



## Particle Dynamics at and near the Electron and Ion Diffusion Regions as a Function of Guide Field

Barbara Giles (1), James Burch (2), Tai Phan (3), James Webster (4), Levon Avanov (1,5), Roy Torbert (6), Li-Jen Chen (1,5), Michael Chandler (7), John Dorelli (1), Robert Ergun (8), Stephen Fuselier (2), Daniel Gershman (1,5), Benoit Lavraud (9), Thomas Moore (1), William Paterson (1), Craig Pollock (10), Christopher Russell (11), Yoshifumi Saito (12), Robert Strangeway (11), Shan Wang (5), and the additional members of the MMS Team  
(1) NASA Goddard Space Flight Center, Greenbelt, MD, United States (barbara.giles@nasa.gov), (2) Southwest Research Institute, San Antonio, TX, United States, (3) University of California, Berkeley, Berkeley, CA, United States, (4) Rice University, Department of Physics and Astronomy, Houston, TX, United States, (5) University of Maryland, College Park, MD, United States, (6) University of New Hampshire, Durham, NH, United States, (7) NASA Marshall Space Flight Center, Huntsville, AL, United States, (8) University of Colorado / Laboratory for Atmospheric & Space Physics, Boulder, CO, United States, (9) Research Institute in Astrophysics and Planetology, Toulouse, France, (10) Denali Scientific, Healy, AK, United States, (11) University of California, Los Angeles, Los Angeles, CA, United States, (12) Institute for Space Science, Sagami-hara, Japan

At the dayside magnetopause, magnetic reconnection often occurs in thin sheets of plasma carrying electrical currents and rotating magnetic fields. Charged particles interact strongly with the magnetic field and simultaneously their motions modify the fields. Researchers are able to simulate the macroscopic interactions between the two plasma domains on both sides of the magnetopause and, for precise results, include individual particle motions to better describe the microscopic scales. Here, observed ion and electron distributions are compared for asymmetric reconnection events with weak-, moderate-, and strong-guide fields. Several of the structures noted have been demonstrated in simulations and others have not been predicted or explained to date. We report on these observations and their persistence. In particular, we highlight counterstreaming low-energy ion distributions that are seen to persist regardless of increasing guide-field. Distributions of this type were first published by Burch and Phan [GRL, 2016] for an 8 Dec 2015 event and by Wang et al. [GRL, 2016] for a 16 Oct 2015 event. Wang et al. showed the distributions were produced by the reflection of magnetosheath ions by the normal electric field at the magnetopause. This report presents further results on the relationship between the counterstreaming ions with electron distributions, which show the ions traversing the magnetosheath, X-line, and in one case the electron stagnation point. We suggest the counterstreaming ions become the source of D-shaped distributions at points where the field line opening is indicated by the electron distributions. In addition, we suggest they become the source of ion crescent distributions that result from acceleration of ions by the reconnection electric field.

Burch, J. L., and T. D. Phan (2016), Magnetic reconnection at the dayside magnetopause: Advances with MMS, *Geophys. Res. Lett.*, 43, 8327–8338, doi:10.1002/2016GL069787.

Wang, S., et al. (2016), Two-scale ion meandering caused by the polarization electric field during asymmetric reconnection, *Geophys. Res. Lett.*, 43, 7831–7839, doi:10.1002/2016GL069842.