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Refining the mechanisms of dust detection by antenna instruments

Ming-Hsueh Shen (1,2), Zoltan Sternovsky (1,2), Libor Nouzak (3), Mihály Horányi (1), David Malaspina (1), Sean Hsu (1), and Sheng-Yi Ye (4)

(1) Laboratory for Atmospheric and Space Physics, University of Colorado, Boulder, CO, United States, (2) Smead Aerospace Engineering Sciences Department, University of Colorado, Boulder, CO, United States, (3) Department of Surface and Plasma Science, Charles University, Prague, Czech Republic, (4) Physics and Astronomy Department, University of Iowa, Iowa City, IA, United States

Antenna-based detection of dust grains by spacecraft is a valuable mechanism to enhance the science return from existing missions and conduct serendipitous observations of interplanetary dust by spacecraft not equipped with dedicated dust instruments. In principle, the detection mechanism is simple: dust grains encountering the spacecraft at high relative speeds undergo ionization upon impact and some fraction of the resulting charged particles is recollected on the spacecraft body or antenna resulting in a measurable signal. However, there is a large uncertainty in calculating the mass of the dust particle from the impact signal. In order to enhance our understanding of the dust impact signals, a series of supporting laboratory measurements have been conducted using the dust accelerator facility at the University of Colorado. These included the measurements of the impact charge yield for common spacecraft materials and reproducing the basic signal-generating mechanisms. We identified three mechanisms (spacecraft charging, antenna charging, and antenna pickup) that also depend on the geometric arrangement, the impact location, as well as the bias potentials of the elements. The effective temperature of the impact plasma is on the order of 1eV for electrons and 5-20 eV for ions and determines the recollected charge, depending on the floating potential of the spacecraft. Our current focus is the study of impact signal generation by the Cassini spacecraft. For this a 20:1 reduced size model of Cassini has been constructed. The three Radio Plasma Wave Science (RPWS) antennas are configured either in a dipole or a monopole mode. The laboratory measurements support the recent suggestion that most dust detection events recorded in the dipole mode are due to antenna hits, as opposed to impacts on the spacecraft body. External magnetic fields, such as those close to Saturn during the proxima orbits, introduce a significant reduction of the dust impact signals when the Larmor radius of the ejected charge is smaller than the characteristic size of the spacecraft.