



Characterization of the refractory organic matter present in the dust particles of 67P/Churyumov-Gerasimenko.

Nicolas Fray (1), Donia Baklouti (2), Anaïs Bardyn (1,3), Christelle Briois (3), Hervé Cottin (1), Cécile Engrand (4), Henning Fischer (5), Martin Hilchenbach (5), Robin Isnard (1), Léna Le Roy (6), Paola Modica (1), John Paquette (5), Jouni Ryno (7), Oliver Stenzel (5), Sandra Siljeström (8), and Laurent Thirkell (3)

(1) LISA, UMR 7583 du CNRS, CRETEIL, France (nicolas.fray@lisa.u-pec.fr), (2) IAS, CNRS / Université Paris Sud, Bâtiment 121, 91405 Orsay, France, (3) LPC2E, CNRS / Université d'Orléans Avenue de la Recherche Scientifique 3A, F-45071 Orléans cedex, France, (4) CSNSM, CNRS/IN2P3, Université Paris Sud, UMR 8609, Université Paris Saclay Bâtiment 104, 91405 Orsay, France, (5) Max-Planck-Institut für Sonnensystemforschung Justus-von-Liebig-Weg 3, D-37077 Göttingen, Germany. , (6) University of Bern Sidlerstr. 5, CH-3012 Bern, Switzerland., (7) Finnish Meteorological Institut, Observation services Erik Palménin aukio 1, 00560 Helsinki, Finland, (8) Departement of Chemistry, Materials and Surfaces, SP Technical Research Institute of Sweden Box, 857, 501 15 Borås, Sweden

The Cometary Secondary Ion Mass Analyser (COSIMA), a miniaturized time-of-flight secondary ion mass spectrometer (ToF-SIMS), is one of the dust particle instruments onboard the orbiter of the Rosetta mission that arrived to comet 67P/Churyumov-Gerasimenko in mid-2014. COSIMA analyses the mineral and organic composition of dust particles that are captured on metal targets exposed to space [1, 2].

The mass spectra acquired by COSIMA show that refractory organic matter is ubiquitous in cometary dust particles [2]. The nature of this carbonaceous material will be discussed. We will highlight the abundance of organic matter in these dust particles as well as the nitrogen to carbon elemental ratio of this refractory organic matter. These results obtained on the dust particles of 67P will be compared to those from other astrophysical objects (carbonaceous chondrites, IDPs, micro-meteorites). This comparison could provide clues on the origin and evolution of the cometary organic matter.

References: [1] Hilchenbach, M. et al. (2016) *ApJ*, 816, L32. [2] Schulz, R. et al. (2015) *Nature*, 518, 216-218. [3] Fray, N. et al. (2016) *Nature*, 538, 72-74.