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InSight JPL

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- The InSight spacecraft is just now finishing its 510th sol on Mars and is operating almost flawlessly.
- All the instruments are operating 24.6/7, as well as or better than designed, with the exception of the HP<sup>3</sup> mole.
  - -SEIS is measuring motions ~10x better than its design requirement.
  - -The mole is mechanically sound, but unexpected soil conditions have thus far prevented it from reaching its planned depth.
- All data is being released through the PDS within 3-6 months of acquisition. Raw images are released within minutes of receipt on the ground.
  - Currently, roughly 60% of this data and >90% of our commands are being relayed through Odyssey.



- The HP<sup>3</sup> was deployed to the surface in mid-February 2019 and immediately began penetration.
- A depth of 35 cm was reached relatively rapidly (within a few hundred strokes); repeated subsequent hammering (~9000 strokes) resulted in no measurable further progress.
- We have subsequently concluded that the mole has lost sufficient hull friction to maintain downward progress due to unexpected soil conditions.
  - Loosely cemented, porous duricrust >15 cm thick
- We developed and tested a recovery plan to using the robotic arm, first to increase the hull friction, and later to provide downward force to the back of the mole.



## First Three Attempts to "Pin and Hammer"





# One Step Forward, Three Steps Back...





Getting into position on Mars – sol 427



#### Most recent hammering session – sol 489

### The Challenge of Operating a Seismometer on Mars

- Need extremely high sensitivity expected (and found!) fewer and smaller quakes than on the Earth
  - Sensitivity target: 2.5x10<sup>-9</sup>m/sec<sup>2</sup>/Hz<sup>1/2</sup>
  - This is equivalent to displacement amplitudes smaller than a hydrogen atom
- The development team worked hard to minimize/compensate for all noise sources:
  - Instrument intrinsic noise
  - Temperature variations
  - Wind
  - Atmospheric pressure variations
  - Magnetic field variations
- Lander vibrations



## Mars Seismic Data: Full-Sol Spectrogram, Sol 185



#### All Seismic Data as of Sol 506) InSight

Feb '19 sols: 72-506

![](_page_8_Figure_2.jpeg)

Local Time

![](_page_9_Picture_0.jpeg)

#### InSight Event List (4/16/2020)

- Currently there are 470 events in the InSight catalog.
  - 2 Quality A
    - » Clear seismic phases (e.g. P and S) and polarization
  - 90 Quality B
    - » Signal clearly observed, clear seismic phases, but no polarization
  - 176 Quality C
    » Signal clearly observed, but no clear phases
  - 202 Quality D
    - » Signal only weakly observed
    - » OR likely not a seismic event
    - » OR signal possibly contaminated by environmental conditions
  - Virtually all of the Qual A and B, and many of the Qual C have been be identified by the InSight MQS as tectonic quakes.
  - Many of the remaining events are also likely tectonic in origin.

#### Clearest Marsquake Signal to Date: S0173a, Magnitude 3.7

![](_page_10_Figure_1.jpeg)

#### Clearest Marsquake Signal to Date: S0173a, Magnitude 3.7

![](_page_11_Figure_1.jpeg)

EGU 2020

#### Sight Comparison of Terrestrial and Martian Quake Signals

![](_page_12_Figure_1.jpeg)

Banerdt, Smrekar et al. 2020

#### Event Alignment Guided by PSD Envelope Similarity

![](_page_13_Figure_1.jpeg)

Giardini et al. 2020

#### Seismicity Map for Mars – 12 "Locatable" Events

![](_page_14_Figure_1.jpeg)

![](_page_15_Picture_0.jpeg)

![](_page_15_Figure_1.jpeg)

Giardini et al., 2020

#### Estimating Seismic Activity Rate

- Mars' activity appears to be close to pre-InSight predictions, perhaps somewhat higher
- However, there may be a deficit of larger marsquakes.
- This is a preliminary estimate based on ~9 months of data; will need the full 2 years of the prime mission for a reliable estimate.

![](_page_16_Figure_4.jpeg)

### Mars Precession and Moment of Inertia from RISE

- Precession rate from the first year of RISE tracking alone is as strong as all previous missions combined.
- RISE precession measurement gives a Moment of Inertia of 0.36342 ± 0.00018
   But...
- This is not a particularly interesting result; core radius and density can't be separated.
- Measurement of the nutation to a precision that will allow the separation of core radius and density is expected from an additional ~year of tracking.

![](_page_17_Figure_5.jpeg)

Kahan et al. in prep.

## First Magnetic Measurements from the Surface of Mars

#### Notable early results include:

- DC field at the landing site ~10X stronger than measured from orbit
  - ⇒ significant crustal variations at spatial scales <150 km.</p>
- Pulsations are observed with that may be used to probe conductivity; at higher frequencies the power decreases and the vertical components are attenuated relative to horizontal, suggesting relatively high conductivity at depth.

![](_page_18_Figure_5.jpeg)

### InSight Meteorology

InSight is returning continuous high-rate pressure, temperature, and wind measurements, providing an unprecedented view of atmospheric behavior at time scales from less than a second to months and seasons.

![](_page_19_Picture_2.jpeg)

205.7

715.4

720.0

705.0

1840.0

![](_page_19_Figure_3.jpeg)

#### Ground Rigidity from Dust Devils

![](_page_20_Figure_1.jpeg)

The InSight pressure sensor detects ~10 pressure drops per day

Associated ground deformation measured by SEIS provides an estimate of ground compliance to ~3-5 m depth

1,500

V<sub>p</sub> (m s<sup>-1</sup>

1.000

 $V_{\rm p}$  regolith (m s<sup>-1</sup>)

500

10

Lognonné et al., 2020

## Pictures! (4360 Images as of Sol 490, 13 April, 2020)

![](_page_21_Picture_1.jpeg)

#### Sunset over Elysium, sol 145