# Integrated Science Operations of CASSIOPE e-POP with the Swarm Constellation for New Studies of Magnetosphere-Ionosphere Coupling

Andrew Yau<sup>1</sup> Stephanie Erion<sup>1</sup> Andrew Howarth<sup>1</sup> Gordon James<sup>2</sup> David Knudsen<sup>1</sup> Richard Langley<sup>3</sup> David Miles<sup>4</sup> <sup>1</sup>University of Calgary <sup>2</sup>Natural Resources Canada <sup>3</sup>University of New Brunswick <sup>4</sup>University of Iowa











# Key Points:

- 1. CASSIOPE e-POP currently in its 7<sup>th</sup> year of operation, as 4<sup>th</sup> component of Swarm
- 2. New/enhanced studies of magnetosphere-ionosphere coupling (MIC) enabled by:
  - Complementary orbital coverage, measurement capabilities, sampling resolutions
  - Integrated e-POP and Swarm science operations
  - Increased e-POP operation duty cycle and full altitude, latitude, local time sampling
- 3. Studies focus on space weather effects of MIC in the ionosphere and thermosphere
  - e.g. anomalous satellite orbit drag; radio scintillation
  - This presentation illustrates studies on small-scale plasma irregularities and orbit drag

## e-POP and Swarm data are highly complementary for studies of M-I coupling

2020/01



2018/01

- e-POP aurora image (FAI), ion composition (IRM), VLF-HF wave (RRI), scintillation (CER) data complement Swarm **B**, **E**, N<sub>e</sub> data
- e-POP data at high sampling rate (10<sup>2</sup> to 10<sup>3</sup> samples/s) enable studies of small-scale MIC structures and processes
- Elliptical e-POP orbit (325 × 1500 km, 81°) provides altitude sampling and frequent multi-spacecraft conjunction orbits

### Small-scale plasma density structure and variability



IRM measures incident plasma current at  $10^3$  samples/s, and uses data to infer small-scale (10-100 m) density structure ( $n_e$ ,  $\Delta n_e/n_e$ )

<u>Objective</u>: Extend Swarm mesoscale observations to small scale - determine occurrence, spectral characteristics, effect on scintillation

- a) Summary data (example): 31-min orbit on 2017/01/27; 2 s avg. plasma current, ion count rate (∝ density), ground track
- b) Original (un-decimated) data showing ~10-m structures
- c) Filter small-scale 'variance' from mesoscale amplitude and width
- d) Altitude distribution of mean variance ( $\Delta n_e$ ) at high, mid, and low latitude; inset: magnetic latitude vs. local time distribution



### **Thermospheric density variations and anomalous low-Earth-orbit decay**



NRLMSISE-00 model density (10<sup>-12</sup> kg/m<sup>3</sup>)

b)

c)



Full-orbit GAP-A data (1-20 Hz) are required to yield spacecraft positions of sub-decimeter RMS uncertainty and to derive *localized* spacecraft acceleration (or deceleration)

Spacecraft deceleration - anomalous orbit decay - is related to mass density near perigee

e-POP perigee is ~125 km (~2.5 scale heights near solar min) lower than Swarm A, C

<u>Objective</u>: More sensitive determination of thermospheric mass density near solar min

- a) GAP-A data coverage before and since 2017/05
- b) MSIS model density along e-POP orbit vs.
- c) F10.7 and Ap