

The large-scale ionospheric transient current system response to upstream solar wind IMF Bz north-south and south-north turnings as seen by the WIND satellite and the full SuperMAG network of ground based magnetometers

Joe Dods (1), Sandra Chapman (1), Jesper Gjerloev (2,3)

(1) CFSA, University of Warwick, Coventry, United Kingdom, (2) Johns Hopkins University Applied Physics Laboratory, Laurel, USA, (3) Department of Physics and Technology, University of Bergen, Bergen, Norway

We characterise the response of the quiet-time (no substorms or storms) large scale ionospheric convection system to north-south and south-north IMF turnings by using a dynamical network of ground-based magnetometers. Canonical correlation between all pairs of SuperMAG magnetometer stations in the northern hemisphere (MLat $50-82^{\circ}$) is used to establish the extent of near-simultaneous magnetic response between regions of MLT-MLat. Parameters and maps that describe spatial-temporal correlation are used to characterise the system and its response to the turnings aggregated over several hundred events. We find that regions that experience large increases in correlation post-turning coincide with typical locations of a two cell convection system and are influenced by the IMF B_y . The time between the turnings reaching the magnetopause and a network response is found to be ~8-10 minutes and correlation in the dayside occurs 2-8 mins before that in the nightside.