

Shanghai Astronomical Observatory Chinese Academy of Sciences



#### (EGU2020-4324)

# Status on Chinese Space Geodesy Network and its Applications

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### GGOS session G2.1

The Global Geodetic Observing System: Improving infrastructure for future science Tuesday, 5 May 2020, 08:30–10:15



**1. Status of Chinese space geodetic networks** 

2. Data processing software update

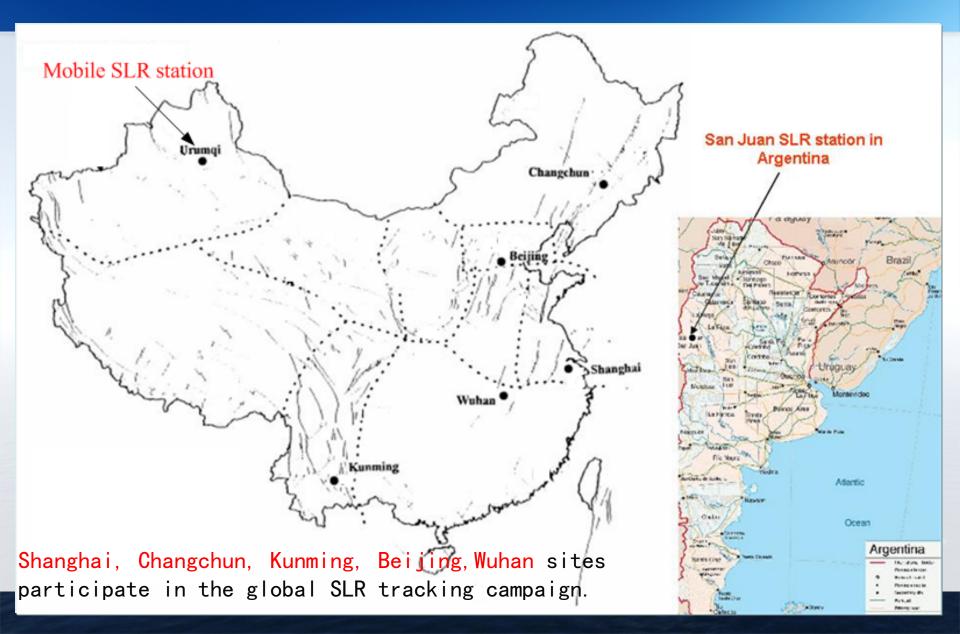
3. CERS, STRF and CERS EOP products

4. Conclusions and future plans

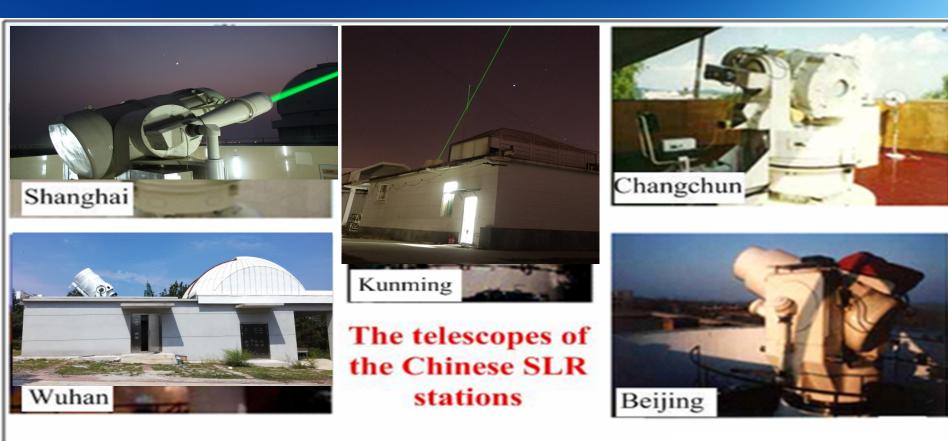
# **1. Status of Chinese space geodetic networks**

# **Status of Chinese SLR network**

# **Distribution of Chinese SLR sites**



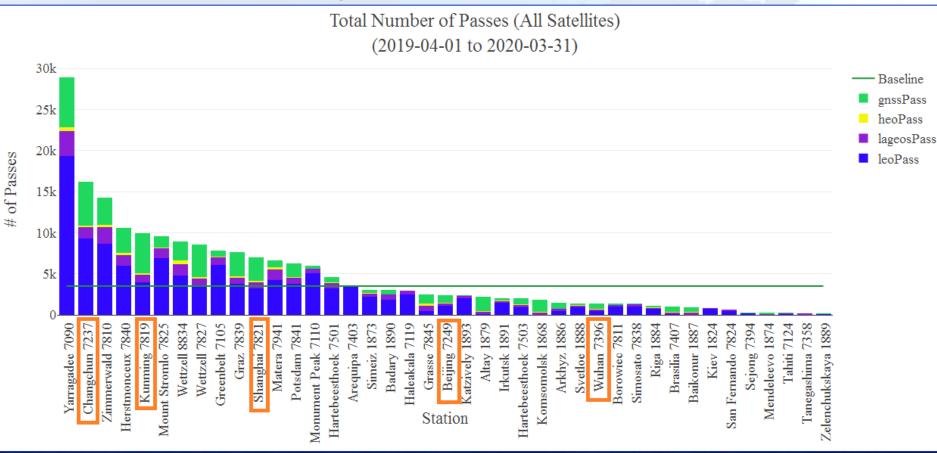
# **Tracking telescopes of SLR network**



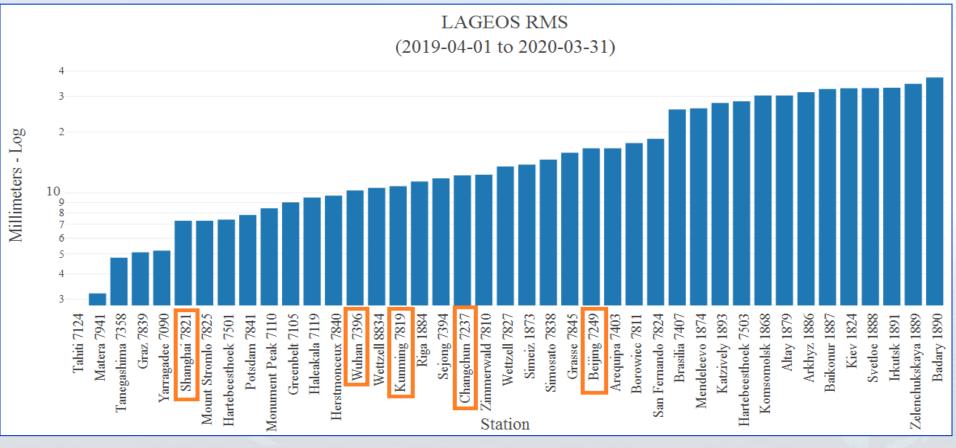




 The routine 1kHz SLR measurements in Shanghai, Changchun, Kunming, Beijing and Wuhan sites and the total passes of laser data per year of over 35k for GEO, GNSS, Lageos, LEO satellites.



• The routine 1kHz SLR measurements with the precision of less than 12mm to Lageos satellites.



The short and long term stability of laser data are less than 20mm and

 The mobile SLR system (TROS) with the aperture of 1 meter has installed in Xinjiang Observatory (Urumqi). The observations to ILRS satellites has been performed since Sep. 19 in 2019.



# The mobile SLR system (TROS) in Xinjiang Observatory (Urumqi)

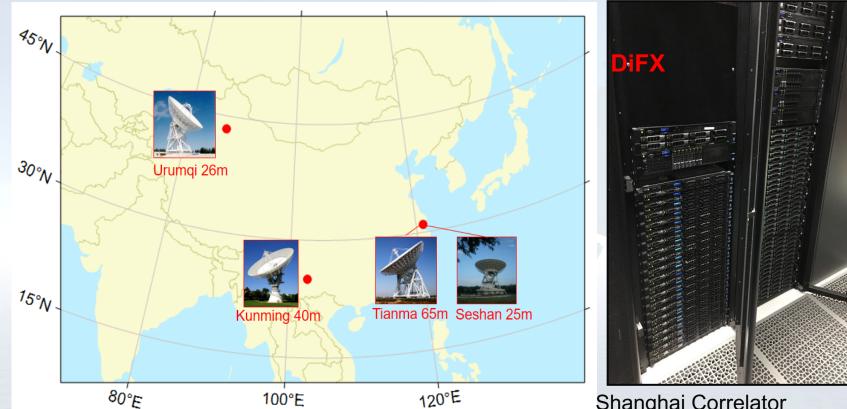
Very significant site in Chinese SLR network

- Shanghai station is developing 10kHz ~ 100kHz rep. rate SLR technology in order to improve the system stability and data quality. Due to high data density, the high rep. rate SLR data are also used to analyse the rotation rate of geodetic satellites.
- Kunming sites has realized the Lunar Laser Ranging with the distance of 380,000km.
- The system updating work in SanJuan station is underway for kHz SLR measurements as one valuable sites in the southern hemisphere.
- The updating 2kHz routine SLR measurements at night and in daylight is also being underway in Chinese network and will be finished in 2022 for geodetic observations.

# **1. Status of Chinese space geodetic networks**

# **Status of Chinese VLBI network**

# Chinese VLBI facilities operational for the IVS



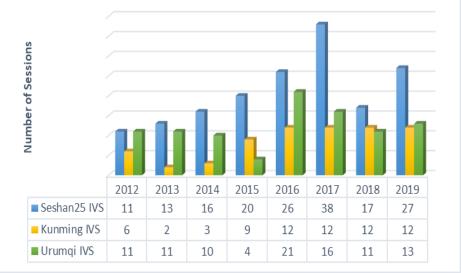
Currently 4 stations regularly participate in the IVS observing program.

Shanghai Correlator 400 CPU cores, 800TB storage, 2Gbps international fiber link

# Participation of IVS observing sessions

- Seshan25 and Urumqi are IVS network stations
- Kunning and Tianma65 are cooperative stations of the IVS
- Tianma 65-m antenna supports 3~6 IVS astrometric sessions on annual basis
- Seshan25 also participates in 1-hour IVS intensive session monthly for UT1 measurement

Statistics of observed IVS sessions



Each IVS regular observing session spans 24 hours

# Mixed-mode geodesy using Chinese VLBI stations

| Date     | Starting<br>time | Duration | Code   | Stations        |
|----------|------------------|----------|--------|-----------------|
| 20150122 | 9:00             | 24       | cn1501 | ShKmUrKsJs      |
| 20150122 | 9:00             | 24       | cdsn02 | ShUrKsJs        |
| 20150321 | 0:00             | 24       | cdsn03 | ShUrKsJs        |
| 20150630 | 17:30            | 24       | cdsn04 | ShKsJs          |
| 20150917 | 10:00            | 24       | cdsn05 | ShKmKsJs        |
| 20151022 | 2:00             | 24       | cn1502 | ShKmUrBjKsJs    |
| 20151113 | 3:00             | 24       | cn1503 | KmUrKsJs        |
| 20151210 | 8:00             | 24       | cn1504 | KmUrKsJs        |
| 20160506 | 6:00             | 24       | cn1601 | ShKmUrKsJs      |
| 20160702 | 6:00             | 24       | cn1602 | ShUrKsJsYgKe    |
| 20160928 | 18:00            | 24       | apsg39 | APSG+KsJs       |
| 20161130 | 6:00             | 24       | cn1603 | KmUrKeYgKsJs-Ur |
| 20161229 | 3:00             | 24       | cn1604 | KmUrKsJs        |
| 20170325 | 06:00            | 24       | cn1701 | ShKmUrKsJs      |
| 20170511 | 06:0             | 24       | cn1702 | ShUrKs          |
| 20170523 | 17:30            | 24       | cn1703 | ShKmUrKsJs      |
| 20170725 | 17:30            | 24       | APSG40 | APSG + KsJsJlSy |
| 20170808 | 17:30            | 24       | cn1703 | T6UrKsJs        |

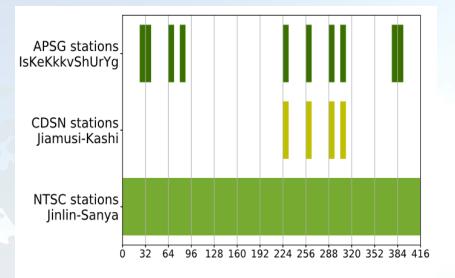


Fig. 2: X-band Frequency sequence wrt 8181MHz.

|     | •        |      | -   |   |   |   |   |   |
|-----|----------|------|---|---|---|---|---|---|
| es  | Jilin    |      | Sanya   |   | Jiamusi   |   | Kashi   |   |
| adj | ustment  | σ    | adjustment                                      | σ   | adjustment  | σ   | adjustment  | σ   |
| -   | 49.76    | 8.91 | -58.10  | 12.58   | -53.59  | 26.56   | 17.23   | 26.70   |
|     | 24.64    | 1.87 | 101.36  | 2.53  | 52.94   | 8.87  | 35.70   | 0.44  |
|     | 24.98    | 1.57 | -20.60  | 2.11  | -0.73   | 6.05  | 106.65  | 5.93  |
|     | adj<br>- | es   | es<br>adjustment σ<br>-49.76 8.91<br>24.64 1.87 | $\begin{array}{c ccccccccccccccccccccccccccccccccccc$ |

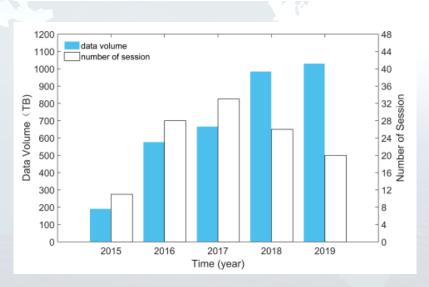
(He et al. 2019, EVGA)

# Shanghai VLBI correlator

118 IVS sessions and 20 CVN sessions have been correlated from 2015 to 2019

- AOV: astrometry of weak ICRF sources
- **APSG**: regional geodesy around Asian and Pacific area
- AUST-AST: Australian astrometric sessions
- CRF/CRD: monitoring CRF sources, more focused on southern hemisphere
- RD: astrometry of Gaia transfer sources
- CVN: measurement of station positions of Chinese stations

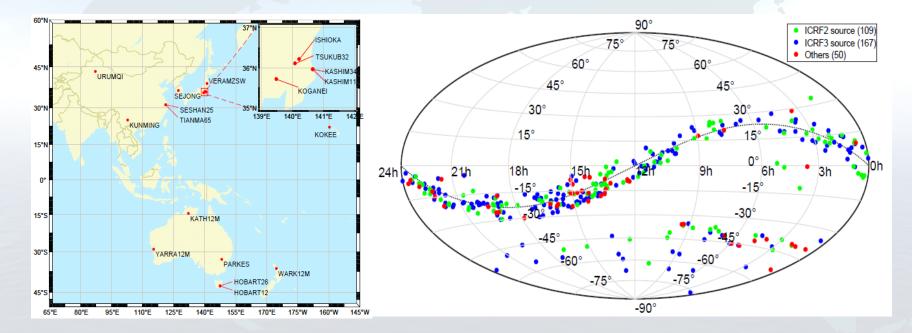
| Session Name | 2015 | 2016 | 2017 | 2018 | 2019 |
|--------------|------|------|------|------|------|
| AOV          | 3    | 3    | 3    | 7    | 6    |
| APSG         | 2    | 2    | 2    | 2    | 2    |
| AUS-AST      | 0    | 13   | 14   | 6    | 0    |
| IVS-CRF/CRD  | 6    | 9    | 9    | 6    | 6    |
| IVS-RD       | 0    | 1    | 5    | 5    | 6    |
| IVS Session  | 11   | 28   | 33   | 26   | 20   |
| CVN Session  | 7    | 5    | 5    | 3    | 0    |



# Astrometry of weak sources with AOV/APSG

AOV is unique to astrometry of weak sources in the middle southern hemisphere and the ecliptic plane

SHAO scheduled and correlated part of AOV observing sessions from 2015 onwards



#### **Tianma VGOS stations**

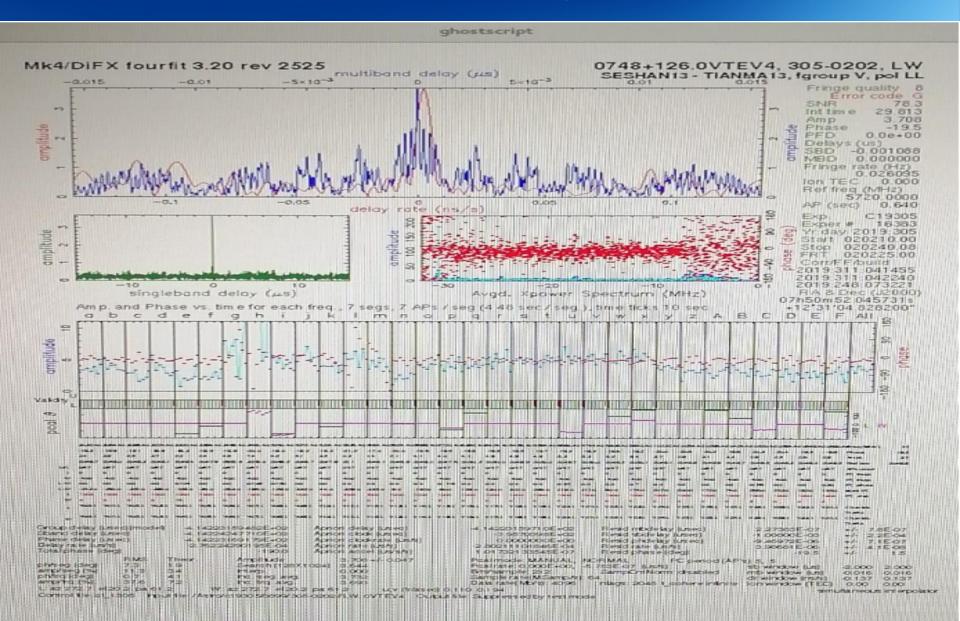
## Seshan VGOS – Tianma VGOS Stations

ShVGOS



| F(Ghz) | EL(°) | SEFD-H(Jy) | SEFD-V(Jy) |
|--------|-------|------------|------------|
| 4      | 36    | 2167.7     | 2058.6     |
| 6      | 35    | 2387.9     | 2185.8     |
| 8      | 33    | 1842.5     | 1702.3     |
| 10     | 31    | 2847.2     | 2377.1     |
| 12     | 29    | 3103.3     | 2812.3     |
| 14     | 27    | 3596.3     | 2966.3     |

## Seshan VGOS – Tianma VGOS fringe test on Nov. 7, 2019

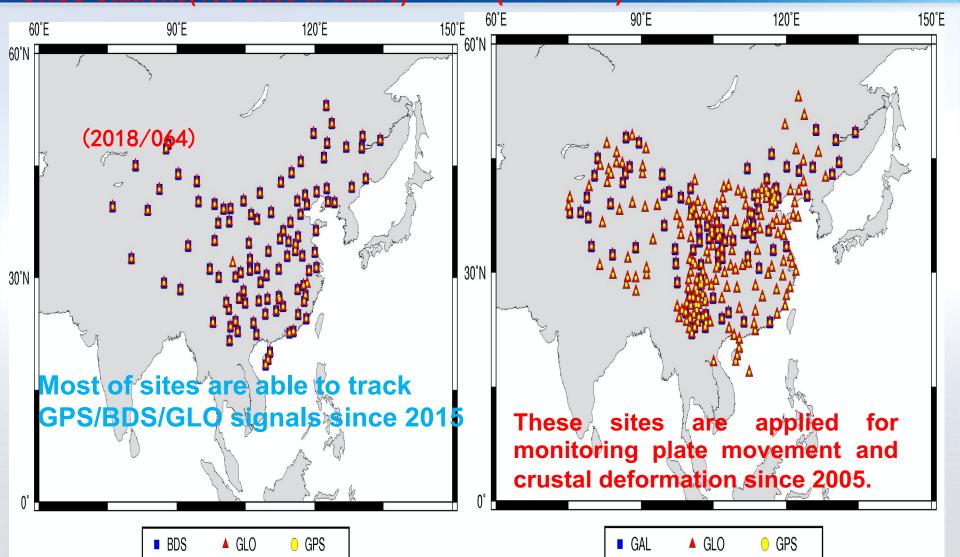


# **1. Status of Chinese space geodetic networks**

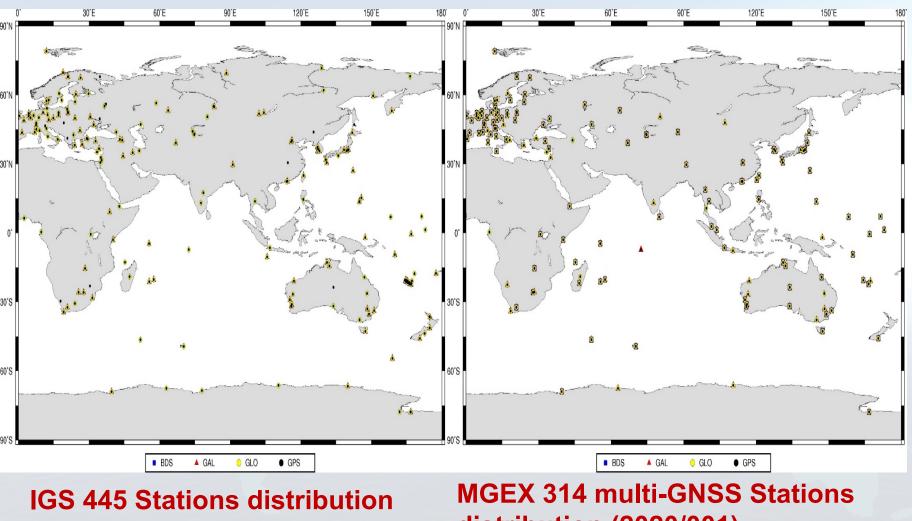
# **Status of Chinese GNSS network**

# **Chinese GNSS network**

Sub-net1 distribution of 103 multi-GNSS Stations(150 sites in future) Sub-net2 distribution of 260 Stations (2020/001)



# **Global GNSS network from IGS and MGEX used**



(2020/001)

distribution (2020/001)

# 2. Data processing software update

# **GNSS data processing software update**

The GNSS data processing soft ware has been updated accord ing to the reco mmended mod els by IERS con ventions(2010). Nowadays, GN SS data proces sing software h as realized mult i-GNSS precise orbit determina tion and the SL **R** orbit validatio n.

Loading effects

on stations

pressure effects

on satellites

Background models

#### Updated models Orbit improvetion Solid earth tide

Ocean loading

Ocean pole tide

S1-S2 Atmospheric pressure loading

Solar radiation pressure(box-wing)

Earth radiation pressure(box-wing)

post-seismic deformation (PSD)

IAU2006/2000PN

time-variable gravity filed(eigen-6s41.gfc)

**DE430** 

High order ionosphere terms for GNSS

> GNSS DSB (CODE/DLR/CAS)

1mm-3mm

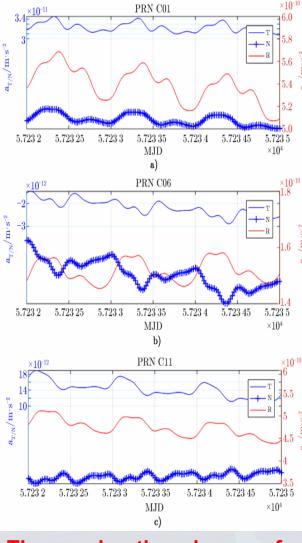
10mm-20mm

3mm-5mm

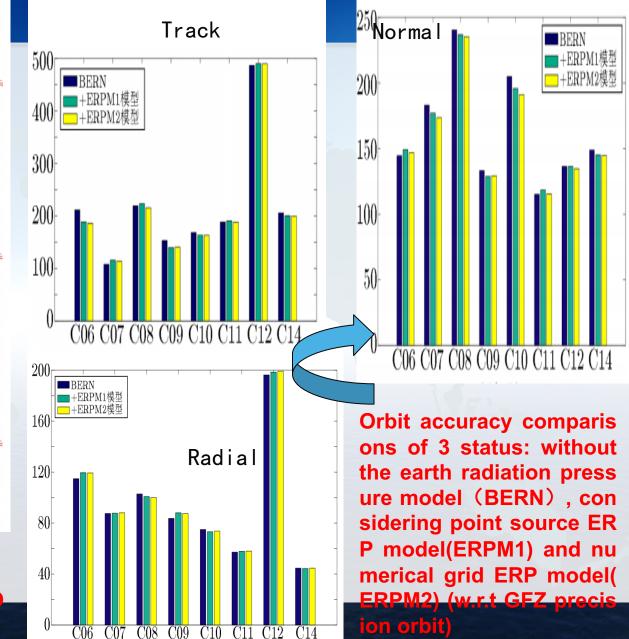
0.5mm-1mm

1mm-3mm

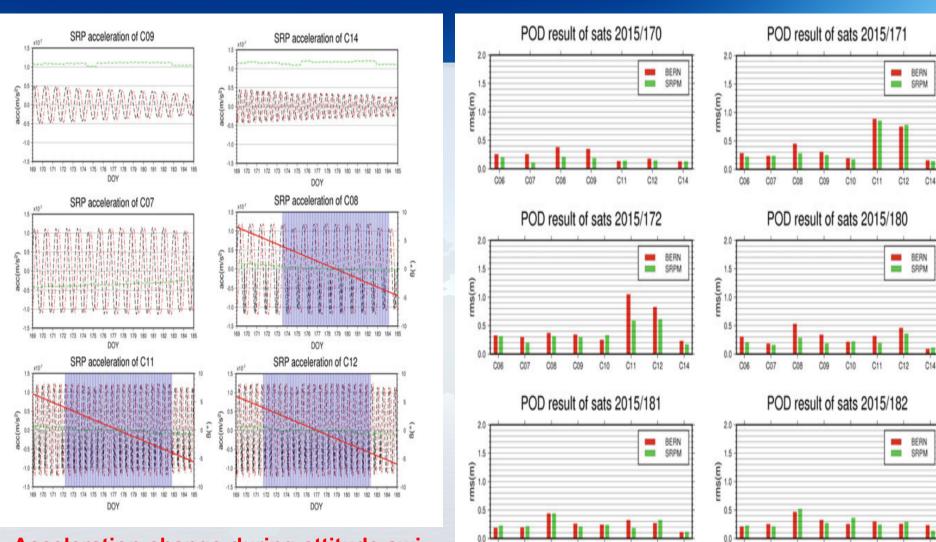
# **Earth Radiation Pressure for BDS**



The acceleration change of earth radiation pressure of BDS GEO (a) / IGSO (b) / MEO (c) satellite in R / T / N.



#### Solar Radiation Pressure model considering attitude switching



C06 C07

C08

C09

C11

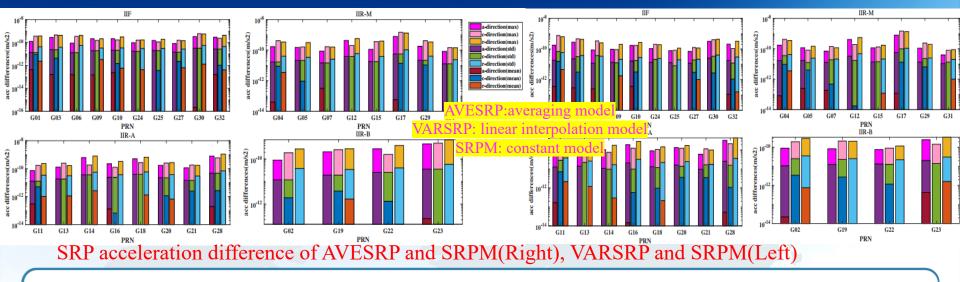
Acceleration change during attitude swi tching period (The R/T/N was plotted in black/red/green. The blue shadow is sat ellite orbit-normal and the sun elevation to drawn with red.)

The orbit comparison between the BERN and SRPM model during attitude switching (w.r.t GFZ precision orbit) in period.

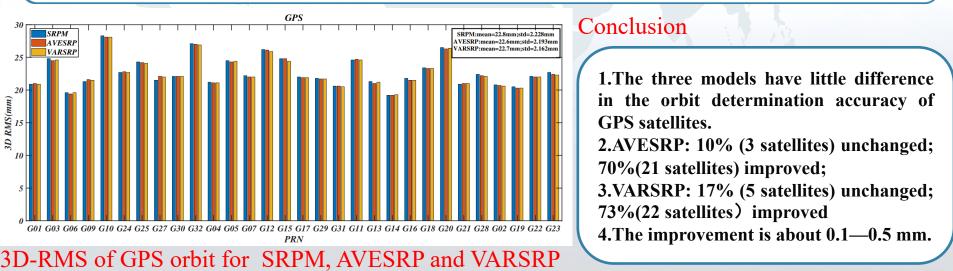
C06 C07 C08 C09

C11

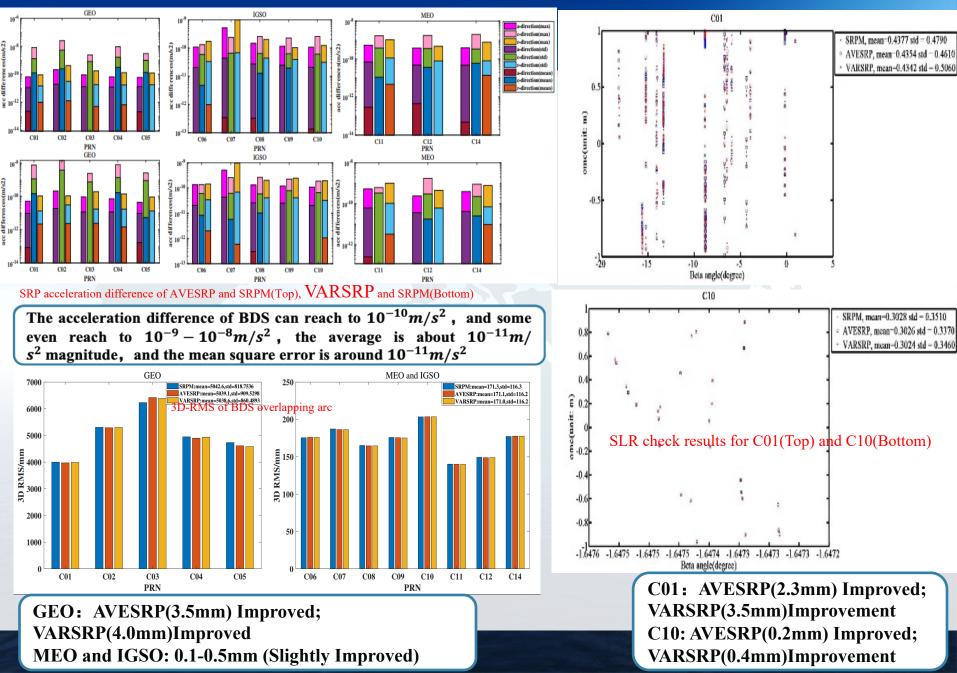
#### SRP model with considering Solar Irradiance Change(GPS test)



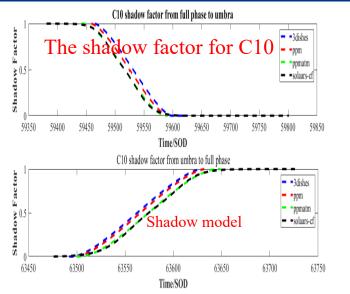
The VARSRP and AVESRP models are more consistent and show common features. The acceleration difference of most GPS satellites is up to  $10^{-10}m/s^2$ , some even reach to  $10^{-9}m/s^2$ , the mean value is around  $10^{-13}m/s^2$ , and the mean square error is about  $10^{-11}m/s^2$ .



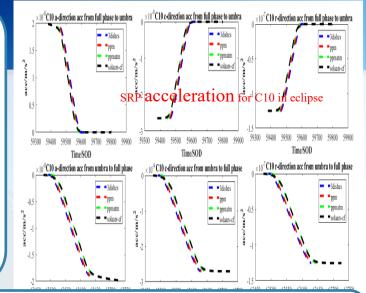
#### SRP model with considering Solar Irradiance Change(BDS test)



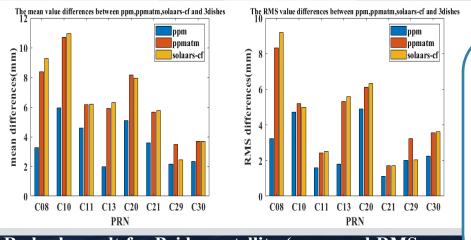
#### The Refinement of Shadow Model for Beidou Satellites



1. Shadow models(PPM, PPMatm, SOLAARS-CF) with earth oblateness and atmosphere effect enter shadow earlier and come out later than 3dishes 2. The SRP acceleration of PPMatm and SOLAARS-CF have significant difference with PPM and 3dishes 3. PPMatm is highly consistent with SOLAARS-CF



3dishes: conical spherical model considering earth shadow and moon shadow PPM: prospective projection model considering earth oblateness(from Li(2018)) PPMatm: linear function describing atmosphere effect based on PPM model(from Li(2018)) SOLAARS-CF: semi-empirical model considering earth oblateness and atmosphere effect(from Robertson(2015))



SLR check result for Beidou satellites(mean and RMS of three models regard to 3dishes)

#### Conclusion

- 1. PPMatm and SOLAARS-CF's performance is better than other two models
- 2. Mean: The maximum improvement of PPMatm and SOLAARS-CF can reach to 10mm, it is generally 2-8mm
- 3. RMS: The maximum improvement of PPMatm and SOLAARS-CF can reach to 8mm, it is generally 2-6mm
- 4. The performance of PPMatm and SOLAARS-CF is comparable

#### The Refinement of Shadow Model for Beidou Satellites

**Eclipse** 

Reception

**Eclipse** 

Launch

**Time period** 

#### The statistics result of Inter-satellite link check

| (DC                                       | DY)   |               | (Yes or No)                  |            | (Yes or No)         |
|---|-------|---------------|------------------------------|------------|---------------------|
| 043-                                      | -070  | C19           | Yes                          | C23        | No                  |
| 043-                                      | -070  | C21           | Yes                          | C23        | No                  |
| 102·                                      | -118  | C27           | Yes                          | C22        | No                  |
| <b>102</b> ·                              | -118  | C30           | Yes                          | C23        | No                  |
| 0.2<br>0 -<br>m -0.2 -<br>0.6 -<br>-0.8 - |       | C19-C23 Inter | Satellite Link check results | in eclipse | 0.2833<br>ns=0.2810 |
| -1<br>58525                               | 58530 | 58535         | 58540 5                      | 8545 58550 | 58555               |
|   |       |               | Time/MJD                     |            |                     |

| sat-li | nk      | Shadow model | Mean(m) | RMS(m) | dif <sub>mean</sub> (m) | dif <sub>RMS</sub> (m) |
|--------|---------|--------------|---------|--------|-------------------------|------------------------|
| C19-C2 | 23      | 3dishes      | -0.2050 | 0.2866 | 0                       | 0                      |
|        |         | PPM          | -0.1995 | 0.2833 | 0.0055                  | 0.0033                 |
|        |         | PPMatm       | -0.1967 | 0.2810 | 0.0083                  | 0.0056                 |
|        |         | SOLAARS-CF   | -0.1969 | 0.2810 | 0.0081                  | 0.0056                 |
| C21-C  | 23      | 3dishes      | -0.2393 | 0.2981 | 0                       | 0                      |
|        |         | PPM          | -0.2375 | 0.2965 | 0.0018                  | 0.0016                 |
| 2      |         | PPMatm       | -0.2357 | 0.2946 | 0.0036                  | 0.0035                 |
|        |         | SOLAARS-CF   | -0.2360 | 0.2947 | 0.0033                  | 0.0034                 |
| C27-C  | C27-C22 | 3dishes      | -0.2412 | 0.3896 | 0                       | 0                      |
|        |         | PPM          | -0.2358 | 0.3859 | 0.0054                  | 0.0037                 |
|        |         | PPMatm       | -0.2331 | 0.3829 | 0.0081                  | 0.0067                 |
|        |         | SOLAARS-CF   | -0.2329 | 0.3834 | 0.0083                  | 0.0062                 |
| C30-C2 | 22      | 3dishes      | -0.1419 | 0.3532 | 0                       | 0                      |
|        |         | PPM          | -0.1385 | 0.3515 | 0.0034                  | 0.0017                 |
|        |         | PPMatm       | -0.1375 | 0.3509 | 0.0044                  | 0.0023                 |
|        |         | SOLAARS-CF   | -0.1384 | 0.3511 | 0.0035                  | 0.0021                 |

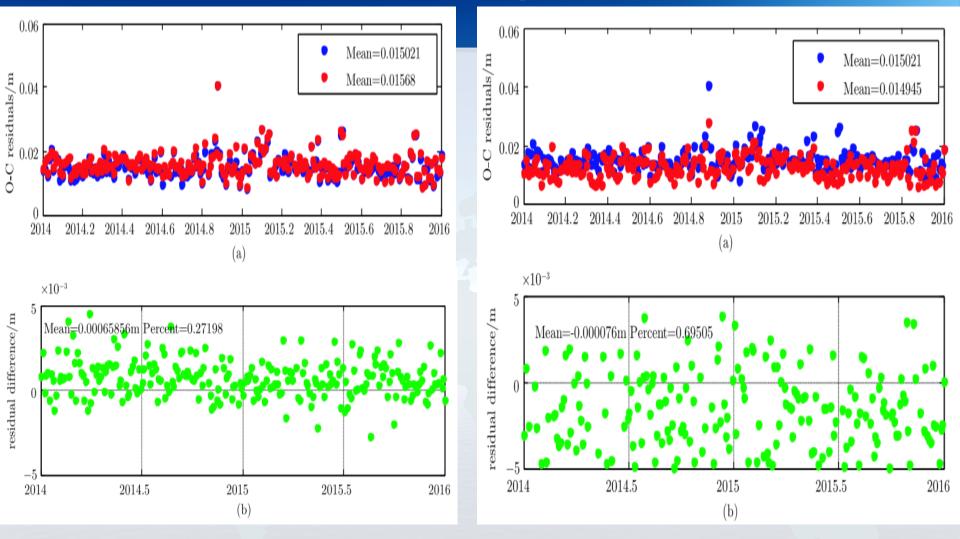
#### Conclusion

- Mean: The maximum improvement of 1. **PPMatm and SOLAARS-CF can reach to** 8mm, it is generally 2-6mm
- **RMS:** The improvement of PPMatm and 2. **SOLAARS-CF** is generally 2-6mm
- 3. The performance of PPMatm and **SOLAARS-CF** is comparable

# 2. Data processing software update

# SLR data processing software update

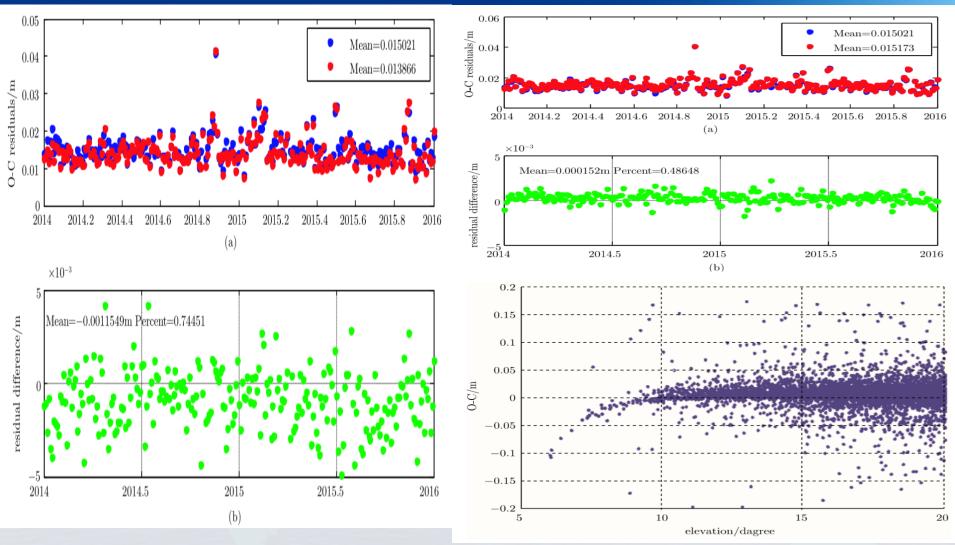
# **Ocean Tide Model Update**



Lageos1(left) and Lageos2(right)'s residual WRMS with CSR3.0(blue) model and FES2004(red) model, and differences between 2 models(down). FES2004 shows better.

### **Reference Frame Update**

## **Troposphere Model Update**

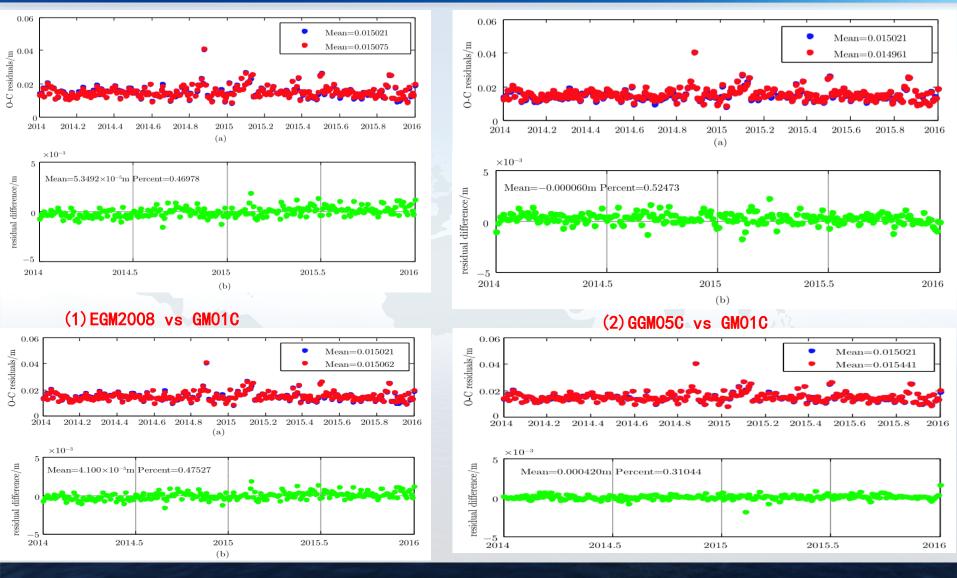


Lageos1's residual WRMS with SLRF2008(blue) and SLRF2014(red), and the difference between 2 models(down). SLRF2014 shows better.

Lageos1's residual WRMS with Marini-Murray model(blue) and Mendes-Pavlis(red) model, and the difference between 2 models. Mendes-Pavlis model performs better in low elevation.

## **Earth Gravity Model update**

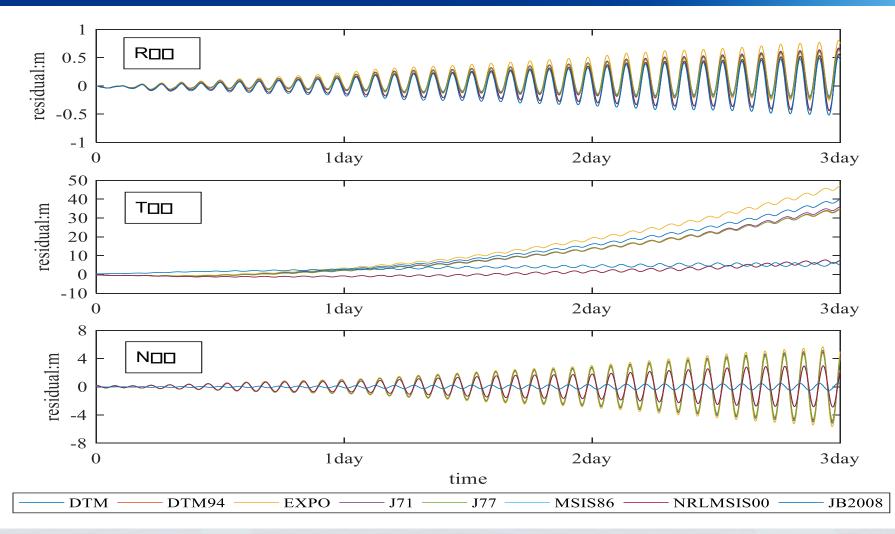
Influence of Different gravity models on Lageos1 orbit( 4 models:EGM2008, GGM05C, GOCO05C, G0CO05S, in blue point) with CM01C model(in red). EGM2008, GGM05C and G0C005S perform better than GM01C. GGM05C shows better.



(3) GOC005C vs GM01C

(4) GOC005S vs GM01C

### **Atmosphere Drag Model Update**

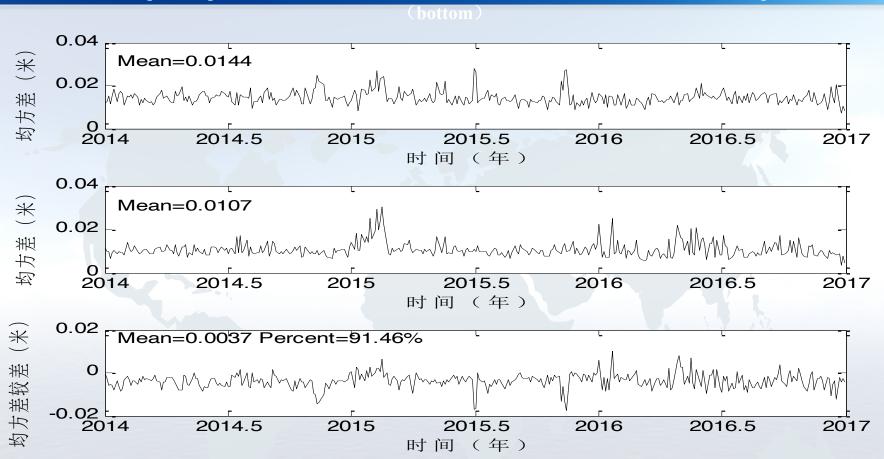


We add 2 new drag models: NRLMSISE00 and JB2008. The RTN difference of JASON2 predicted orbit and precise orbit with 8 atmosphere drag model(exponential model, DTM, DTM94, J71, J77, JB2008, MSIS86, NRLMSIS00). NRLMSISE00/MSIS86 drag models

show better.

## Weighting strategy Update

LAGEOS 1 orbit precision rms by old experience weighting strategy (up), by modified FCM weighting strategy with SLR seasonal quick report information (middle), the difference of above two and the improved arc ratio.



The observed residual rms was improved by 3.7mm on average and 91.46% of the arc segment was improved. The residual RMS of all the stations involved in the calculation was reduced.

# Weighting strategy Update

SLR core sites residual RMS and NPT number with different weighting strategy

| SLR site | <b>Old</b> experience w | eighting strategy | Modified FCM weighting strategy |              |  |
|----------|-------------------------|-------------------|---------------------------------|--------------|--|
| SEN SILE | NPT Number              | Residual RMS(m)   | NPT Number                      | Residual RMS |  |
| 7080     | 1497                    | 0.0189            | 1500                            | 0.0176       |  |
| 7090     | 34835                   | 0.0176            | 34919                           | 0.0172       |  |
| 7105     | 14973                   | 0.0235            | 15116                           | 0.0158       |  |
| 7110     | 10258                   | 0.0204            | 9739                            | 0.0140       |  |
| 7501     | 9066                    | 0.0195            | 8862                            | 0.0178       |  |
| 7810     | 14999                   | 0.0108            | 15016                           | 0.0090       |  |
| 7825     | 13491                   | 0.0169            | 13983                           | 0.0128       |  |
| 7839     | 6466                    | 0.0125            | 6417                            | 0.0107       |  |
| 7840     | 12991                   | 0.0102            | 13067                           | 0.0098       |  |
| 7941     | 13970                   | 0.0154            | 14210                           | 0.0095       |  |
| 8834     | 4227                    | 0.0208            | 4289                            | 0.0201       |  |

The residual RMS of all SLR sites are reduced and the NPT number is improved.

# 3. CERS, STRF and CERS EOP products

Homepage

Homepage

OrganizationInfo

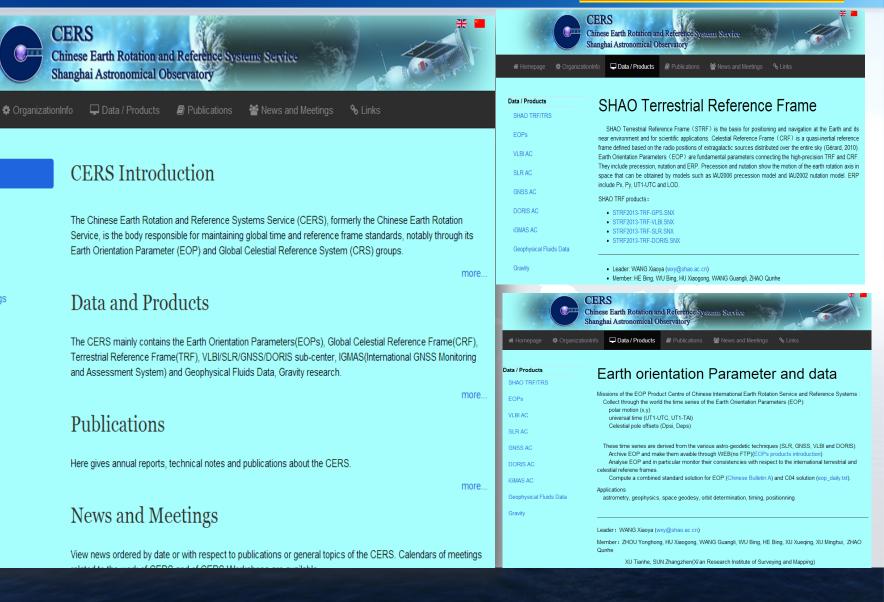
Data / Products

News and Meetings

Publications

Links

#### http://cers.shao.ac.cn

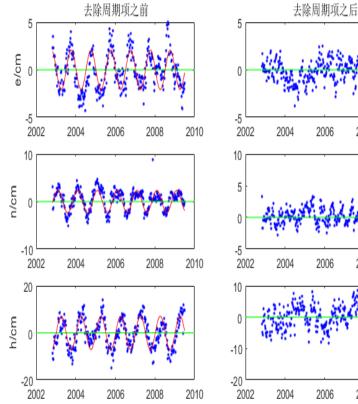


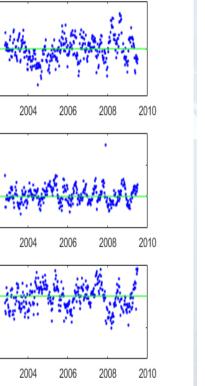
# Nonline TRF establishment

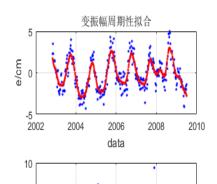
 $X_{PSD}(t) = X(t_0) + \dot{X}(t - t_0) + \sum_{i=1}^{2} (a_i \cos\left(\frac{wt}{i} - \phi_i\right) + b_i \sin\left(\frac{wt}{i} - \phi_i\right)) + \delta X_{PSD}(t)$ 

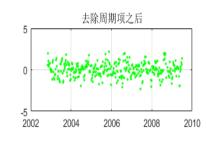
n/cm

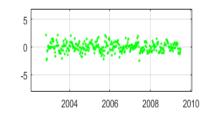
-10 -2002

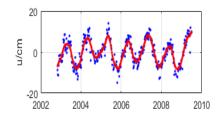






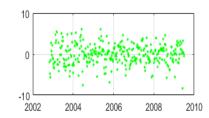






2006

2004



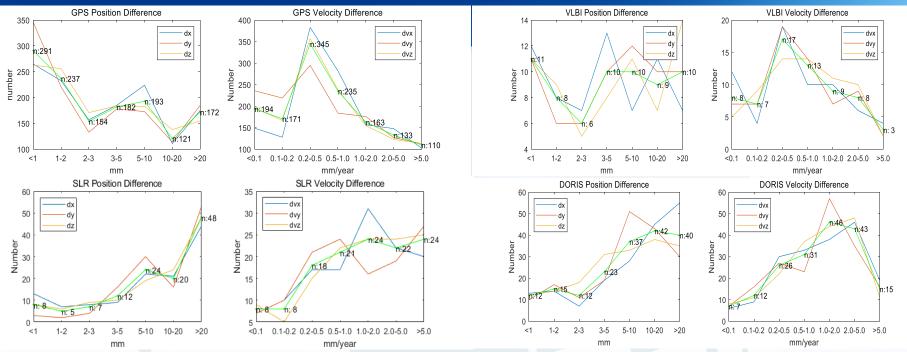
#### Period fitting based on least square

#### SSA fitting

2010

2008

# **STRF Products accuracy evaluation**

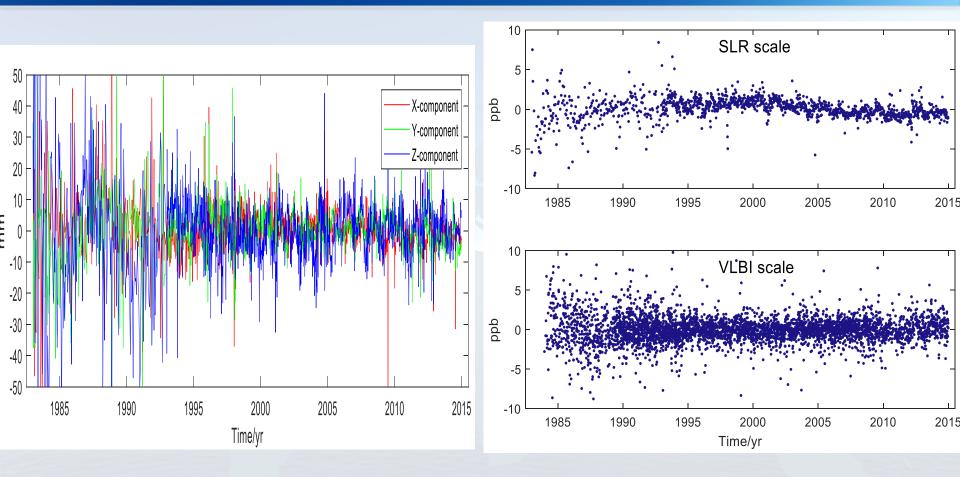


# Site number statistics with different accuracy range of STRF2014 (SHAO nonlinear Terrestrial Reference Frame) for GPS/VLBI/SLR/DORIS w.r.t ITRF2014.

Comparison result statistics of the coordinates and velocities for four techniques between SHAO nonlinear Terrestrial reference frame (STRF) and ITRF2014

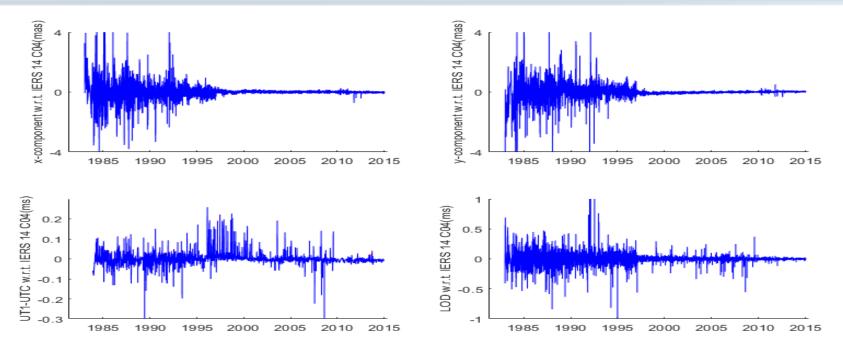
| Technique      | Coordinate accuracy statistics<br>(mm) |       | Velocity accuracy statistics<br>(mm/yr) |           |
|----------------|--|-------|---|-----------|
| accuracy range | <1                                     | (1,2) | <0.1                                    | (0.1,0.2) |
| GNSS           | 21.6%                                  | 17.6% | 14.4%                                   | 12.7%     |
| VLBI           | 17.2%                                  | 12.5% | 12.5%                                   | 10.9%     |
| SLR            | 6.5%                                   | 4.0%  | 6.5%                                    | 6.5%      |
| DORIS          | 6.6%                                   | 8.3%  | 3.7%                                    | 6.6%      |

# **STRF Products accuracy evaluation**



Translation parameters and Scale factors of STRF2014.

## CERS CO4 EOP Products accuracy evaluation



CERS CO4 EOP products comparison with IERS 14 CO4

| EOD Deremetere                    | WRMS            |  |  |
|-----------------------------------|-----------------|--|--|
| EOP Parameters                    | w.r.t. IERS C04 |  |  |
| x-component of Polar Motion (mas) | 0.0564          |  |  |
| y-component of Polar Motion (mas) | 0.0576          |  |  |
| UT1-UTC (ms)                      | 0.0103          |  |  |
| LOD (ms)                          | 0.0109          |  |  |

# 4. Conclusions and future plans

- Chinese space geodetic network has contributed to global space geodesy and GGOS. They will be developed with SLR equipment modification, more GNSS sites and more VGOS sites in future.
- ② The accuracy and stability of TRF will be improved with multiple efforts from different countries. The observation model and dynamics model of each technology still need to be improved. At present, basic tests of SLR and GNSS have been completed, and VLBI and DORIS also need to be improved.
- ③ There is a space for improvement in the solution strategy. SLR has been successfully tested, and similar tests are needed for other technologies.
- ④ The results of the international combination centres for various technologies also need to be compared and tested. The third combination centre is necessary.
- (5) STRF2020 is under considered with dense Chinese GNSS network and multi-GNSS data analysis.

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# Thank you for your attention!