



The longitudinal distribution of solar energetic particles

Raúl Gómez-Herrero (1), Nina Dresing (2), Andreas Klassen (2), Bernd Heber (2), Radoslav Bučík (3), David Lario (4), Daniel Pacheco (5), Angels Aran (5), Neus Agueda (5), Javier Rodríguez-Pacheco (1), Juan José Blanco (1), Ignacio Cernuda (1), and Francisco Espinosa Lara (1)

(1) Universidad de Alcalá, Departamento de Física y Matemáticas, Alcalá de Henares, Spain (raul.gomezh@uah.es), (2) Institut für Experimentelle und Angewandte Physik, University of Kiel, D-24118, Kiel, Germany, (3) Institut für Astrophysik, Georg-August-Universität Göttingen, D-37077, Göttingen, Germany, (4) The Johns Hopkins University, Applied Physics Laboratory, Laurel, MD 20723, USA, (5) Dep. Física Quàntica i Astrofísica, Institut de Ciències del Cosmos (ICCUB), Universitat de Barcelona, Martí i Franquès 1, 08028 Barcelona, Spain

Multi-spacecraft observations provide a powerful tool for the investigation of the spatial distribution of energetic particles in the heliosphere. Early observations from the Helios mission showed that solar energetic particles originating from a single solar event sometimes spread over very broad longitudinal spans. Ulysses observations evidenced they can also reach high heliographic latitudes. During the last decade, the unprecedented combination of multi-point remote-sensing and in-situ observations from the STEREO mission and the fleet of near-Earth spacecraft has greatly enhanced our understanding of the longitudinal distribution of solar energetic particle events in the ecliptic plane. These observations have illustrated a rich variety of events, ranging from narrow impulsive events sometimes showing unexpected longitudinal distribution patterns, to wide-spread events filling almost the entire space within the Earth orbit. We present a revision of these observations, the possible underlying physical processes involved, and the prospects for upcoming missions as Solar Orbiter and Parker Solar Probe.