



Uranus and Neptune thermal evolution

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Uranus and Neptune have highly different intrinsic heat fluxes. This is surprising since they share a large number of very similar observed values such as mean density, surface temperature and atmospheric composition. Here, we present new Uranus and Neptune evolution calculations for a wide range of varied input parameters such as water equation of state, albedo or solar irradiation and show that those parameters can have considerable influence on the cooling behavior. We also discuss the influence of a possible radiative boundary layer between the H/He-rich outer envelope and the ice rich inner envelop on the resulting present heat flux.