The Spacecraft Potential's Influence on the FOV of Rosetta-ICA at Low Ion Energies

Sofia Bergman⁽¹⁾, Gabriella Stenberg Wieser⁽¹⁾, Martin Wieser⁽¹⁾, Fredrik Leffe Johansson⁽²⁾ and Anders Eriksson⁽²⁾

<u>sofia.bergman@irf.se</u>

(1) Swedish Institute of Space Physics, Kiruna, Sweden(2) Swedish Institute of Space Physics, Uppsala, Sweden

EGU Sharing Geoscience Online, 4-8 May 2020





Background

- Spacecraft charging is problematic when doing low-energy ion measurements, since the particles are either attracted to or repelled from the spacecraft prior to detection
- This distorts the effective FOV of the instrument
- The Ion Composition Analyzer (ICA) on board Rosetta was measuring positive ions down to energies of a few eV
- Rosetta was commonly charged to a negative potential of -10 to -20 V, which distorted the lowenergy part of the ICA data
- In this study we use the Spacecraft Plasma Interaction Software (SPIS) to model the distortion of the effective FOV of ICA at low energies

Trajectories of high-energy ions entering ICA

Trajectories of lowenergy (a few eV) ions entering ICA





[Bergman et al., 2020a]





ICA Instrument Description

- Mass resolving ion spectrometer
- Positive ions
- Energy range: a few eV/q 40 keV/q

Nominal FOV:

- Total FOV of 360° x 90°
- In azimuthal direction: 16 sectors of 22.5°
- In elevation: 16 elevation steps of 5.6°



[Bergman et al., 2020b]





SPIS Simulations

- PIC approach to model the environment
- Test particles traced from the instrument
- Plasma model (valid close to perihelion):

H ₂ O ⁺	e -
n = 1000 cm ⁻³ T = 0.5 eV v = 4 km/s	n = 1000 cm ⁻³ T = 8 eV

• \rightarrow Spacecraft potential of -21 V



[[]Bergman et al., 2020a]





Results - Example

Results for one pixel of the instrument E = 5-10 eV (at infinity)

[Bergman et al., 2020a]



Nominal FOV



- Dashed square: Nominal FOV of the studied pixel
- Color scale: Flux of particles at the simulation boundary that are reaching the studied pixel
- The colored area hence corresponds to the effective FOV of this pixel





Results – Effect on Different Elevation Angles



[•] Sector 3, E=5-10 eV

•

Extreme elevation angles are much more severely distorted than angles close to the aperture plane







Results – Different Debye Lengths of the Surrounding Plasma



 The sensitivity to a changing Debye length varies, however, b different instrument pixels





Results – Varying Spacecraft Potentials

Question: Does the FOV distortion scale linearly with the spacecraft potential?

 $U_{s/c} = -21 V$ E = 5-10 eV



 $U_{s/c} = -42 V$

22.5°

0°

360°

azimuth

- The Debye length is the same in both cases
- If the ion energy is scaled according to the potential of the spacecraft, the FOV distortion is not necessarily the same
- The small nonlinearity observed is mainly caused by the photoemission and bulk flow of the cometary plasma





Conclusions

sofia.bergman@irf.se

- The FOV of ICA is distorted at low ion energies
- The distortion varies between different instrument pixels
- The distortion is sensitive to changes in the Debye length, but the sensitivity varies between different instrument pixels
- A small nonlinearity exists in the relation between FOV distortion and spacecraft potential, mainly caused by the photoemission and bulk flow of the cometary plasma
- The distortion is considered insignificant at an energy above twice the spacecraft potential

Publications:

Bergman et al. (2020a). *JGR: Space Physics*. <u>https://doi.org/10.1029/2019JA027478</u> Bergman et al. (2020b). *JGR: Space Physics*. <u>https://doi.org/10.1029/2020JA027870</u>





Publications

Bergman, S., Stenberg Wieser, G., Wieser, M., Johansson, F. L. & Eriksson, A. (2020a). The influence of spacecraft charging on low-energy ion measurements made by RPC-ICA on Rosetta. *Journal of Geophysical Research: Space Physics*, 125. <u>https://doi.org/10.1029/2019JA027478</u>

Bergman, S., Stenberg Wieser, G., Wieser, M., Johansson, F. L. & Eriksson, A. (2020b). The influence of varying spacecraft potentials and Debye lengths on in situ low-energy ion measurements. Submitted to *Journal of Geophysical Research: Space Physics*. <u>https://doi.org/10.1029/2020JA027870</u>

sofia.bergman@irf.se



