

specMACS Observations during EUREC⁴A

Cloud Droplet Size Distributions (from Observations of Glory and Cloudbow) and Cloudmask

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OVERVIEW

How is specMACS designed?

- downward measurements; mounted onboard research aircraft HALO
- two hyperspectral line cameras (400 nm 2500 nm) and two 2D polarization cameras

Which products will be derived from the measurements?

- Identify clouds in the observations (cloudmask) slide 2
 - · cloud detection algorithm based on brightness threshold and water vapor (WV) absorption in the shortwave infrared (SWIR) solar spectrum \rightarrow works also if sunglint present
- Retrieve microphysical parameters sides 3-3
 (polarized) observations of the backscatter glory and the cloudbow depend on effective radius $r_{\rm eff}$ and width σ of cloud droplet size distribution (CDSD)

Which radiative transfer models are used?

- cloudmask: DISORT (included in the libRadtran package)
- CDSD: MYSTIC (included in the libRadtran package,
 - [Mayer and Kylling, 2005, Emde et al., 2016])

CLOUD MASK [GÖDDE, 2018]

Challenges

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- common cloud detection approaches use contrast of bright clouds in front of dark background
- approach fails in the presence of sunglint over ocean (sunglint increases brightness of ocean)

Solution: Use absorption by water vapor!

- absorption by WV in the shortwave infrared solar spectrum (SWIR)
- ► no cloud → SWIR radiation reflected from ocean's surface → higher WV absorption than radiation reflected at cloud (due to longer path through the atmosphere and higher WV concentrations near surface)

cloud mask 2020-01-31

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CLOUD MICROPHYSICS

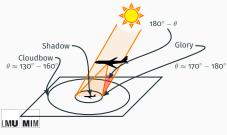
Challenges

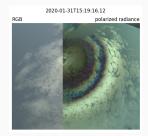
traditional bi-spectral approaches (e.g., Nakajima-King technique)...

- ... retrieve cloud optical thickness and cloud effective radius
- ... suffer from 3D radiative transfer effects (e.g., shadows)
- … cannot retrieve width of droplet size distribution

Solution: Fit MYSTIC simulations to observations of glory and cloudbow!

Formation: Single scattering by liquid cloud droplets (\rightarrow not affected by 3D effects); scattering angle: θ





visibility of glory and cloudbow in polarized measurements (right) enhanced due to removal of "Auti multiple-scattering background

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RETRIEVAL

Dependence on cloud droplet size distribution?

- angular radius of glory decreases with effective radius while the radius of the cloudbow increases
- narrow size distribution enhances contrast of the rings (not shown)

cross-section of a backscatter glory (SZA = 10°) effective radius 0.30 backscatte direction Reflectivity optical thickness width of size distribution first rings $\tau = 5$, $r_{eff} = 10 \mu m$, $\sigma = 0.98 \mu m$ 0.00 Viewing zenith angle [degree] ∆z (from Kölling, 2020

Aggregation of observations

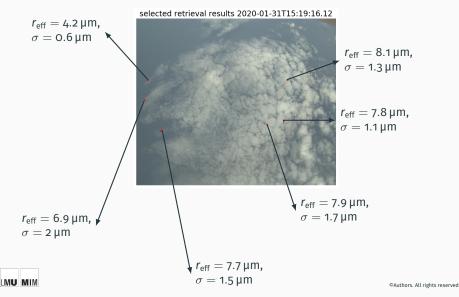
- ► observation of same cloud element in successive images → observation from different angles
- necessary: cloud top height (from [Kölling, 2020])

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Results

pre-calculated MYSTIC simulations are fitted to the aggregated observations [Mayer et al., 2004] \rightarrow retrieve $r_{\rm eff}$, σ with a spatial resolution of about 20 m



Thank you and stay healthy!



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