



Gravity science investigation of Ceres from Dawn

Ryan Park (1), Alexander Konopliv (1), Bruce Bills (1), Julie Castillo-Rogez (1), Sami Asmar (1), Nicolas Rambaux (2), Carol Raymond (1), Christopher Russell (3), Maria Zuber (4), Anton Ermakov (4), Scott King (5), and Marc Rayman (1)

(1) Jet Propulsion Laboratory, Solar Systems Dynamics, Pasadena, United States (ryan.s.park@jpl.nasa.gov), (2) IMCCE, Observatoire de Paris, Paris, France, (3) UCLA, Los Angeles, United States, (4) Massachusetts Institute of Technology, Cambridge, MA, United States, (5) Department of Geoscience, Virginia Tech, Blacksburg, VA, United States

The Dawn gravity science investigation utilizes the DSN radiometric tracking of the spacecraft and on-board framing camera images to determine the global shape and gravity field of Ceres. The gravity science data collected during Approach, Survey, and High-Altitude Mapping Orbit phases were processed. Currently, the latest gravity field called CERES08A is available, which is globally accurate to degree and order 5. Combining the gravity and shape data gives the bulk density of $2162.5 \pm 8 \text{ kg/m}^3$. The low Bouguer gravity at high topography area, or vice versa, indicates that the surface of Ceres is likely compensated and that its interior presents a low-viscosity layer at depth. The degree 2 gravity harmonics show that the rotation of Ceres is very nearly about a principal axis. This consistent with hydrostatic equilibrium at 1% level, and infers a mean moment of inertia of Ceres is 0.36, implying some degree of central condensation. Based on a simple two-layer model of Ceres and assuming carbonaceous chondrites and hydrostatic equilibrium, the core size is expected to be $\sim 280 \text{ km}$ with corresponding average thickness of the outer shell of $\sim 190 \text{ km}$.