



# A new instrument in earthquake early warning system by detection and modeling of prompt gravity signals

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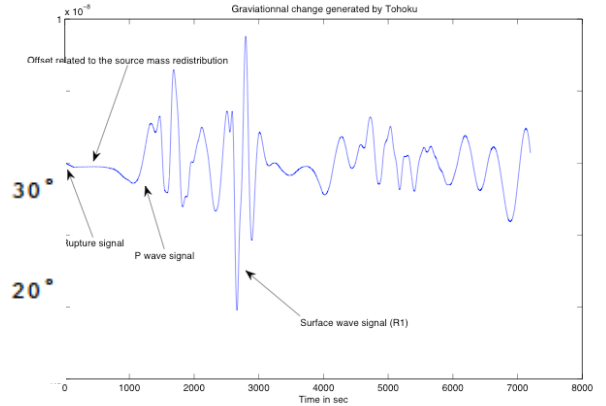
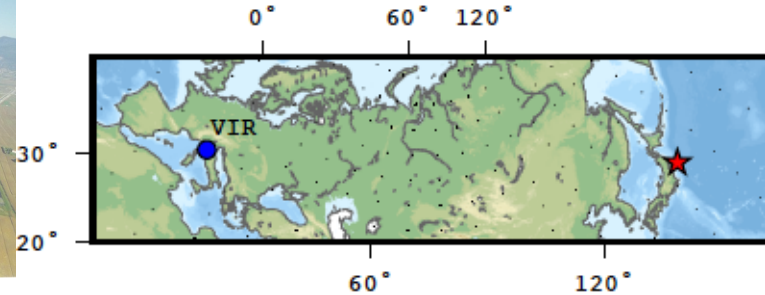
# Outline: The story and the approach

- **Motivation (2012)**

VIRGO  
(*Gravitational wave Interferometer*)



## normal mode calculation

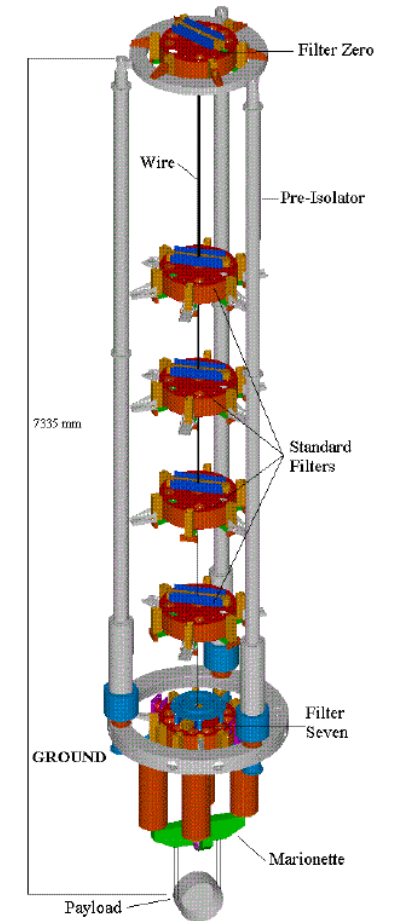
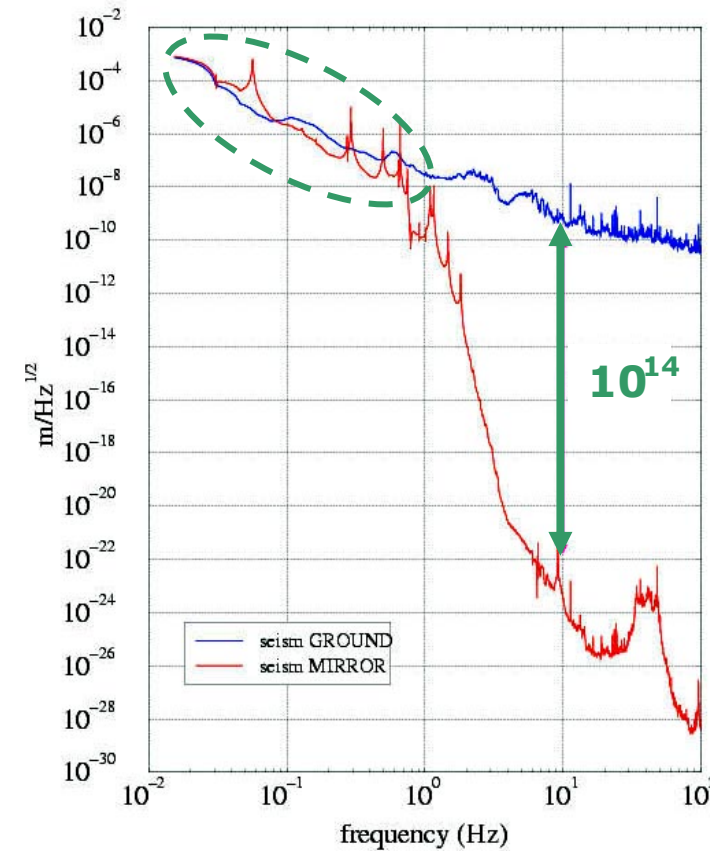


- **Search for « instantaneous » gravity signal**
  - a) Very Large earthquakes: Japan-Tohoku (9.0:-; 11/03/2011)
  - b) Instruments: Superconducting gravimeter, very broadband seismometers STS1
- **EEWS (Earthquake Early Warning Systems)**
- **Search for a prompt gravity signal:** Kamioka SG data very broadband seismic data  
Montagner et al., Nature Com., 2016, Vallée et al., Science, 2017, Vallée and Juhel, JGR, 2019)
- Controversy (Heaton, Nature Com., 2017)
- **Need for New Instruments: TORSION Bars**
- **Perspectives and ongoing projects : EEWS (Earthquake Early Warning Systems)** (Juhel et al., JGR, 2018)
- **PEGASEWS: Prompt Earthquake GrAvity Signals: Early Warning Systems**

# LABEX UnivEarthS (2012)- Geophysics

## Gravitational wave interferometers: VIRGO

VIRGO (Italian – French Gravitational wave detector)

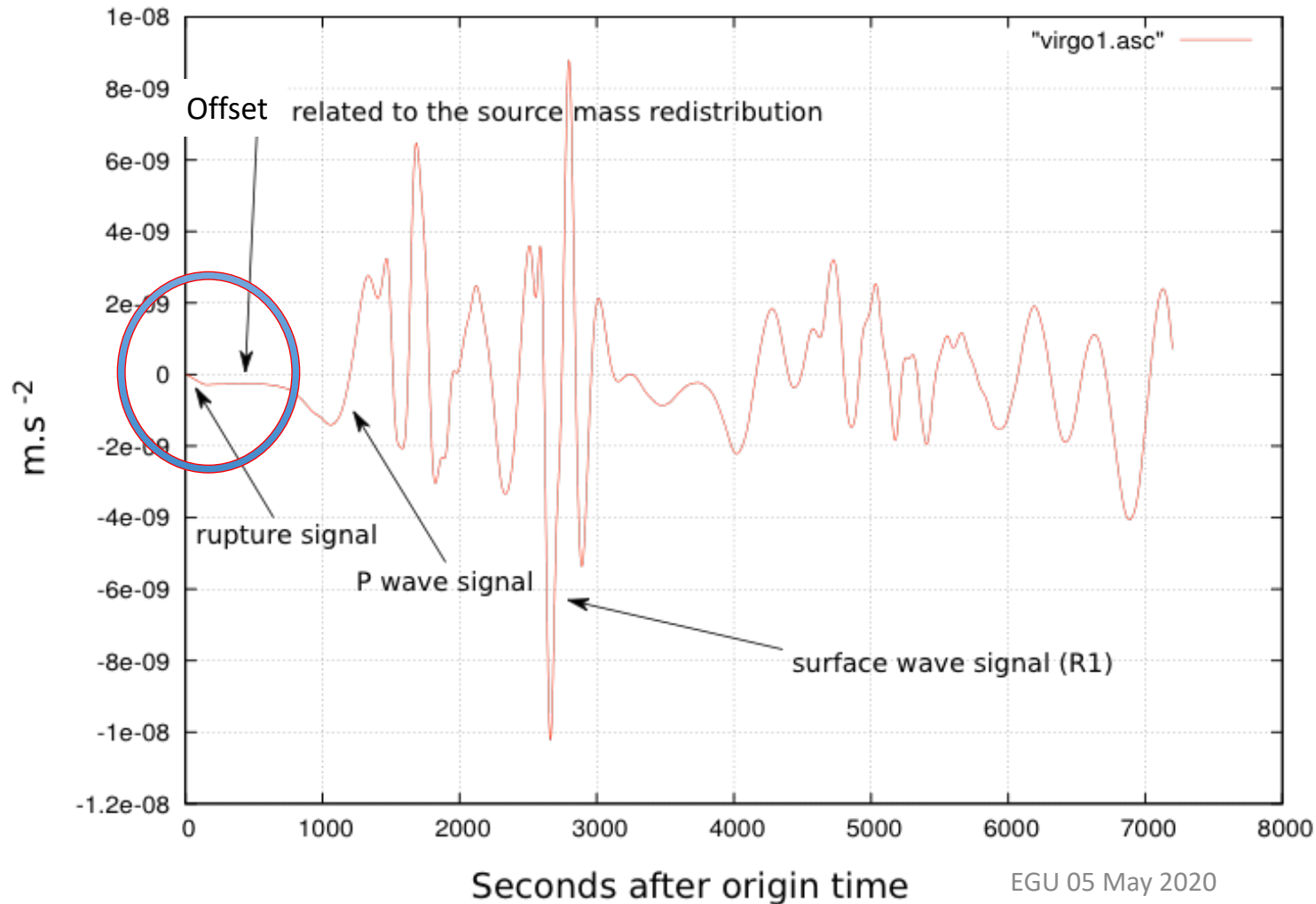


**Seismic wall at  $f > 1$  Hz**

# Motivation

## NORMAL MODE THEORY

Simulated gravitational change at the Virgo site . Source: Tohoku  
Z component low-pass filtered at 100s



Theoretical signal  
Before P-arrival:  
 $0.03 \mu\text{gal}$   
Very small

$$1 \mu\text{gal} = 10^{-8} \text{ m/s}^2 = 10 \text{ nm/s}^2$$

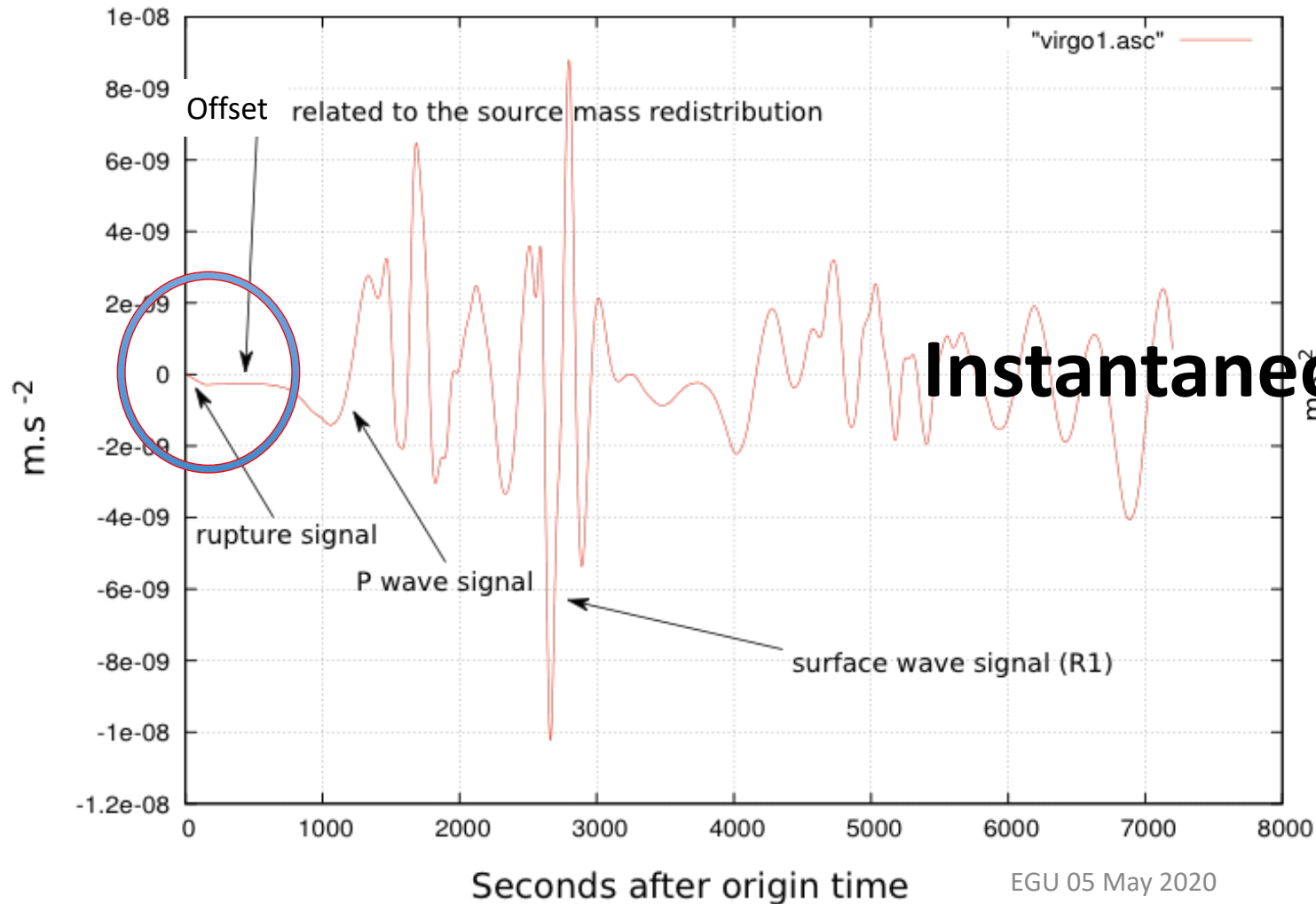


# Motivation

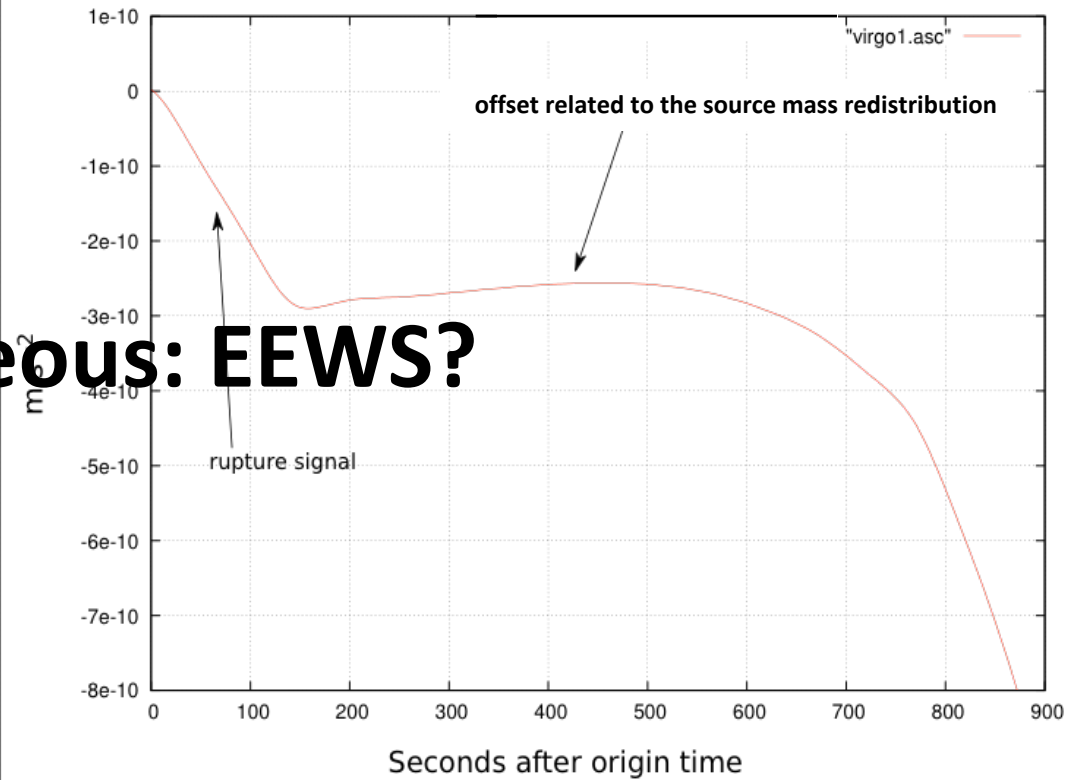
## NORMAL MODE THEORY

Simulated gravitational change at the Virgo site . Source: Tohoku  
Z component low-pass filtered at 100s

$\Delta \approx 90^\circ$  ( $\approx 10\,000\text{km}$ )

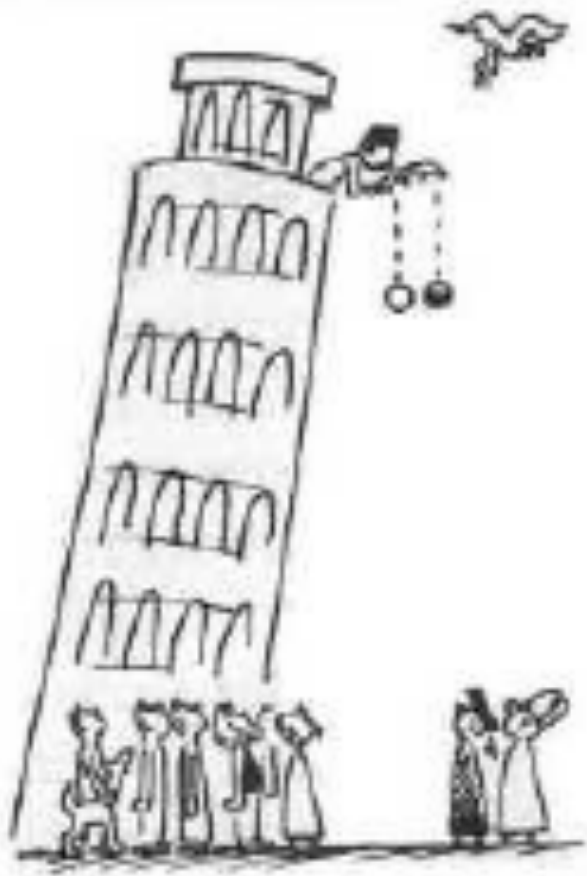


Simulated gravitational change at the Virgo site . Source: Tohoku  
Z component



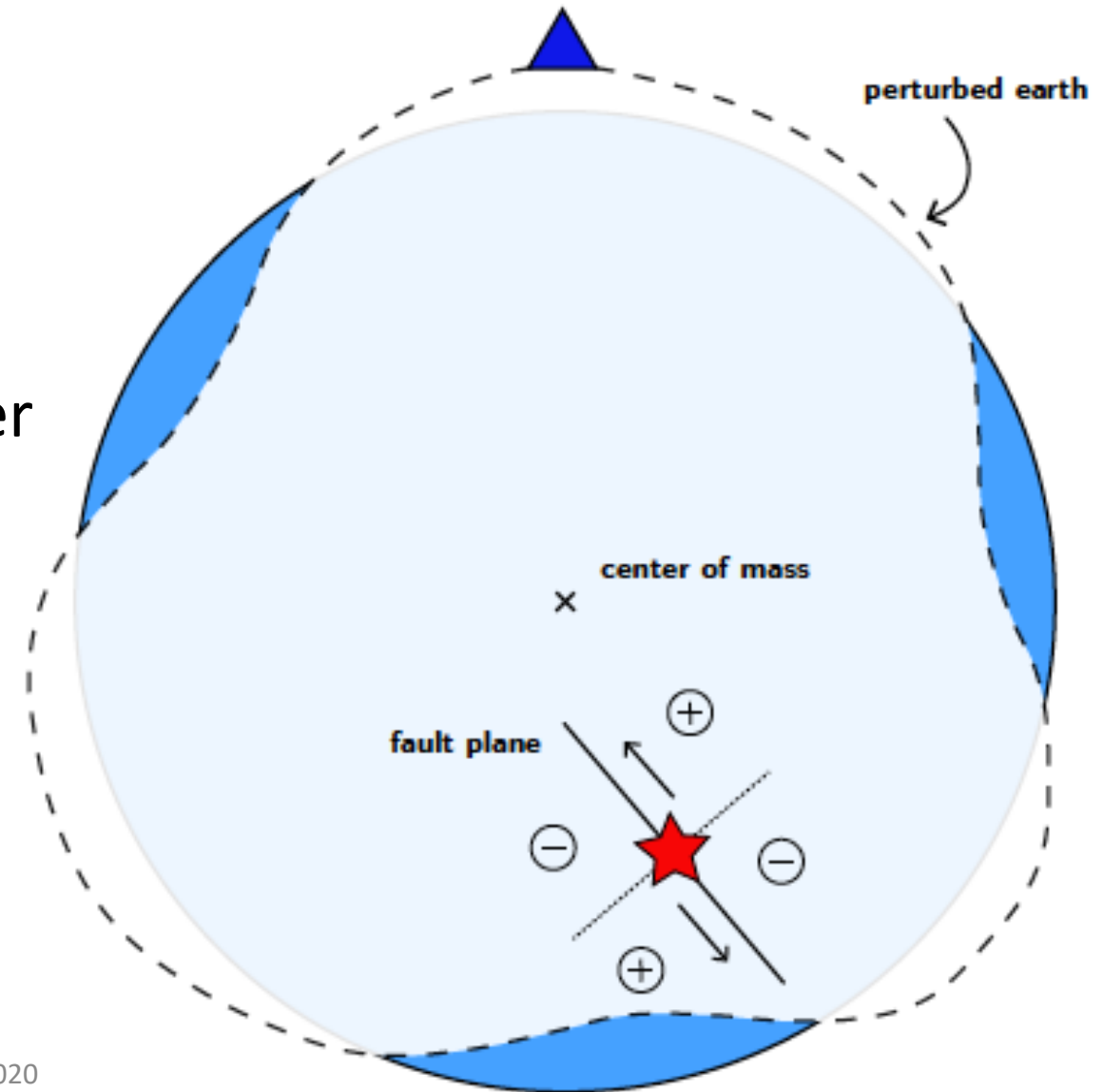
**Instantaneous: EEWS?**

# Gravity perturbations induced by earthquakes?



Galileo Galilei  
Free fall at Pisa Tower  
(1604)

Isaac Newton:  
universal gravitation law (1685)

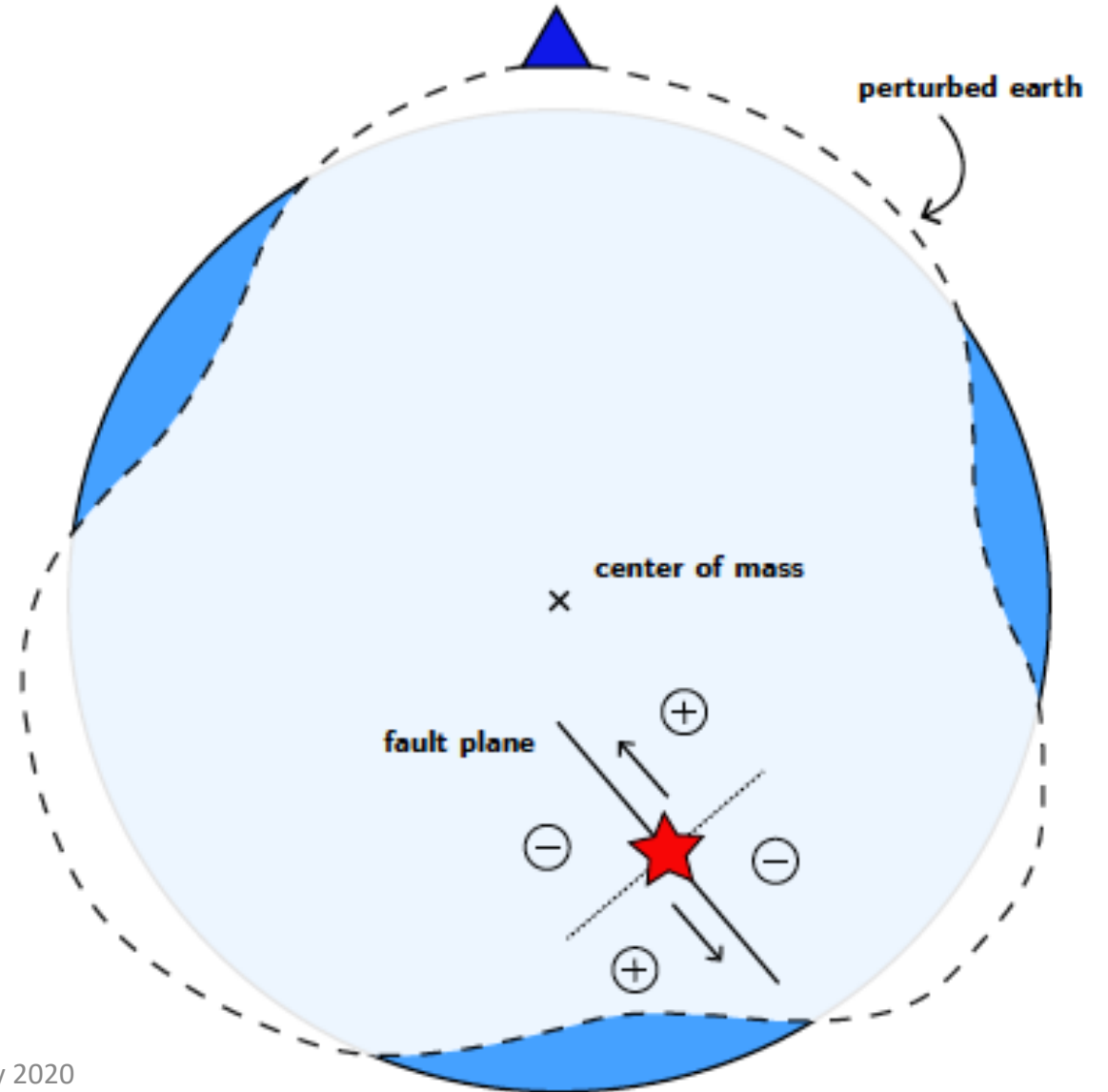


# Gravity perturbations induced by earthquakes

- Mass redistribution  $-\nabla(\rho_0 \mathbf{u})$
  - Free air gravity anomaly: perturbation of the Earth surface
- Okubo, GJI, 1991, JGR, 1992, ...

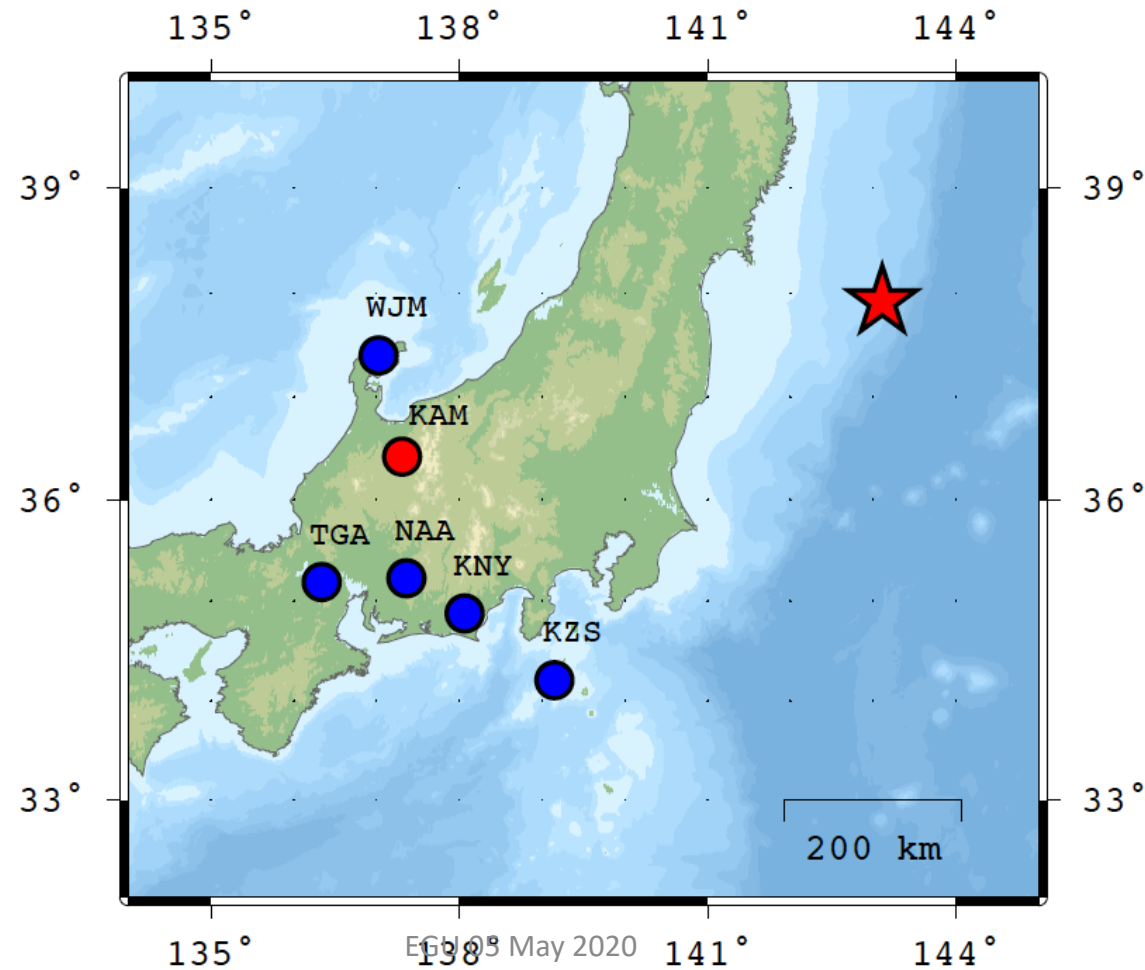
## Theoretical Approach

- direct numerical calculations (plane case):  
(Harms et al., GJI, 2015, 2016)  
(Vallée et al., Science, 2018)
- Normal mode Theory (spherical case)  
(Juhel et al., GJI, 2018)

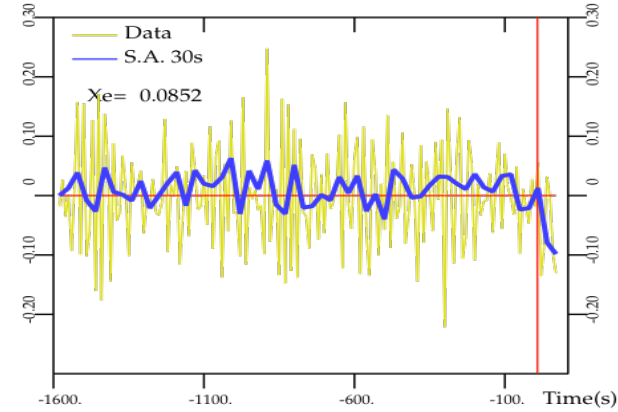


# From Static to Dynamic gravity changes induced by earthquakes

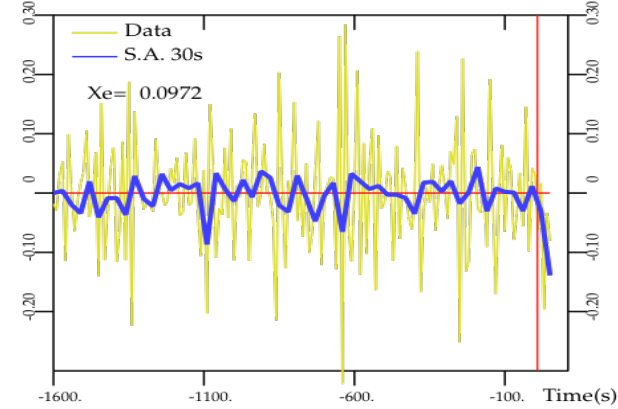
***Superconducting gravimeter Kamioka  
+ Broadband Japanese network F-NET (STS1, STS2...)***



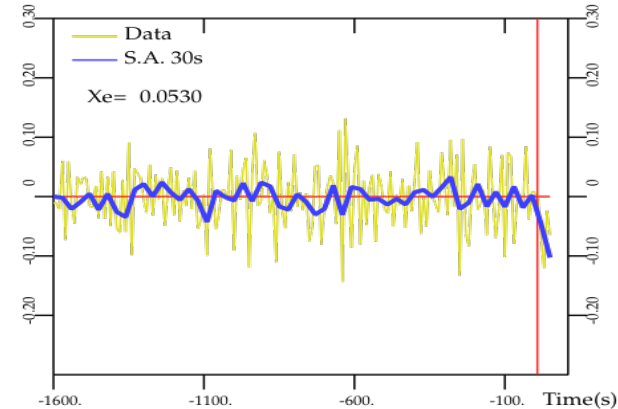


KNM- time dependence seismic anom. 0.04 40pt sigmap0.5  
dG microGal

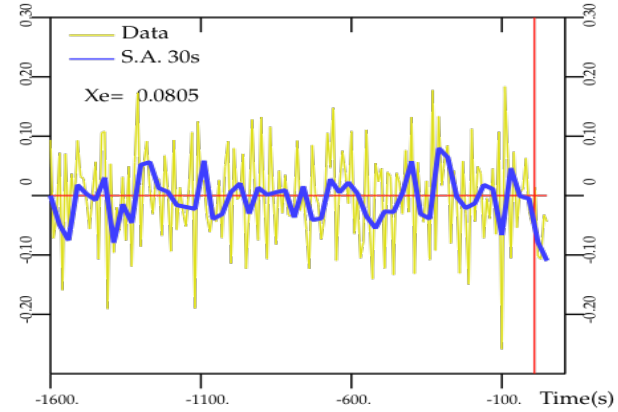
KZK- time dependence seismic anom. 0.04 40pt sigmap0.5  
dG microGal



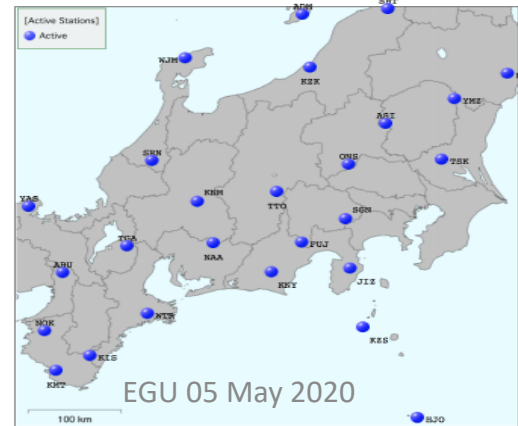
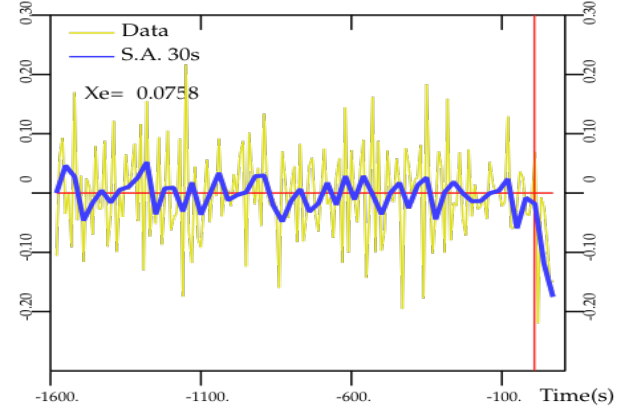
STACK 4 stations FNET 0.04 40pt sigmap0.5  
dG microGal

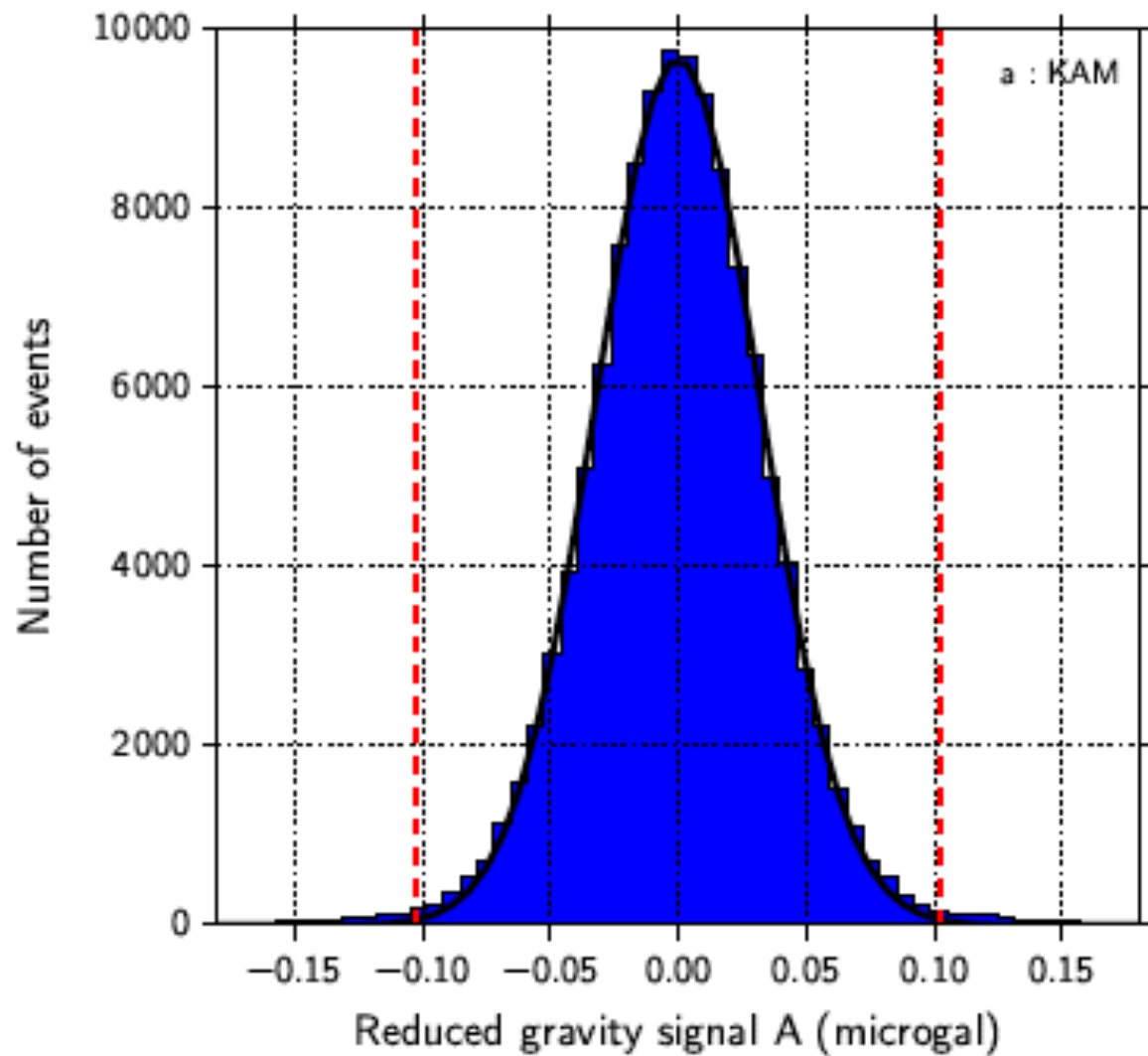


ONS- time dependence seismic anom. 0.04 40pt sigmap0.5  
dG microGal



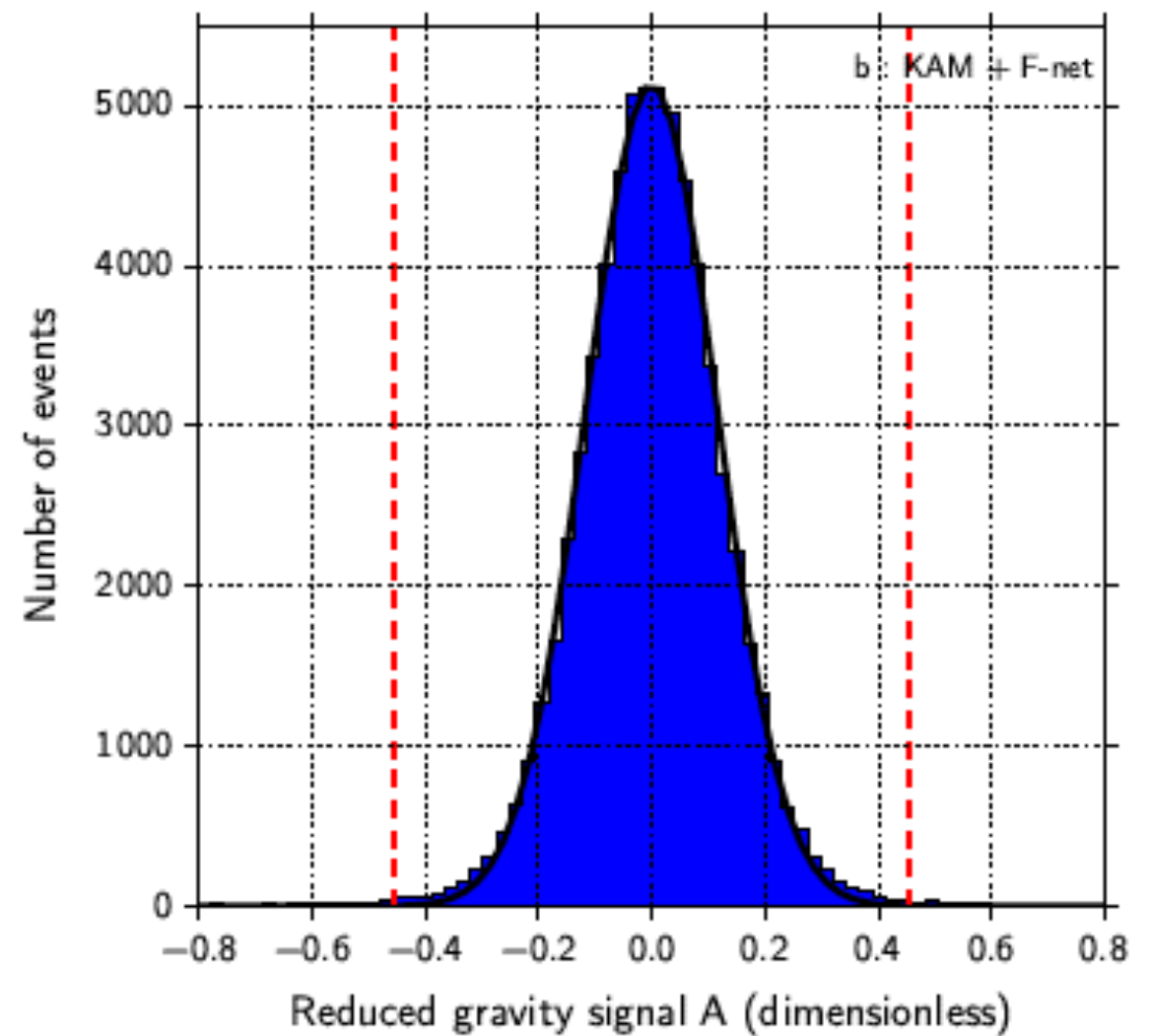
WJM- time dependence seismic anom. 0.04 40pt sigmap0.5  
dG microGal





For the Kamioka station only :

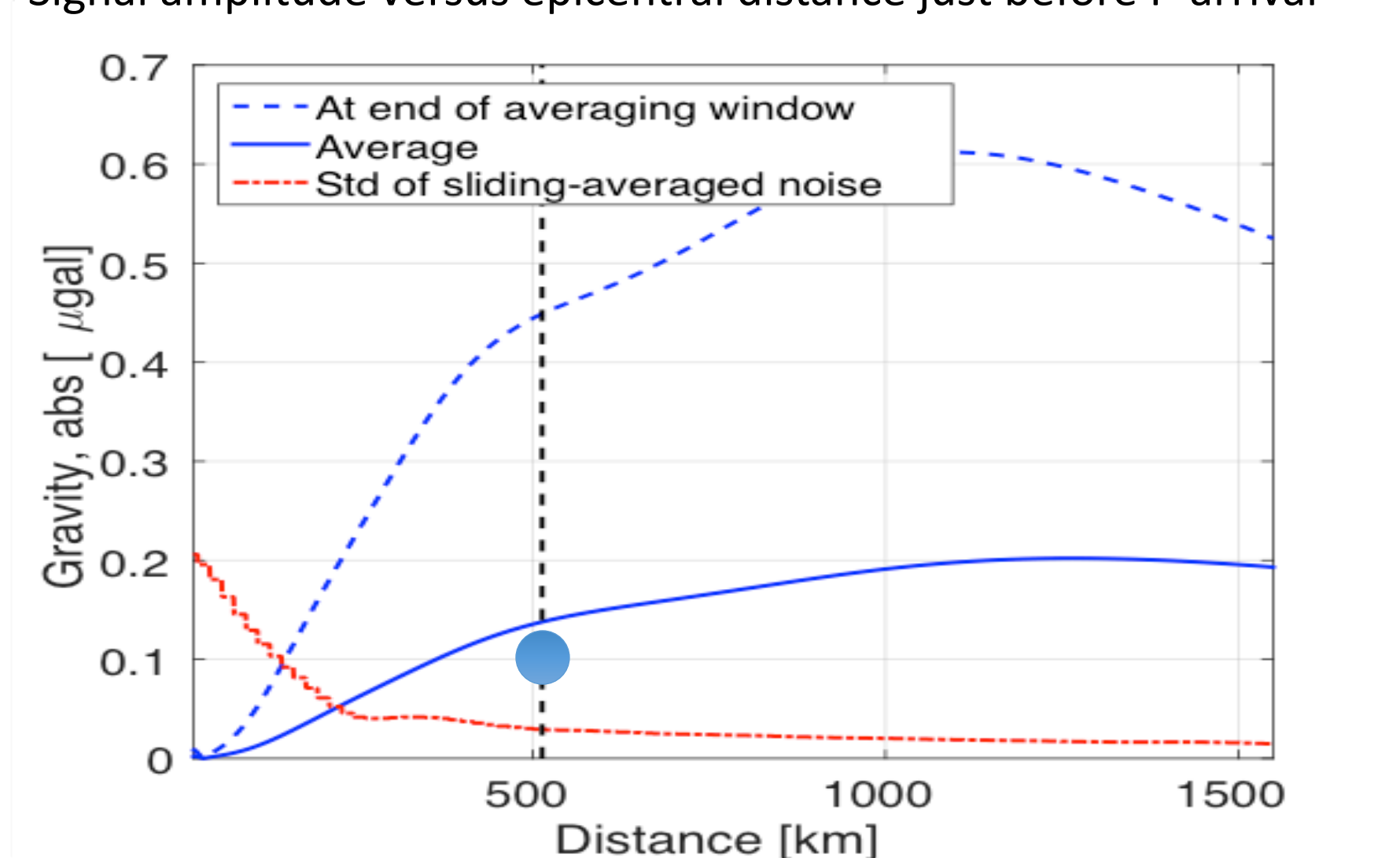
$$p(A > A_{Tohoku}) = 1.6\%$$

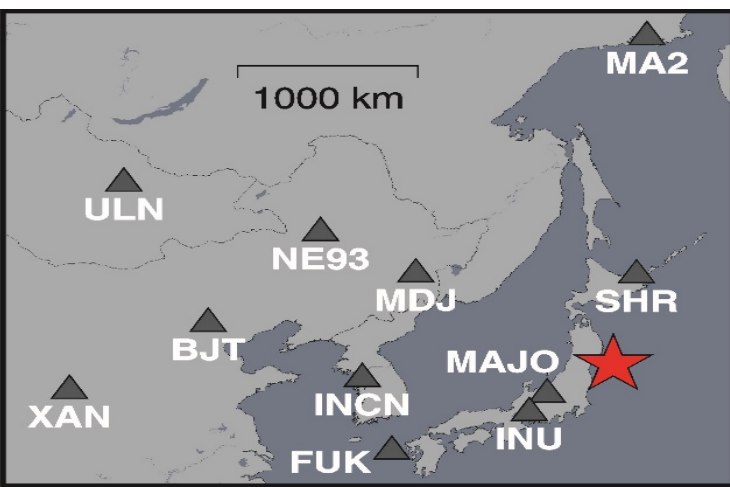


For a stacked waveform :  
Kamioka + 5 F-net

$$p(A > A_{Tohoku}) = 0.82\%$$

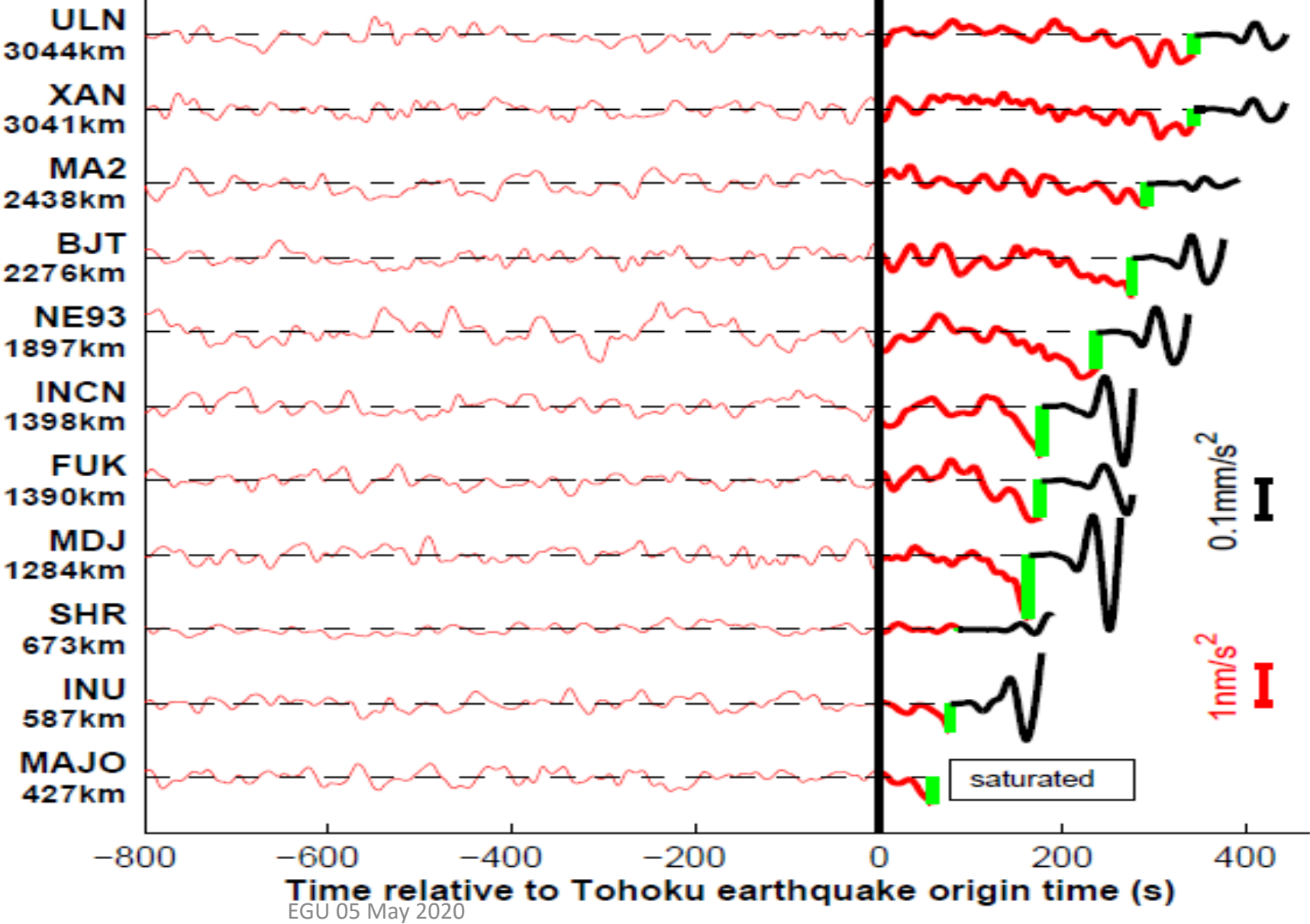
## Signal amplitude versus epicentral distance just before P-arrival





With seismometers, such a tiny signal requires excellent stations to be recorded

Green tick : P wave arrival



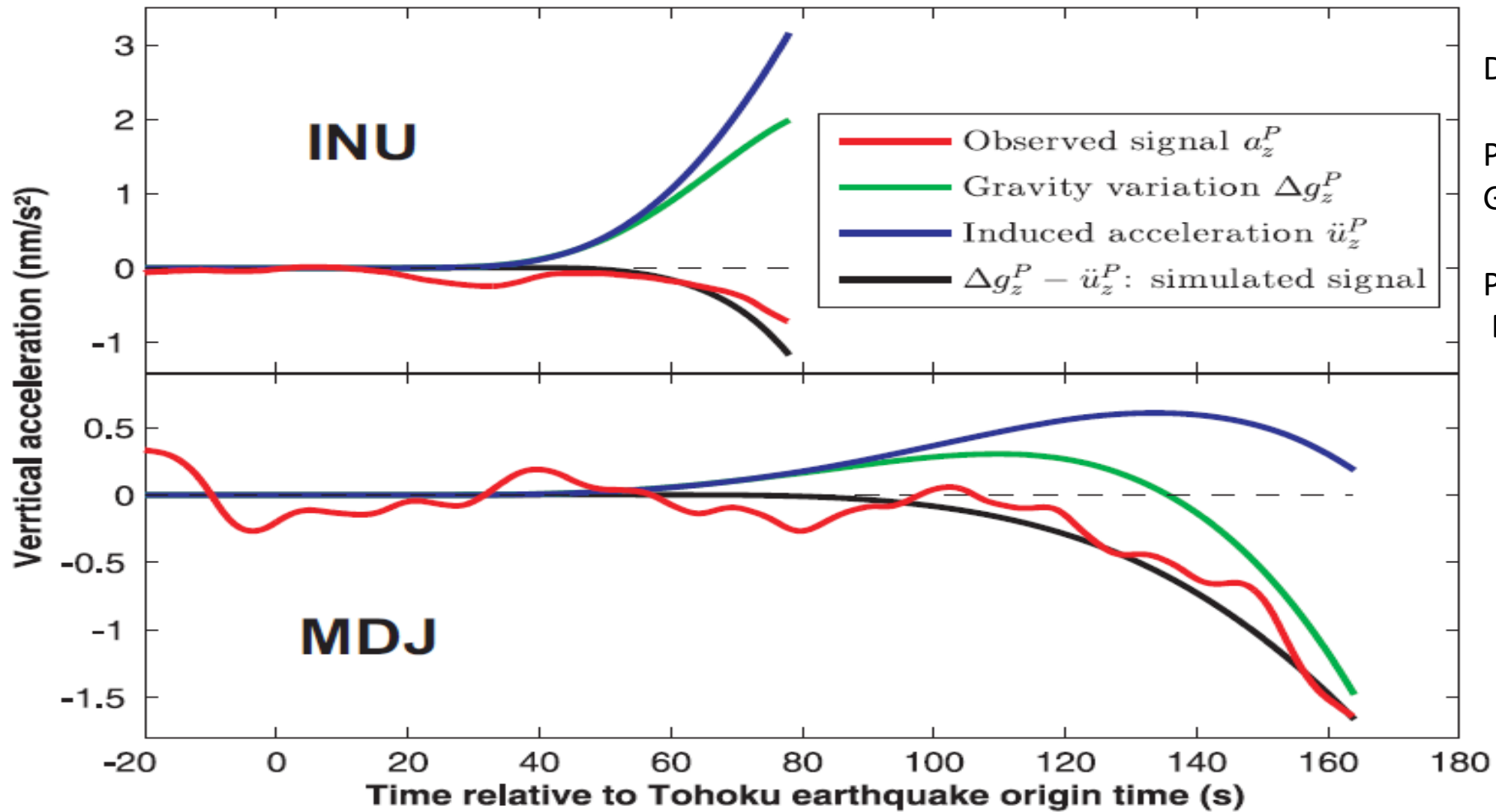
FDSN stations (IRIS or GEOSCOPE) + F-NET (Japan)  
*0.002-0.03Hz frequency range*

Relative amplitudes between the pre-P and the post-P signals

Pre-P signals are 10<sup>5</sup> to 10<sup>6</sup> smaller



## Data & simulations at INU (GEOSCOPE, G) and MDJ (IRIS-China, IC)



DISTINCTION

PGS: Prompt  
Gravity Signal

PEGS: Prompt  
Elasto-Gravity Signal

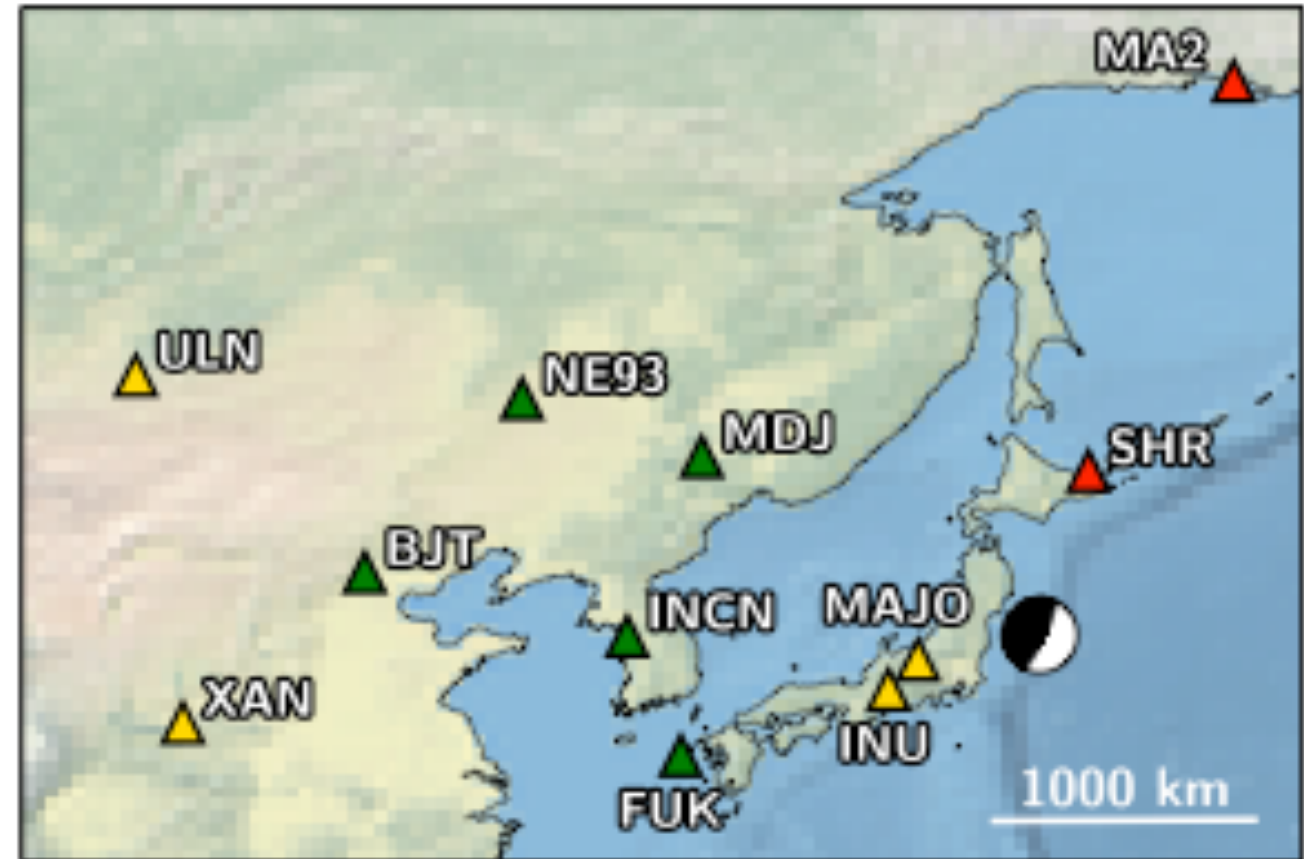
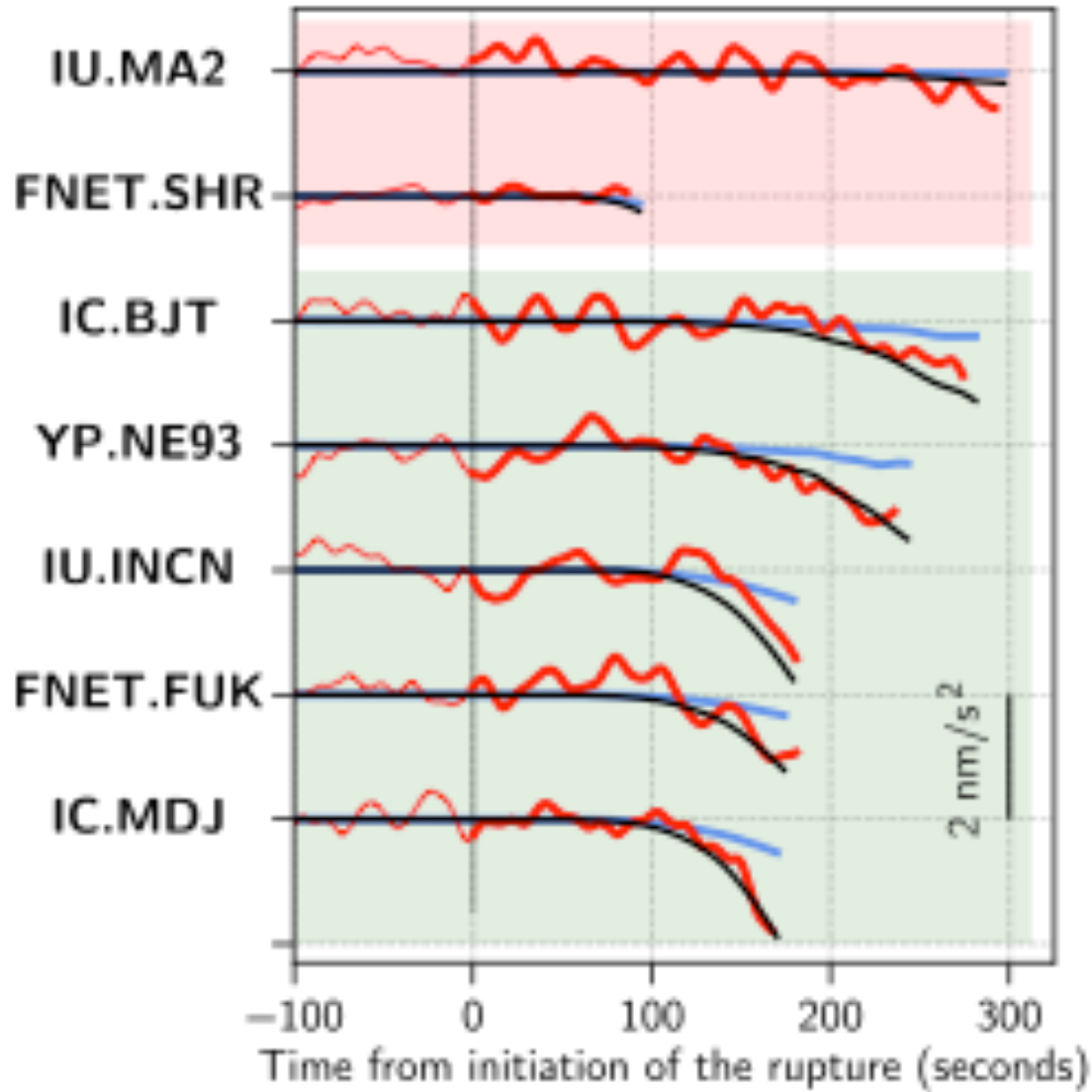
*Heaton, Nature Com.,  
2017*

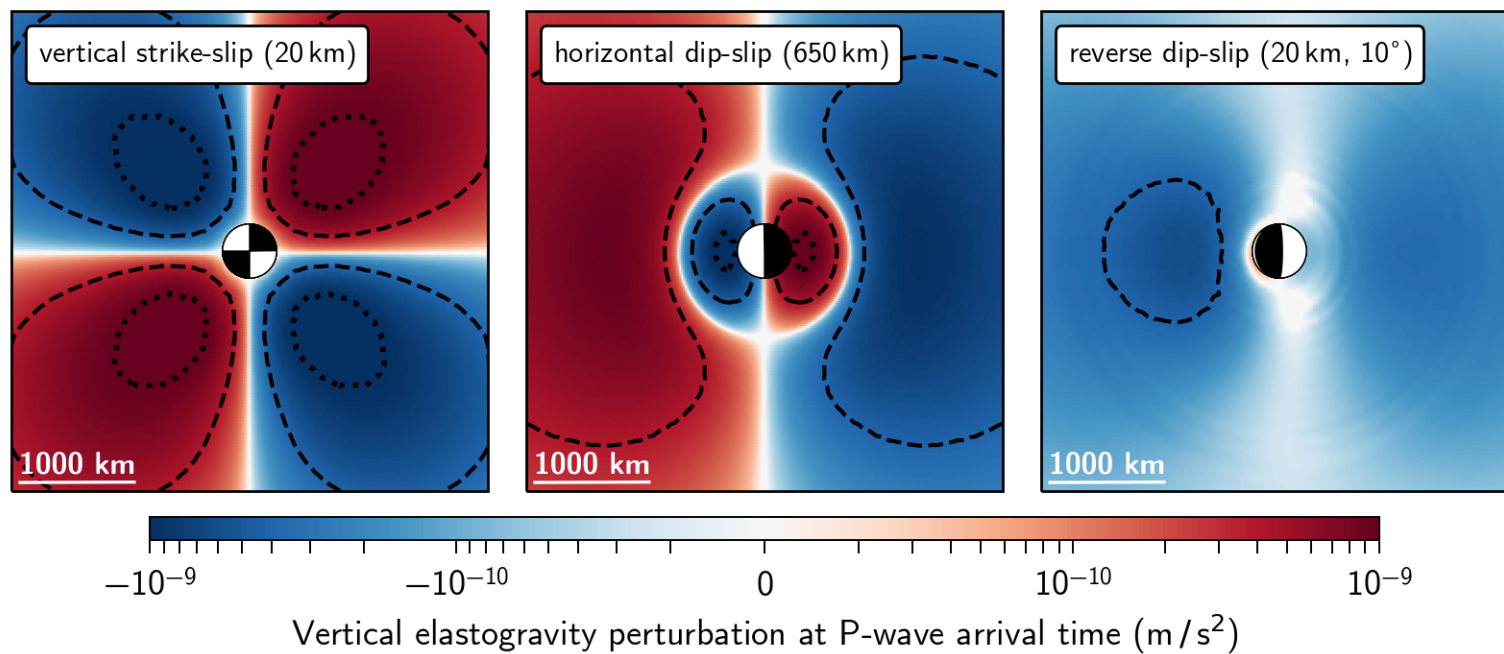
*Vallée et al., Science,  
2017*

*Juhel et al., G.J.I.,  
2018*

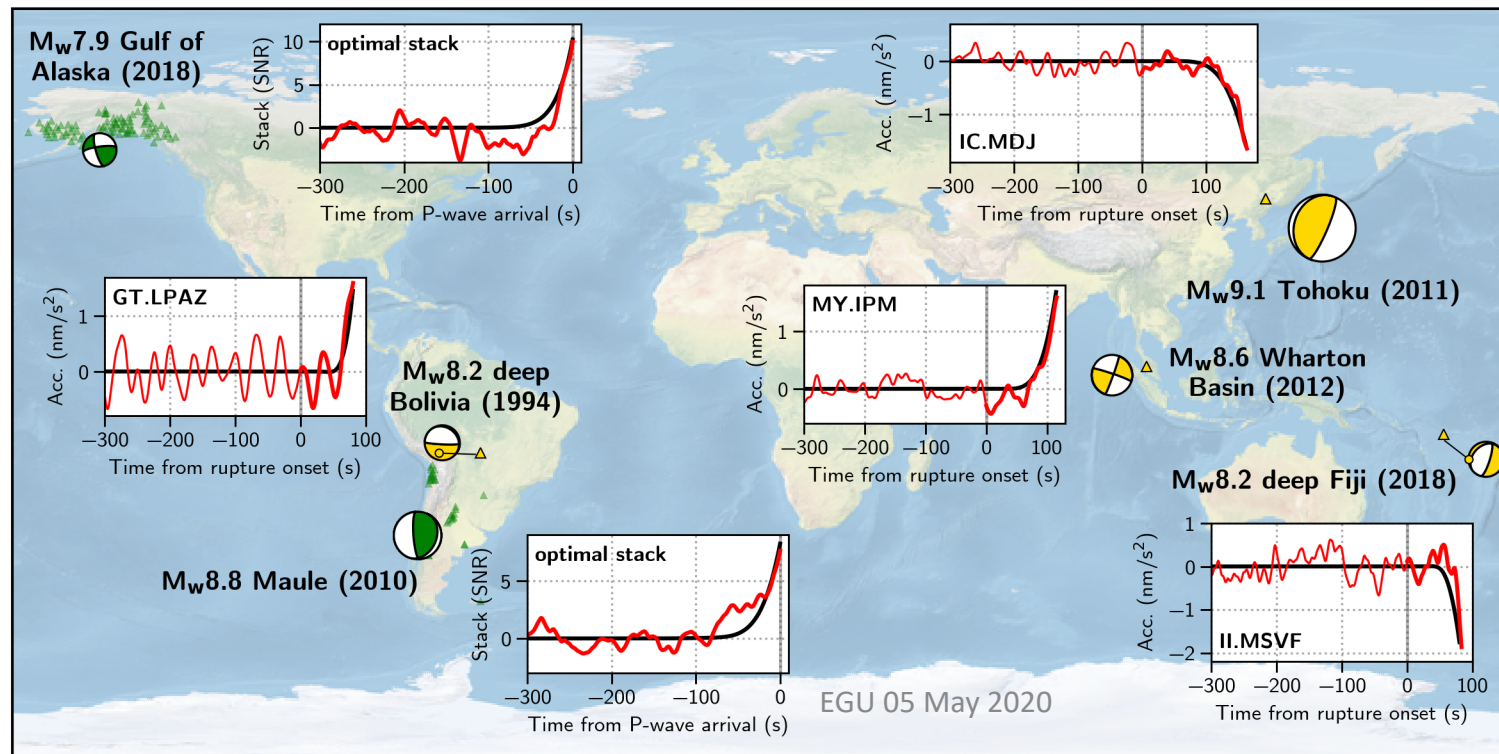
**Prompt gravity signal and inertial acceleration do not cancel**

# Complete simulation at all stations



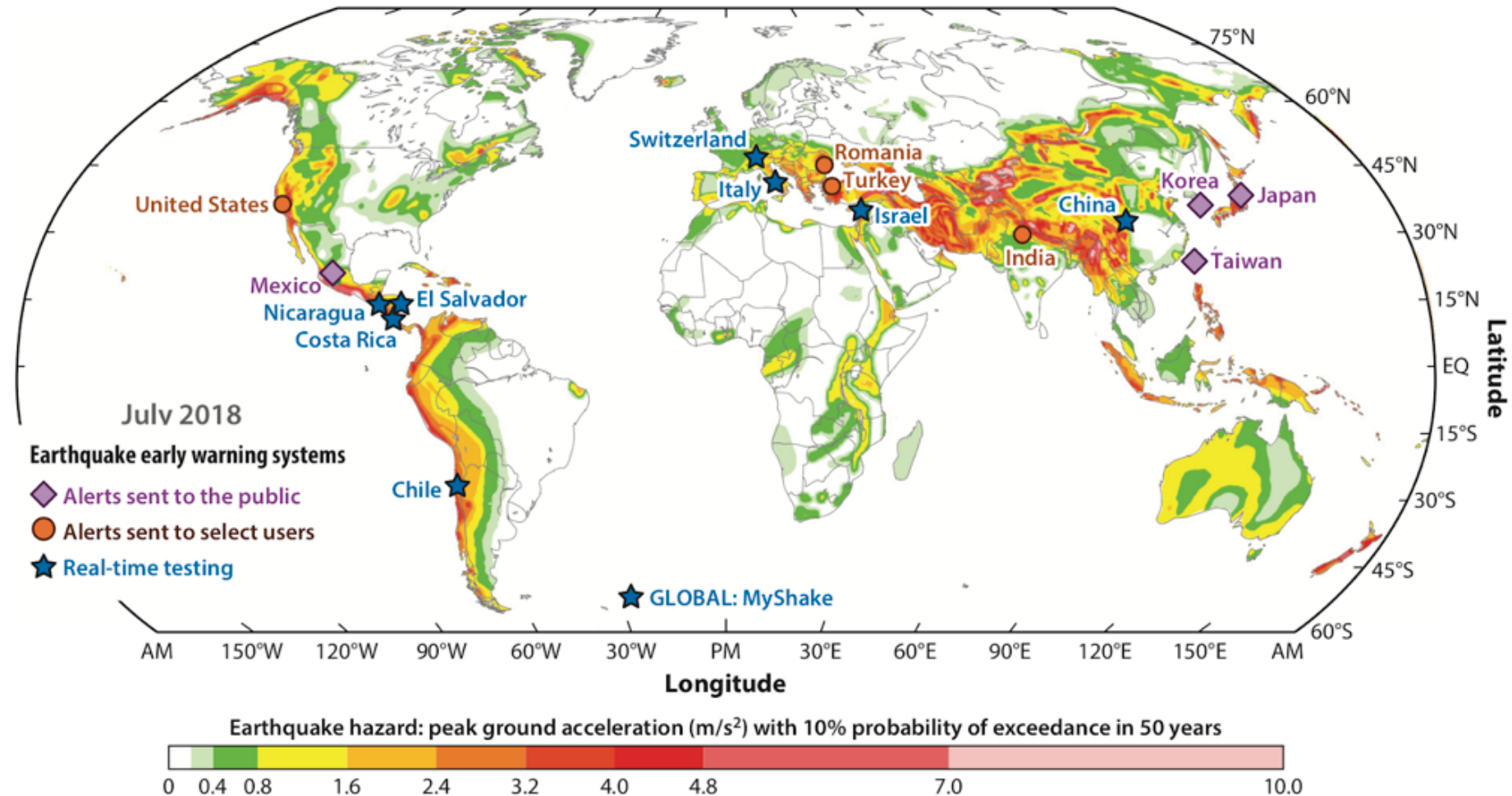


New detection of PEGS  
(prompt elasto-gravity  
signals) for other  
earthquakes



Vallée & Juhel, 2019

# Earthquake Early Warning Systems in the World Based on P-wave detection



Status of EEWS around the world up to July 2018. Shown are the locations of systems that broadcast alerts to all members of the public (purple), systems distributing alerts to select users (orange), and systems undergoing real-time development and testing (blue). The background colors indicate the seismic hazard (see colorbar legend) (From Allen et al., 2019).



# NEED FOR NEW INSTRUMENTS

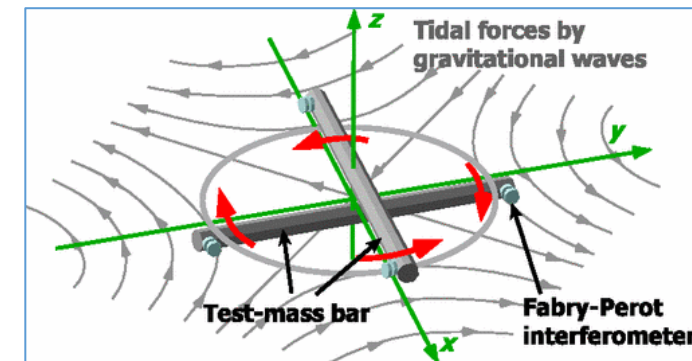
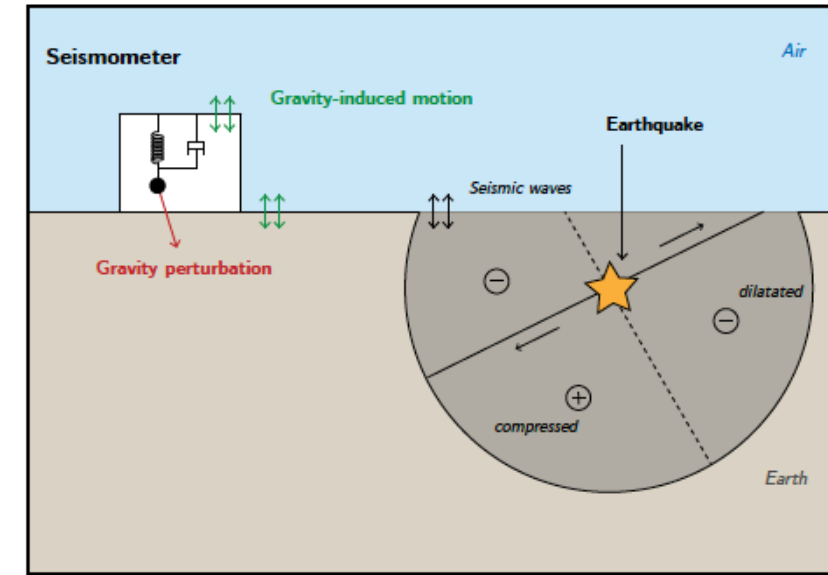
(gradient of the gravity field)

Sub-Hz gravitational –wave detection

- 1) Superconducting gradiometers
- 2) Atom interferometers
- 3) Torsion bar antennas  
(Collaboration with Univ. Tokyo)

## PEGASEWS detector

Prompt Earthquake GrAvity Signals- Early Warning System



# NEED FOR NEW INSTRUMENTS

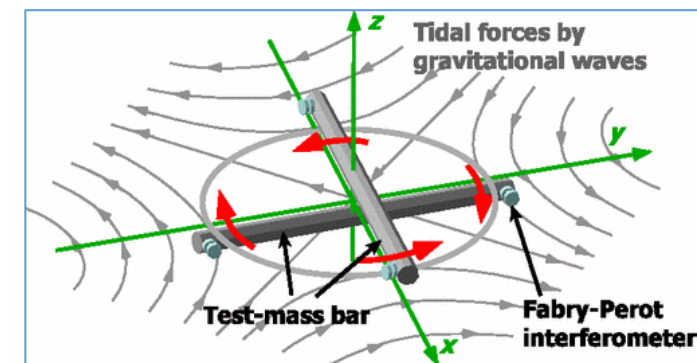
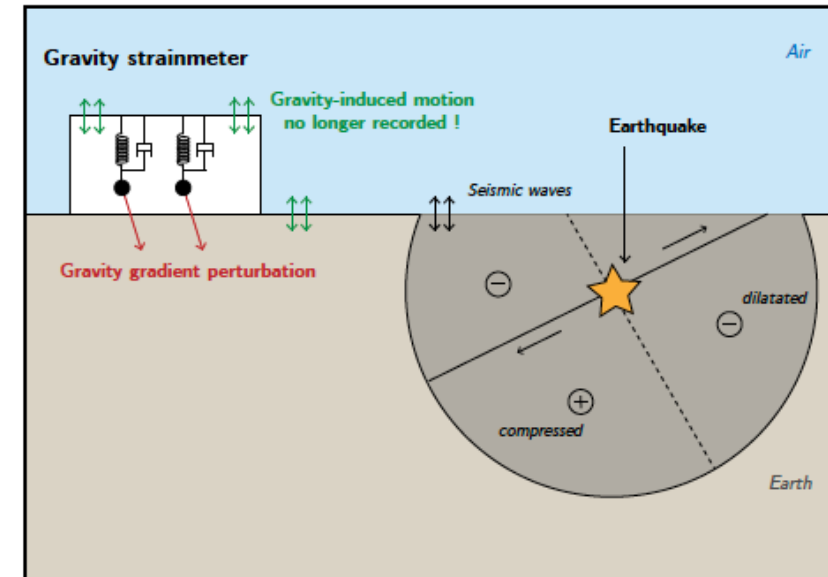
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Sub-Hz gravitational –wave detection

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## PEGASEWS detector

Prompt Earthquake GrAvity Signals- Early Warning System

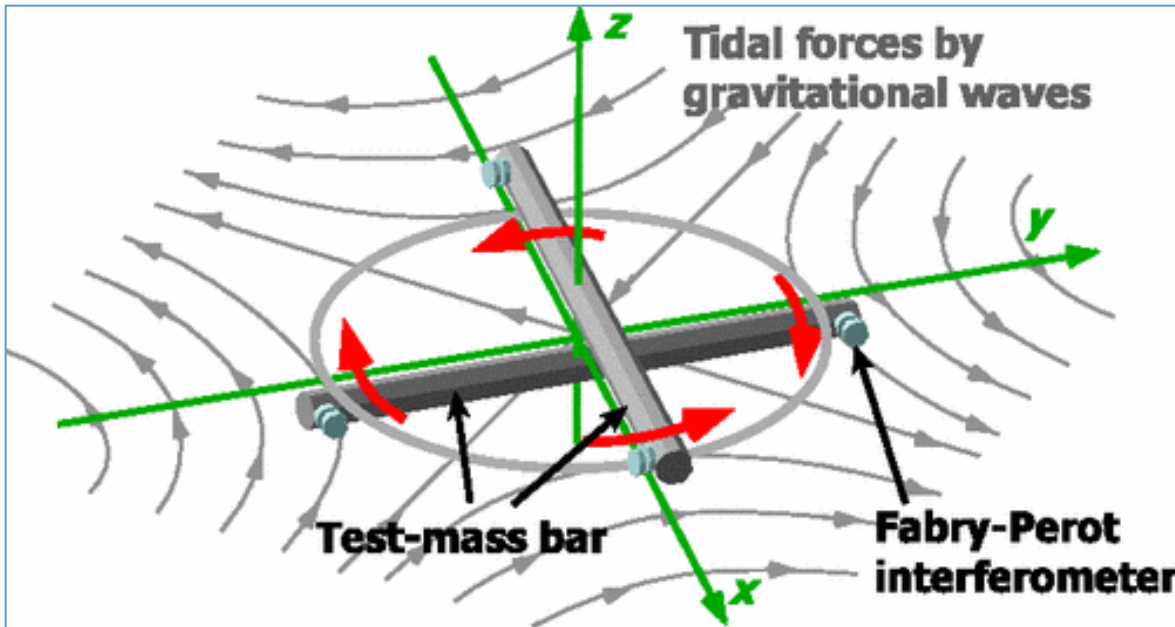


# Sub-Hz Gravitational wave detectors: TOBA

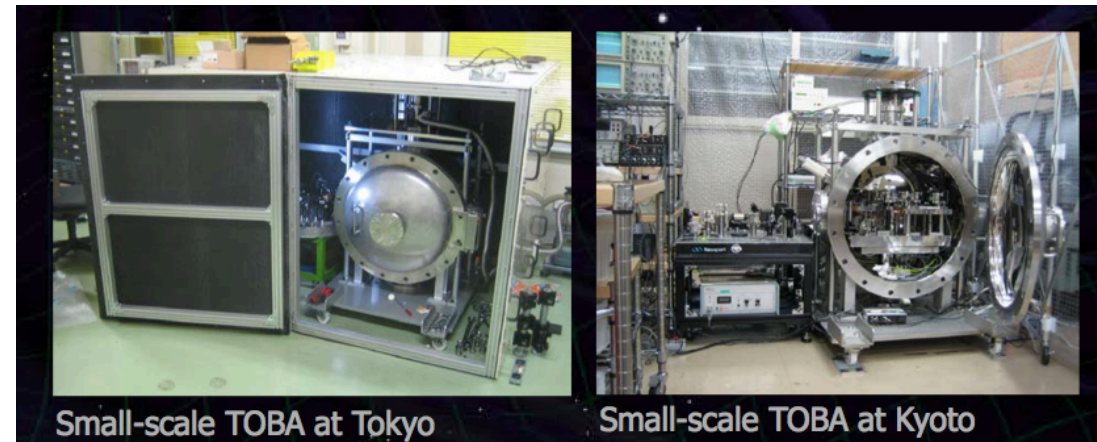
Devices designed to measure **gravitational waves**,  
minute distortions of space-time that are predicted by  
Einstein's theory of **general relativity**  
**Max Sensitivity 0.1Hz (seismic band)**

Present sensitivity  $10^{-8}$

=> goal **PEGASEWS**  $10^{-15}$   $\nu$ Hz



TOBA concept (torsion-bar antenna)-  
University of Tokyo (Ando et al., 2010)

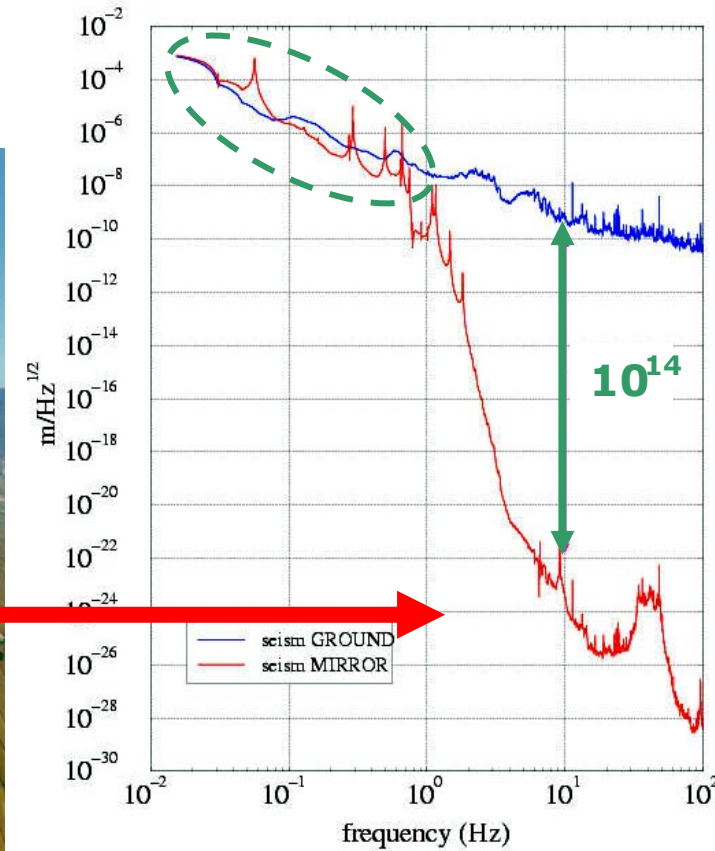
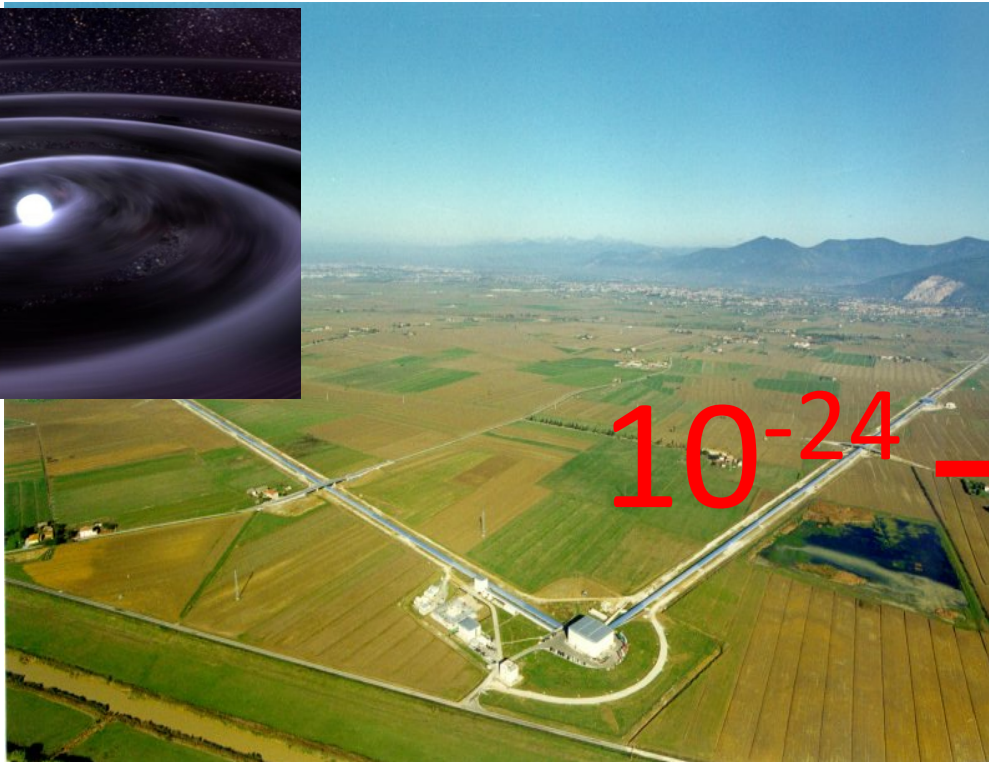
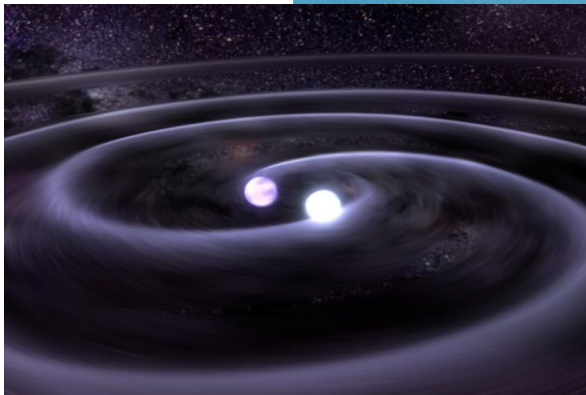


# LABEX UnivEarthS (2012)- Geophysics

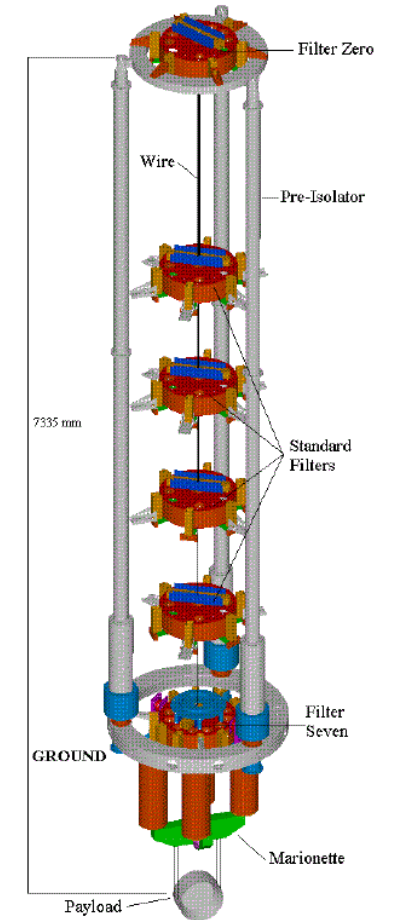
## Gravitational wave interferometers: VIRGO

VIRGO (Italian – French Gravitational wave detector)

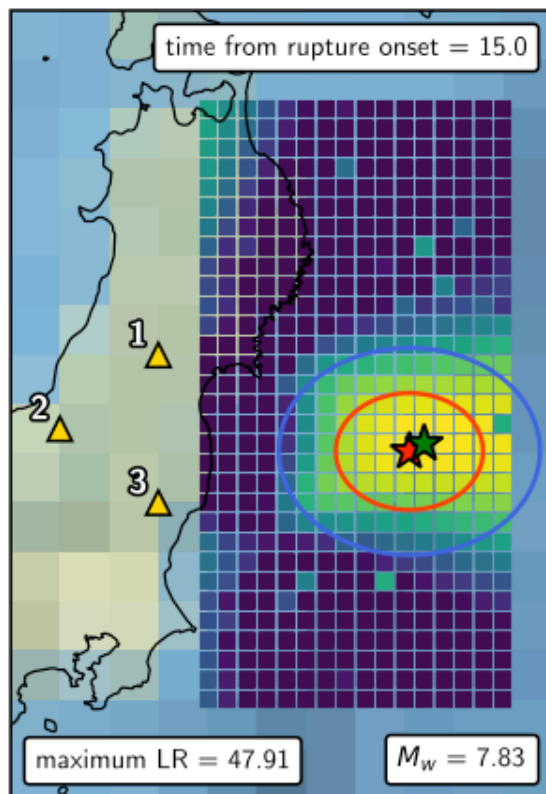
$10^{-18}$  ->  $10^{-24}$



Seismic wall at  $f > 1$  Hz





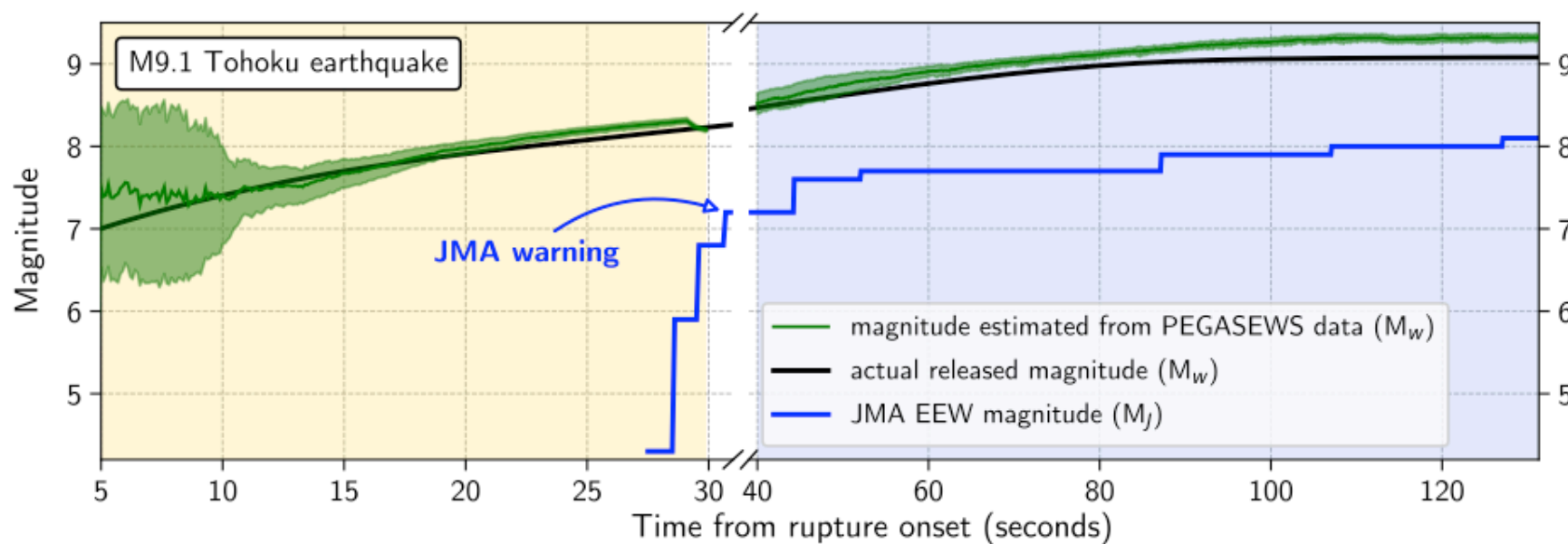
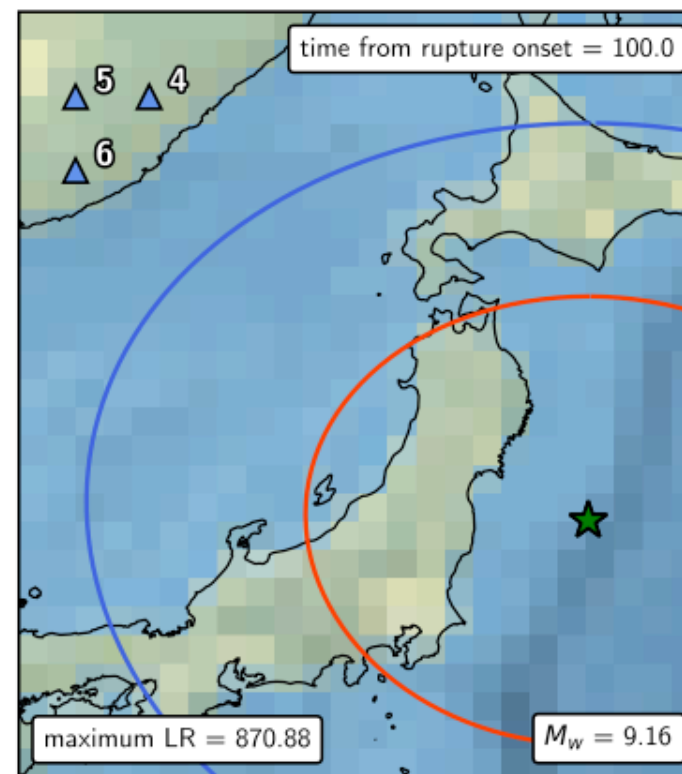


## Earthquake early detection

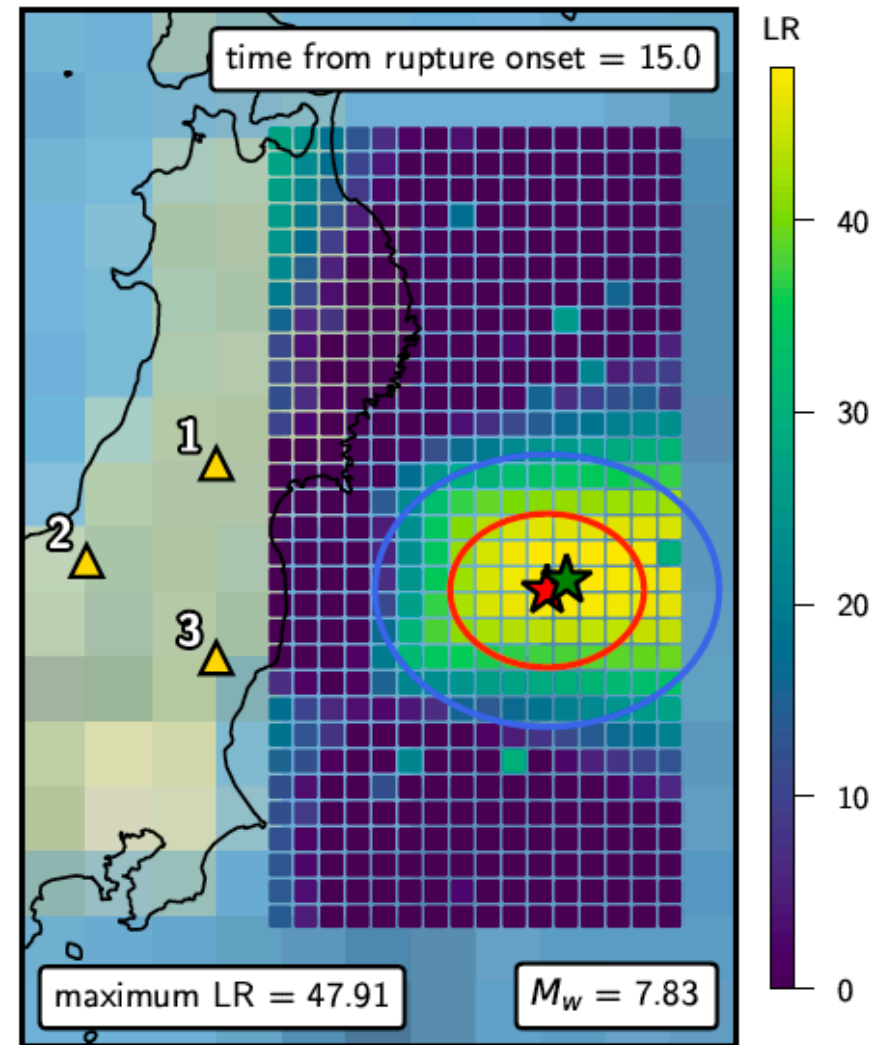
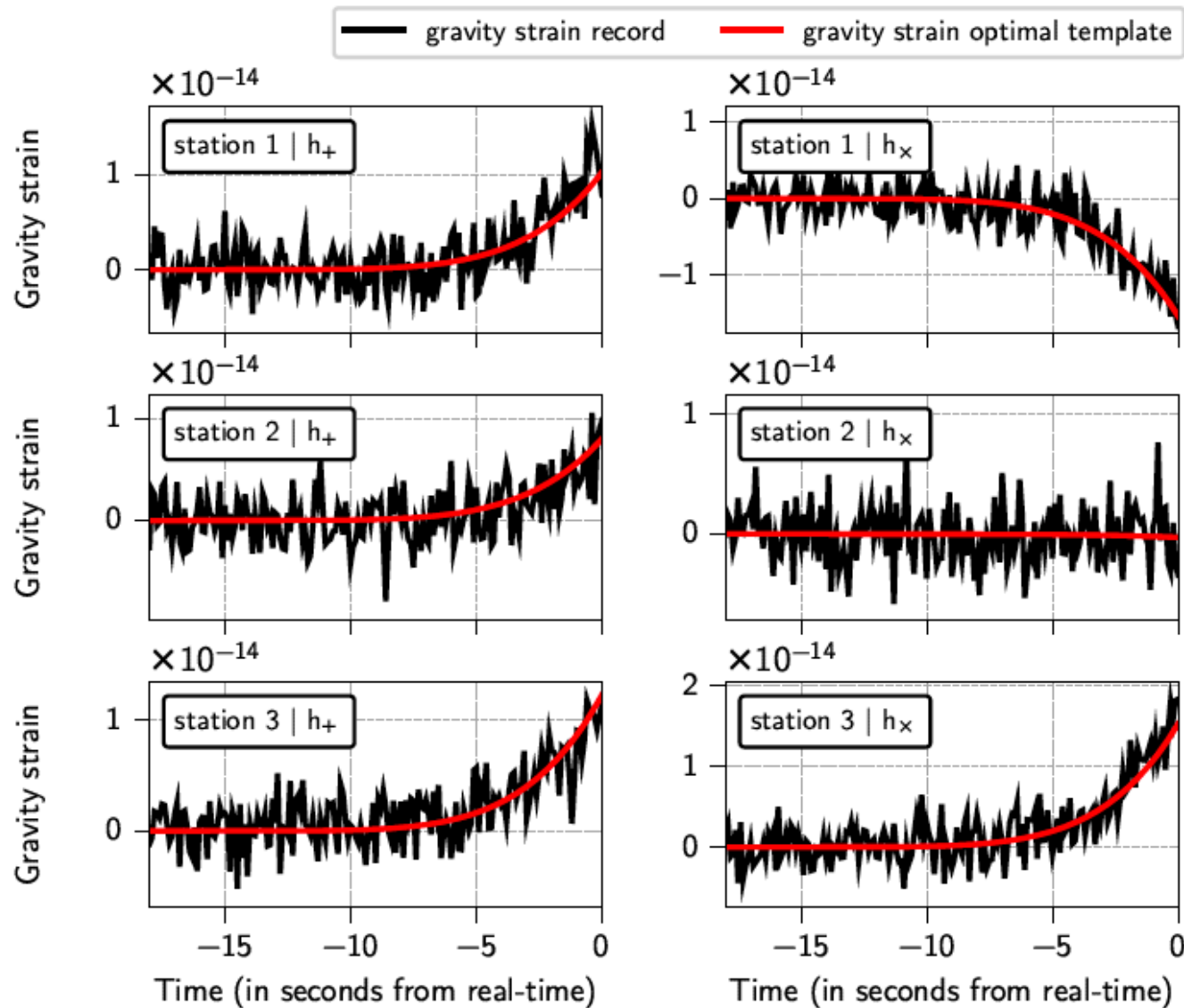
Near-source PEGASEWS detectors

Final magnitude estimation

Regional PEGASEWS detectors



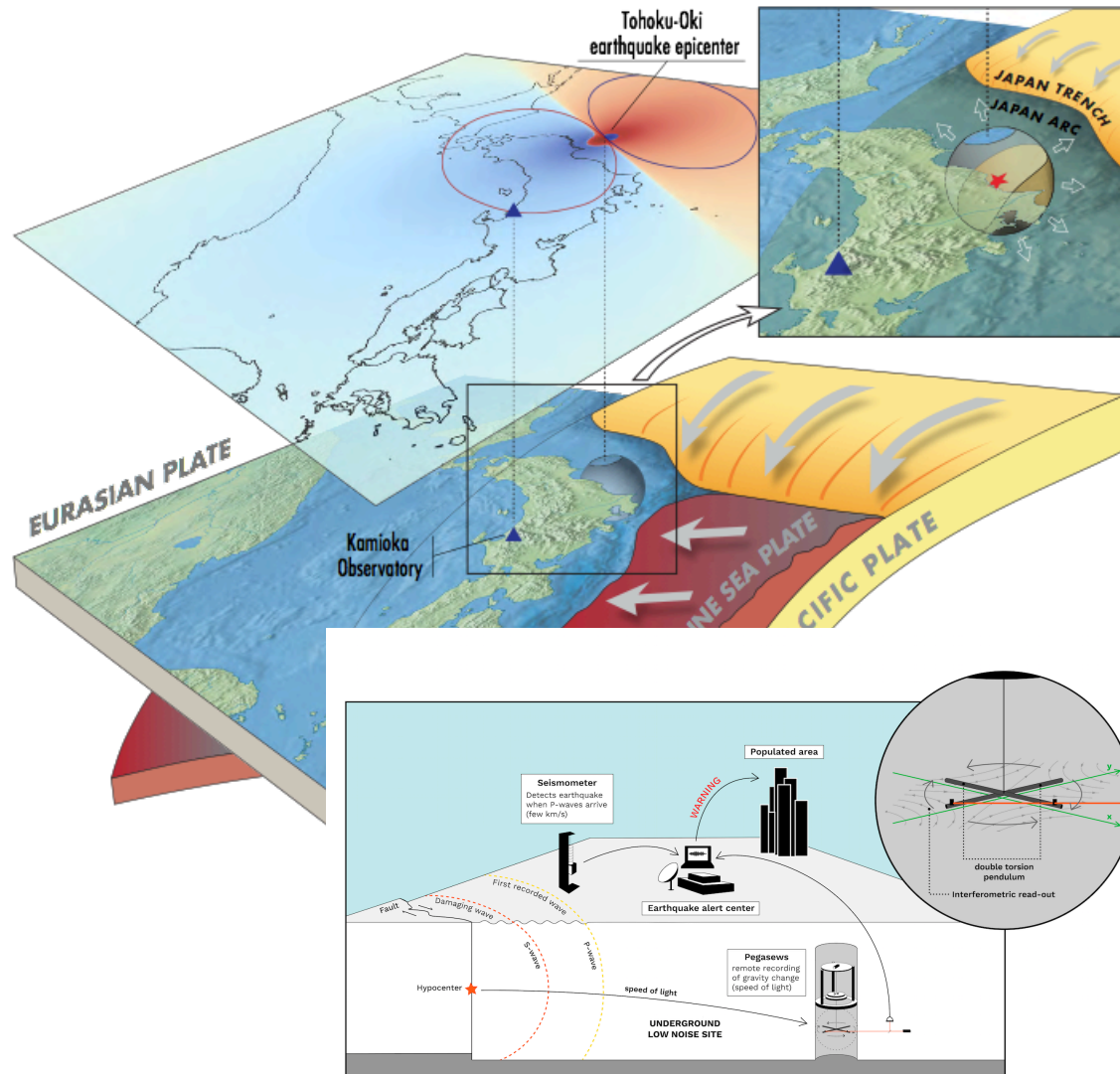
# What can we expect with a network of 3 PEGASEWS detectors?





# Conclusions:

## From Gravity field to Earthquake Early Warning Systems



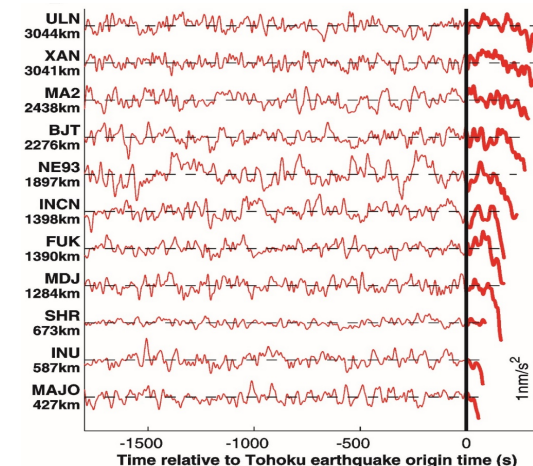
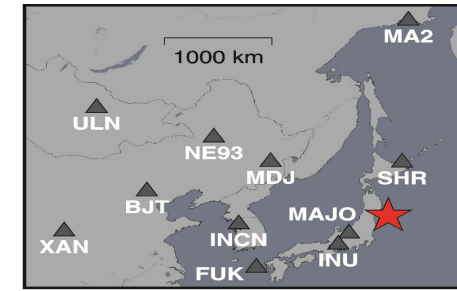
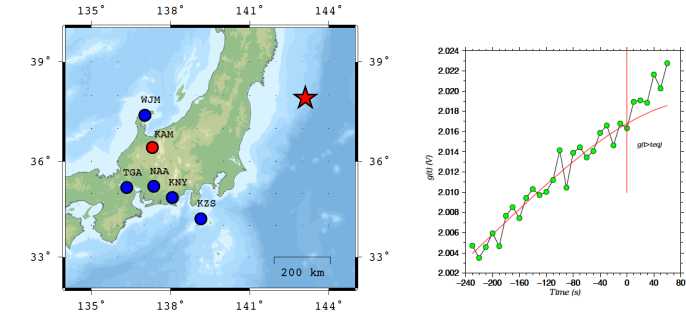
-Detection of a prompt gravity signal for Tohoku eq.:  
Very, very small  $<10^{-9} \text{ m/s}^2$

-Detection on  
Superconducting gravimeter,  
VBB stations in Japan and Eastern Asia  
 $\approx -0.1\text{-}0.2 \mu\text{Gal}$  at 500-1000km  
(Montagner et al., 2016; Vallée et al., 2017)

-For other earthquakes (Vallée & Juhel, 2019)

**-EEWS: magnitude estimate**

-Need for new gravity  
Instruments (TOBA, Atom  
Interferometers,  
superconducting gravimeters)  
In the frequency range 0.01-1Hz



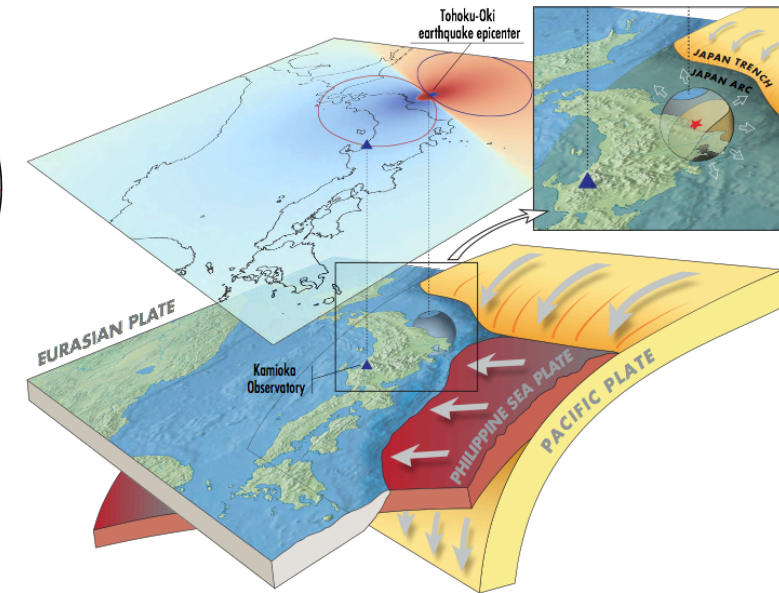
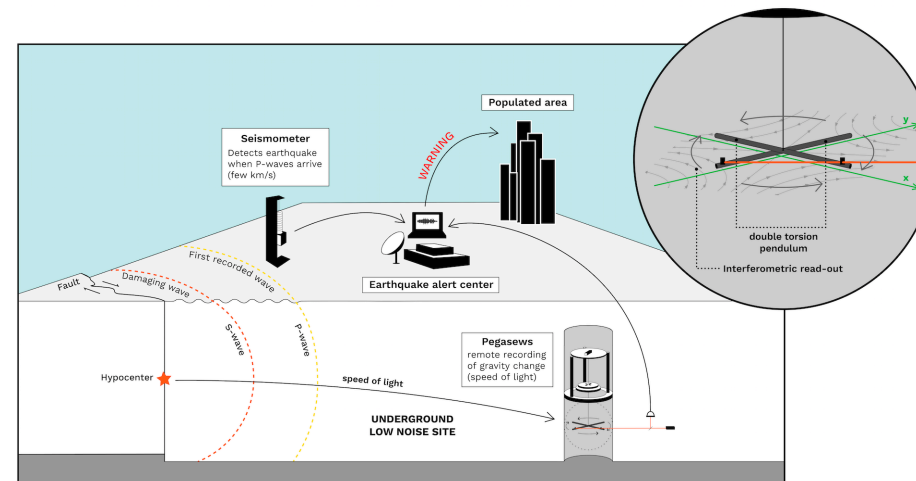
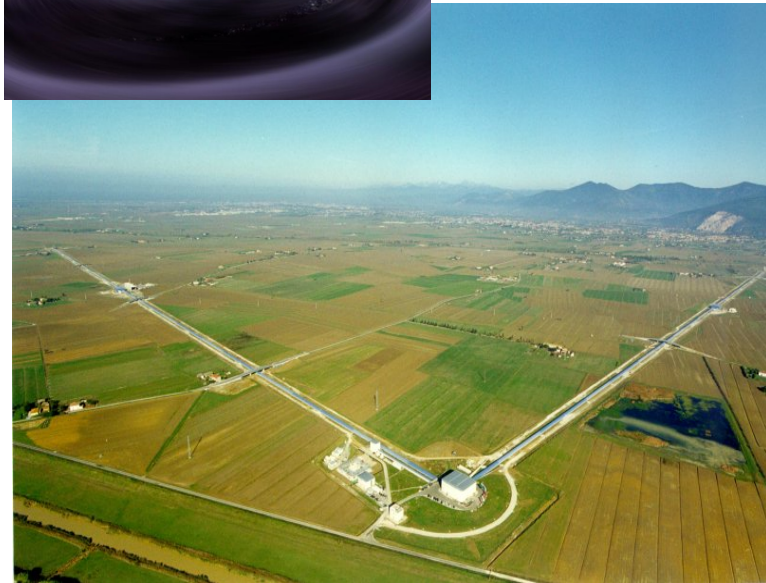
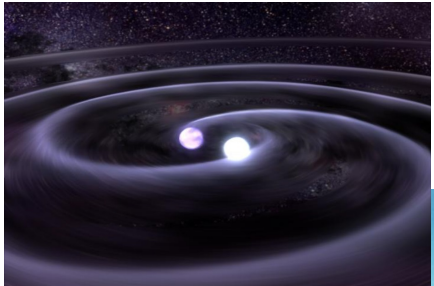


# Conclusions

## From gravitational waves to Earthquake Early Warning Systems

### Speed of light seismology

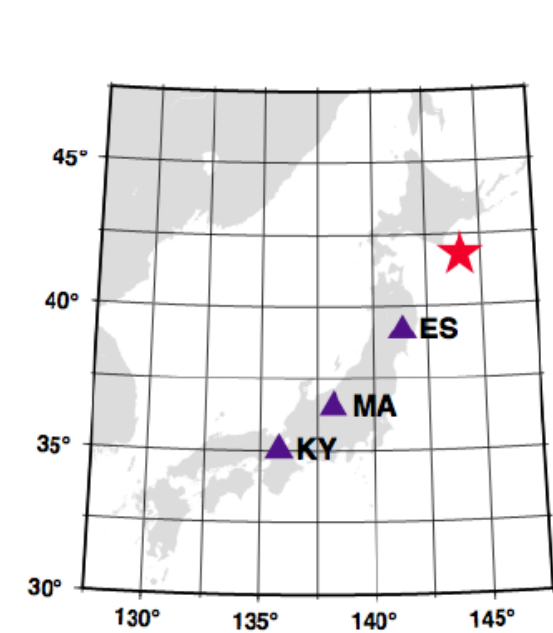
#### PEGASEWS project



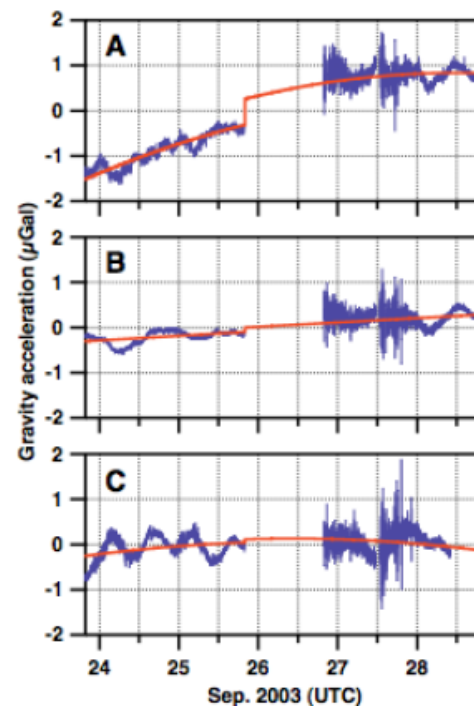


# Static gravity changes induced by earthquakes measured

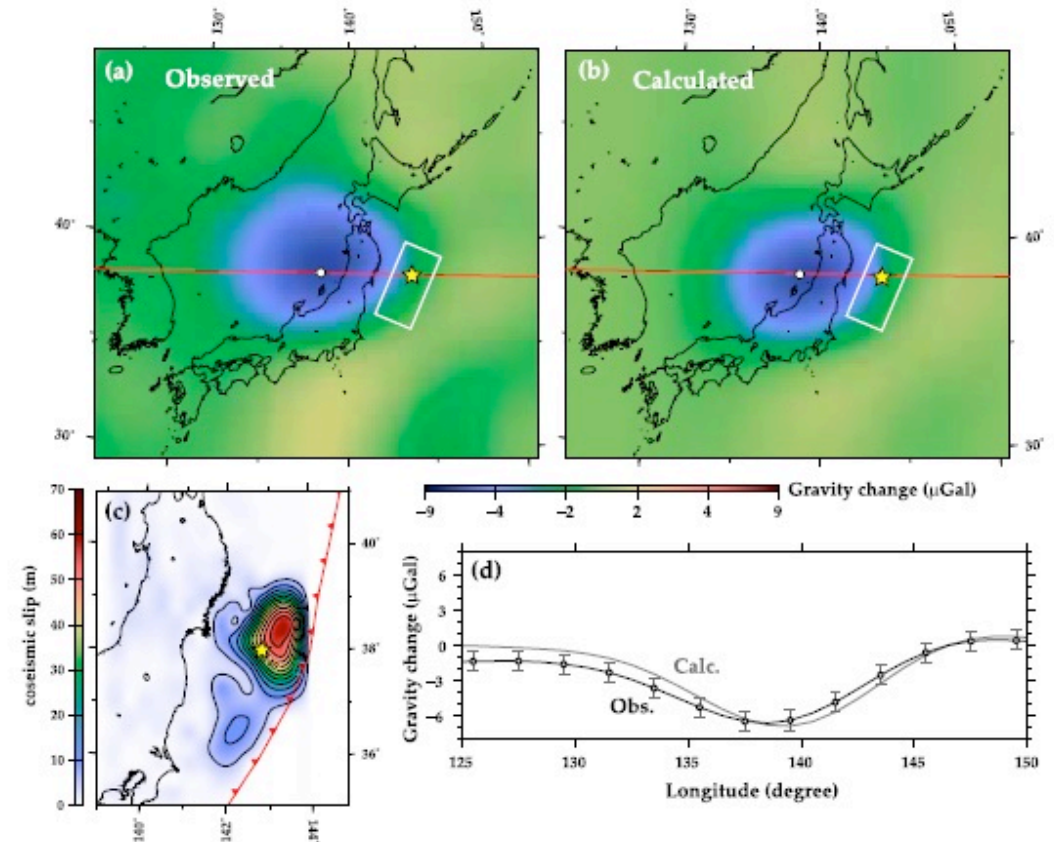
- Ground-based stations: Superconducting gravimeters after large earthquakes (2003 M=8.0 Tokachi-oki earthquake)



Imanishi et al. (Science, 2004)



- GRACE / GOCE satellites gravity changes after versus before large earthquakes (2011 M9.0 Tohoku-oki earthquake)



Matsuo and Heki (2011)

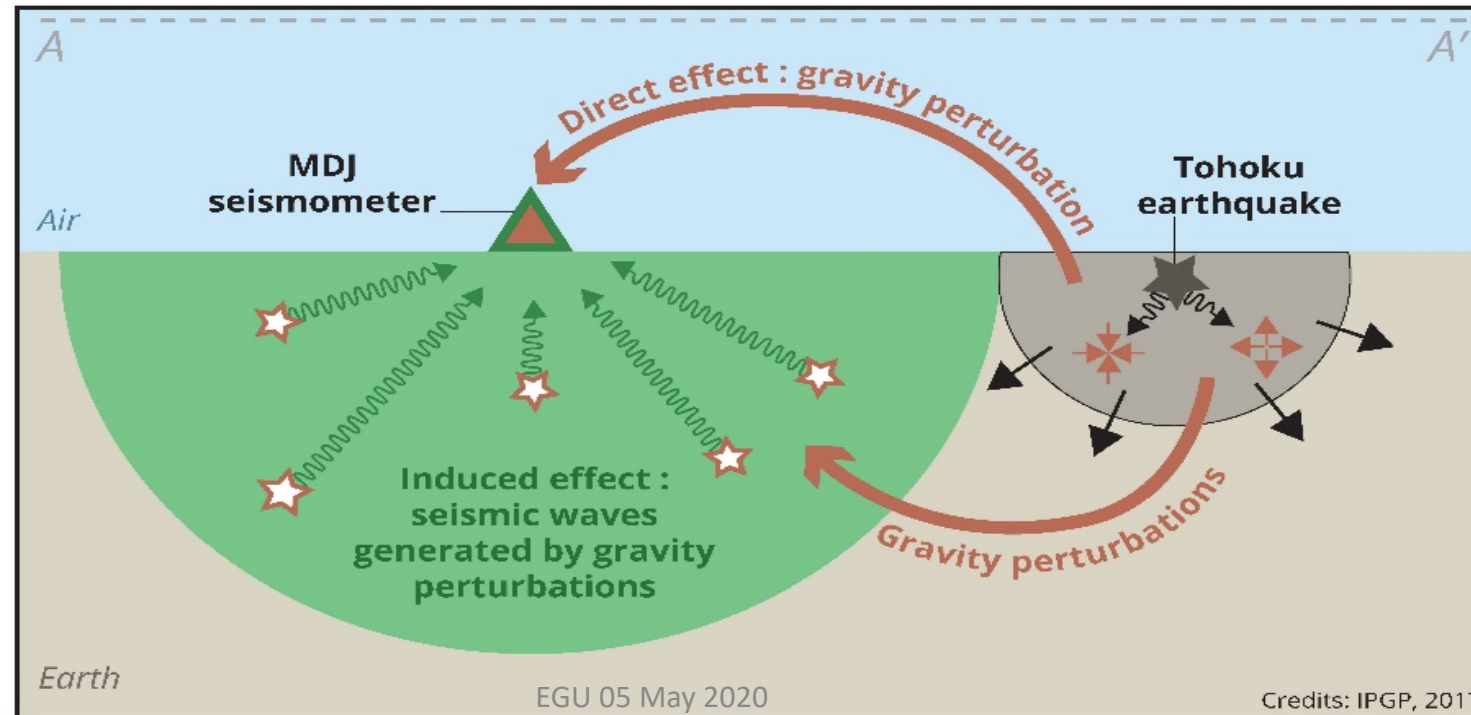
# Insights from the set of self-gravitating equations

$$\begin{cases} \rho_0 \ddot{\mathbf{u}} &= \nabla \cdot \boldsymbol{\sigma} + \Delta \rho \mathbf{g}_0 + \mathbf{f} + \rho_0 \Delta \mathbf{g}, & \text{Force balance equation} \\ \nabla \cdot \Delta \mathbf{g} &= -4\pi G \Delta \rho, & \text{Poisson equation} \\ \Delta \rho &= -\nabla \cdot (\rho_0 \mathbf{u}), & \text{Continuity equation} \end{cases}$$

➔ In this general formulation, there is a **full coupling** between the gravitational perturbation  $\Delta \mathbf{g}$  and the displacement  $\mathbf{u}$

However  $\Delta \mathbf{g}$  and  $\Delta \rho$  are coupled to the **displacement** field while **the force term** directly influences the **acceleration**

Illustration of the modeling approach (Vallée et al., 2017)



Courtesy of  
Martin Vallée