

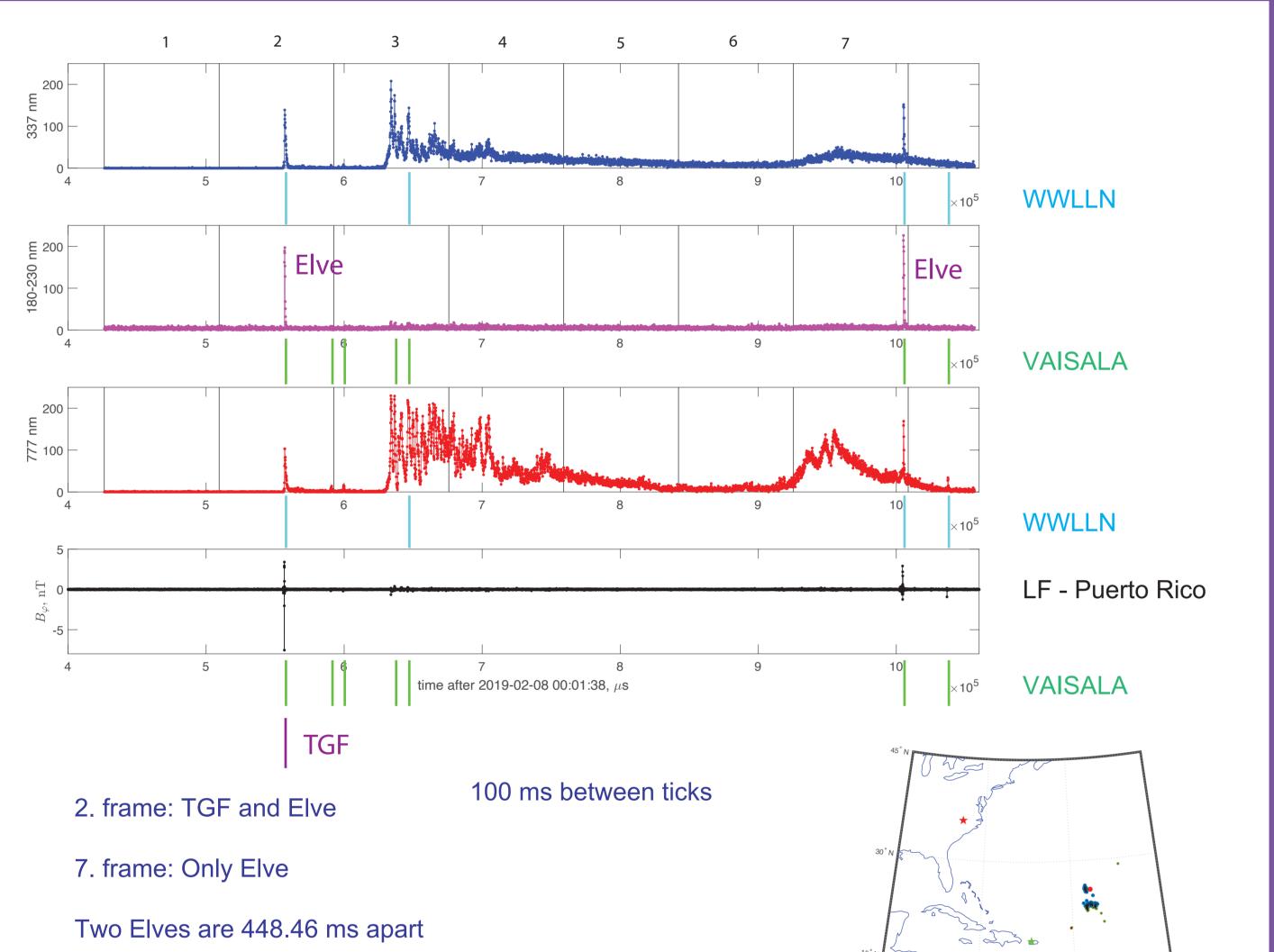
BIRKELAND CENTRE FOR SPACE SCIENCE

Abstract

On February 8, 2019 the Atmosphere-Space Interaction Monitor (ASIM) passed a thunderstorm system north east of Puerto Rico and observed a TGF and an Elve from the same lightning stroke at the very beginning of a lightning flash. A second Elve was observed 456 ms later but without any signature of a TGF about 300 km south-east of the first Elve. The strokes associated with the two Elve events were detected by WWLLN and Vaisala, which allows for an absolute timing accuracy of the ASIM measurements of at least 100 us. Images of the lighting strokes support the source locations for the Elves and TGF. Low Frequency radio measurements from Puerto Rico indicate that the first stroke was an intracloud positive while the latter was a cloud-to-ground stroke. The UV emissions from the Elves preceded the optical emissions in 777 nm by 50 us and 90 us, respectively. Modelling of the scatter and absorption of the 777 nm emissions within the

cloud can explain these delays as well as the duration of the emissions. Current moments derived from radio measurements at Puerto Rico and Duke University indicate a fast (30 us) and large (200 kA) current pulse in the IC+ stroke emitting an electromagnetic wave that produces the Elve and a slow (1-2 ms) current that also could have contributed to the optical signals.

Overview of the lightning activity



· Martin

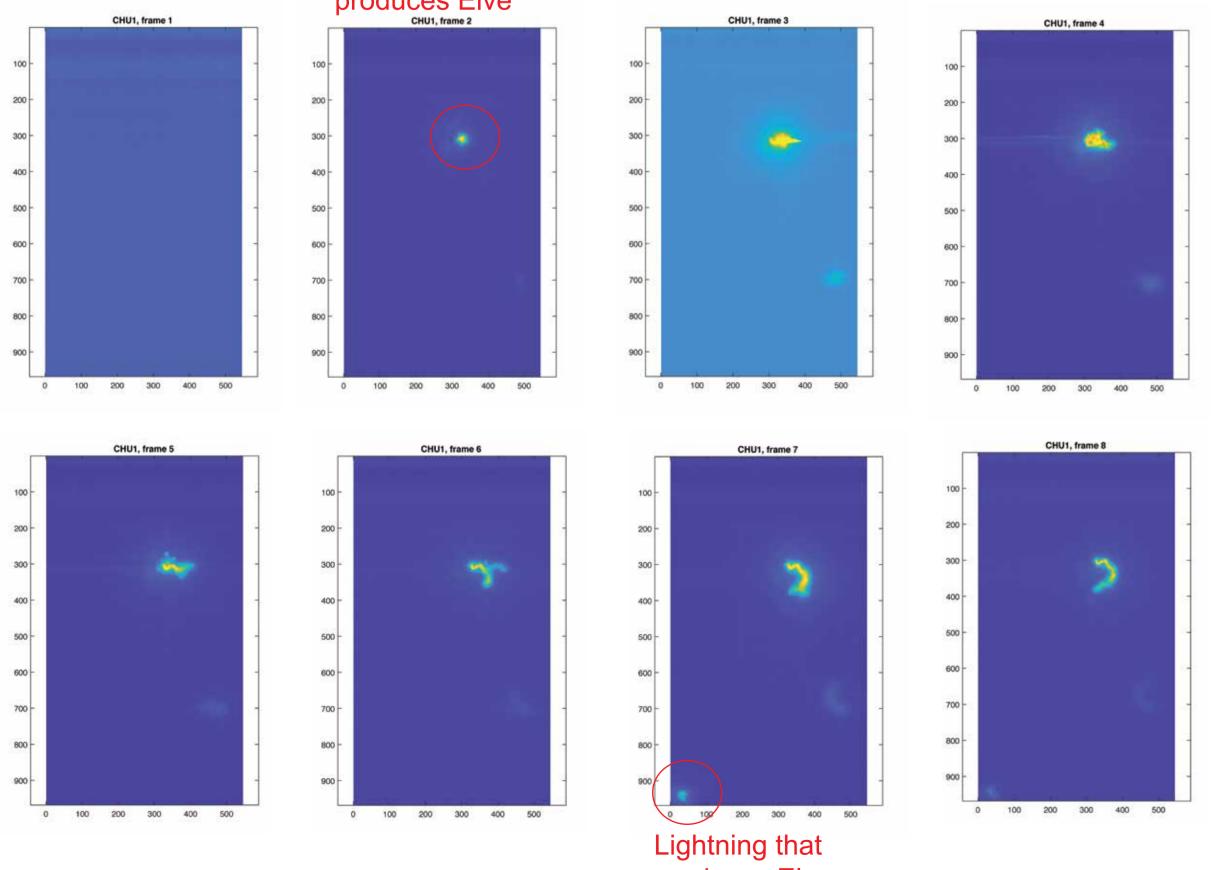
Two Elves are 448.46 ms apart

4. WWLLN and VAISALA also 448.46 ms apart

5. WWLLN and VAISALA gives absolute timing

Location of lightning activity

Lightning that produces Elve



produces Elve

The simultaneous WWLNN and VAISALA data locate the two Elves about 300 km apart

One TGF and two Elves produced by the same thunderstorm system

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				,	3) University 4) Duke Universi de Astrofisica de Andalucia	
					6) Department of Physics, I (Nikolai.Ostgaard@	
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	1		23			
 The sequence of events (see 1. Weak signals in 337 r Leader steps are seer Interpretation: Leader initi 	nm and 777 nm 1 n in the radio (LF	ms before) - starting -	-1 ms before th		n below):	
TGF starts same time Interpretation: Fast and la	' emission (20 us nolecules that em (+/- 80 us)	s delayed to hits in the LH e simultened	HB. Rise time is ous with the on	s ~100 us a set of TGFa	nd duration ~1ms.	
3. Optical pulse in 777 n Interpretation: The delay the cloud (see below the stro	can be explained modeling). The	d by scatteri	ng of 777 nm e	emission wit		
Signal in the 337 nm starts wi the large pulse as in 777 nm.	th a small jump s	simultaneou	s with the UV a	and then a s	slower pulse similar to	
Fast and large current pulse of	vives the electror	naghnetic p	ulse (dj/dt).			
The radio indicates an IC+		0				
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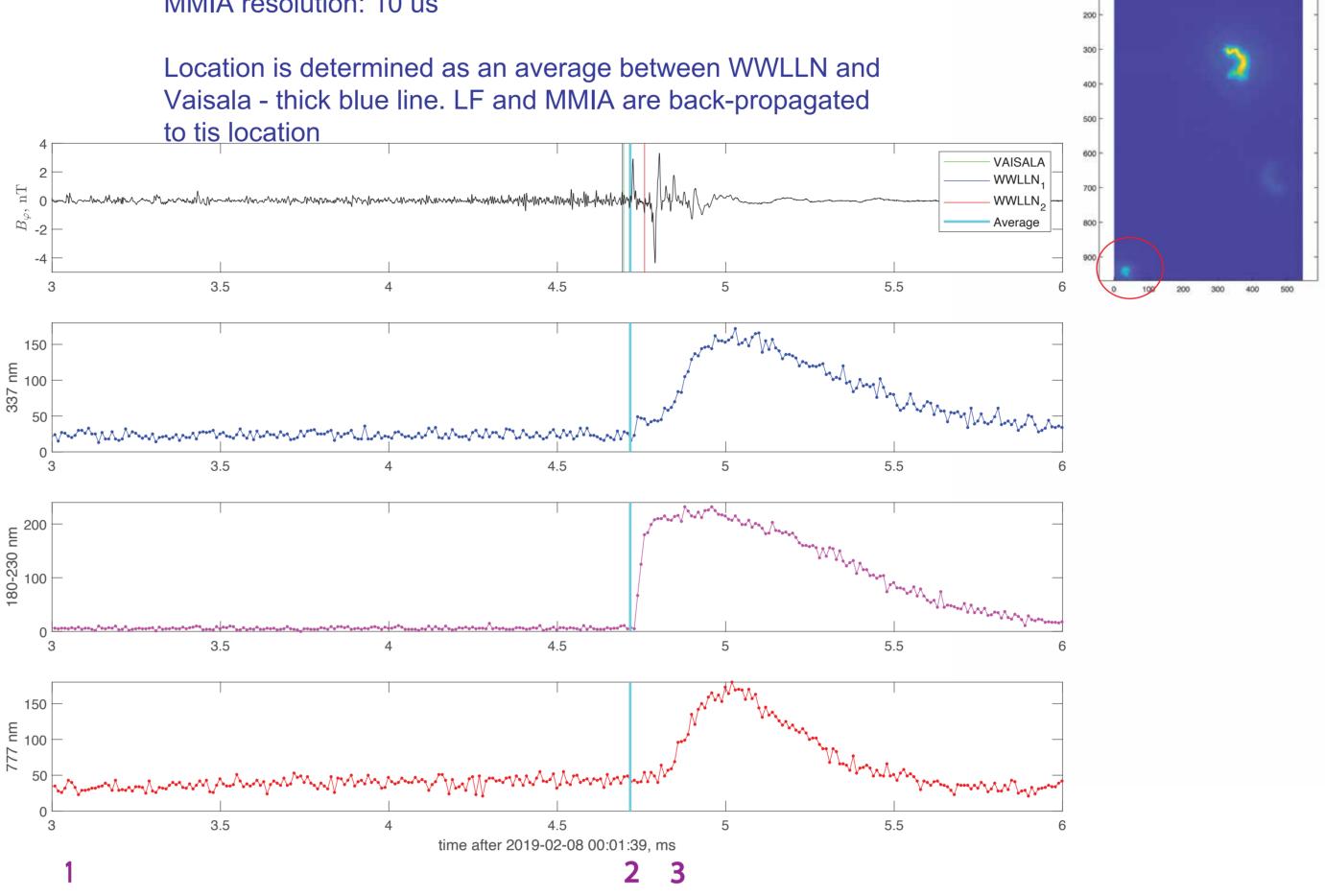
5.564 5.566 time after 2019-02-08 00:01:38, µs

uib.no [+47 5558 2794])

Second Elve from CG-

500 us between ticks

MMIA resolution: 10 us



The sequence of events:

- Continuous activity seen in radio (LF) 2 ms before the Elve Interpretation: Leader initiation and propagation
- Fast and large amplitude current is seen in radio Elve is seen in the UV emission (delayed 12 us to LF onset) with rise time: ~70 us and lasts for ~1ms
- Optical pulse in 777 nm ~100 us delayed to LF. Rise time: ~190 us, and a duration of 800 us. Interpretation: The delay and duration of the 777 nm emission can be explained by scattering through the cloud, but without any scattering from cloud below (CG-lightning) - see modeling

Signal in the 337 nm starts simultaneous with the UV and then similar to the large pulse as in 777 nm

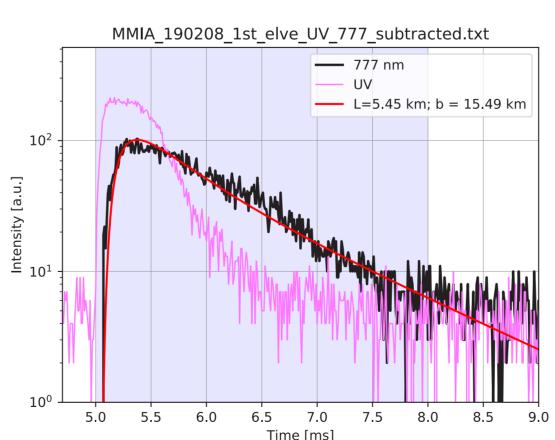
Fast and large current pulse gives the electromaghnetic pulse (dj/dt).

The LF radio from Puerto Rico indicates an CG- Opposite polarity than the first one

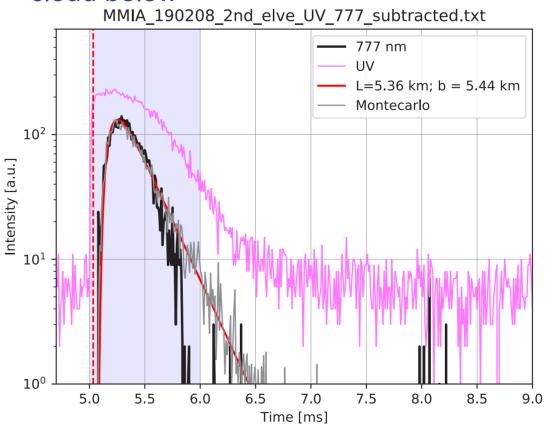
The scattering of 777 nm emissions within the cloud

Homogeneous cloud bounded by two parallel planes, droplets of 10 um, and density of 100/cm3 L: distance from cloud top to source b: is the thickness of cloud

1st Elve: If cloud top at 15-16 km, the IC+ is at ~10 km, reasonable for IC+. Long duration is due to scattering from below



2nd Elve: If cloud top at 10 km, the CG- is at ~5 km, but more important: It is from the bottom of the cloud, consistent with CG-. Short duration is due to no scattering from cloud below



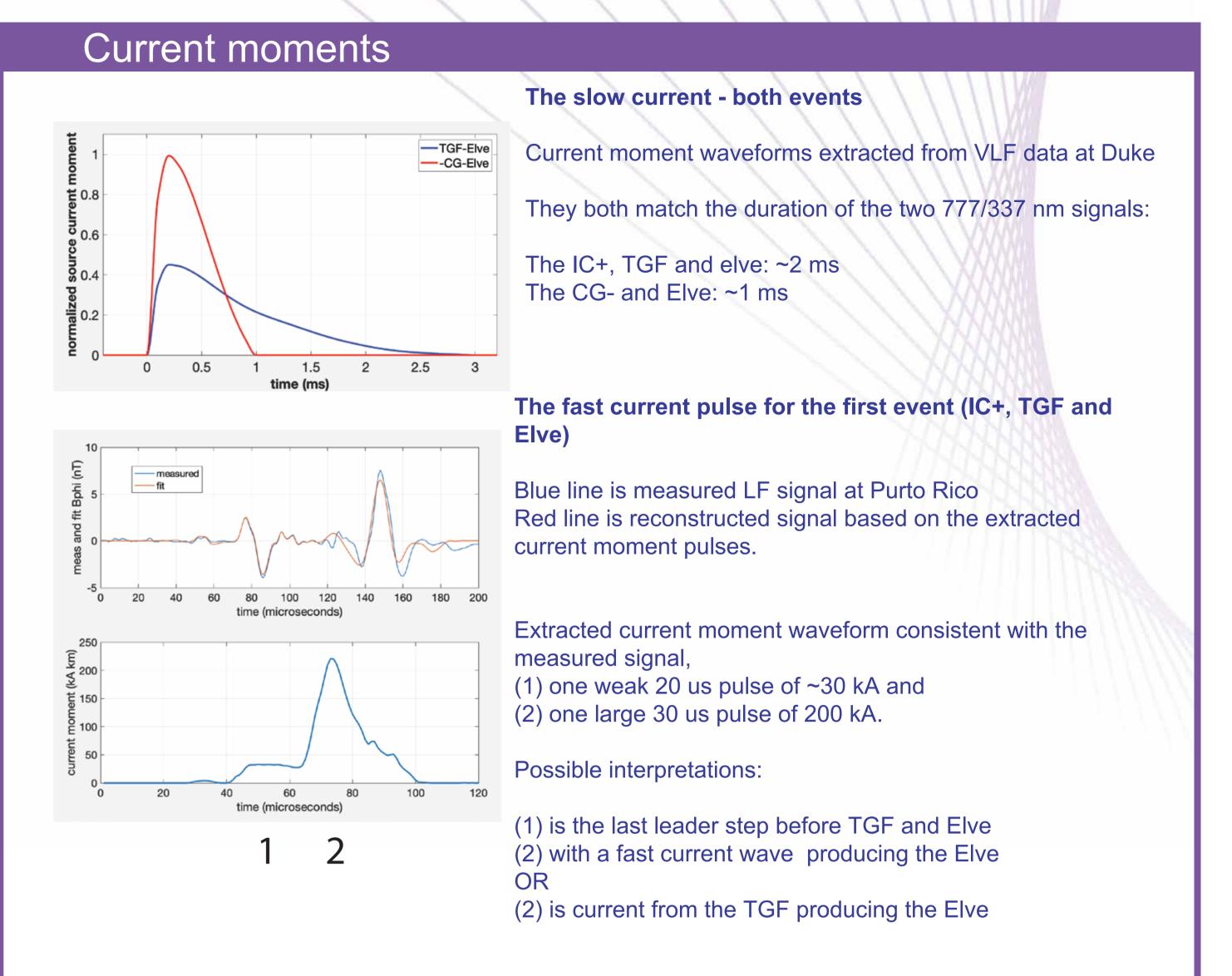
Acknowledgement

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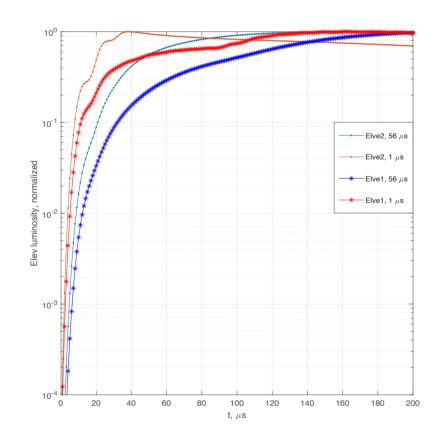
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Modeling Elve signal in UV for the IC+

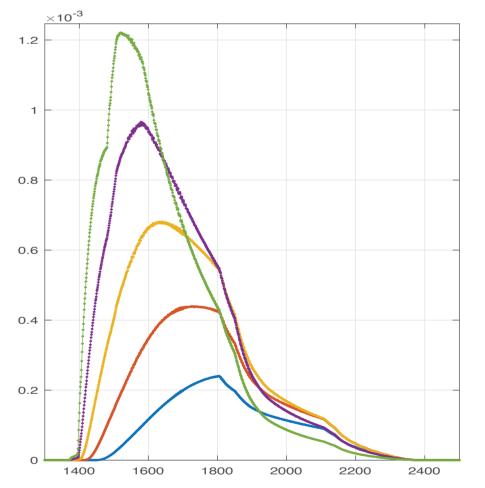
Input for modeling: The extracted 30 us current pulse from LF Puerto Rico data

The lifetime of excited states of N2 molecules of 56 us which gives a delay relative to direct EM wave of 20us and 12 us for 1st and 2nd Elve, respectively (1.5 order of magnitude down from peak)



The different colors are for different directional diagrams meaning different angles of the main lobe

Green curve has rise time (100 us) consistent with measured UV. All have compatible duration of ~1 ms



Summary

1) Two Elves were observed ~456 ms and 300 km apart.

2) IC+ produced an Elve and a TGF

3) 456 ms later a CG- produced an Elve but no TGF was observed from space

4) LF radio measurements of the IC+ show that the weak optical signals (337 nm and 777 nm) before the Elve/TGF are associated with upward negative leader steps.

5) LF measurements for 2nd Elve show it is an CG-, preceeded by continuous downward leader propagation

6) The UV signal is delayed 20 us and 12 us to the onset of the LF pulse, due to geometry and lifetime of excited states of N" molecules.

7) The optical emissions in 777 nm are delayed to the onset of the LF pulse by 90 us and 100 us, respectively. Both the delays and duration of 777 nm emissions can be explained by scattering within the cloud.

8) Reconstructed current moments give:

a) one fast (30 us) and large (200 kA) pulse emitting an electromagnetic wave that produces an Elve.

b) one slow (1-2 ms) current pulse, that may contribute to the optical signal.

9) Open questions:

For the IC+: Is the fast and large current a fast current wave from a leader step that produces both electromagnetic wave and a TGF, OR: is the current from the TGF itself?