



Ionospheric ions in the reconnection region

Wenya Li (1), Mats André (1), Yuri Khotyaintsey (1), Andris Vaivads (1), Sergio Toledo-Redondo (2), Daniel Graham (1), Andrey Divin (1), Stephen Fuselier (3), David Young (3), Drew Turner (4), Per-Arne Lindqvist (5), Robert Ergun (6), Christopher Russell (7), Werner Magnes (8), Roy Torbert (9), Barbara Giles (10), Thomas Moore (10), Michael Chandler (11), and Jim Burch (3)

(1) Swedish Institute of Space Physics, Uppsala, Sweden (wyli@irfu.se), (2) Science directorate, European Space Agency, ESAC, Madrid, Spain, (3) Southwest Research Institute, San Antonio, Texas, USA, (4) Space Sciences Department, The Aerospace Corporation, El Segundo, California, USA, (5) Royal Institute of Technology, Stockholm, Sweden, (6) Laboratory of Atmospheric and Space Physics, University of Colorado, Boulder, CO, USA, (7) Department of Earth and Space Sciences, University of California, Los Angeles, California, USA, (8) Space Research Institute, Austrian Academy of Science, Graz, Austria, (9) Space Science Center, University of New Hampshire, Durham, New Hampshire, USA, (10) NASA Goddard Space Flight Center, Greenbelt, Maryland, USA, (11) NASA Marshall Space Flight Center, Huntsville, AL, USA

Magnetosheath plasma usually determines many properties of the asymmetric magnetic reconnection at the subsonic region of Earth's magnetopause. Cold plasma originating from the ionosphere can also reach the magnetopause, and modify the kinetic physics of the asymmetric reconnection. On 1st November 2015, the Magnetospheric Multiscale (MMS) spacecraft observed a magnetopause crossing with high-density ($10 - 60 \text{ cm}^{-3}$) cold plasma in the magnetosphere and reconnection jets. Jets of ionospheric and magnetosheath origin, with different velocities, are found at the same time between the current sheet and magnetosheath separatrix. The ionospheric-ion jets consisting of H^+ , He^+ , and O^+ are accelerated to about 300 km/s along the magnetic field. The magnetopause crossing is estimated to be about 300 ion-inertial lengths away from the magnetic reconnection X-line. The ionospheric-ion jet here is suggested to be from the cold ion inflow close the X-line.