# Advancement of the AMSSP for Ground-Based Measurements of the Complete Stokes Vector



## $(\mathbf{i})$

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### **1** Introduction

Aerosol scattering changes the state of sun light in terms of spectral intensity and polarization. Thus, polarization can be used to retrieve aerosol properties such as optical properties (AOT, refractive index,...), particle number concentration and other aerosol related parameters. The knowledge of aerosol parameters helps to examine and understand the effects of aerosols on the climate. One way to describe polarization of light is the Stokes vector. The multispectral polarimeter AMSSP (airborne multispectral sunphotometer and polarimeter) measures four intensities with different sets of polarizing components simultaneously. Those intensities and a proper characterization of the components of the instrument are needed to calculate the complete Stokes vector. All angles and further parameters of the polarization sensitive optical components were determined during laboratory calibrations. On the basis of experimental experiences, the measurement setup of the AMSSP will be improved to get more stable polarization information.

### **3** Laboratory calibration



### Data processing

- Working fit only with very good alignment of setup and accurate rotational stage of analyzer (rotating polarizer)
- Pre-processing: offset correction, averaging, wavelength calibration, time normalization
- Fit measured intensity with theoretical intensity derived using Mueller matrices of polarizer  $(M_p)$  and wave plates  $(M_w)$  [2] and an input Stokes vector ( $\vec{S}_0 = [1, 0, 0, 0]$ ) of unpolarized light

### Results

- Angles and retardations of optical elements determined with fit
- Additional fit parameter: amplitude factor, used for relative calibration
- Matrix A and  $A^{-1}$  calculated to determine Stokes vector
- Complete polarized Stokes vector could be calculated using  $\vec{S} = A^{-1}\vec{I}$ shown in upper plot for 633 nm
- Absolute difference ( $\Delta$ ) between calculated Stokes parameter and theoretical values for each wavelength as mean over all analyzer angles in lower plot

Fit function to find angle of polarizer:

$$I = \vec{S}[0]$$
  
=  $(M_p(\phi_p)M_p(\phi_{analyzer})\vec{S}_0)[0]$ 

Fit function to find parameter of waveplate:

$$I = \vec{S}[0]$$
  
=  $(M_w(\phi_w, \delta_w)M_p(\phi_p)M_p(\phi_{analyzer})\vec{S}_0)[0]$ 









 $\Delta A$ 

### 2 AMSSP and measurement principle

Fig. 1: Setup of AMSSP

• Stokes vector  $ec{S} = [I,Q,U,V]$  of the light in front of the instrument is calculated by:  $\vec{S} = A^{-1}\vec{I}$  with  $\vec{I}$  the measured intensities and A the matrix which characterizes the effect of the optical components of the instrument on the light [1]

## 4 Uncertainty estimation and consitency

### **Uncertainty estimation**

$$\left(\Delta \frac{S_i}{S_0}\right)^2 = \sum_{j=0}^3 \left(\frac{\partial \frac{S_i}{S_0}}{\partial A_{ij}^{-1}} \Delta A_{ij}^{-1}\right)^2 + \left(\frac{\partial \frac{S_i}{S_0}}{\partial A_{0j}^{-1}} \Delta A_{0j}^{-1}\right)^2 + \left(\frac{\partial \frac{S_i}{S_0}}{\partial I_j} \Delta I_j\right)^2$$

with  $\Delta A^{-1}$  calculated using difference quotients:

$$=\sqrt{\left(\frac{\partial A^{-1}}{\partial \phi_{w0}}\Delta \phi_{w0}\right)^{2}+\ldots+\left(\frac{\partial A^{-1}}{\partial \delta_{0}}\Delta \delta_{0}\right)^{2}+\ldots+\left(\frac{\partial A^{-1}}{\partial \phi_{p3}}\Delta \phi_{p3}\right)^{2}}$$

and uncertainty of measured intensities

$$\Delta I_{j} = \sqrt{\Delta I_{dc}^{2} + \Delta I_{lin.Regr.}^{2} + \Delta I_{\Delta 180^{o}}^{2}}$$

- with uncertainties originating from the data processing: dark current correction  $(\Delta I_{dc})$ , averaging  $(\Delta I_{lin.Regr.})$ and from fluctuations of setup  $(\Delta I_{\Delta 180^o})$ . Consistency
- Consitency check as shown by Immler et al. [3]:

$$||m_1 - m_2|| > k\sqrt{u_1^2 + u_2^2}.$$

• k-value for all estimated Stokes parameters at the different angles of the analyzer:



### **5** Advancement

	Fig. 2: Op
6 F	André Hollstein et (2009), pp. 4767–4
[2]	Russell A Chipman,
[3]	FJ Immler et al. "R Measurement Techniq

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• On four paths through the instrument, the light is changed by one polarizer and one wave plate

• Matrix A depends on angles of the polarizer  $\phi_p$ , angles of the wave plates  $\phi_w$  and the retardation  $\delta_w$ 

• New operating platform:

- rotation of complete instrument towards desired viewing direction

 no mirrors needed which could change polarization properties of light indistinct

 direct sun measurements for adjustment and measurement of AOT

 simultaneous measurements of atmospheric trace gases with Pandora-2S

 Adjustable entrance optics to ensure same viewing direction of four spectrometer

 New waveplates with less wavelength dependence to make calibration routine more robust



perating platform before installation of instruments.

"Optimization of system parameters for a complete multispectral polarimeter". In: Applied optics 48.24

Nai Sze Tiffany Lam, and Garam Young. Polarized Light and Optical Systems. CRC Press, 2018, pp. 169,175. erence quality upper-air measurements: Guidance for developing GRUAN data products". In: Atmospheric s 3.5 (2010), pp. 1217–1231