

3D Radiative Transfer for ICON-LEM

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Convective Organization Depends on Radiative Transfer Solver

Our results in [1] suggest that solar radiative heating has the potential to organize clouds depending on the sun's incidence angle. Over land, these 3D effects may significantly alter the boundary layer dynamics and cloud development on a timescale of half an hour. Studying these in a more realistic setup is our motivation to implement 3D radiative transfer into ICON-LEM.

1D Radiative Transfer

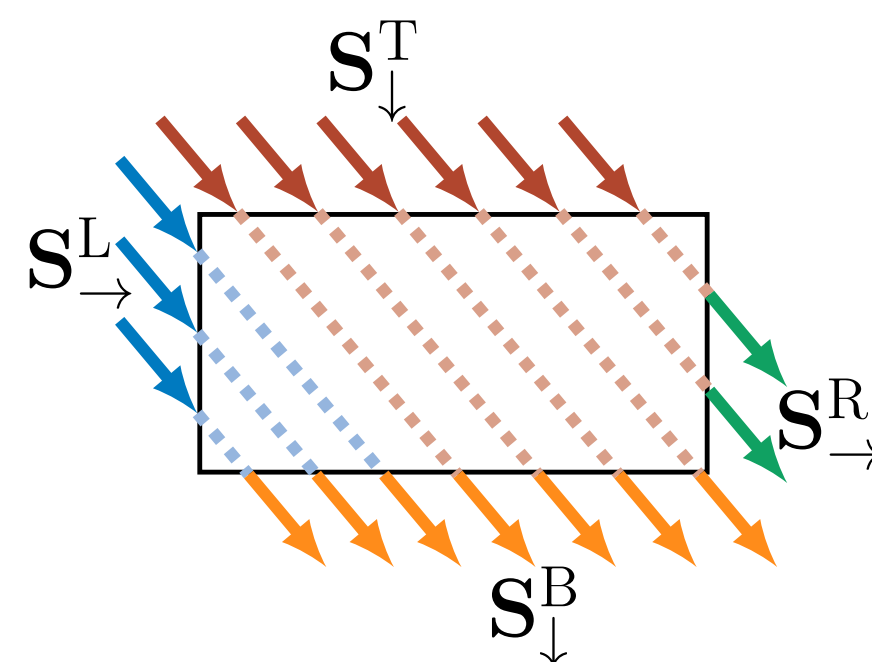


3D Radiative Transfer

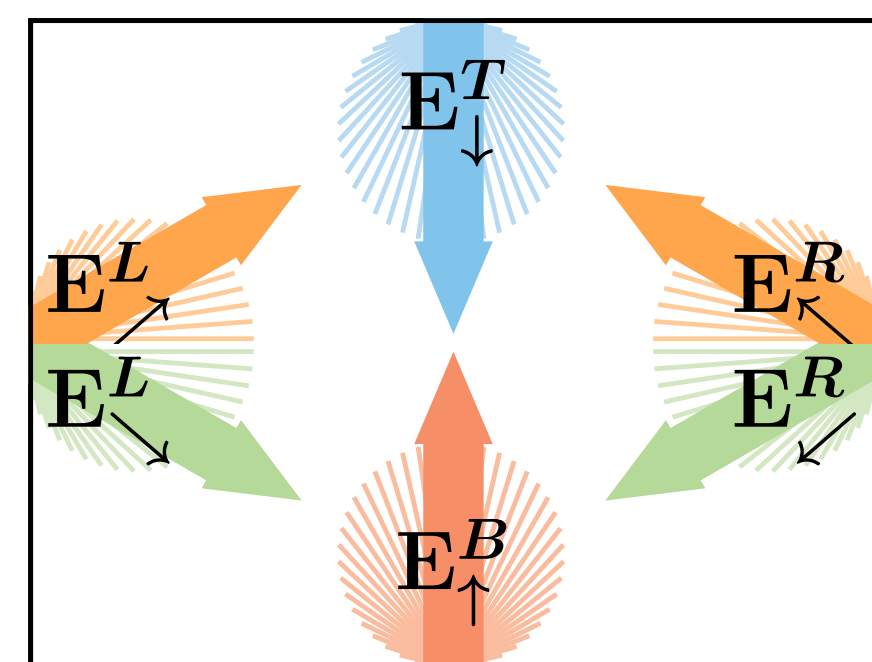


Virtual photographs of LES simulations, as seen from a cruising altitude of 15 km. The cloud simulations either use 3D or 1D Radiative-Transfer calculations and show differences with respect to cloud size distribution and the organization in cloud streets. Both visualizations are performed with MYSTIC (physically correct MonteCarlo renderer in libRadtran [6, 7]).

Concept for a new Solver



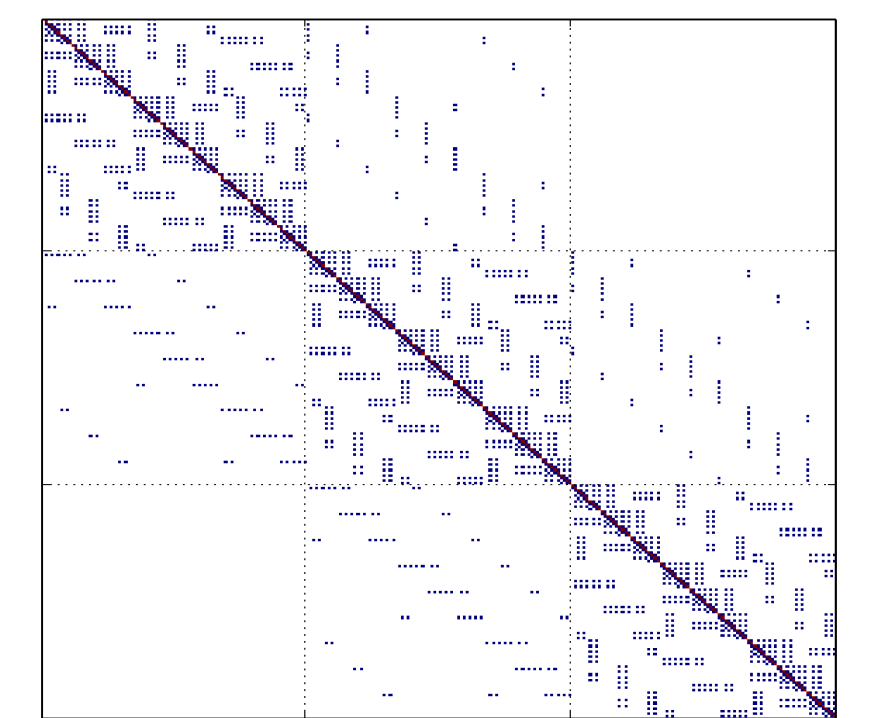
Discretize radiation streams angularly and spatially → at least 5 streams (one for each face) for direct radiation (fixed angle)



Diffuse angular discretization assumes Lambertian surfaces. Two streams (up- and downward), as in 1D solver. Additional streams sideways allow for horizontal energy transfer

$$\begin{bmatrix} E_{\downarrow}^T \\ E_{\downarrow}^B \\ E_{\downarrow}^L \\ E_{\downarrow}^R \\ E_{\uparrow}^L \\ E_{\uparrow}^R \\ S_{\downarrow}^B \\ S_{\downarrow}^L \\ S_{\downarrow}^R \end{bmatrix} = \begin{bmatrix} a_1 a_2 a_3 a_4 a_5 a_6 a_7 a_8 a_9 a_{10} a_{11} \\ a_2 a_1 a_4 a_3 a_5 a_6 a_7 a_8 a_9 a_{10} a_{12} \\ a_5 a_6 a_7 a_8 a_9 a_{10} a_{11} b_{13} \\ a_5 a_6 a_8 a_7 a_{10} a_9 b_{14} b_{13} \\ a_6 a_5 a_9 a_{10} a_7 a_8 b_{15} b_{13} \\ a_6 a_5 a_{10} a_9 a_8 a_7 b_{16} b_{15} \\ 0 & 0 & 0 & 0 & 0 & 0 & c_{10} & c_{11} \\ 0 & 0 & 0 & 0 & 0 & 0 & c_{11} & c_{10} \end{bmatrix} \begin{bmatrix} E_{\uparrow}^T \\ E_{\uparrow}^B \\ E_{\uparrow}^L \\ E_{\uparrow}^R \\ E_{\downarrow}^L \\ E_{\downarrow}^R \\ S_{\uparrow}^B \\ S_{\uparrow}^L \\ S_{\uparrow}^R \end{bmatrix}$$

Use MonteCarlo to compute transport coefficients for single boxes and couple to neighbors → huge, but sparse matrix

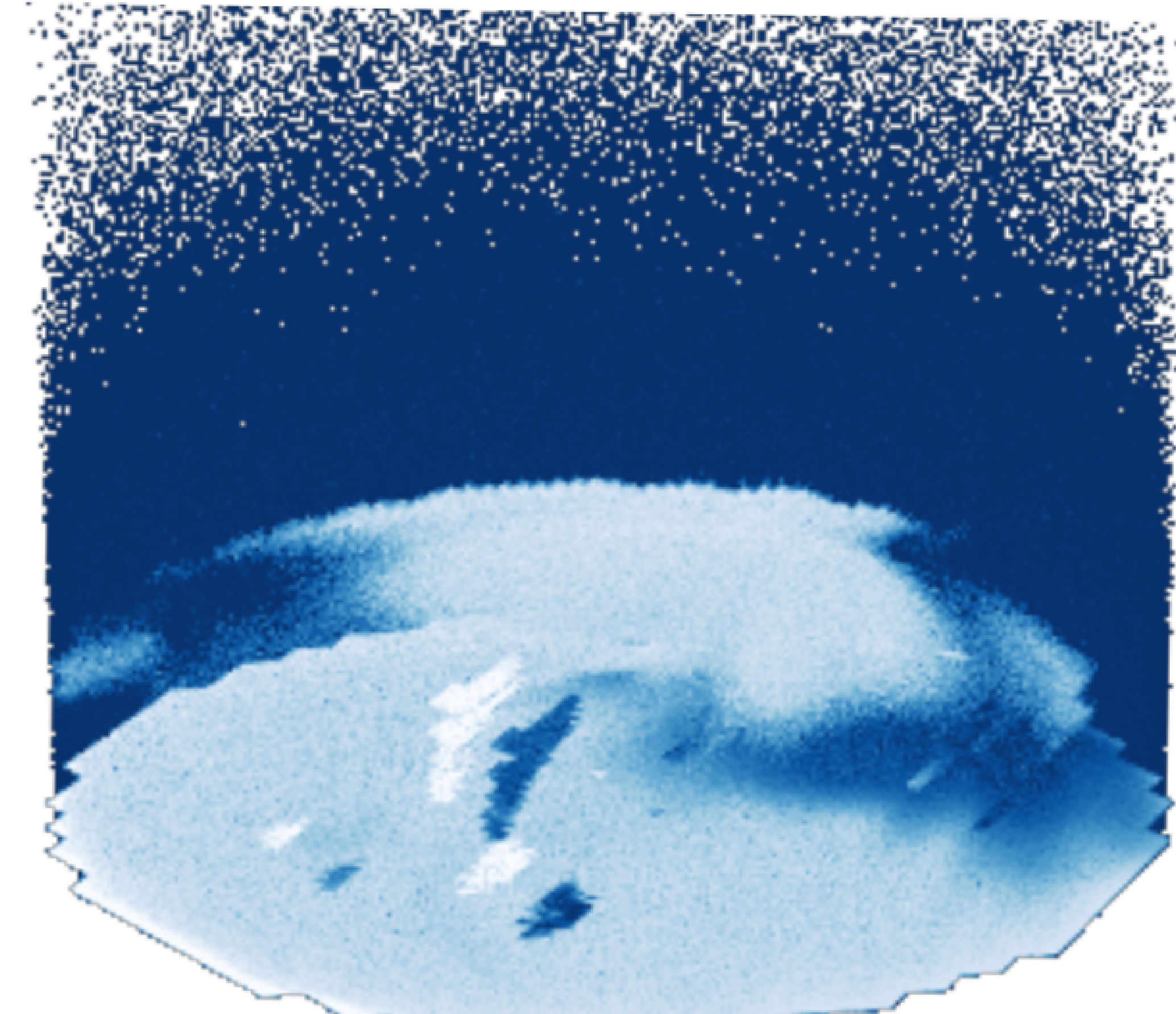


Solve linear equation system with parallel iterative methods (PETSc):
ASM + SOR or GMRES + Algebraic Multigrid (GAMG)

RayLi

The Ray based Light Simulator (RAYLI) is a 3D MonteCarlo ray-tracing library which can be used on the native ICON-Mesh. The concept of a MonteCarlo solver for radiative transfer is the idea that we trace photon paths through the atmosphere and statistically solve the integral in the radiative transfer equation. The beauty of MonteCarlo raytracing is the flexibility and accuracy. We can consider complex atmospheric scenes without the need for any simplifications.

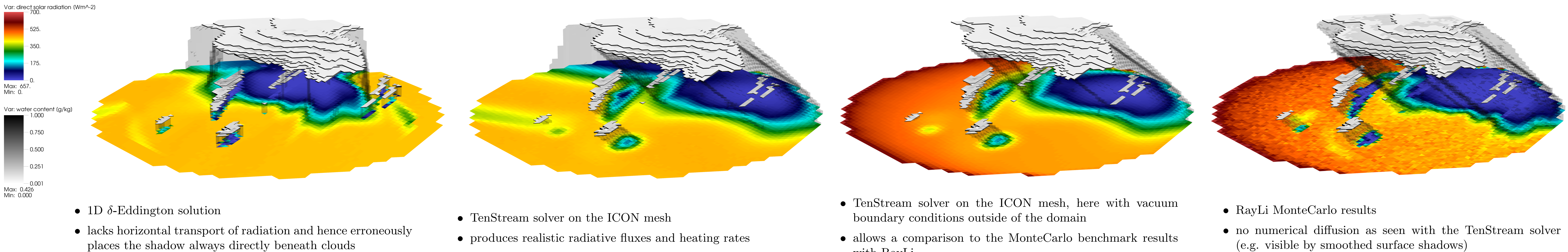
However, MonteCarlo methods are computationally too demanding to be run in NWP or LES models directly. Yet, they are superb candidates to benchmark radiative transfer parameterizations. One particular feature of the RayLi library is the abstraction of the underlying mesh which allows a straightforward coupling with the TenStream library and the ICON-LEM grid. In the following, we will use RayLi as a benchmark model to assess the performance of the TenStream solver on unstructured meshes.



- virtual photograph of an ICON-LEM scene, rendered with RayLi within the TenStream

3D Effects of Direct Solar Radiation in ICON-LEM and Validation against "RayLi"

This panel depicts results for direct solar radiation for several radiative transfer solvers in an ICON-LEM simulation with a horizontal resolution (edge length) of 1.2 km (solar zenith angle is 61°, surface albedo is constant).



References

- [1] F. Jakub and B. Mayer. "The Role of 1D and 3D Radiative Heating on the Organization of Shallow Cumulus Convection and the Formation of Cloud Streets" ACP (2017).
- [2] C. Klinger et. al. "Effects of 3D Thermal Radiation on Cloud Development" ACP (2017).
- [3] F. Jakub, "On the Impact of Three Dimensional Radiative Transfer on Cloud Evolution" LMU (2016).
- [4] F. Jakub and B. Mayer. "A Three-Dimensional Parallel Radiative Transfer Model for Atmospheric Heating Rates for use in Cloud Resolving Models-the TenStream solver." JQSRT (2015).
- [5] F. Jakub and B. Mayer. "3D Radiative Transfer in Large-Eddy Simulations - Experiences coupling the TenStream solver to the UCLA-LES" GMD (2016).
- [6] C. Emde et. al. "The libRadtran Software Package for Radiative Transfer Calculations" GMD (2015).
- [7] B. Mayer. "Radiative transfer in the cloudy atmosphere." EPJ (2009).