Geophysical Research Abstracts Vol. 20, EGU2018-5193, 2018 EGU General Assembly 2018 © Author(s) 2018. CC Attribution 4.0 license.



Magnetic reconnection signatures of dissipating flux ropes associated with dipolarization events: A MMS Case Study

Gangkai Poh (1), James Slavin (1), San Lu (3,8), Guan Le (2), Paul Cassak (4), Jonathan Eastwood (5), Dogacan Ozturk (1), Shasha Zou (1), Rumi Nakamura (6), Wolfgang Baumjohann (6), Christopher Russell (3,8), Daniel Gershman (2,9), Barbara Giles (2), Craig Pollock (2), Thomas Moore (2), Roy Torbert (10), and James Burch (7) (1) Department of Climate and Space Sciences and Engineering, University of Michigan, Ann Arbor, Michigan, United States, (2) NASA Goddard Space Flight Center, Greenbelt, Maryland, United States, (3) Department of Earth and Space Sciences, University of California, Los Angeles, California, United States, (4) Department of Physics and Astronomy, West Virginia University, Morgantown, West Virginia, United States, (5) The Blackett Laboratory, Imperial College London, London, SW7 2AZ, United Kingdom, (6) Space Research Institute, Austrian Academy of Sciences, Graz, Austria, (7) Southwest Research Institute, San Antonio, Texas, United States, (8) Institute of Geophysics and Planetary Physics, Los Angeles, California, United States, (9) Department of Astronomy, University of Maryland, College Park, Maryland, United States, (10) University of New Hampshire, Durham, New Hampshire, United States

Three-dimensional global hybrid simulations and THEMIS observations have shown that earthward-moving flux ropes can undergo magnetic reconnection (or "re-reconnection") with the near-Earth dipole field in the downtail region between the Near Earth Neutral Line and the near-Earth dipole field. Slavin et al., [2003] proposed that earthward flux ropes dissipate through reconnection as the flux ropes push up against the northward geomagnetic field in the inner tail. This "re-reconnection" causes the southward BZ field in the leading edge of the flux rope to "erode" while continuous reconnection of lobe magnetic flux at the X-line tailward of the flux rope causes a "pileup" of northward BZ field, resulting in the asymmetric south-north BZ signature in earthward flux ropes and approximately one-third of DFs with negative BZ dips prior to their observation. In this study, we analyzed magnetic reconnection signatures preceding three earthward-moving flux ropes embedded within DF encounters, which were observed sequentially during MMS Phase 2B tail season. We observed electron outflow jets of ~ 1000 km/s for these events in the north-south direction consistent with the MMS location relative to the center of the cross-tail current sheet. The characteristic quadrupolar Hall magnetic field in the ion diffusion region and intense currents of $\sim 30-60$ nA/m2 were also observed when MMS encounters the reconnection region. J•E' > 0 in the reconnection region is consistent with magnetic reconnection being an energy dissipative process. Lastly, we will also discuss the electron distribution functions observed by the Fast Plasma Instrument and ionospheric responses due to dissipating flux ropes/DFs.