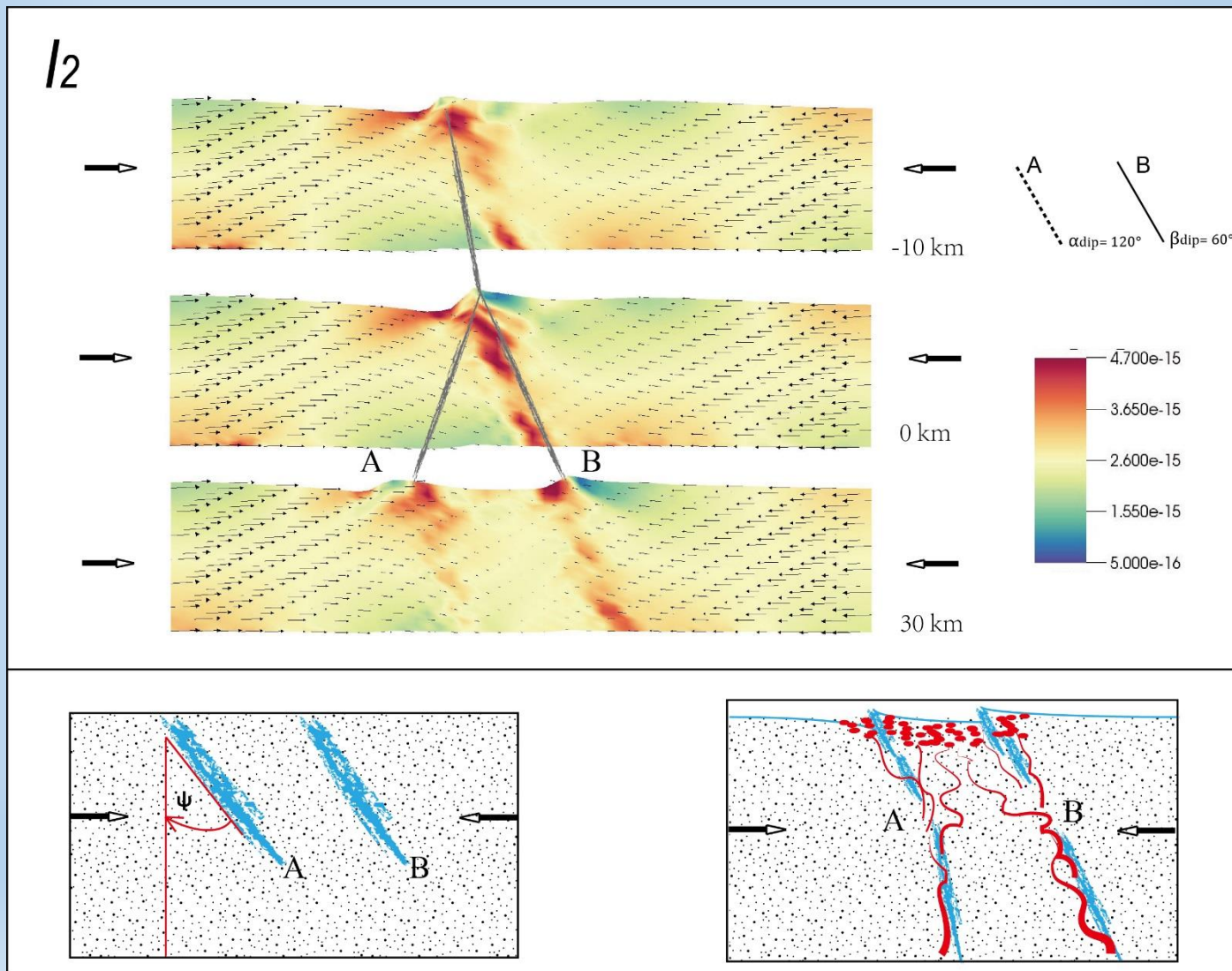


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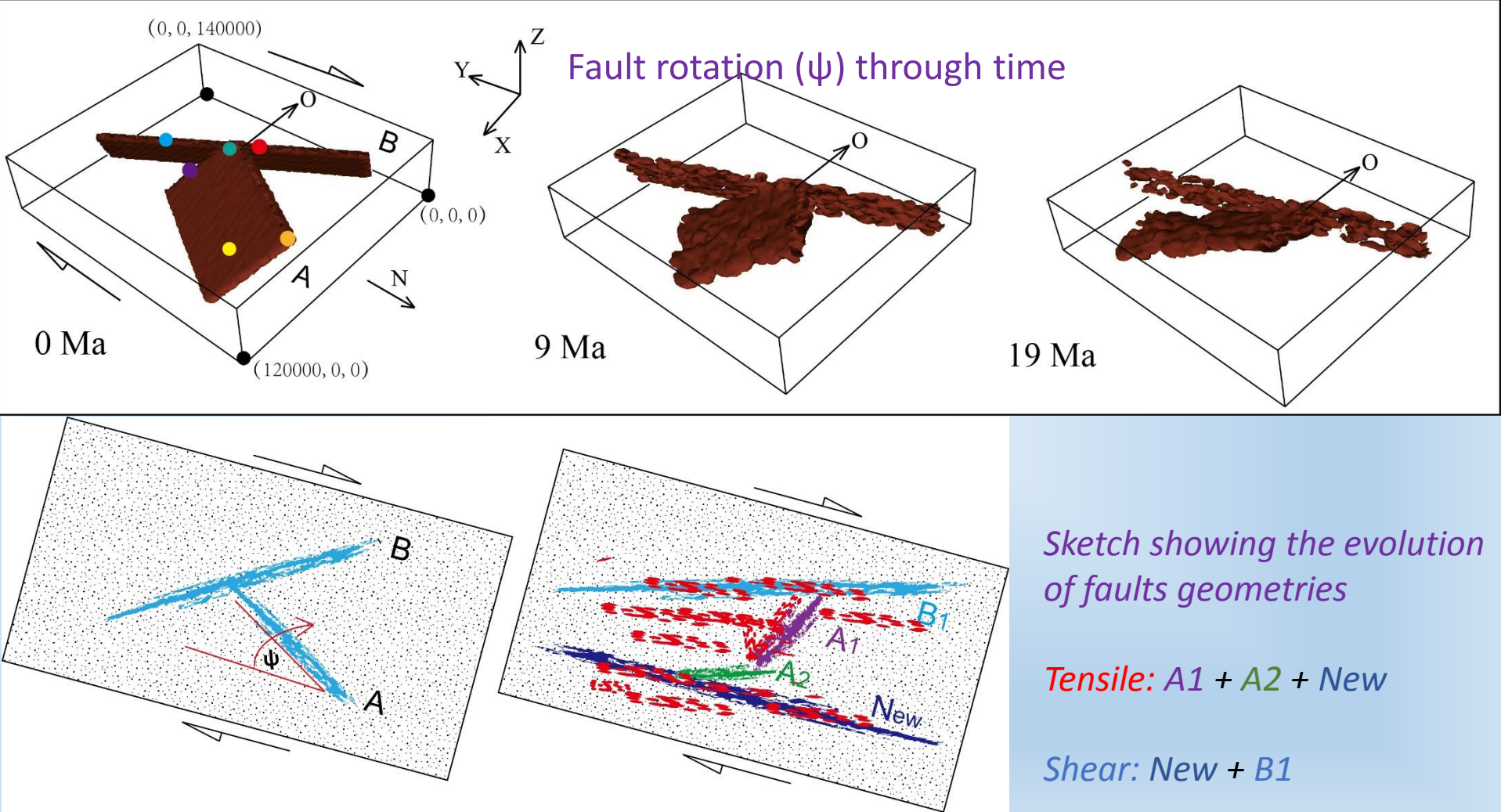
3. Result

Faults rotation under compressional boundary conditions.



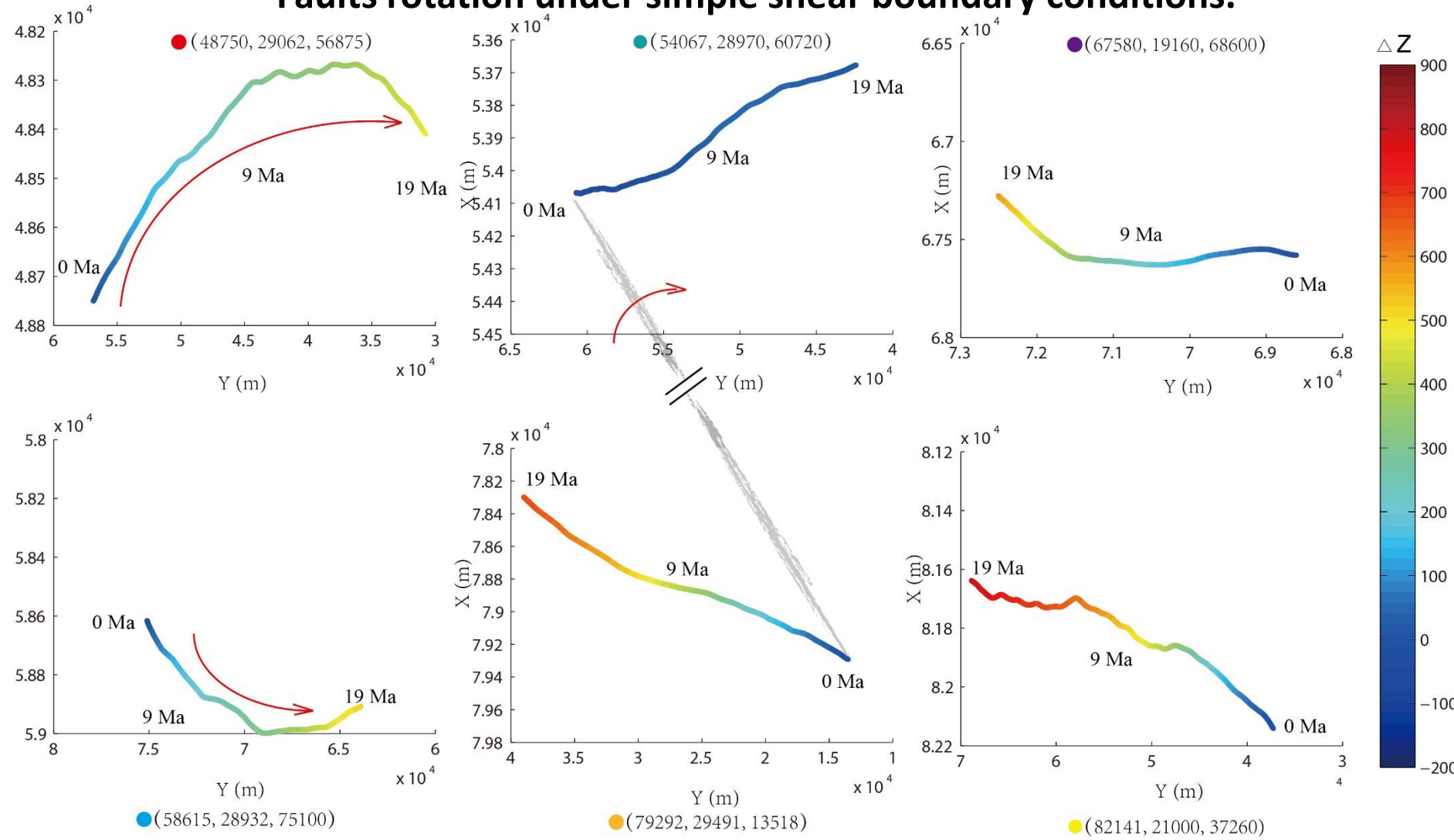
3. Result

Faults rotation under simple shear boundary conditions.



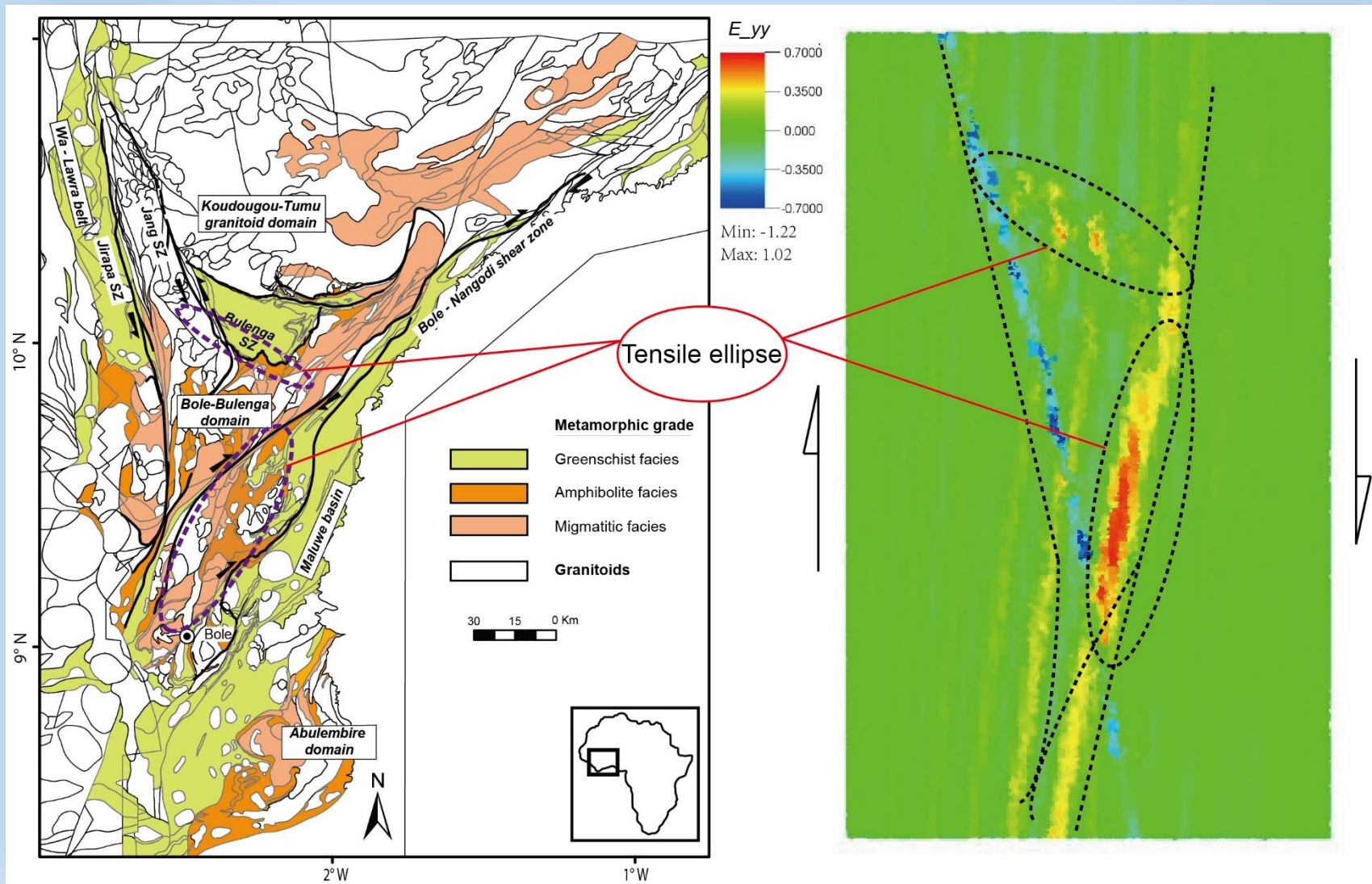
3. Result

Faults rotation under simple shear boundary conditions.



3. Result

Finite strain zones
Tensile in Red, Compressional in Blue

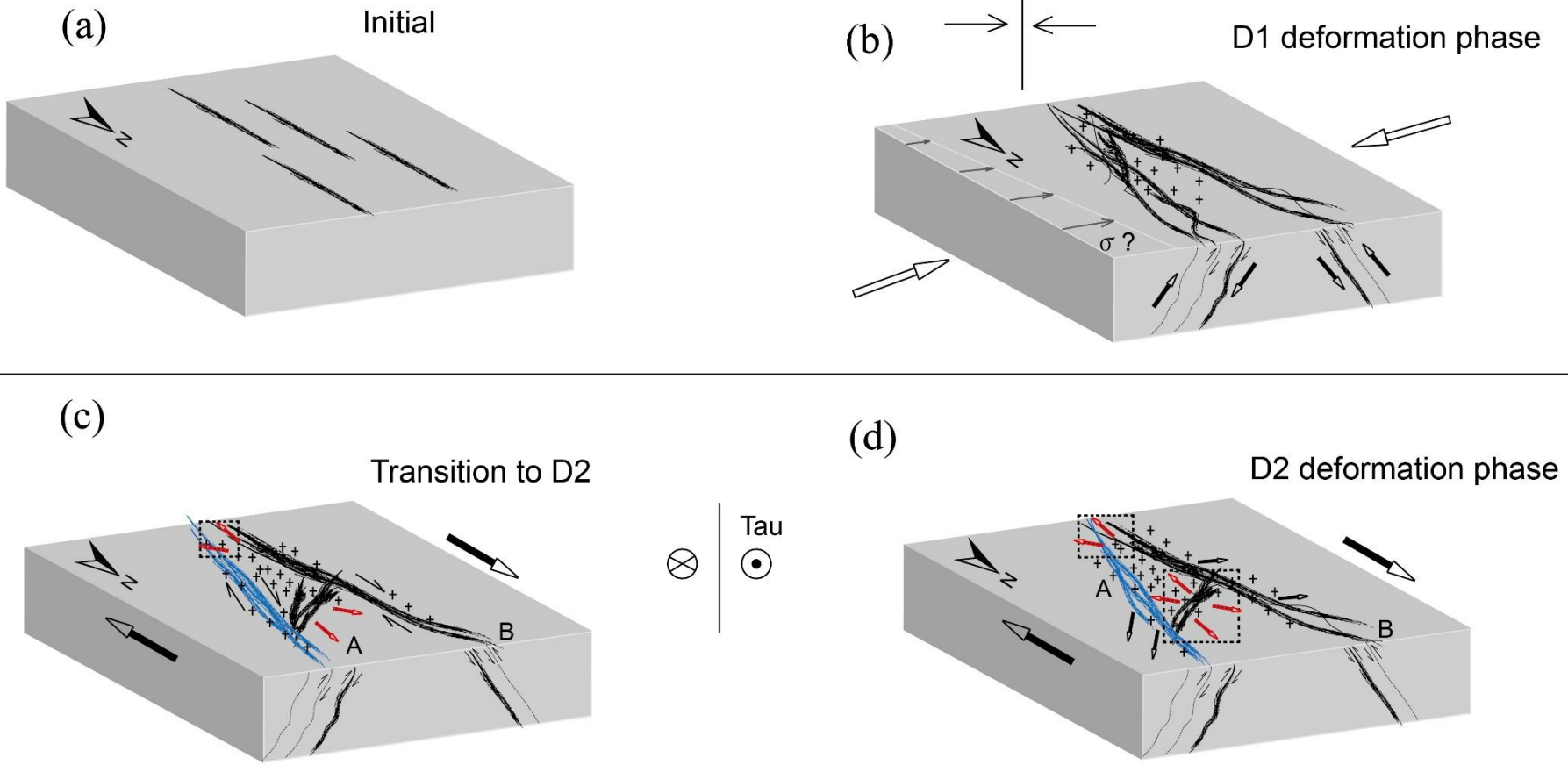


4. Reconstruction

D1 deformation phase: Compression

D2 deformation phase: Local extension

Tectonic scenario for the studied area



Thank you!