



# Large Scale TIDs climatology over Europe using HF Interferometry method

### Estefania Blanch<sup>1</sup>, Antoni Segarra<sup>1</sup>, David Altadill<sup>1</sup>, Vadym Paznukhov<sup>2</sup>, J. Miguel Juan<sup>3</sup>

<sup>1</sup> Observatori de l'Ebre, (OE), CSIC - Universitat Ramon Llull, Roquetes, Spain <sup>2</sup> Boston College, Chestnut Hill, United States <sup>3</sup> Research Group of Astronomy and Geomatics (gAGE) Universitat Politècnica de Catalunya (UPC), Barcelona, Spain











- **Traveling Ionospheric Disturbaces**
- **TechTIDE & MIRA projects**
- **HF Interferomtry method to detect LSTIDs**
- Results: Climatology of LSTIDs for 2014-2019
  - Seasonal & diurnal Occurrence.
  - Seasonal & diurnal Activity.
  - Prevailing direction of propagation.
- Summary & Conclusions.







EGU2020: Sharing Geoscience Online, 4 - 8 May 2020





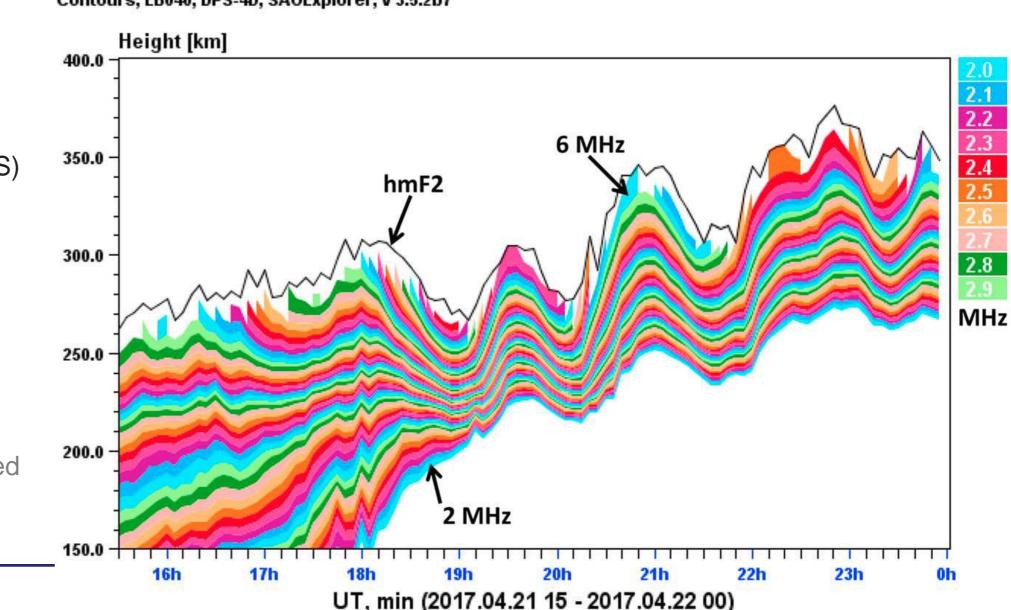
**Travelling Ionospheric Disturbances (TIDs)** are plasma density fluctuations that propagate as waves through the ionosphere at a wide range of velocities and frequencies.

TIDs constitute a threat for operational systems using predictable ionospheric characteristics as they can impose significant disturbances in the ambient electron density and Doppler frequency shifts on HF signals. Contours, EB040, DPS-4D, SAOExplorer, v 3.5.2b7

Systems and services that can be affected by TIDs:

- European Geostationary Navigation Overlay Service (EGNOS)
- Network Real-Time Kinematic (N-RTK)
- High Frequency (HF) communications
- Radio reconnaissance operations
- Very High Frequency Ultra High Frequency (VHF-UHF) radiowave propagation.

Figure: TID signature in the electron isodensity contours observed by the EB040 Digisonde (Spain) during 21–22 April 2017 as the result of moderate auroral activity.









### **TechTIDE project**

**Objective:** To design and test new viable TID mitigation impact strategies for the technologies affected and close in collaboration with operators of these technologies.

**Our Tasks:** Development of new methodology (HF Interferometry method) to detect LSTIDs in near real time using ionospheric characteristics from dense enough networks of digisondes (e.g. Europe and South Africa).



<u>http://www.tech-tide.eu/</u> <u>techtide.project@gmail.com</u> <u>https://twitter.com/Tech\_TIDE</u>

### **MIRA project**

**Objective:** To provide new knowledge and applicability on the climatological behavior of the ionosphere and, specially, on the short term ionospheric disturbances or irregularities.

**Our Tasks:** Improvement of the functionality and application of the HF Interferometry method (developed under TechTIDE project) to perform climatological analysis and characterization of LSTIDs with the aim to obtain an activity index of LSTIDs.



detection, Monitoring and modelling of Ionospheric irRegulArities PGC2018-096774-B-I00 (MCIU, AEI, FEDER) <u>website</u> <u>eblanch@obsebre.es</u> <u>https://twitter.com/obsebre</u>

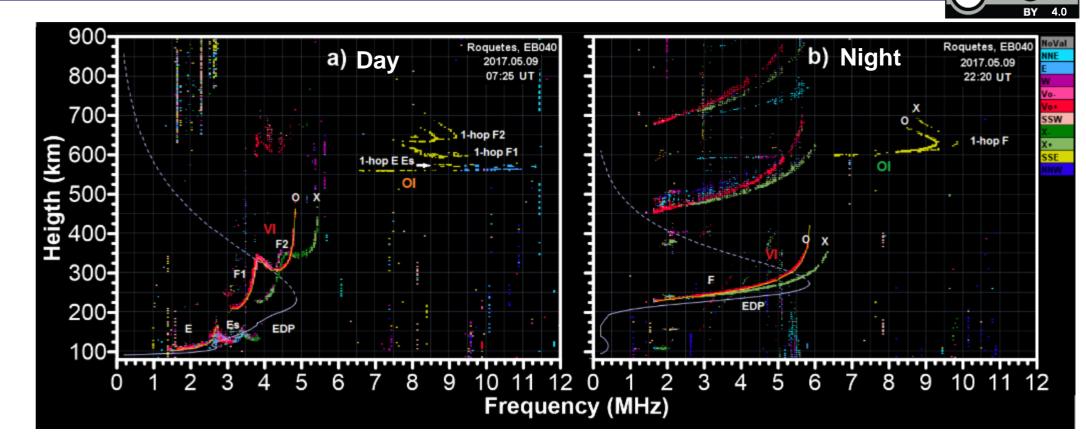


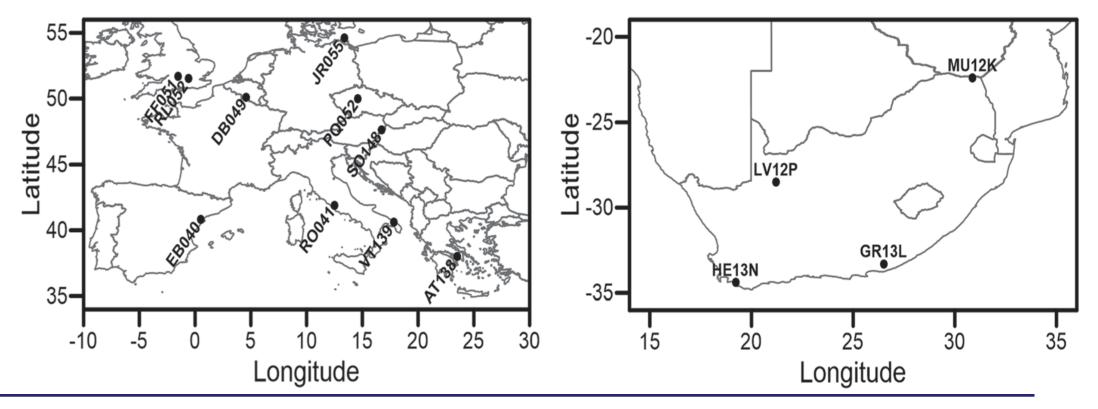


### HF Interferometry method General 2020

## **Data and Network**

- Data from Vertical Incidence Ionospheric sounding (MUF(3000)F2). NRT
- Network of DPS4D with TechTIDE stations working synchronized. European and South Africa networks.
- South Africa: Limitation by small number of stations. Not all operate in NRT





EGU2020: Sharing Geoscience Online, 4 - 8 May 2020



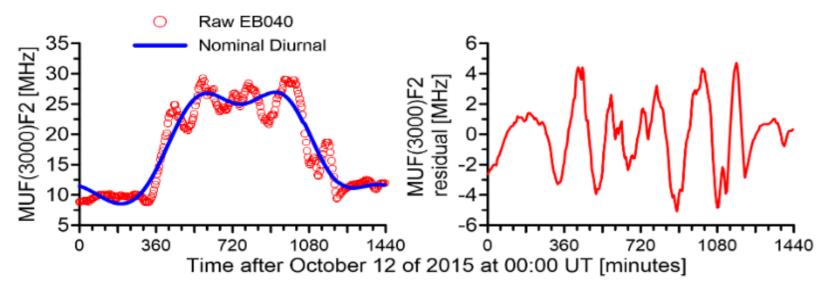


# HF Interferometry method

## **Method and Concept**

### 1.Obtain nominal daily variation and residuals at different sites.

spectral analysis



3.Estimate time delays for different sites by cross-correlation.

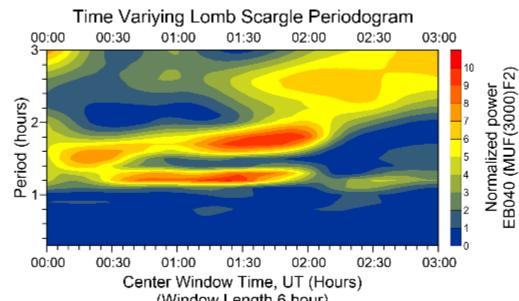
2 MHz 2 MHz MUF(3000)F2 21 22 23 24 23 EGU2020-784 26 22 24 25 25 Hours starting at day 285 of 2015 Hours starting at day 285 of 2015 correlation  $\Delta TM$ .

 $\Delta TM_i$  –

**It provides:** TID occurrence, amplitude, period, spectral contribution, propagation velocity (speed and azimuth)



### 2.Detect coherent TID-like variations by



(Window Length 6 hour)

4. Estimate propagation velocity of irregularities  $\vec{v}$  by time delays of maximum

$$\vec{s} \cdot \Delta \vec{r}_i = 0$$
 ;  $\vec{v} = \frac{\vec{s}}{s^2}$ 





# HF Interferometry method (EGUGeneral 2020)

AT139

0:00

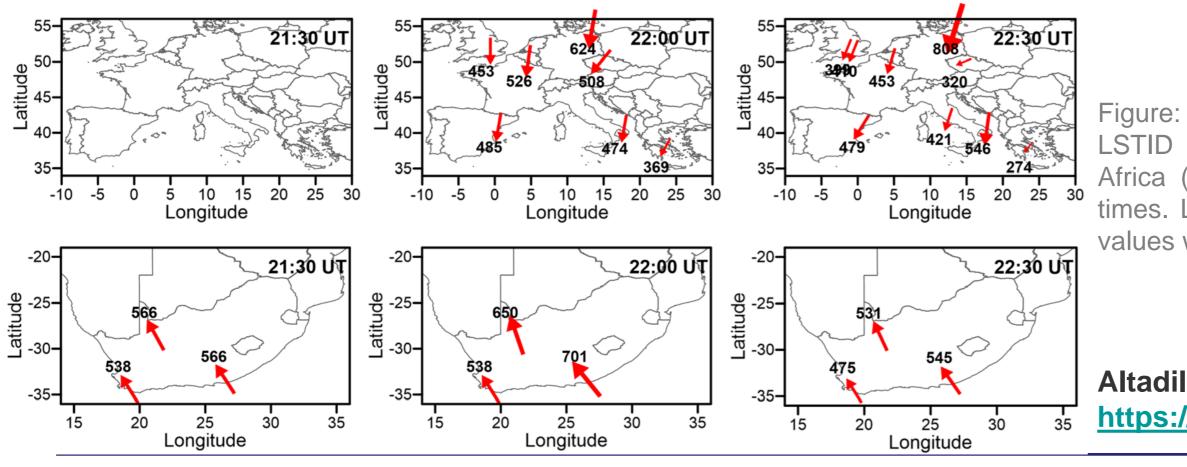
Time [h]

4:00

8:00

## Performance for 12-13 Oct 2015

Figure: Results of detection of the LSTID activity by the HF-Int over two stations located in Europe (left) and South 80 Africa (right) for the time interval from 12UT on October 12 40 2000 to 12UT on October 13 of 2015. Red dots show velocity, [s/ш] blue dots show azimuth, black dots show period and elocity | green dots show **spectral energy contribution** (SEC).



EGU2020-7842, Thursday, 07 May 2020, 10:45-12:30

12:00

16:00

2015-10-12

20:00

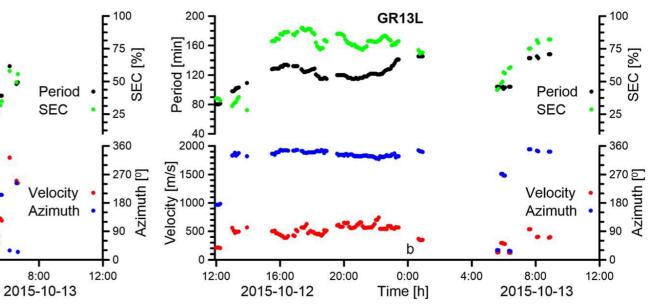


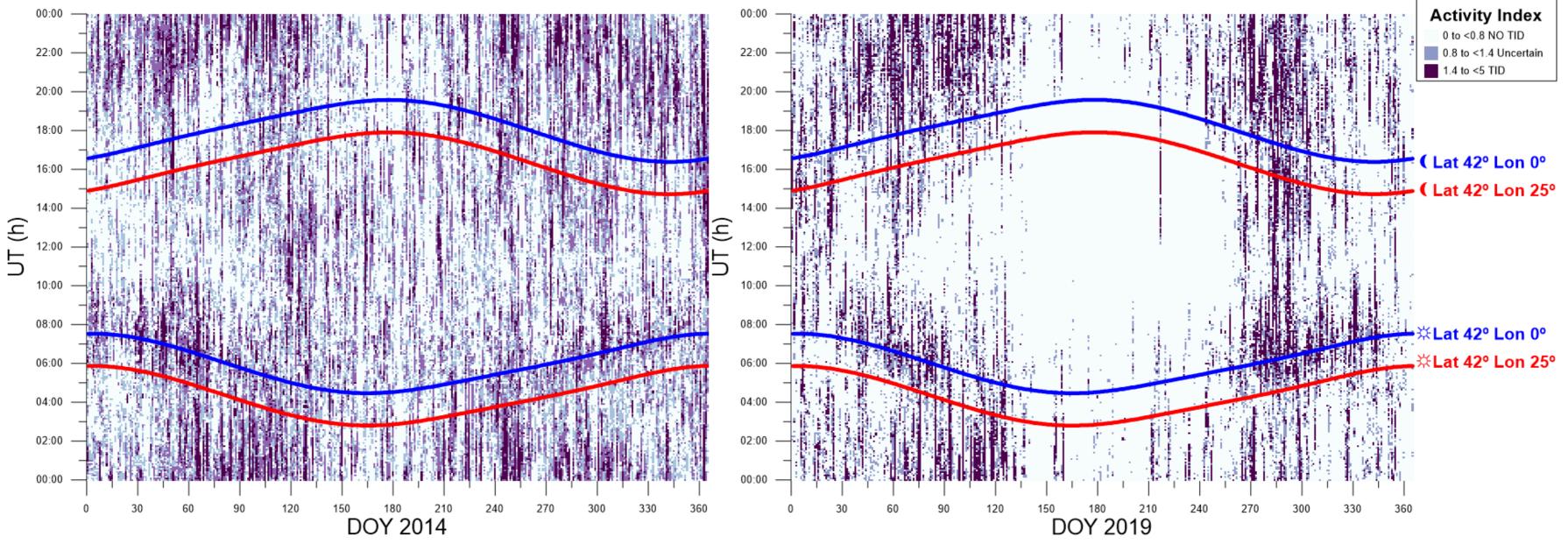
Figure: Maps of the estimated velocity vectors of the LSTID event detected over Europe (up) and South Africa (down) on 12 October 2015 for the indicated times. Length of arrows is proportional to the velocity values which is show in the labels.

### Altadill et al., 2020. J. Space Weather Space Clim., https://doi.org/10.1051/swsc/2019042





## Solar activity & Seasonal & Diurnal LSTIDs occurrence



Poor data availability for summer at low solar activity due to Es layer screening. Larger occurrence during nighttime and near solar terminator.

Enhanced nigh-time occurrence for equinoxes.

EGU General Assembly 2020

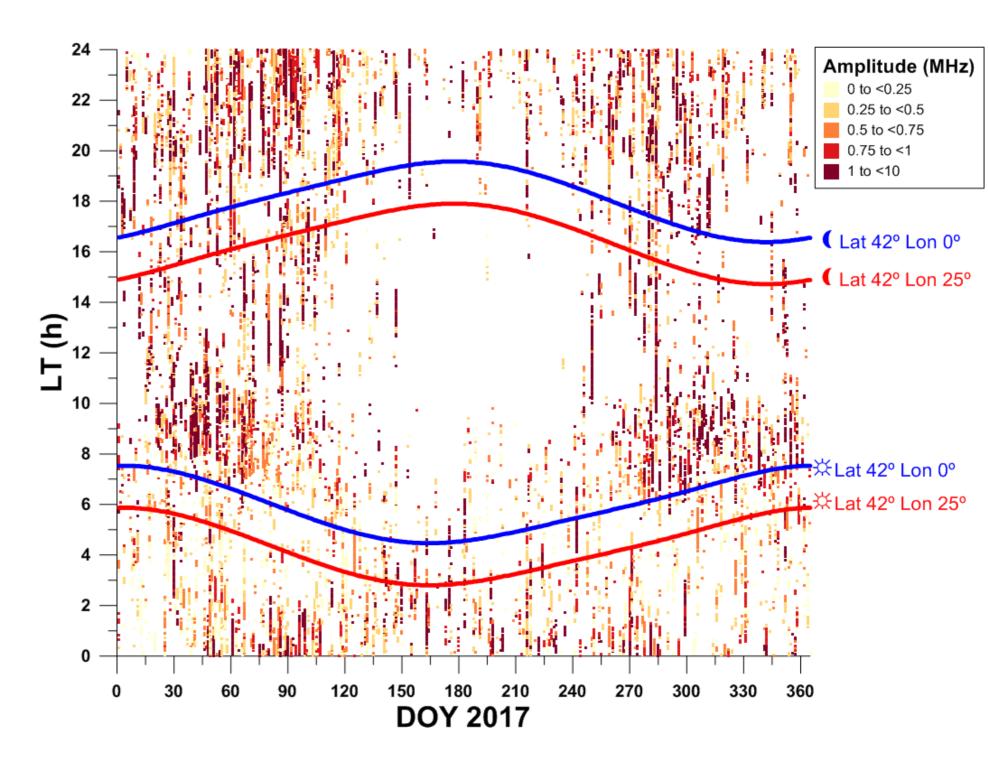




## Seasonal & diurnal occurrence

TéchTIDE

- Poor data availability for summer: Es layer screening.
- Larger occurrence near solar terminator and night-time.
- Enhanced nigh-time occurrence for equinoxes.



atology EGU General 2020

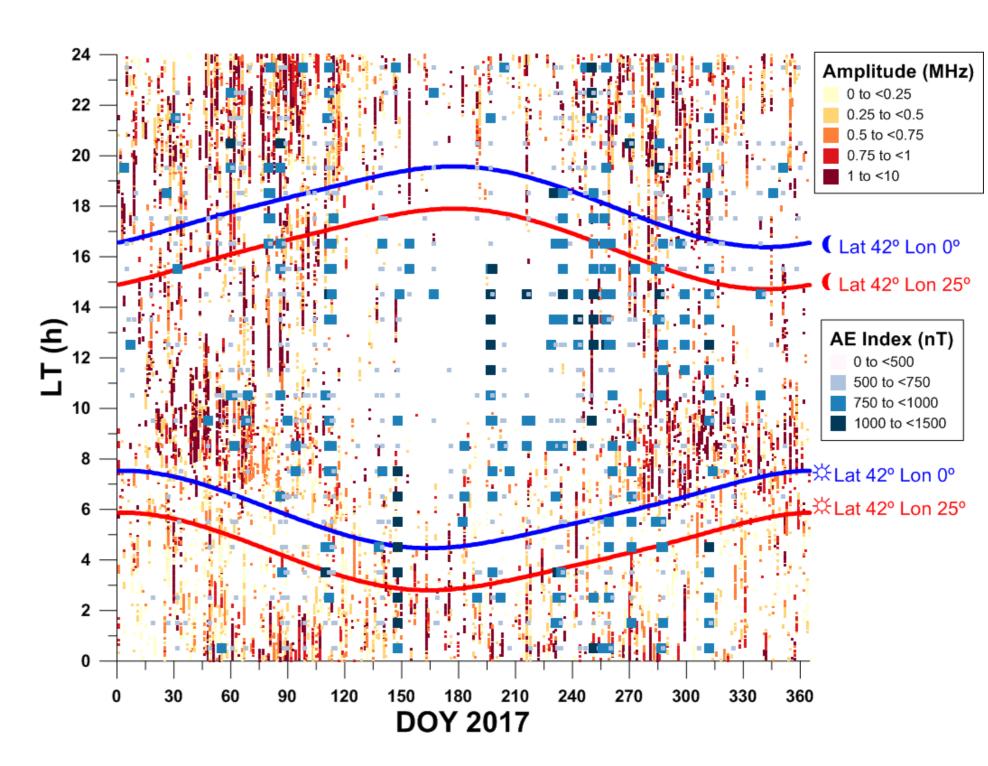




## **Seasonal & diurnal occurrence**

**Téch**TIDE

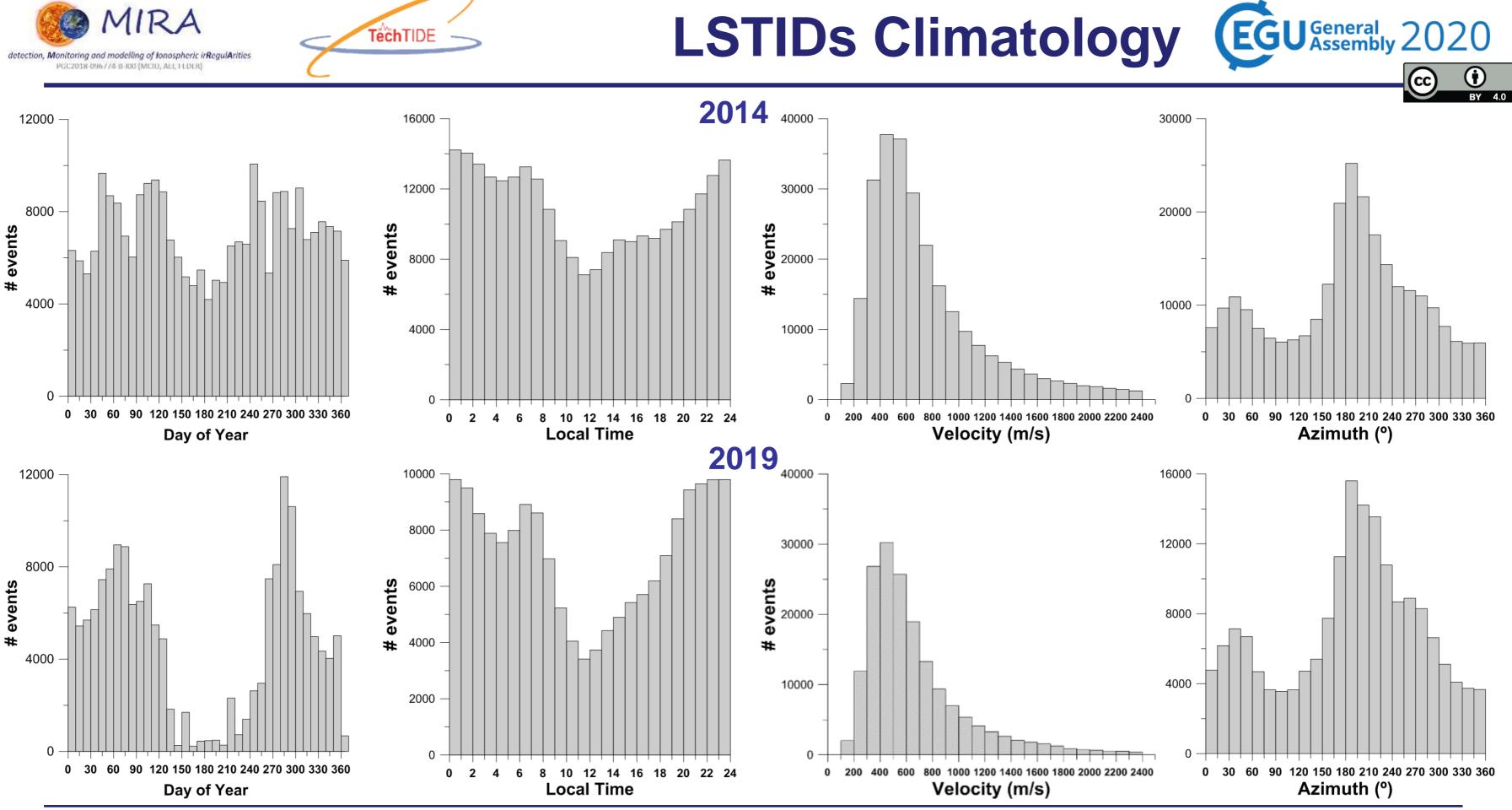
- Poor data availability for summer: **Es** layer screening.
- Larger occurrence near solar terminator and night-time.
- Enhanced nigh-time occurrence for equinoxes.
- Equinoctial occurrence of LSTIDs lacksquarecorrelates with Auroral activity.



EGU General Assembly 2020







EGU2020-7842, Thursday, 07 May 2020, 10:45-12:30

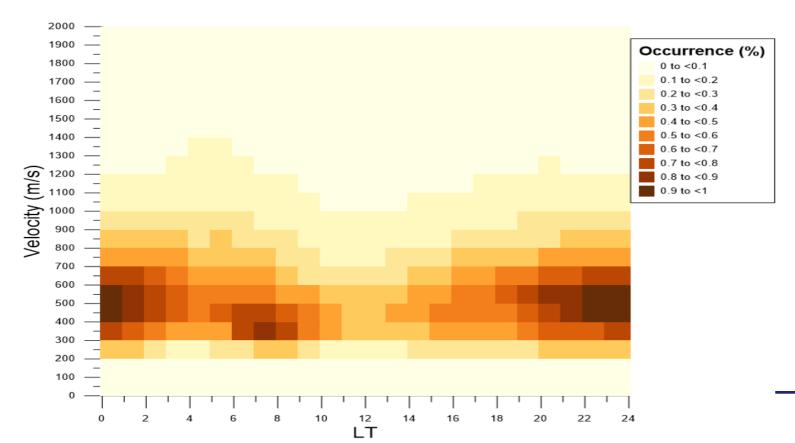
EGU2020: Sharing Geoscience Online, 4 - 8 May 2020

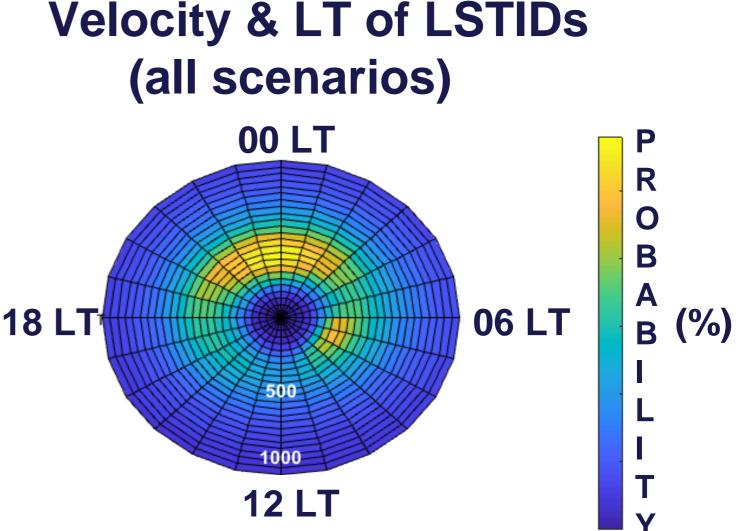




## **Diurnal Velocity & LT analysis**

- Activity tend to concentrate during night time and near Sun Rise (SR).
- Results suggest origin of the LSTIDs lacksquareby the solar terminator and auroral activity.
- Nighttime events related with auroral activity reach larger velocities.







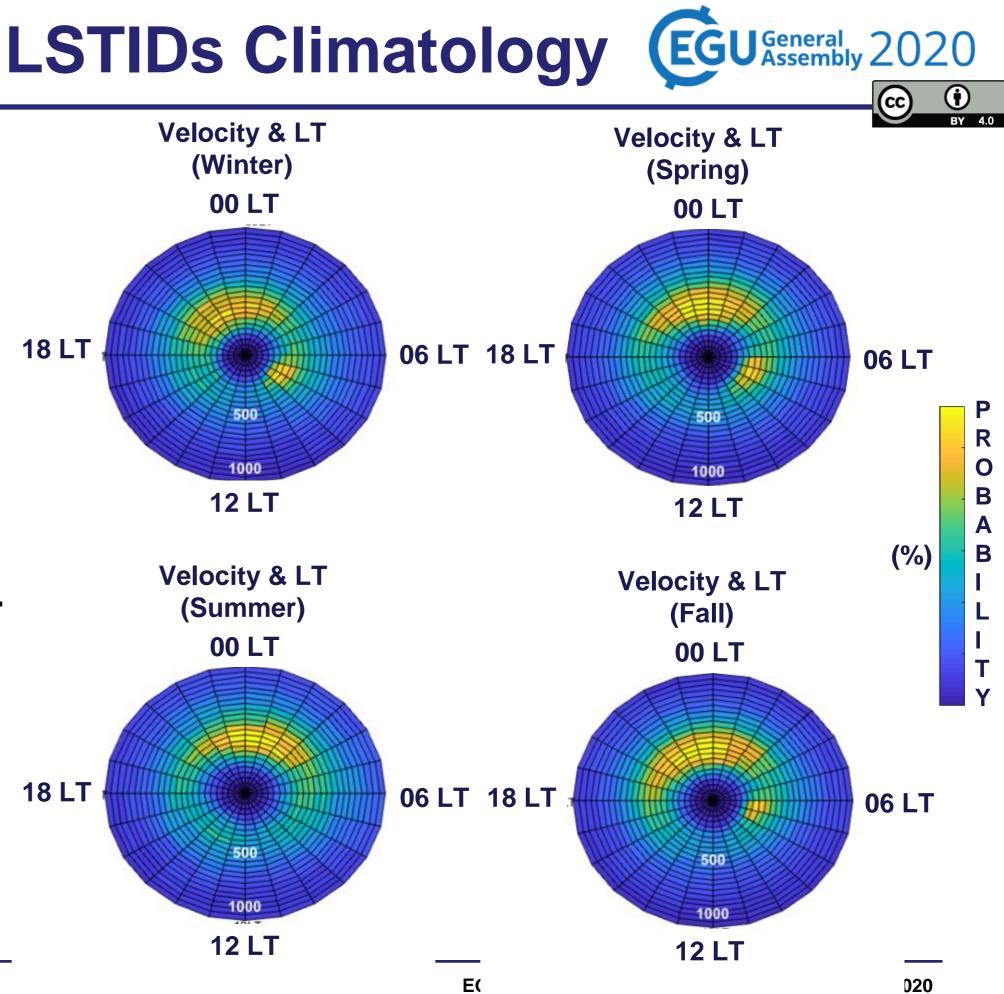




## **Seasonal & Diurnal Velocity &** LT analysis

Similar behavior for all seasons except for summer:

- Activity tend to concentrate during • night time and near Sun Rise (SR) with larger velocities during nighttime.
- Different behavior for summer can be due to Es layer screening.

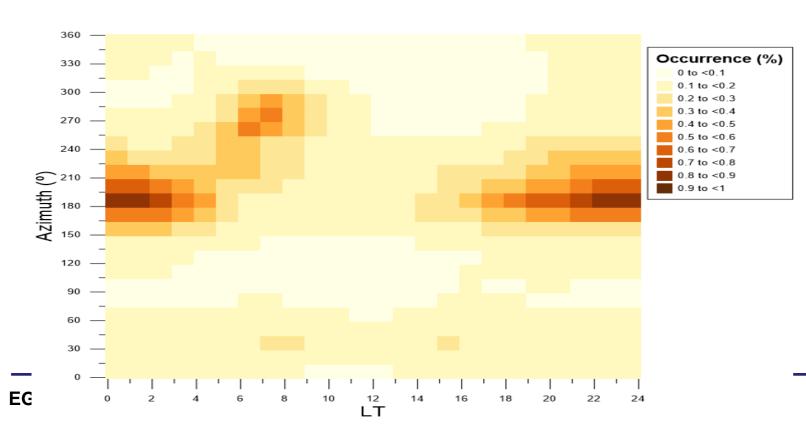


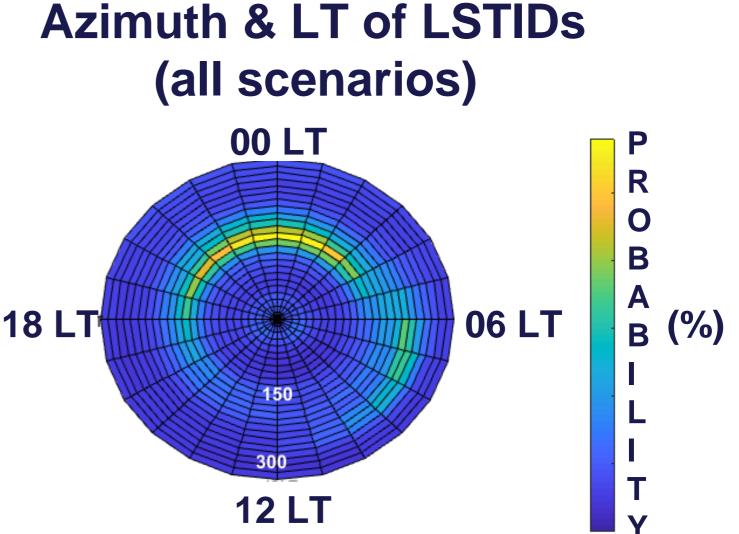




## **Diurnal Azimuth & LT analysis**

- Most of the events present southward lacksquarepropagation (180°) suggesting origin of the LSTIDs by auroral activity.
- Westward propagation (270°) during lacksquaremorning time suggesting origin of LSTIDs by **sunrise** effect.







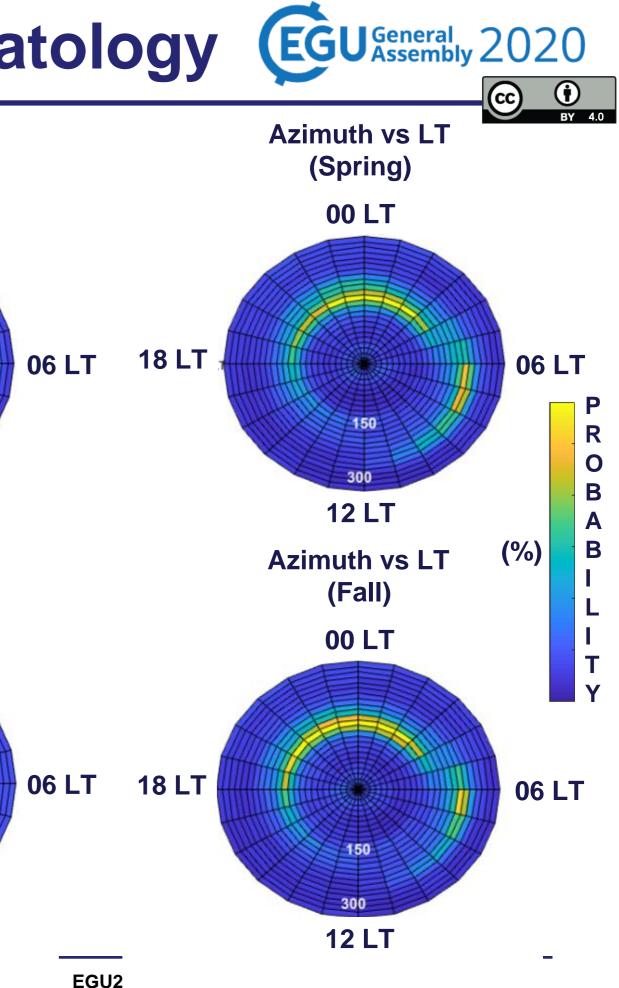


lacksquare



# **LSTIDs Climatology**

### Azimuth vs LT Seasonal & Diurnal Azimuth & (Winter) LT analysis 00 LT Activity tend to concentrate during nigh 18 LT time and near Sun Rise. This behavior 150 is not observed during summer due to 300 Es layer screening effect. 12 LT Azimuth vs LT (Summer) • We observe southward propagation at 00 LT nighttime and westward propagation during morning time. 18 LT 150 300 12 LT





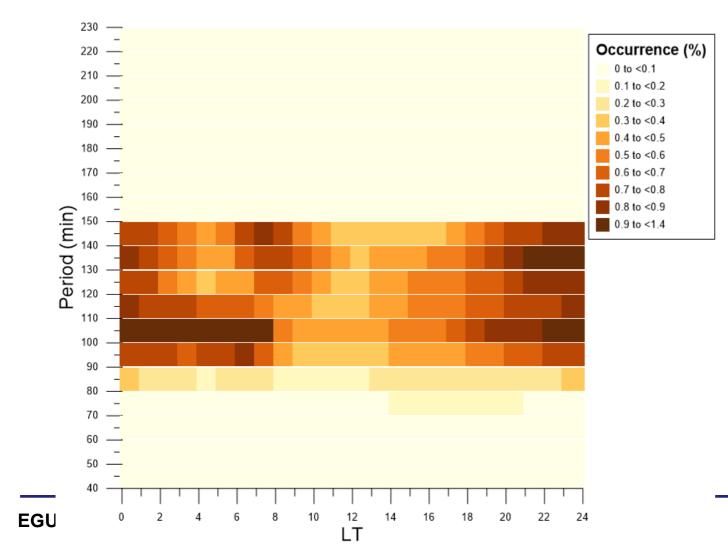


## **Diurnal Period & LT analysis**

Detected LSTIDs present coherent periods from 80 to 140 min, prevailing those cases with 100-130 min during the night.







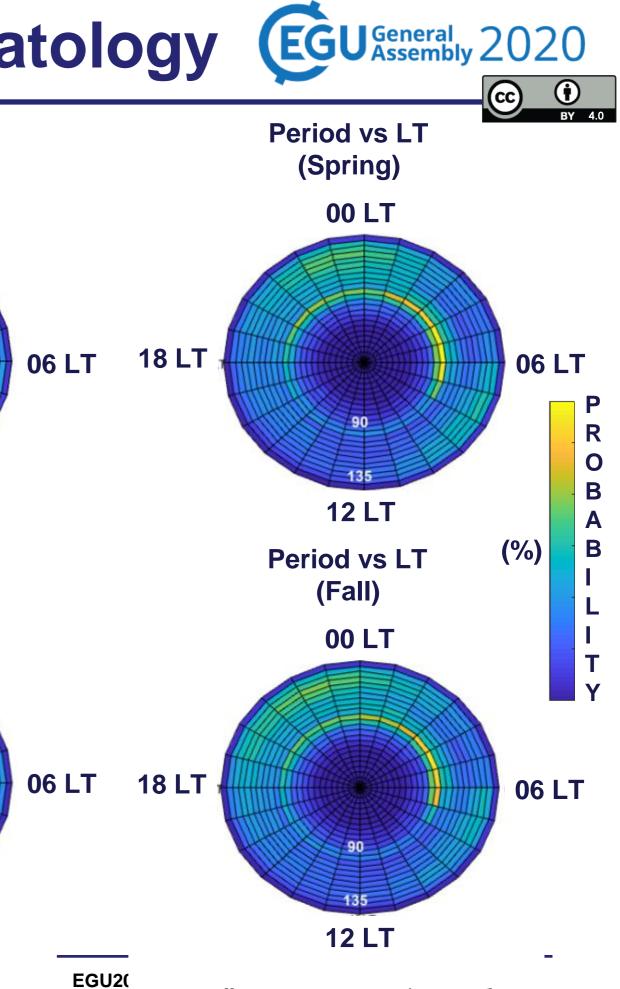


### **Period & LT of LSTIDs** (all scenarios) **00 LT** Ρ R 0 B A **06 LT** (%) Β 90 Т 12 LT Y





### Period vs LT Seasonal & Diurnal Period & LT (Winter) analysis 00 LT Detected LSTIDs tend to have a period 18 LT of 90-120 min, specially when they occur from midnight to sunrise. Larger periods occur before midnight. 12 LT Period vs LT (Summer) 00 LT 18 LT 90 12 LT





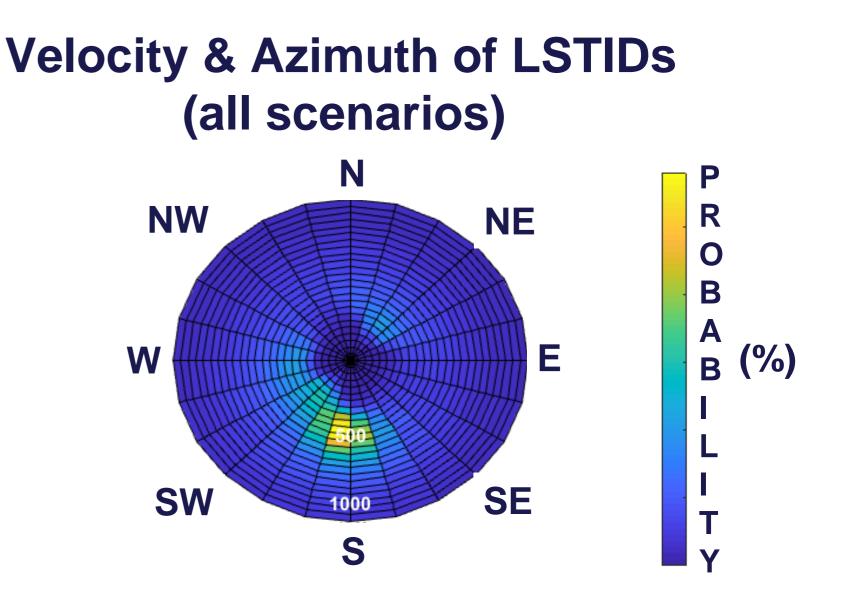


## **Velocity & Azimuth analysis**

 Most of the events present southward propagation (180°) suggesting origin of the LSTIDs by auroral activity.

TéchTIDE

 Fewer events are observed to occur to other directions, mostly from SW to NW suggesting origin of LSTIDs by sunrise effect.



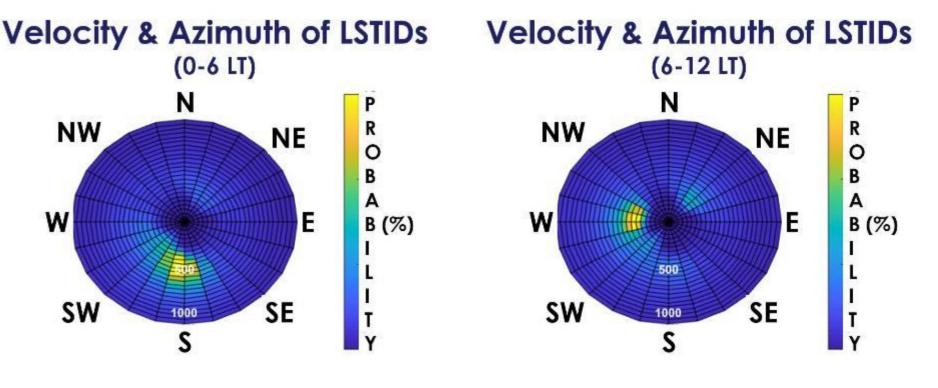


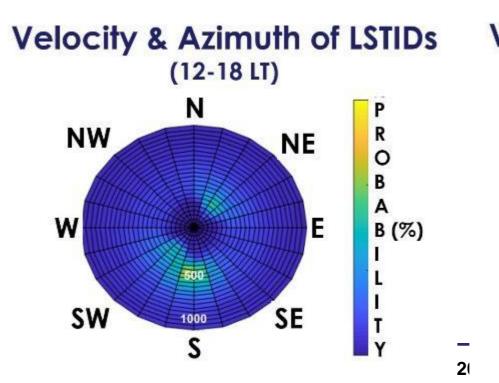


## **Diurnal Velocity & Azimuth analysis**

Propagation azimuth depends of the LT:

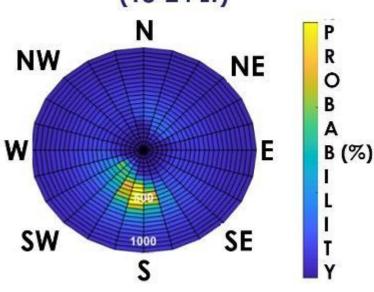
- **Night time (18-6h LT):** propagation is mostly southward that suggest origin of the LSTID by auroral activity.
- Moring time (6-12h LT): propagation is mostly westward that suggest origin of the LSTID by the solar terminator effect due to sunrise.
- Afternoon time (12-18h LT): Fewer number of events occur and dominant propagation direction is S-SW.







Velocity & Azimuth of LSTIDs (18-24 LT)



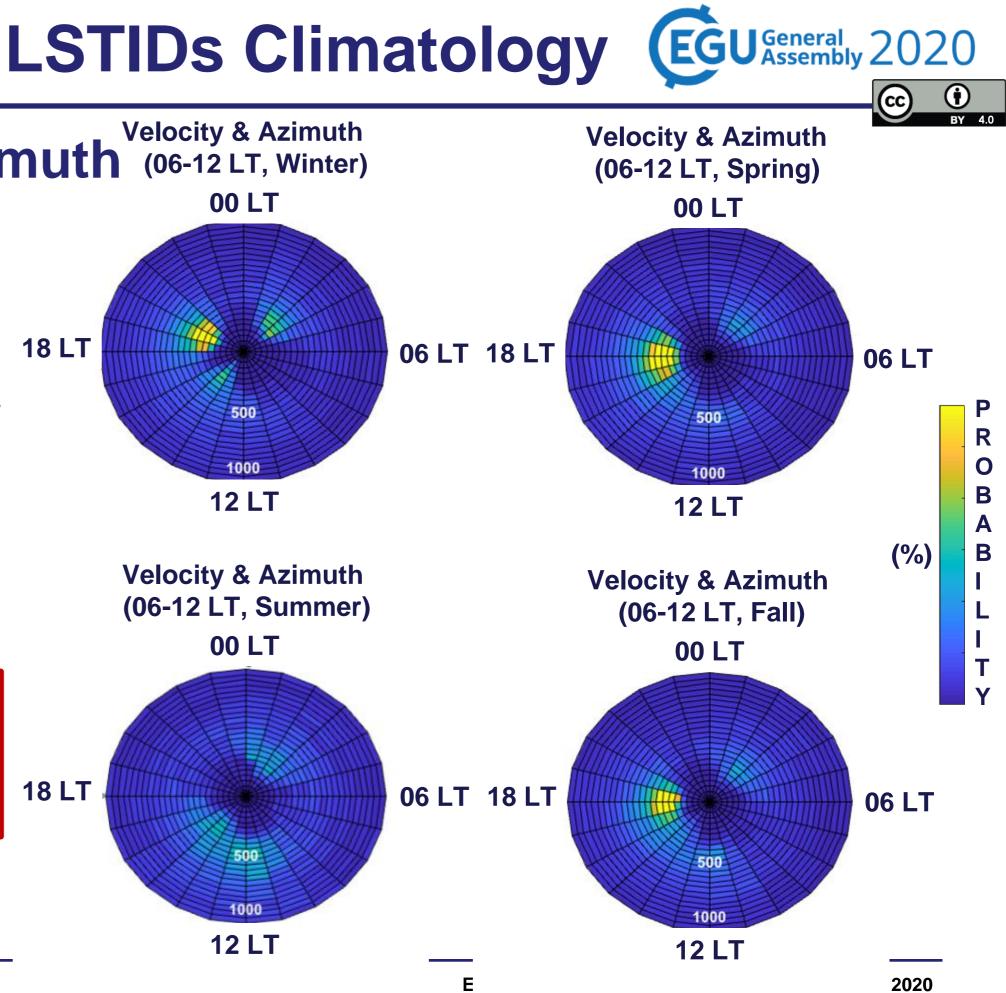




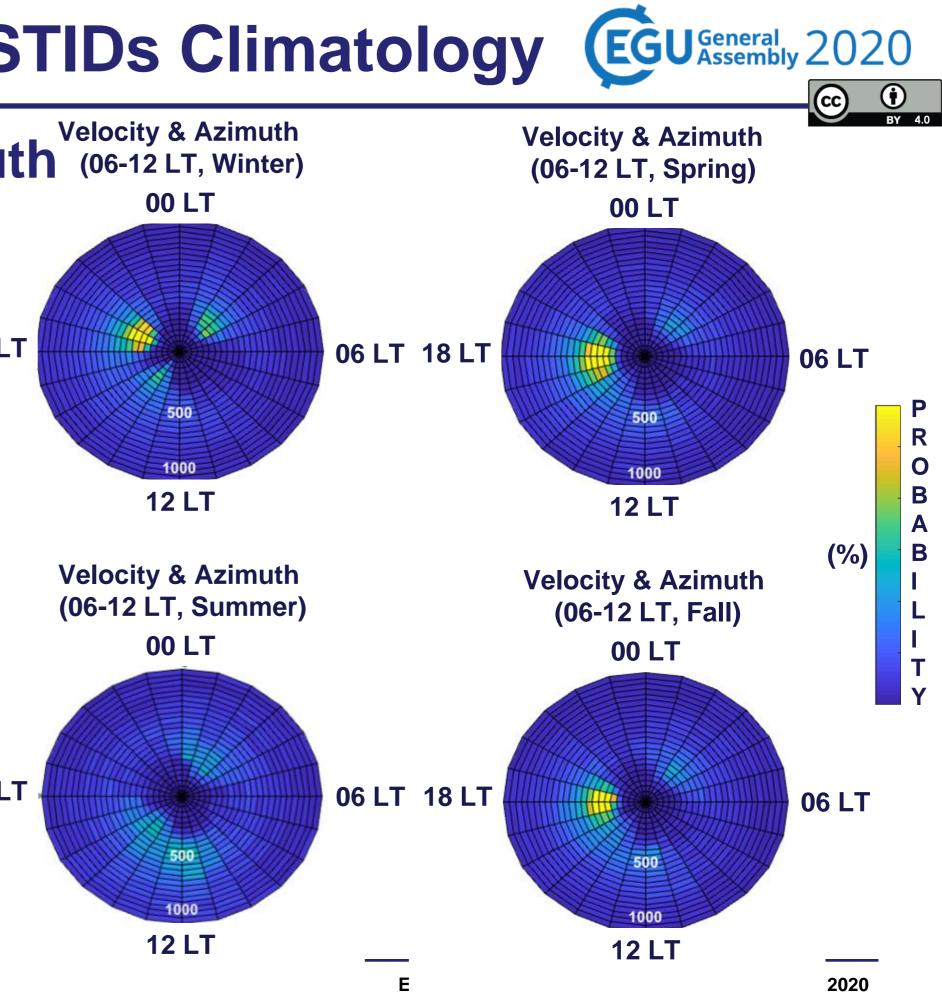
### Morning sector Velocity & Azimuth (06-12 LT, Winter) **00 LT** seasonal analysis

- Northwestward dominant propagation for Winter.
- **Westward** dominant propagation for lacksquareEquinoxes.
- **Southwestward** dominant propagation for **Summer**.

Results suggest **Solar Terminator** and Prevailing **Thermospheric** wind influence.











- LSTIDs detection based ionosonde data work efficiently in NRT for European Region.
- The method provides dominant Period, Amplitude and vector Velocity of propagation.
- The **method** have been **tested** for geomagnetically **disturbed periods**.
- Presence of **distinct LSTID** signatures with dominant **periods** of **90-120 minutes** and southward velocities of 400 - 800 m/s.
- **Climatology** observes **dominant activity** near Sun rise and at night. **Morning** sector  $\bullet$ observes prevailing westward propagation and evening and nighttime observe southward.
- Results suggest origin by the solar terminator and auroral activity and that thermospheric wind influence prevailing propagation.







### This work has been partly developed under the TechTIDE project and MIRA project.



detection, Monitoring and modelling of Ionospheric irRegulArities PGC2018-096774-B-I00 (MCIU, AEI, FEDER)



# Thank you for you attention

# Keep safe and see you in 2021!



EGU2020: Sharing Geoscience Online, 4 - 8 May 2020