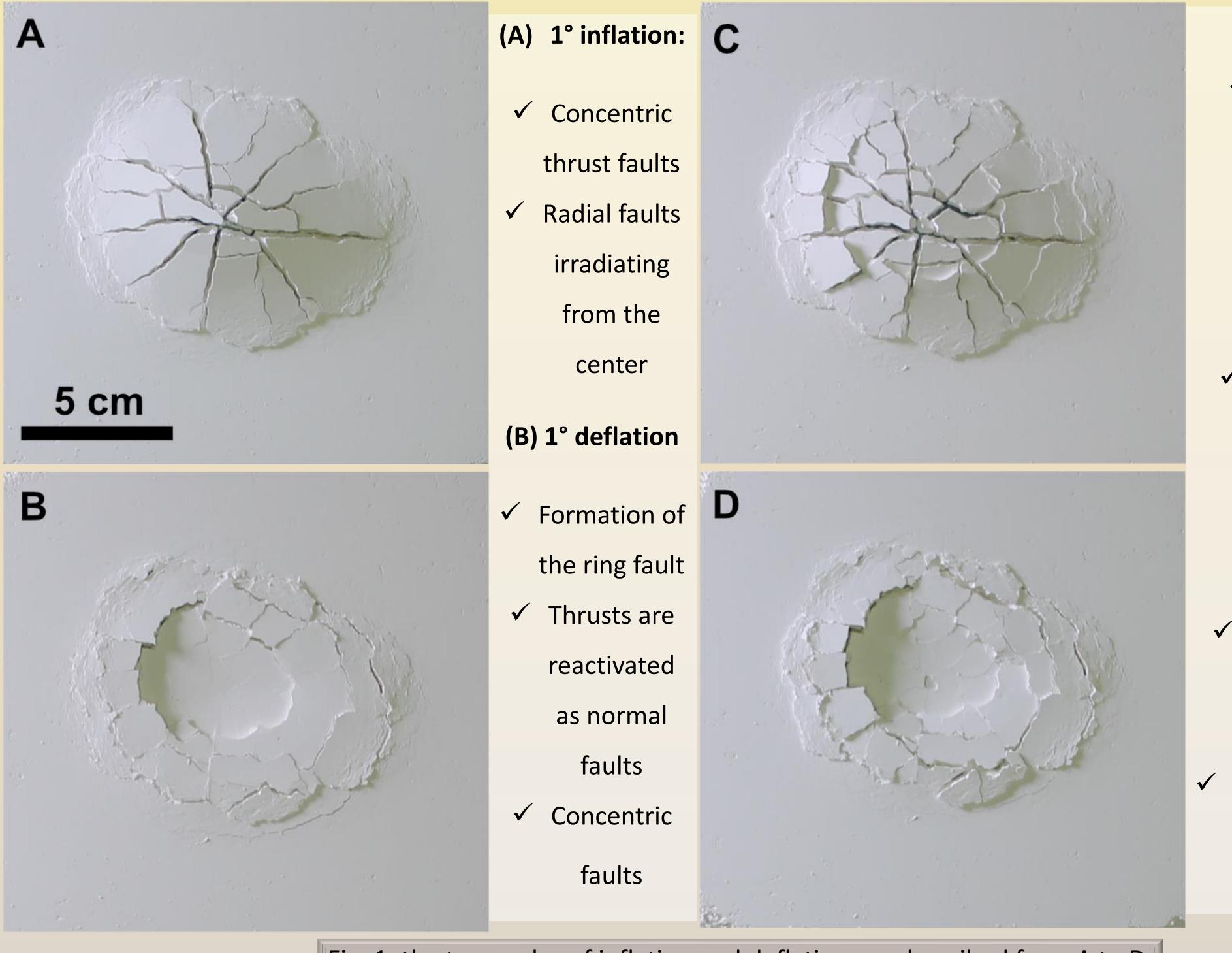


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Our analogue experiment

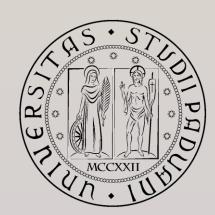
We performed a series of experiments (Fig. 1) using different setups to test the hypothesis that the peculiar features of chaotic terrains on Mars could have been formed by a particular type of caldera collapse known as *piecemeal* or chaotic (Branney & Kokelaar, 1994; Moore & Kokelaar, 1998; Roche et al., 2000; Scandone, 1990; Troll et al., 2002). An analogue magma chamber, where a polyglicerine imitating the magma was repeatedly intruded and extruded, was positioned under K-feldspatic sands simulating the brittle crust. Here the resulting structures formed on the analogue brittle crust after two cycles of inflation and deflation:



References

Branney, M. J., & Kokelaar, P. (1994). Geological Society of America Bulletin, 106(4), 507–530. // Dickson, J. et al. (2018). Lunar and Planetary Science Conference. // Moore, I., & Kokelaar, P. (1998). Geological Society of America Bulletin, 110(11), 1448. // Roche, O. et al. (2000). JGR: Solid Earth, 105(B1), 395–416. // Scandone, R. (1990). Journal of Volcanology and Geothermal Research, 42(3), 285–302. // Troll, V. R. et al. (2002). Geology, 30(2), 135–138.







Chaotic caldera collapse: a new interpretation for the origin of chaotic terrains on Mars

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Fig. 1: the two cycles of inflation and deflation are described from A to D.

(C) 2° inflation:

✓ Development of further radial and concentric faults (their intersection forms the polygonal blocks)

 \checkmark The top is already collapsed, with a minor inflation

(D) 2° deflation:

 \checkmark The caldera floor is collapsed broken into polygonal blocks. The subsidence of the ring fault increased within the 2° cycle.

7°30' S 10°0' S 🗕 10°50' S 🗕

> Fig. 2: The structures occurring within chaotic terrains (and particularly our case of study Arsinoes - Pyrrhae Chaos) fit very well with our experiment: a depressed basin bordered by a ring fault, polygonal blocks formed by the intersection of radial and concentric faults, and a caldera floor that is disrupted following the polygonal pattern. A) Blended MOLA-HRSC DEM; B) and C) CTX mosaic by (Dickson et al., 2018)





Comparison with chaotic terrains on Mars

The resulting structures formed on the analogue brittle crust resemble strongly the faults disrupting the basaltic bedrock of chaotic terrains on Mars. The analogue magma chamber of the experiment was set based on the characteristics of Arsinoes Chaos (Fig. 2), but most of the features are recurrent also in the other chaotic terrains.

Our experiment demonstrated that this process alone can explain already the enigmatic characteristics of chaotic terrains. Other factors (e.g. criosphere melting, groundwater overpressure) might have contributed causing further distruption and a more complex geological setting.

