

# ***Towards spaceborne monitoring of localized CO<sub>2</sub> emissions: an instrument concept and first performance assessment***

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A large, high-resolution image of the Earth's surface as seen from space, showing the curvature of the planet, blue oceans, white clouds, and green landmasses. The text "Knowledge for Tomorrow" is overlaid on the right side of this image.

Knowledge for Tomorrow

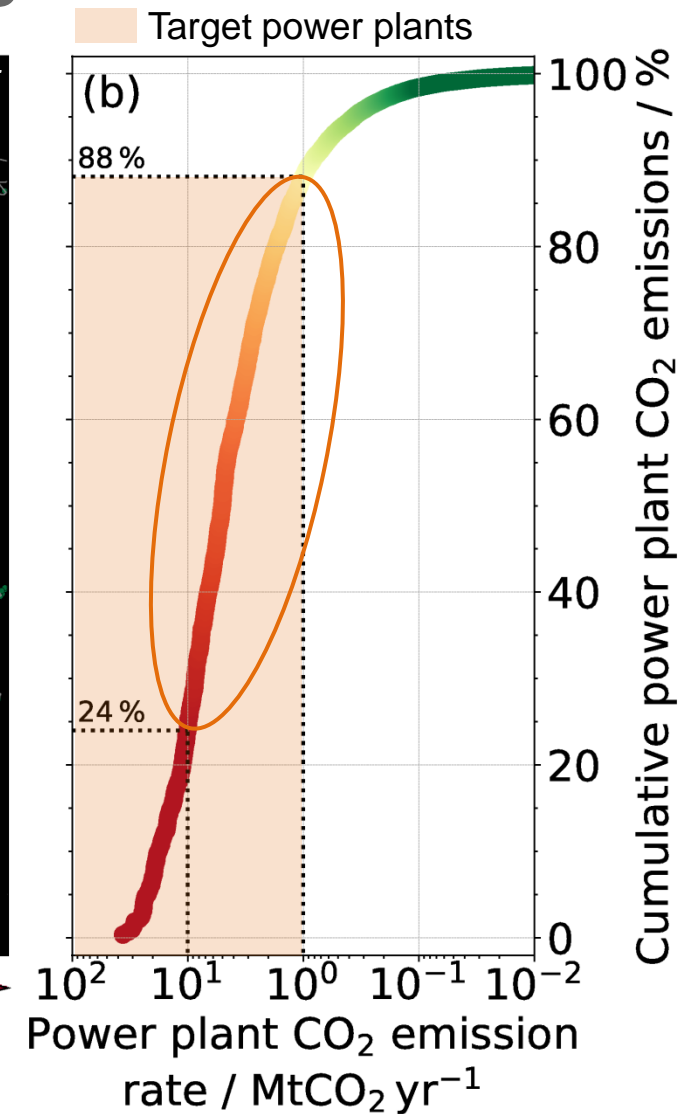
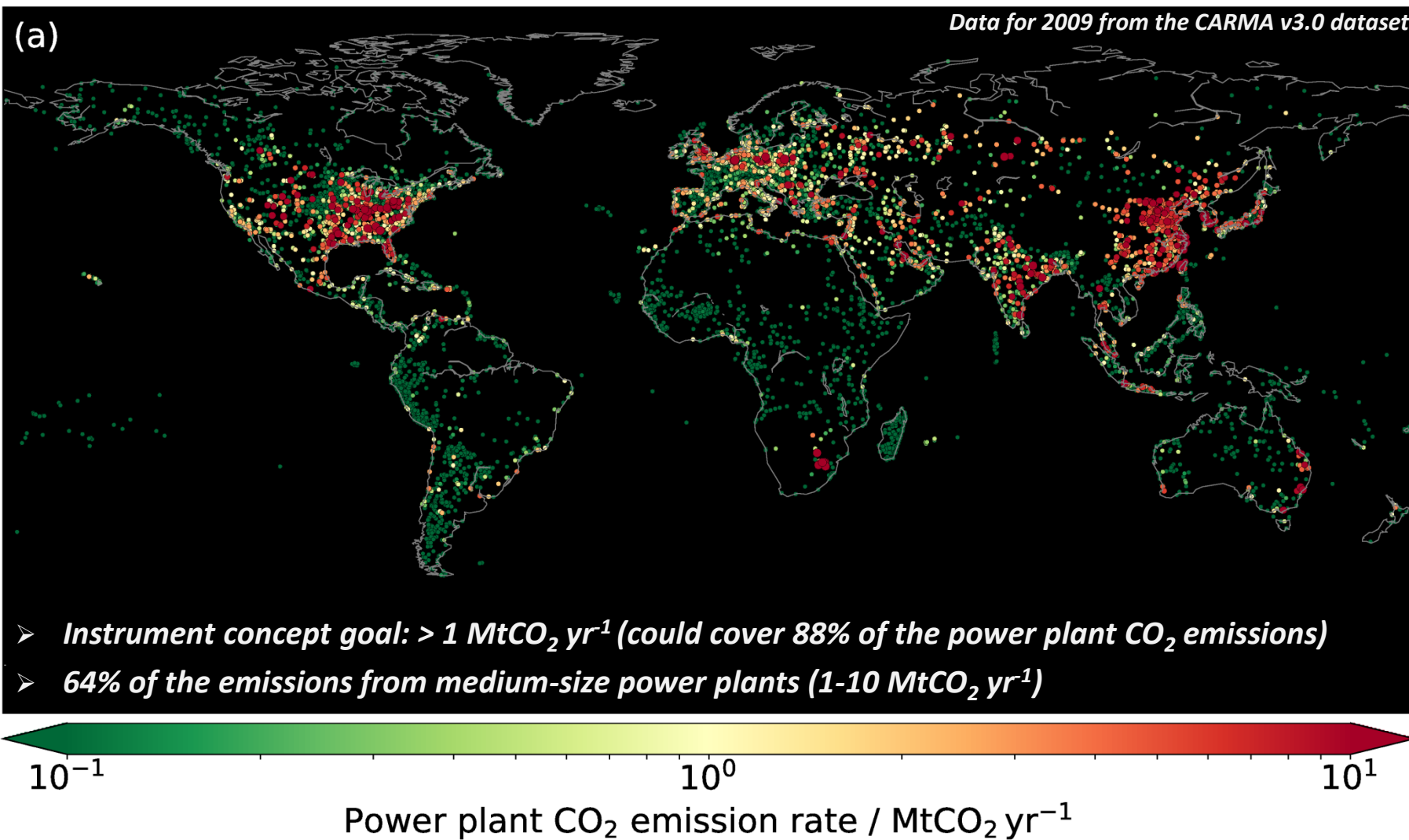
# *Why monitor localized CO<sub>2</sub> emissions?*

## *How would the proposed instrument concept contribute?*

- *UN Framework Convention on Climate Change* requires every country to report their CO<sub>2</sub> emissions
- The *Paris climate agreement* requires *independent verification* of the reported emissions
- *Power plants* are the main source and represent *approx. one third* of all anthropogenic CO<sub>2</sub> emissions
- We propose the concept of an imaging spectrometer *optimized for sub-plume resolution (50×50 m<sup>2</sup>)* to also resolve emissions from *medium-size power plants* (1-10 MtCO<sub>2</sub> yr<sup>1</sup>), currently not targeted by other satellite missions
- Such an instrument would be a *valuable companion and complement* to the fleet of current and planned satellite missions measuring atmospheric CO<sub>2</sub> column concentrations

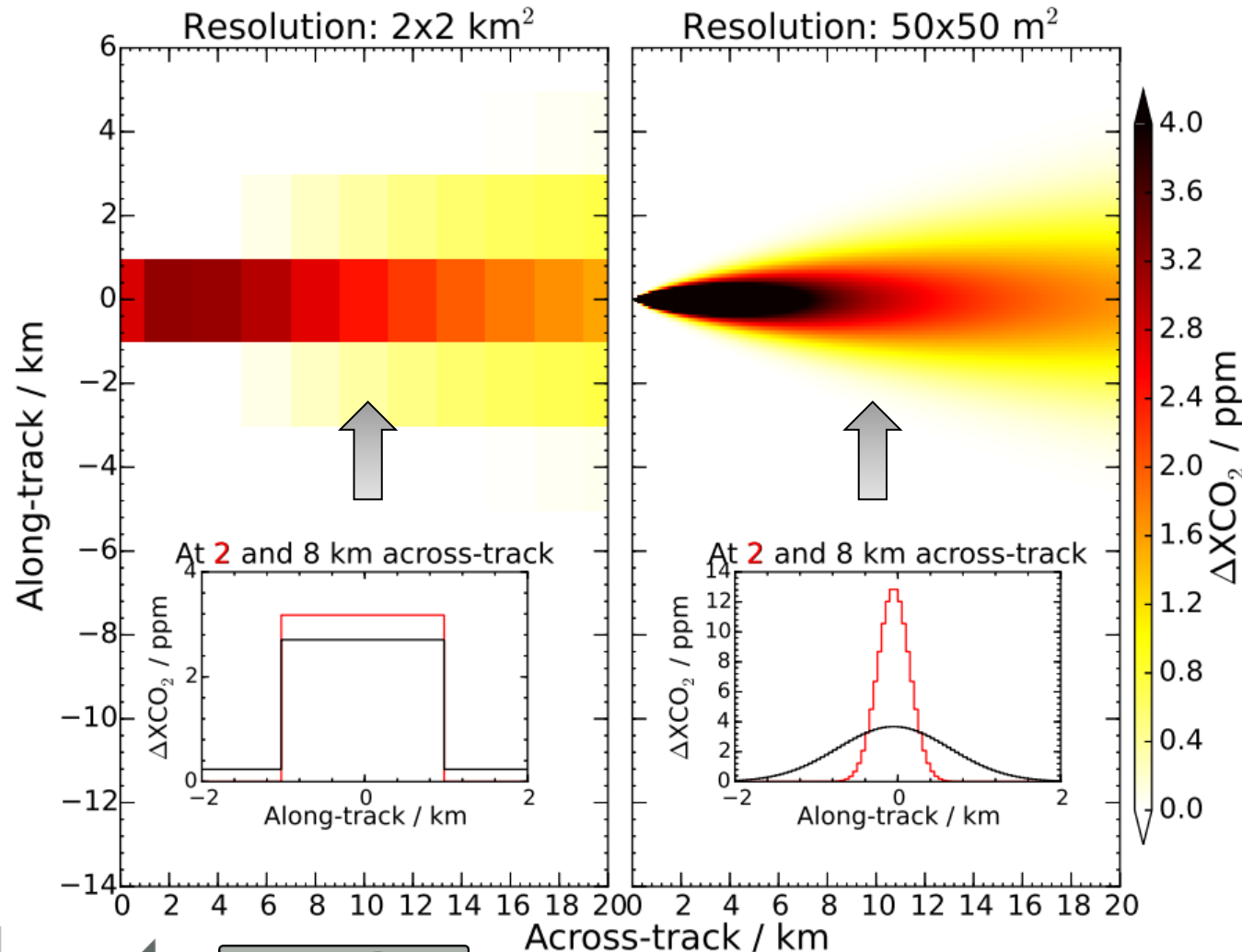


# Global distribution of power plant CO<sub>2</sub> emissions



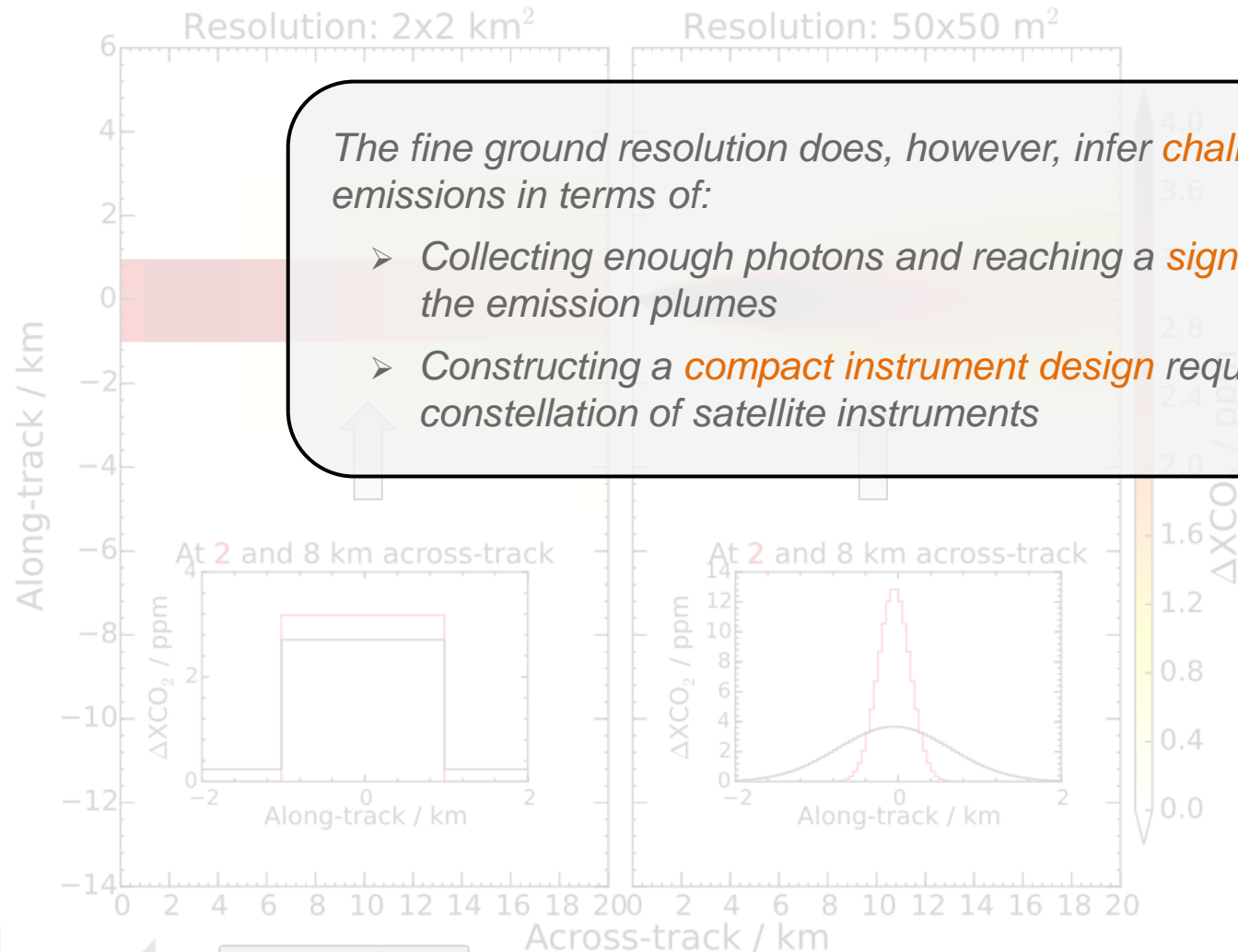


# Fine ground resolution key to measure localized CO<sub>2</sub> emissions



- Power plant emissions cause **large column concentration enhancements in the vicinity of the source**
- But, these enhancements become small when averaging with the background in km-scale ground pixels
- Point source detection and quantification of medium-size power plant emissions need **fine ground resolution**

# Fine ground resolution key to measure localized CO<sub>2</sub> emissions



The fine ground resolution does, however, infer **challenges** for the monitoring of localized CO<sub>2</sub> emissions in terms of:

- Collecting enough photons and reaching a **signal-to-noise ratio** (SNR) sufficient to resolve the emission plumes
- Constructing a **compact instrument design** required for global monitoring through a constellation of satellite instruments

- Power plants emissions cause **large** enhancements in the vicinity of the source
- But, these enhancements become small when averaging with the background in km-scale ground pixels
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# ***Proof of concept and spectral sizing using spectrally degraded GOSAT measurements***

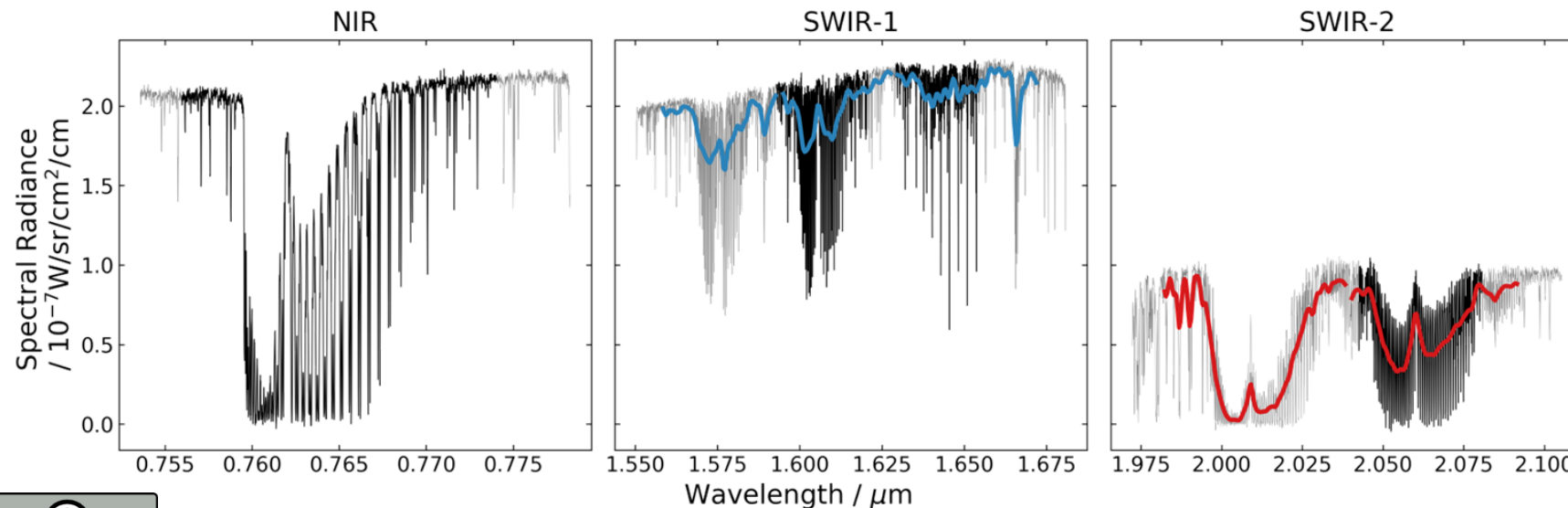


Knowledge for Tomorrow



# Spectral sizing of proposed instrument concept

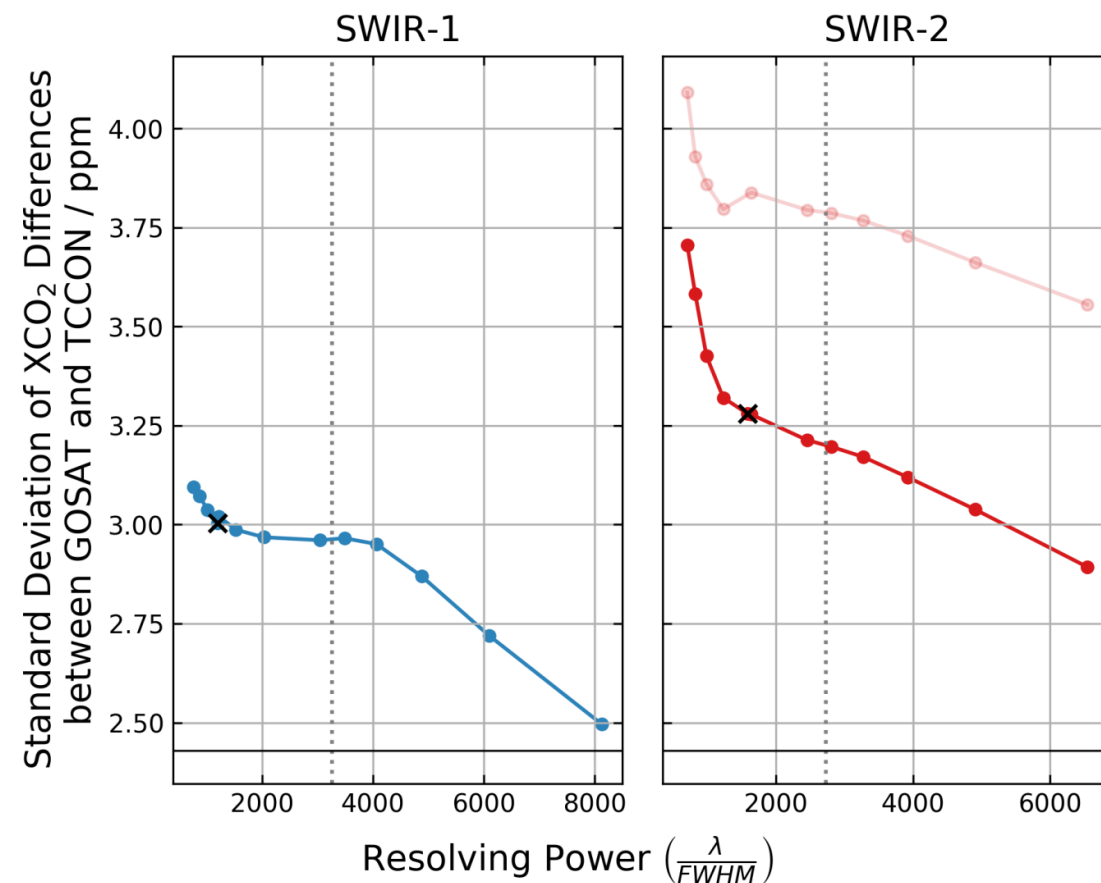
- For a compact instrument design **a single spectral window** shall be used
  - Two alternative spectral windows, SWIR-1 and SWIR-2, are investigated
- Measured GOSAT spectra (grey thin lines) are **convolved to mimic spectra of proposed instrument concept** (**blue** and **red** bold lines)
- XCO<sub>2</sub> is retrieved using the RemoTeC algorithm and the convolved spectra
- **Black thin lines** show the spectra used for native GOSAT XCO<sub>2</sub> retrievals with RemoTeC





# Spectral sizing of proposed instrument concept

- Increasing level of spectral degradation is performed in order to **determine the appropriate spectral resolution** for the proposed instrument
- XCO<sub>2</sub> retrieved from spectrally degraded GOSAT spectra is compared to reference data from TCCON<sup>1</sup>
- Modifications of the light path due to **scattering aerosol are accounted for in retrievals from SWIR-2 spectra**. (Light red line shows results for SWIR-2 retrievals when scattering is neglected)
- Due to low sensitivity to aerosol, **retrievals from SWIR-1 spectra neglect scattering**
- Black crosses mark the **spectral resolution chosen for further analysis**

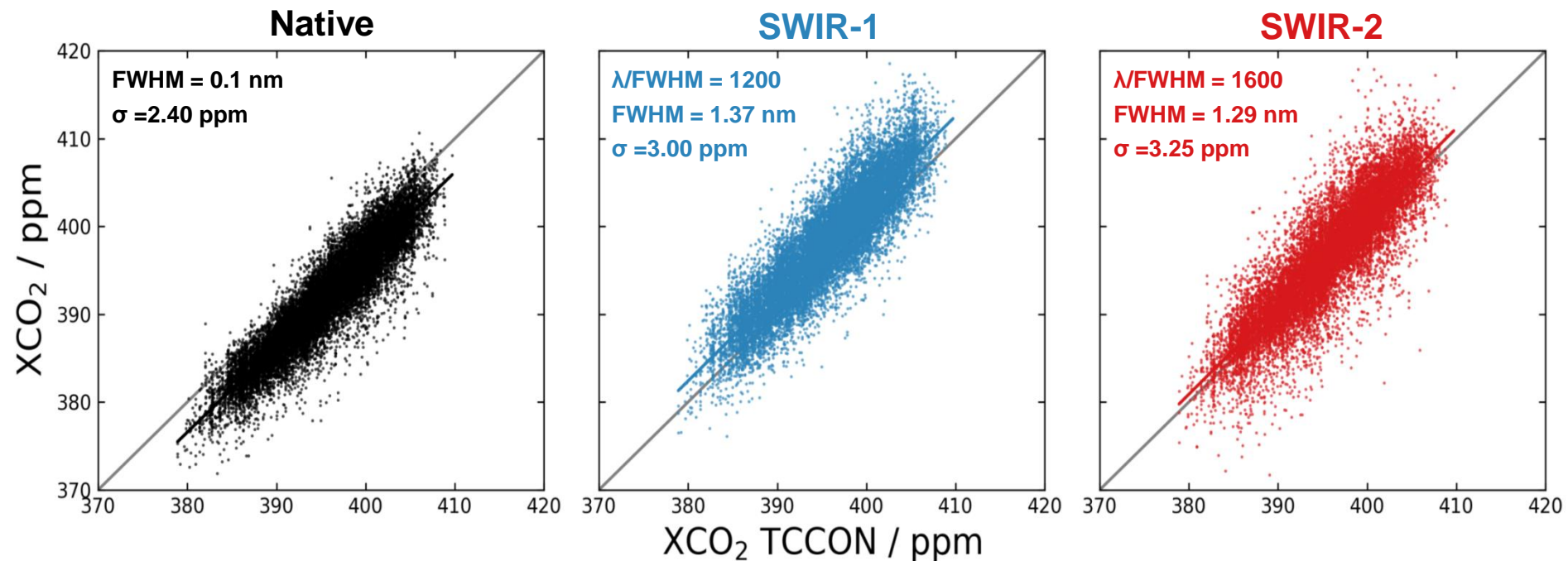


<sup>1</sup>Total Carbon Column Observing Network



# Validation with TCCON measurements

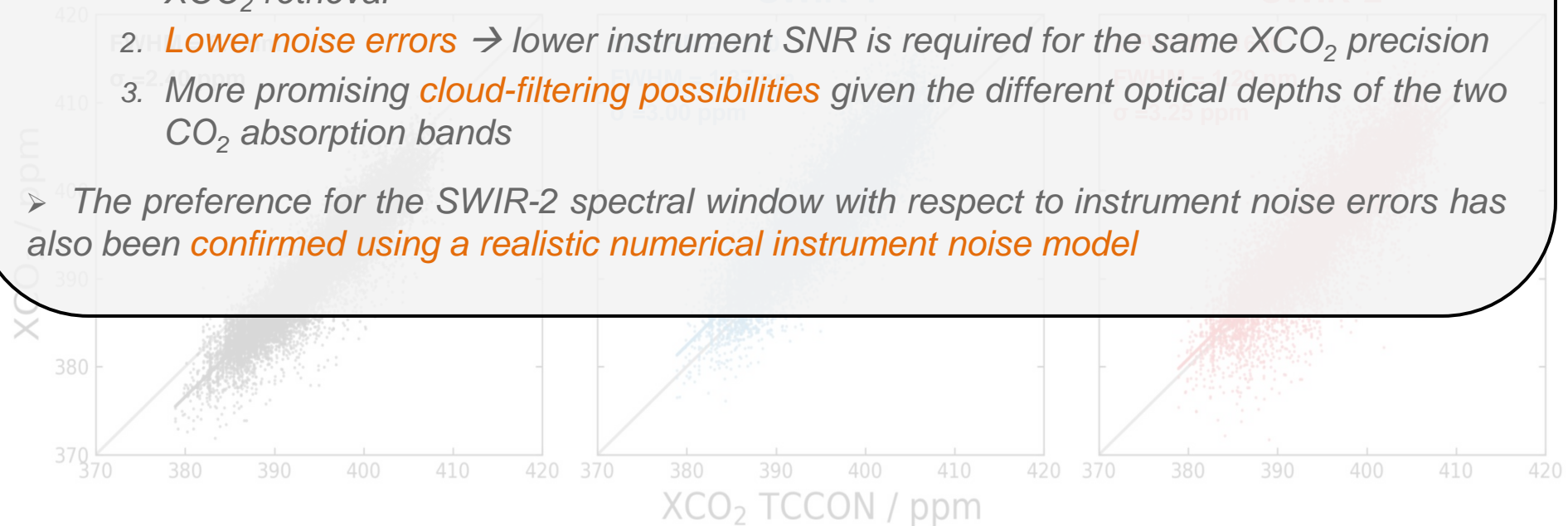
- Precision of  $XCO_2$  retrieved from spectrally degraded GOSAT spectra below 1% for both SWIR-1 (3.00 ppm) and SWIR-2 (3.25 ppm) with respect to reference TCCON measurements
- Precision decreases only moderately compared to native GOSAT retrievals (2.40 ppm)



# Validation with TCCON measurements

## Conclusion

- Analysis using spectrally degraded GOSAT measurements shows **similar precision for the two alternative spectral windows** SWIR-1 and SWIR-2
- However, the **SWIR-2 spectral window is preferred** for the proposed instrument concept, mainly due to three reasons:
  1. Higher sensitivity to **atmospheric aerosol** → scattering can be better accounted for in XCO<sub>2</sub> retrieval
  2. **Lower noise errors** → lower instrument SNR is required for the same XCO<sub>2</sub> precision
  3. More promising **cloud-filtering possibilities** given the different optical depths of the two CO<sub>2</sub> absorption bands
- The preference for the SWIR-2 spectral window with respect to instrument noise errors has also been **confirmed using a realistic numerical instrument noise model**



# *Assessing the proposed instrument's CO<sub>2</sub> monitoring performance through simulations*



# Step 1: Simulating synthetic measurements

## Preliminary instrument design

Orbit, spectral sizing, optical design  
and detector properties → SNR

Orbit	600 km, sun-synchronous
Mass / kg	90
Swath / km	50
Spatial resolution / m <sup>2</sup>	50 × 50
Spectral range / nm	1982–2092
FWHM (2.5 pix) / nm	1.29
Resolving power / -	1600
Aperture diameter / cm	15.0
f-number ( $f_{\text{num}}$ ) / -	2.4
Optical efficiency ( $\eta$ ) / -	0.48
Integration time ( $t_{\text{int}}$ ) / ms	70
Detector pixel area ( $A_{\text{det}}$ ) / $\mu\text{m}^2$	900
Quantum efficiency ( $Q_e$ ) / e <sup>-</sup> photon <sup>-1</sup>	0.8
Dark current ( $I_{\text{dc}}$ ) / fA pix <sup>-1</sup> s <sup>-1</sup>	1.6
Readout-noise / e <sup>-</sup>	100
Quantization noise / e <sup>-</sup>	40



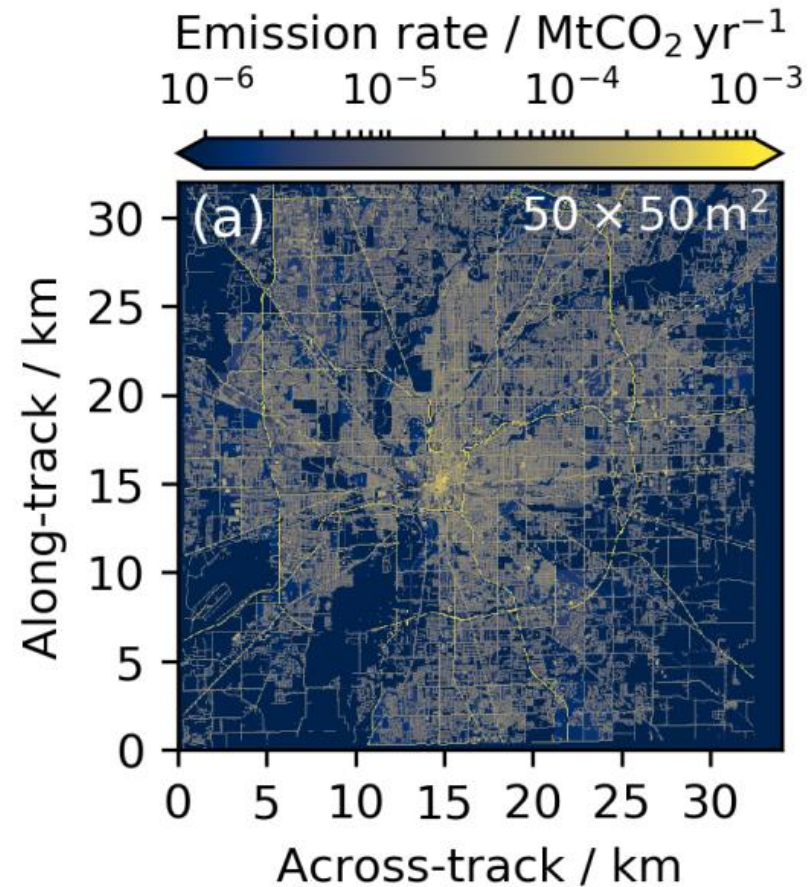
# Step 1: Simulating synthetic measurements

## Preliminary instrument design

Orbit, spectral sizing, optical design  
and detector properties → SNR

## Realistic urban emission scenario

Hestia data for Indianapolis



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## Preliminary instrument design

Orbit, spectral sizing, optical design  
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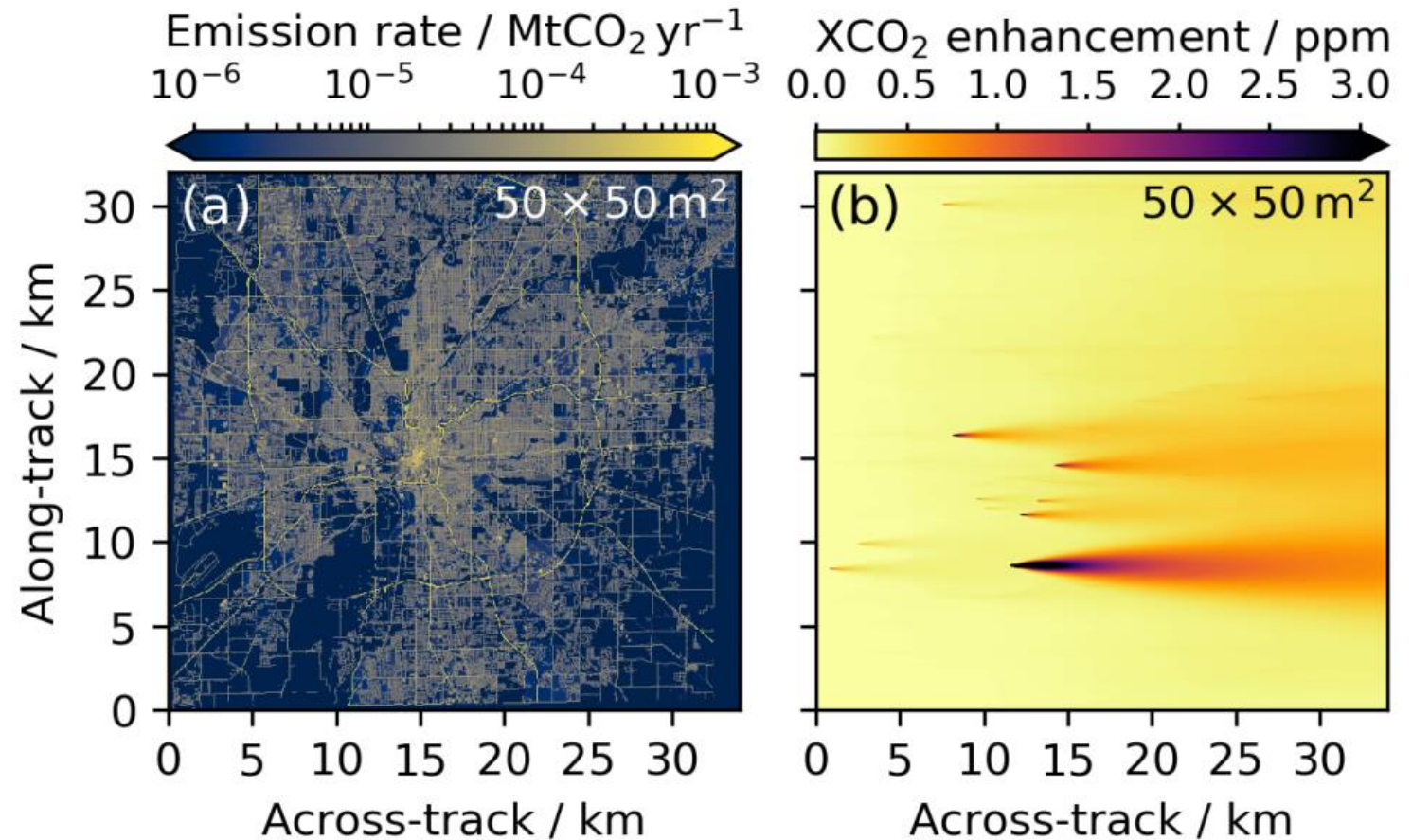
## Realistic urban emission scenario

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## 3D CO<sub>2</sub> concentration field

Gaussian plume model



# Step 1: Simulating synthetic measurements

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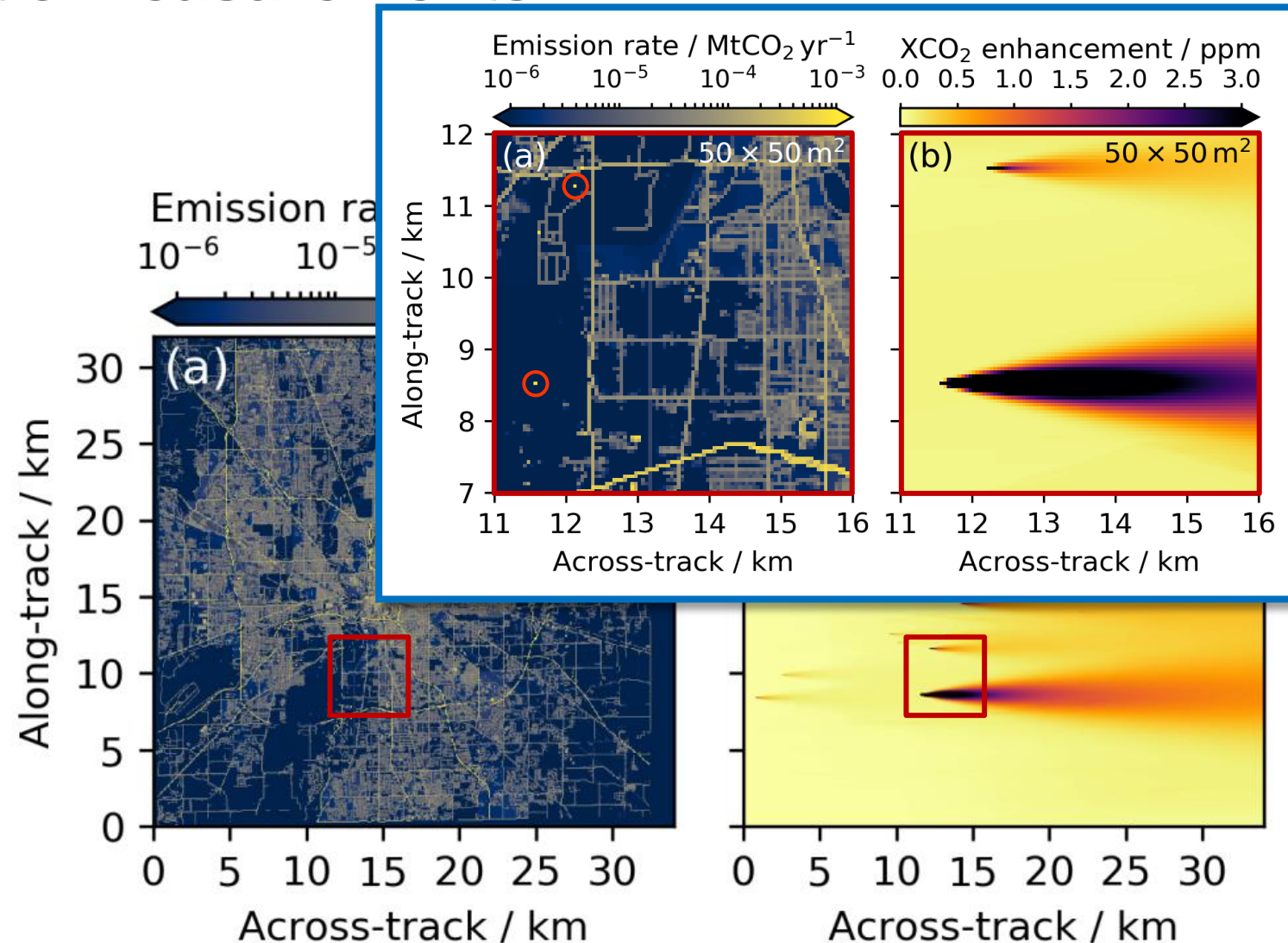
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# *Step 1: Simulating synthetic measurements*

## **Preliminary instrument design**

Orbit, spectral sizing, optical design  
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## **3D CO<sub>2</sub> concentration field**

Gaussian plume model

## **Background meteorological data**

CarbonTracker model





# Step 1: Simulating synthetic measurements

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## 3D CO<sub>2</sub> concentration field

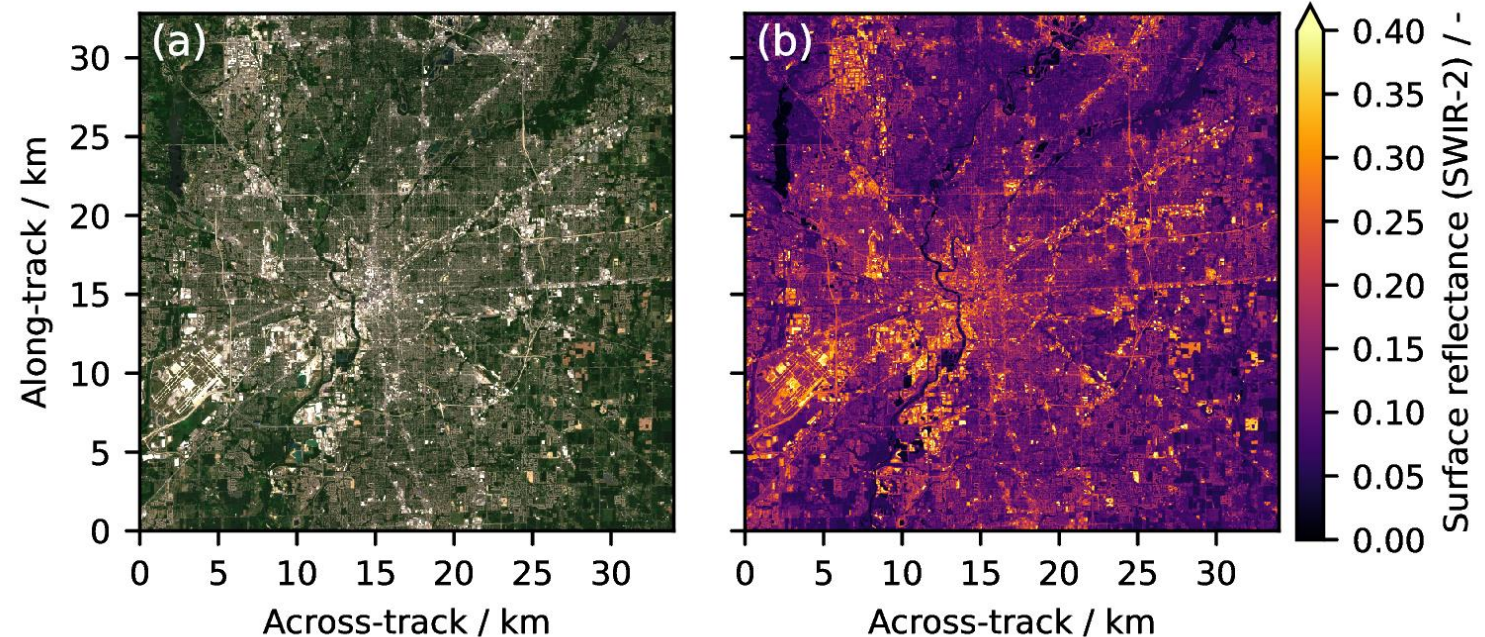
Gaussian plume model

## Background meteorological data

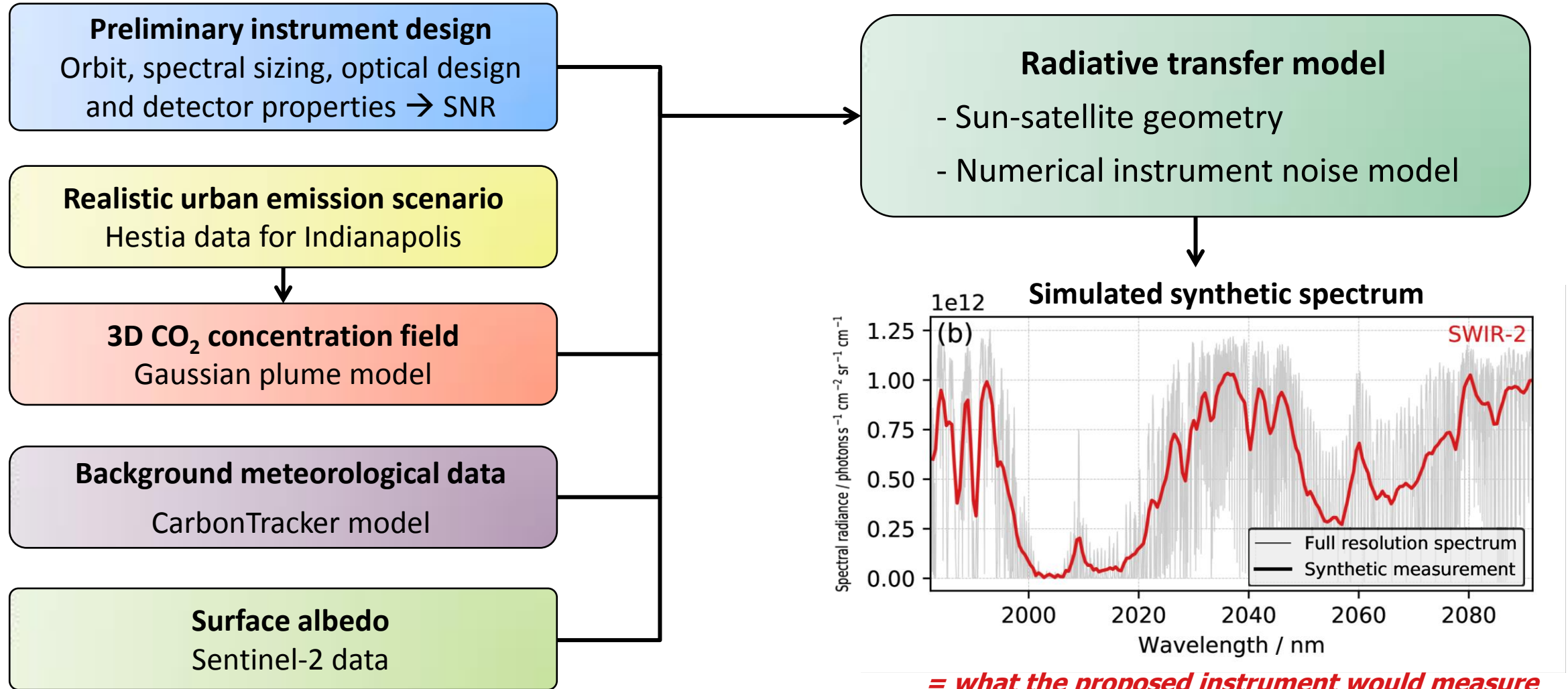
CarbonTracker model

## Surface albedo

Sentinel-2 data



# Step 1: Simulating synthetic measurements

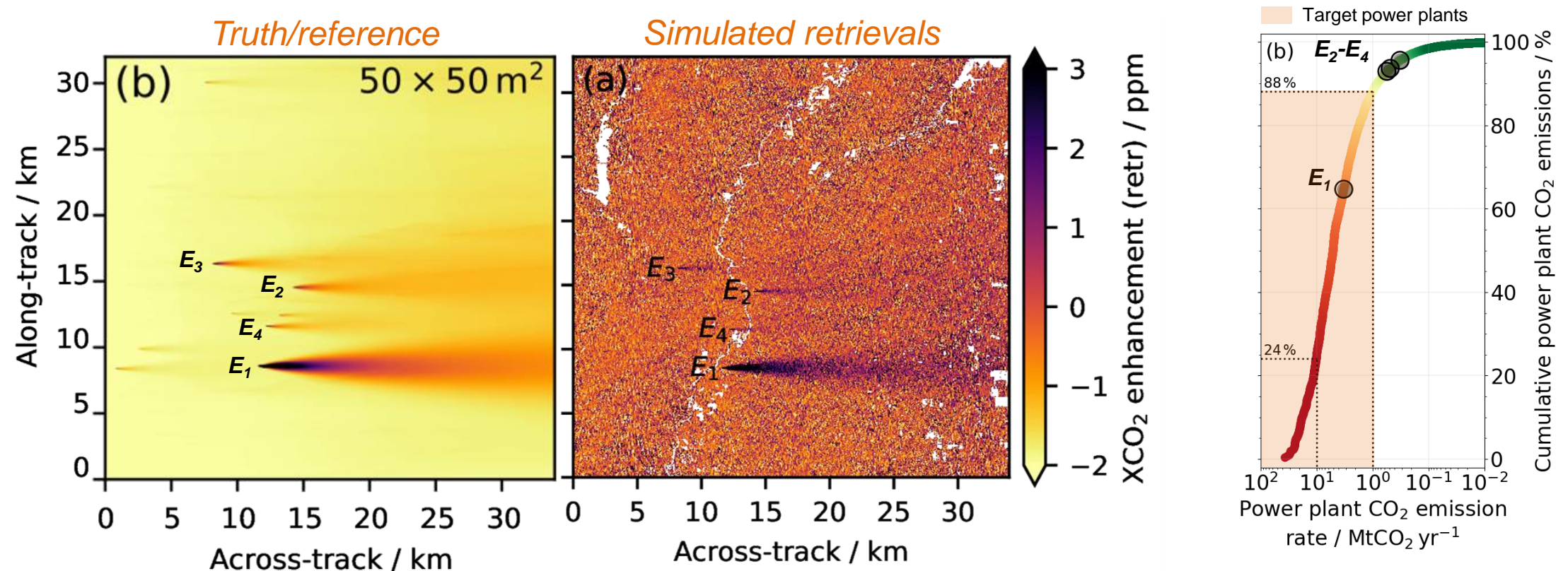


**= what the proposed instrument would measure  
for the given scenario and instrument design**



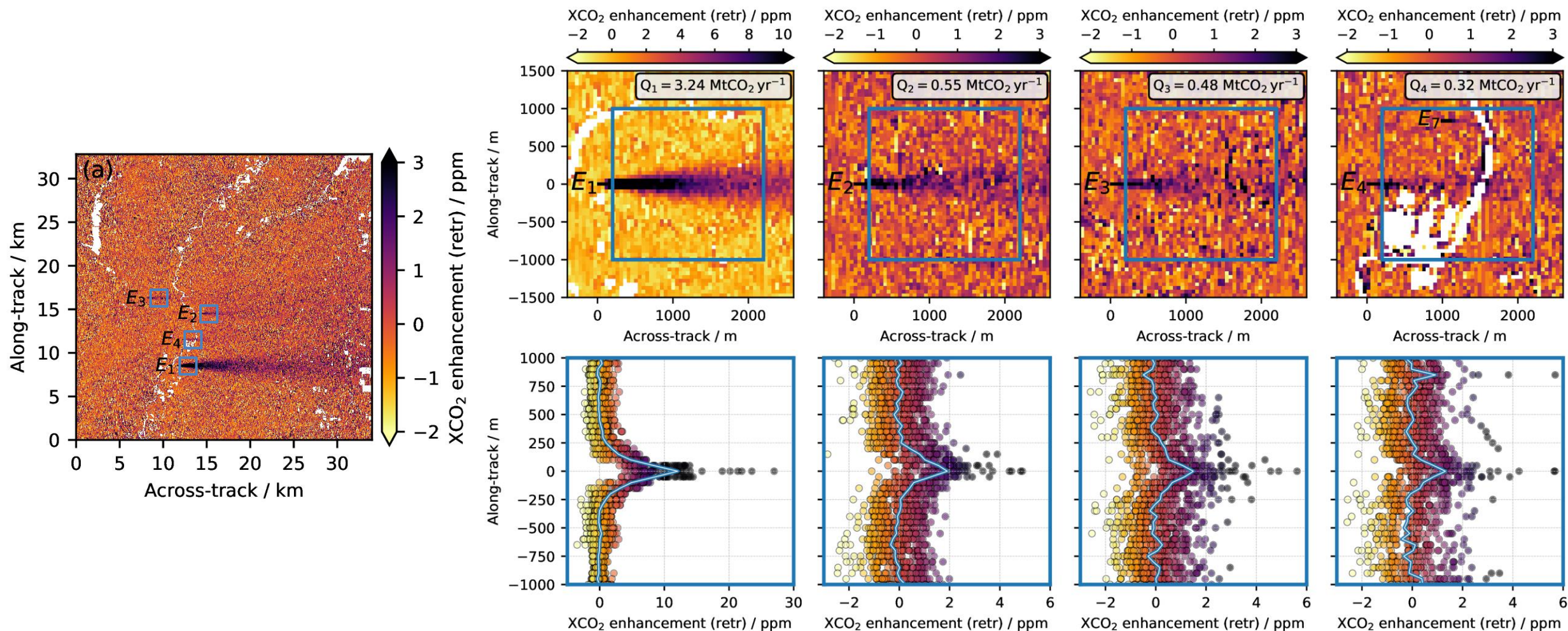
## Step 2: Retrieve corresponding $XCO_2$ from the synthetic measurements

- During the retrieval, a set of *variables in the state vector* ( $XCO_2$ ,  $H_2O$ , albedo, solar/spectral shift) *are fitted*
- The ability to find the true parameters is *limited by the instrument's noise level* (SNR)
- Hence, we can *evaluate the instrument's  $CO_2$  monitoring capabilities* by looking at the retrieved (noisy)  $XCO_2$  field





# Step 2: Retrieve corresponding $\text{XCO}_2$ from the synthetic measurements



**Upper panels** Retrieved two-dimensional fields of  $\text{XCO}_2$  enhancements in the vicinity of the four strongest  $\text{CO}_2$  emitters  $E_1$ ,  $E_2$ ,  $E_3$  and  $E_4$  within the Hestia Indianapolis dataset. **Lower panels** Corresponding per-pixel (circles) and average (solid lines) along-track  $\text{XCO}_2$  enhancements within the area 200 to 2200 m downwind and -1000 to 1000 across-wind of the respective emitters. The blue rectangles in the upper panels show the areas from which the corresponding per-pixel and average along-track  $\text{XCO}_2$  enhancements, depicted in the respective lower panels, are extracted and calculated. The color of the circles follow the color bars in the respective upper panels.



# Summary and conclusions

- Systems for the independent verification of reported CO<sub>2</sub> emissions are needed
- Emissions from **medium-size power plants** (1-10 MtCO<sub>2</sub> yr<sup>1</sup>) are responsible for a significant part of anthropogenic CO<sub>2</sub> emissions in general and from the power plant sector in particular
- In order to independently monitor such emissions from space **we propose an imaging spectrometer with a fine ground resolution of 50 x 50 m<sup>2</sup>**
- **Measurements near 2.0 μm** (SWIR-2 window) are most promising for this task
- Simulations using a **realistic instrument design, emission scenario and surface albedo** show that when the instrument is only limited by its own noise, plumes from emitters with a source strength **down to 0.3 MtCO<sub>2</sub> yr<sup>1</sup>** can be resolved
  - **Significant margin for additional error sources** (e.g. aerosols, meteorology)
- The **compact instrument design**, with a single spectral window, would allow for a **constellation of satellites**, hence increasing the spatial coverage and temporal resolution



# Summary and conclusions

➤ Systems for the independent verification of reported CO<sub>2</sub> emissions are needed

➤ Emissions from medium size power plants (1-10 MtCO<sub>2</sub> yr<sup>-1</sup>) are responsible for a significant part of anthropogenic CO<sub>2</sub> emissions in general and from the power plant sector in particular

➤ In order to independently verify CO<sub>2</sub> emissions, a satellite instrument with a fine ground resolution of 50 x 50 m<sup>2</sup>

➤ Measurements near 2µm (SWIR) are the optimal for this task

➤ Simulate CO<sub>2</sub> plume dispersion with LES, rather than Gaussian, modeling

➤ Include scattering effects by atmospheric aerosol in the radiative transfer simulations also at local/urban scale

➤ Quantify the ability to inversely determine the corresponding CO<sub>2</sub> emission rates under various conditions representing for example different emission source strengths, seasons, surface albedo, meteorological conditions etc.

➤ Simulations using a realistic instrument design, emission scenario and surface albedo show that when the instrument is only limited by its own noise, plumes from emitters with a source strength down to 0.3 MtCO<sub>2</sub> yr<sup>-1</sup> can be resolved

→ Significant margin for additional error sources (e.g. aerosols, meteorology)

➤ The compact instrument design, with a single spectral window, would allow for a constellation of satellites, hence increasing the spatial coverage and temporal resolution



# ***Thank you for you interest!***

*Please feel free to contact me for any questions*

*Email: anytime*

*EGU live chat: 2020/05/06, 08:30-10:15 CEST*

[Johan.Strandgren@dlr.de](mailto:Johan.Strandgren@dlr.de)



*Further details and references can also be found in the following publications:*

[Wilzewski et al., AMT, 2020](#)

[Strandgren et al., AMTD, 2020](#)

