



Mapping of the source regions of the dust jets on comet 67P

Wing-Huen Ip (1,2), Ian-Lin Lai (2), Jui-Chi Lee (3), Zhong-Yi Lin (1), Yu Chen (4), Ting-Wei Chien (5), Yu-Min Chen (5), Jean-Baptiste Vincent (6), Holger Sierks (6), Cesare Barbieri (7,8), Philippe. L. Lamy (9), Rafael Rodrigo (10,11), Detlef Koschny (12), Hans Rickman (13,14), Uwe Keller (15), and Osiris Team (6)

(1) Institute of Astronomy, National Central University, Taoyuan City Taiwan (wingip@astro.ncu.edu.tw), (2) Institute of Space Sciences, National Central University, Taoyuan City Taiwan, (3) Department of Earth Sciences, National Central University, Taoyuan City, Taiwan, (4) Department of Atmosphere Sciences, National Central University, Taoyuan City, Taiwan, (5) Department of Physics, National Central University, Taoyuan City, Taiwan, (6) Max-Planck Institute for Solar System Research, Goettingen, Germany, (7) Department of Physics and Astronomy "G. Galilei", University of Padova, Padova, Italy, (8) Centro de Ateneo di Studi ed Attivita Spaziali "Giuseppe Golombo", University of Padova, Padova, Italy, (9) Laboratoire d'Astro-physique de Marseille, Marseille, France, (10) Centro de Astrobiologia (INTA-CSIC), Madrid, Spain, (11) International Space Science Institute, Bern, Switzerland, (12) Research and Scientific Support Department, European Space Agency, Noordwijk, The Netherlands, (13) Department of Physics and Astronomy, Uppsala University, Uppsala, Sweden, (14) PAS Space Research center, Warszawa, Poland, (15) Institute for Geophysics and Exterrestrial Physics, Technical University of Braunschweig, Braunschweig, Germany

Because of the inclination of the rotational axis to the orbital plane and the orbital motion around the Sun, the sunlit regions on the nucleus surface of comet 67P/Churyumov-Gerasimenko moved from the northern hemisphere to the southern hemisphere between August, 2014 and October, 2015. From the comparison of the dust jet features in images taken at successive time series, the footprints of these jets can be identified by a projection method. The distributions of the corresponding source regions can be compared with the geomorphology of the nucleus surface from inbound to outbound. The correlation of the dust jet activity with the volatile outgassing phenomenon as monitored by different scientific instruments onboard Rosetta will provide important information on the sublimation process.