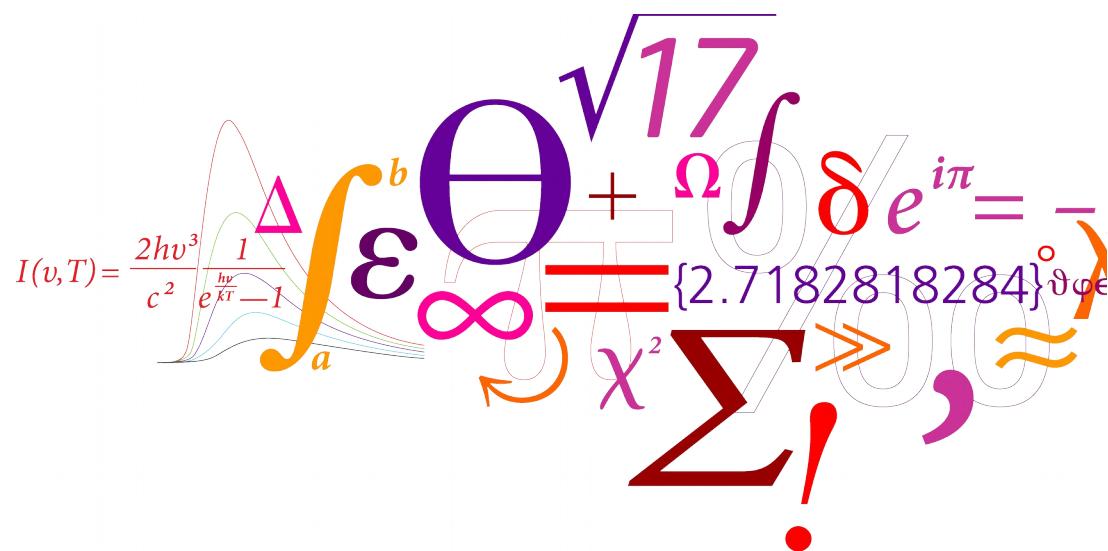


Streamer discharges in the atmosphere of Primordial Earth

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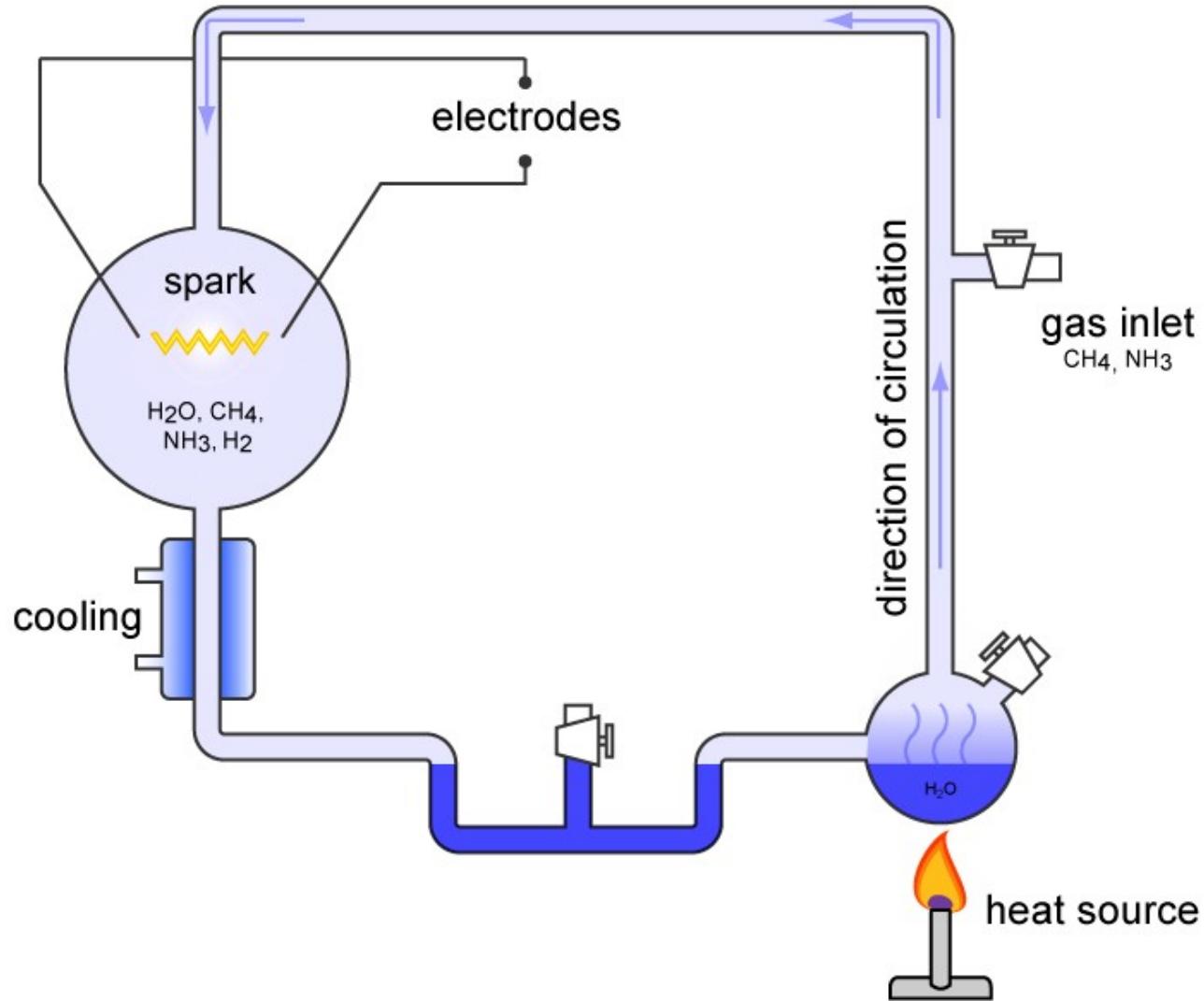
$$I(v, T) = \frac{2hv^3}{c^2} \frac{1}{e^{\frac{hv}{kT}} - 1}$$


The collage includes a graph of a function $I(v, T)$ with multiple curves, a large purple circle with a horizontal line through it, a purple infinity symbol, a purple square root symbol, a purple integral with variables a and b , a purple summation symbol with an exclamation mark, and a purple equals sign with a red arrow pointing to the right.

Outline

- Introduction
 - Life on primordial Earth
 - Electric discharges and streamers
- Electric breakdown field and swarm parameters
- Photoionization vs. background ionization
- Results
 - Electron densities
 - Front velocities
 - Electron energies
 - Inception fields
- Conclusions and outlook

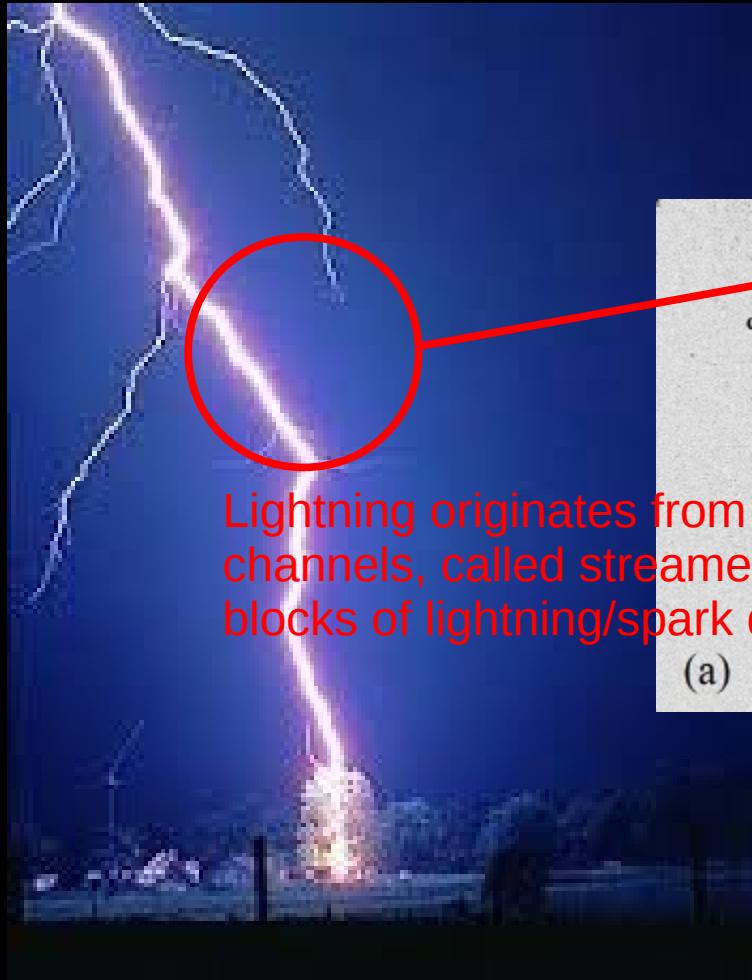
Introduction: Life on Primordial Earth



Miller-Urey experiments suggests that the formation of amino and carbon acids was catalyzed by electric discharges (lightning?) in atmosphere

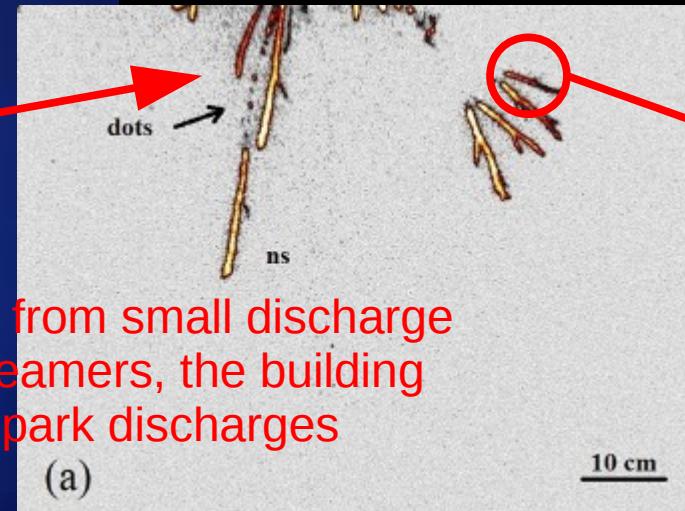
=> what about inception of discharges in the atmosphere of Primordial Earth???

Introduction: Streamer discharges

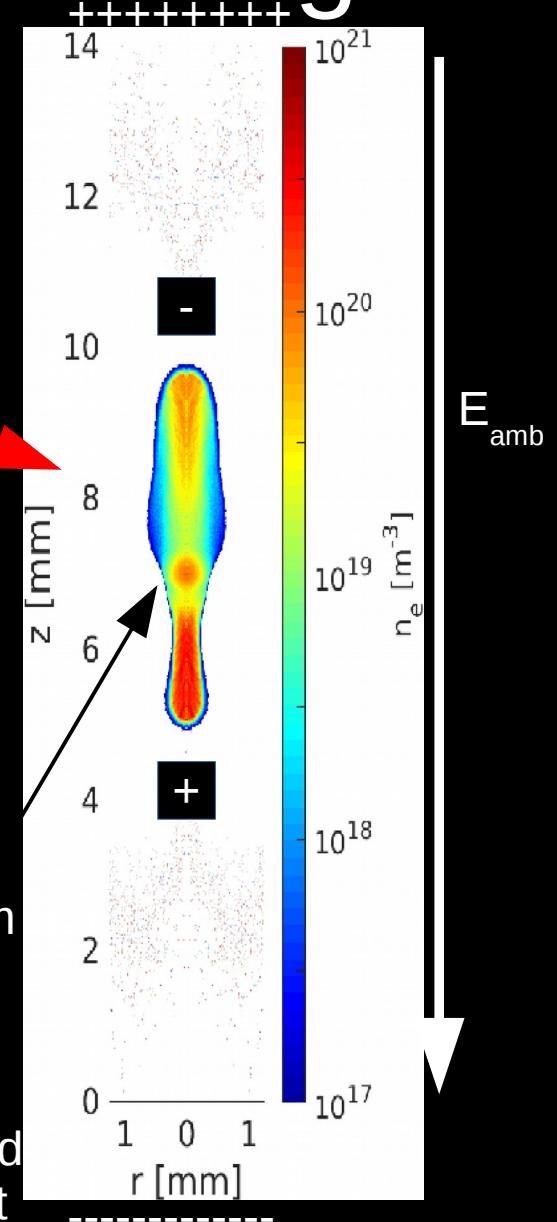


Lightning originates from small discharge channels, called streamers, the building blocks of lightning/spark discharges

(a)

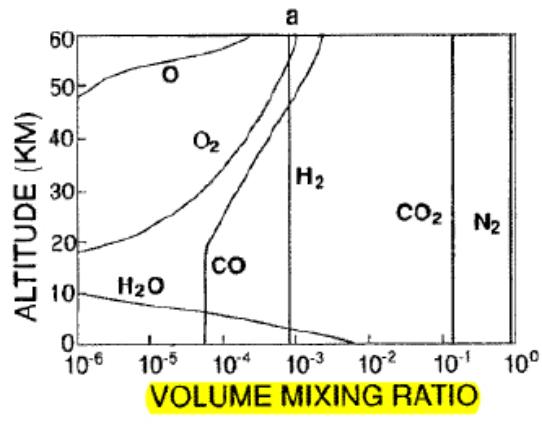


Streamers originate from electron-ion patches in an ambient electric field E_{amb} and evolve by electron acceleration and the ionization of ambient gas (high fields at the tips)



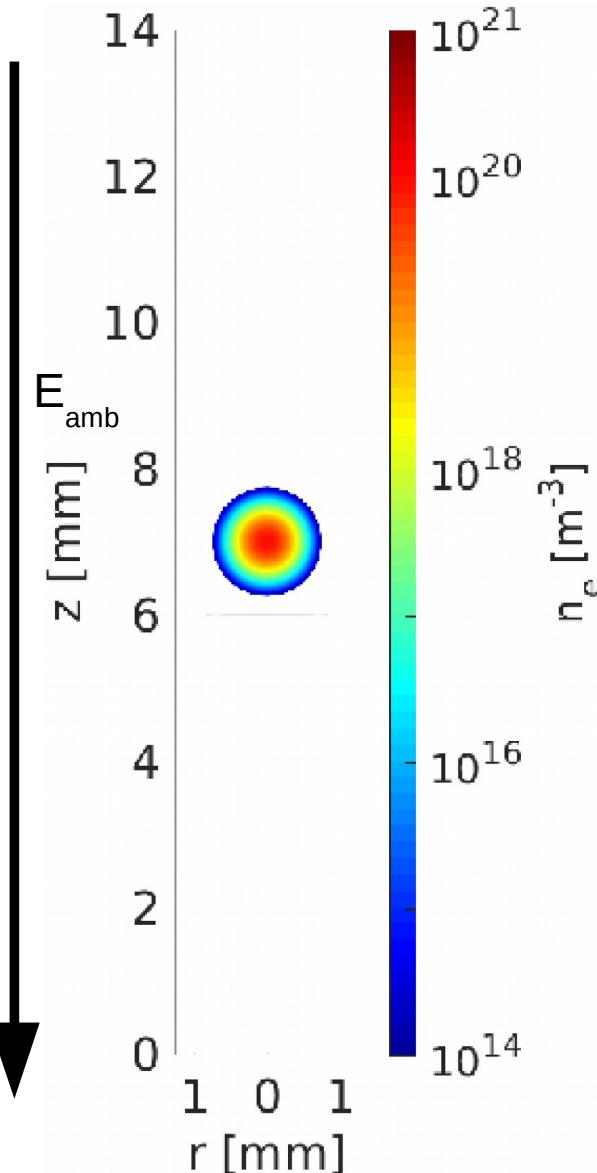
Introduction: Atmospheres on Primordial Earth

- To understand streamers on Primordial Earth and associated chemistry, one needs to know composition of atmosphere
- Miller and Urey used $\text{H}_2\text{O}:\text{CH}_4:\text{NH}_3:\text{H}_2 = 37.5\%:25\%:25\%:12.5\%$ (reducing atmosphere)
- More recent review [Kasting, 1990. Orig. Life Evol. Biosph., 20, 199–231]:



$\text{N}_2:\text{CO}_2:\text{H}_2\text{O}:\text{H}_2:\text{CO} =$
80%:18.89%:1%:0.1%:0.01%
(oxidizing atmosphere)

Modelling of streamers



Ingredients:

- initial electron/ion patch $n_{e,i}(t=0, r, z) = n_{e,0} e^{-\frac{r^2 + (z - z_0)^2}{\lambda^2}}$
($n_{e,0} = 10^{20} \text{ m}^{-3}$, $\lambda = 0.2 \text{ mm}$, $z_0 = 7 \text{ mm}$)
- ambient field E_{amb}
- ambient gas mixture (Urey-Miller/UM, Kasting)
- gas pressure/number density of ambient gas
(here use $2.5 \cdot 10^{25} \text{ m}^{-3}$ as for Modern Earth)

Modelling:

- 2.5D particle code (2D in space, 3D in velocity)
- tracing individual electrons (incl. collision with ambient gas)
- cross section sets as input data for electron scattering in individual gases
- vary mixture and ambient field => study inception of streamer discharges

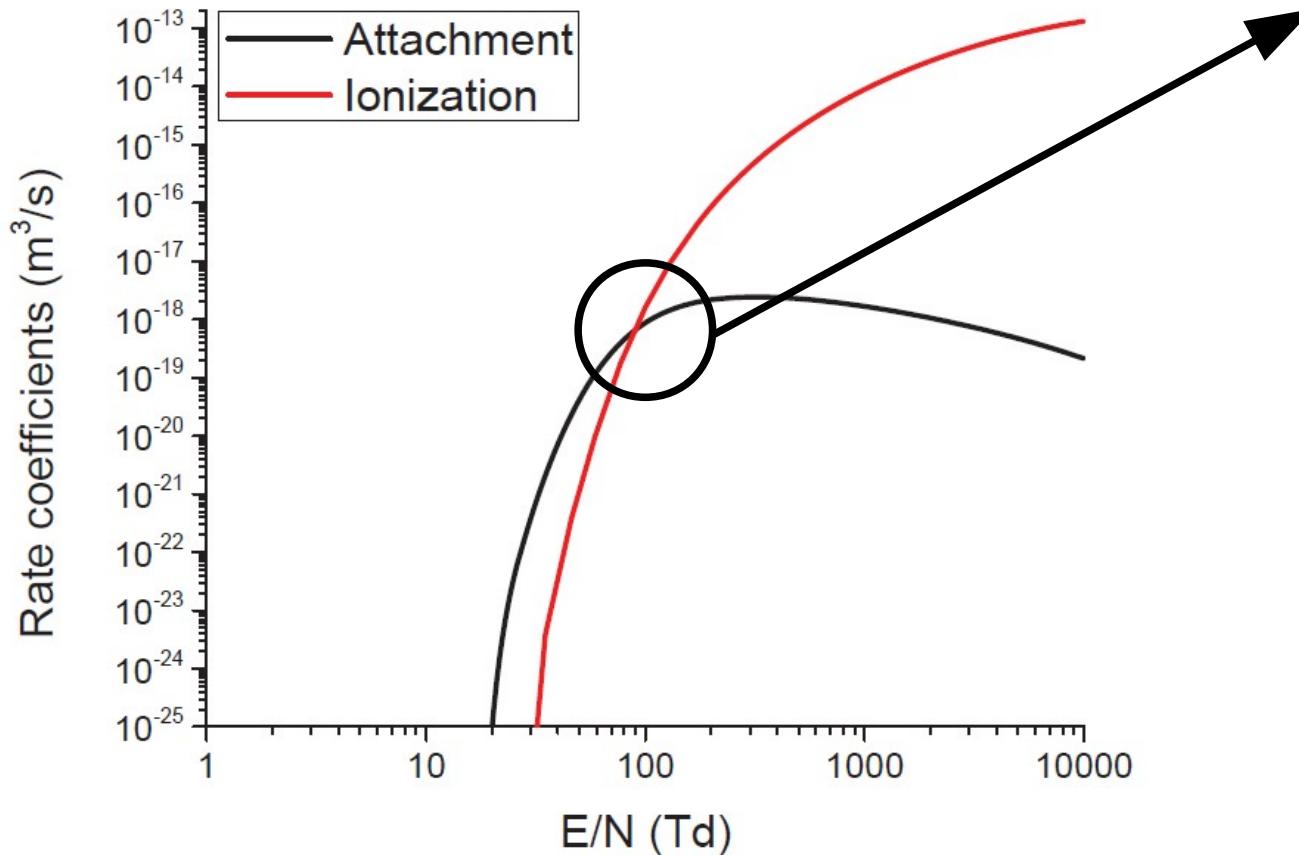
Breakdown fields I

- Electron avalanche/streamer evolution depends on two competing mechanisms (depending on the electric field)
 - Ionization: $A + e^- \rightarrow A^+ + 2e^-$
 - Attachment, e.g. $A + e^- \rightarrow A^-$
- Equilibrium of rates determines the breakdown (or classical) electric field $E_k \Rightarrow \alpha_{\text{ion}}(E_k) = \alpha_{\text{att}}(E_k)$

(α_{ion} : ionization rate
 α_{att} : attachment rate
 E_k : classical electric field indicating (not equal to) field when avalanche-to-streamer transition occurs)

Breakdown fields II

Rates for Kasting's mixture:



$$E_{k,\text{Kasting}} = 90.3 \text{ Td} = 2.3 \text{ MV m}^{-1}$$

Similarly:

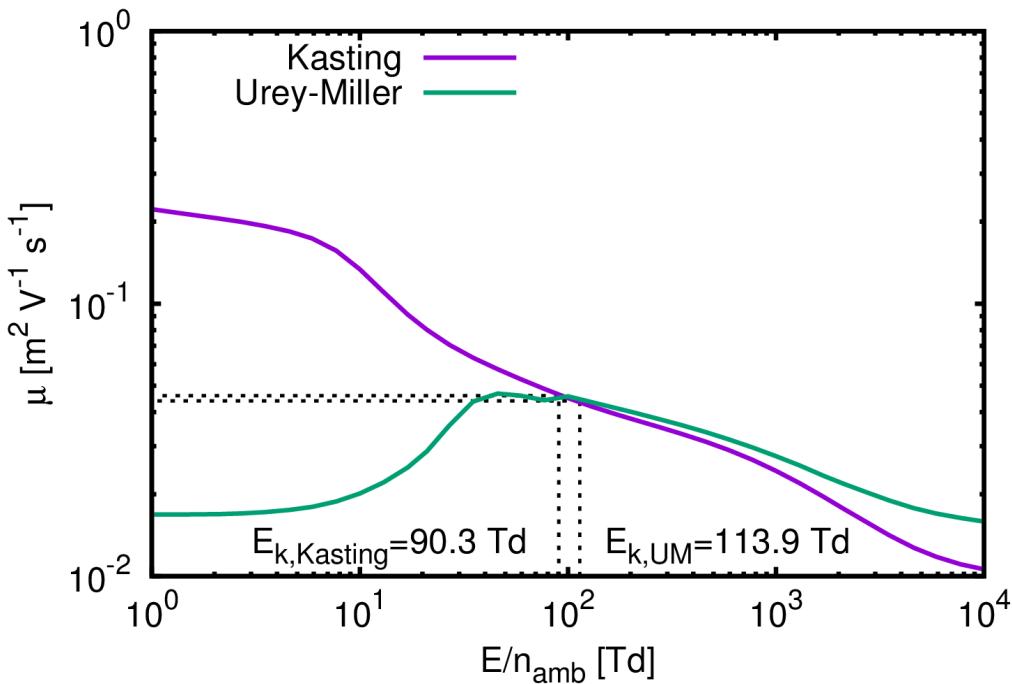
$$E_{k,\text{UM}} = 113.9 \text{ Td} = 2.9 \text{ MV m}^{-1}$$

$$E_{k,\text{ME}} = 125.6 \text{ Td} = 3.2 \text{ MV m}^{-1}$$

Reduced electric field [Td] as ratio of electric field and ambient gas density, here $n_{\text{amb}} = 2.547 \cdot 10^{25} \text{ m}^{-3}$
(avalanche/discharge properties scale with gas density)

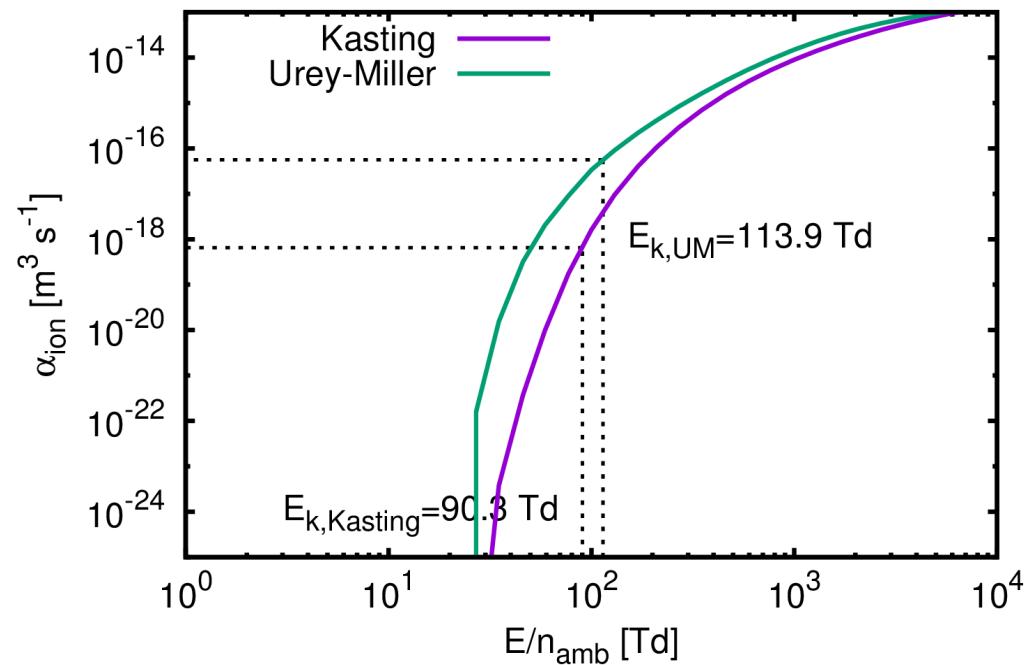
Swarm parameters

Electron mobility



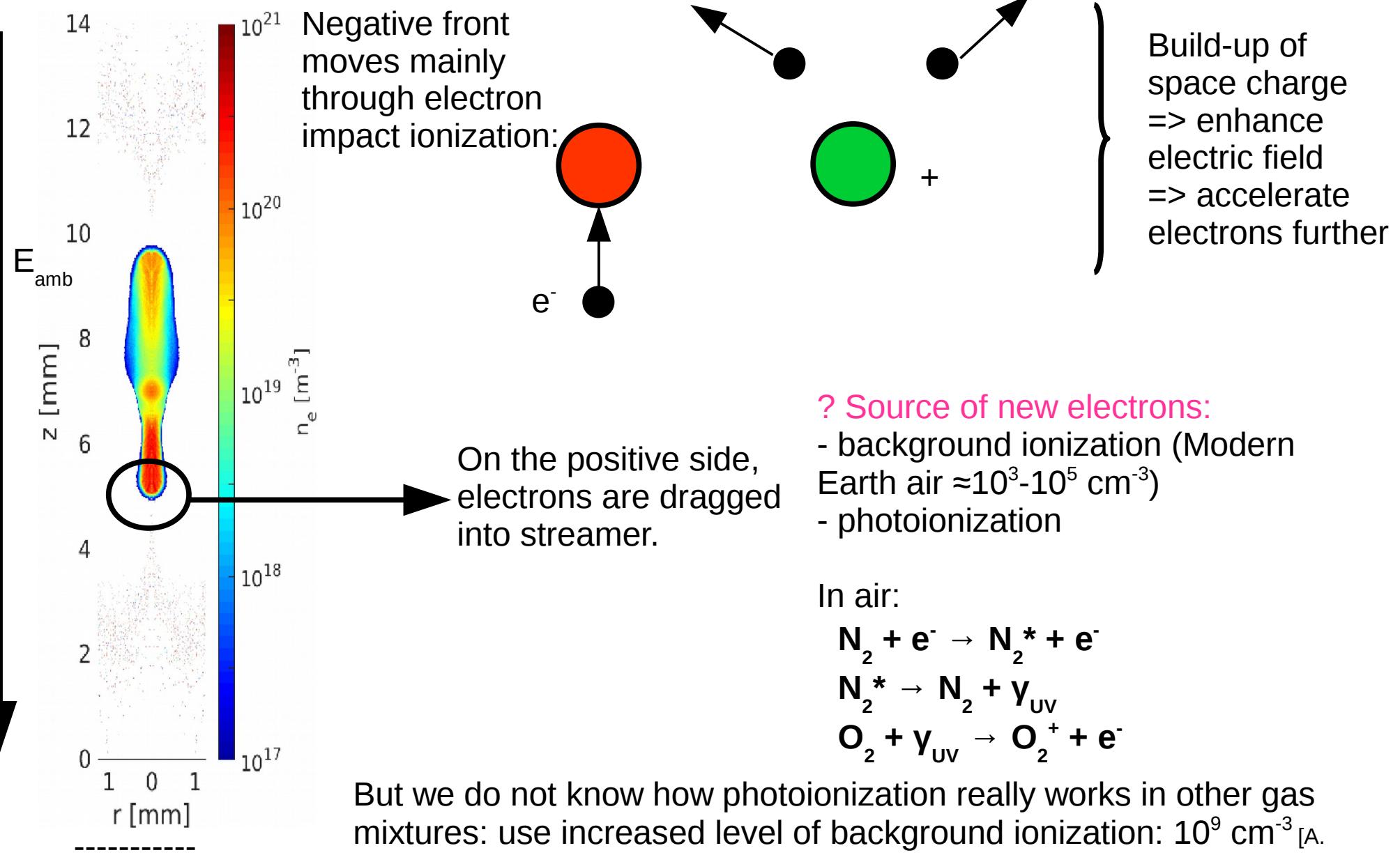
above breakdown: electrons more mobile in Urey-Miller mixture

Ionization rate



ionization more efficient in Urey-Miller mixture

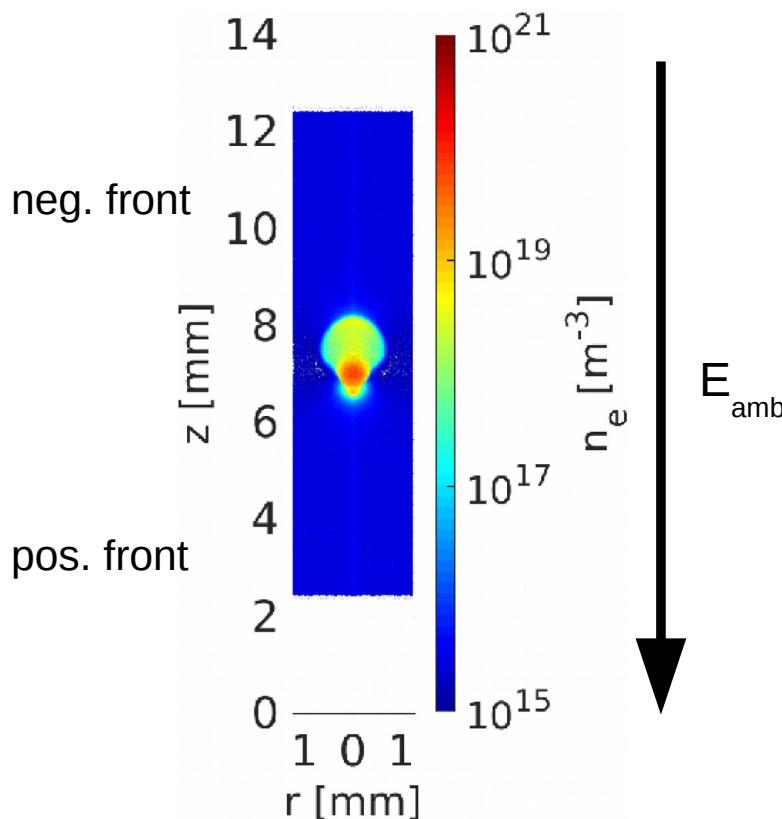
Sources of ionization: Photoionization and background ionization



Electron densities I

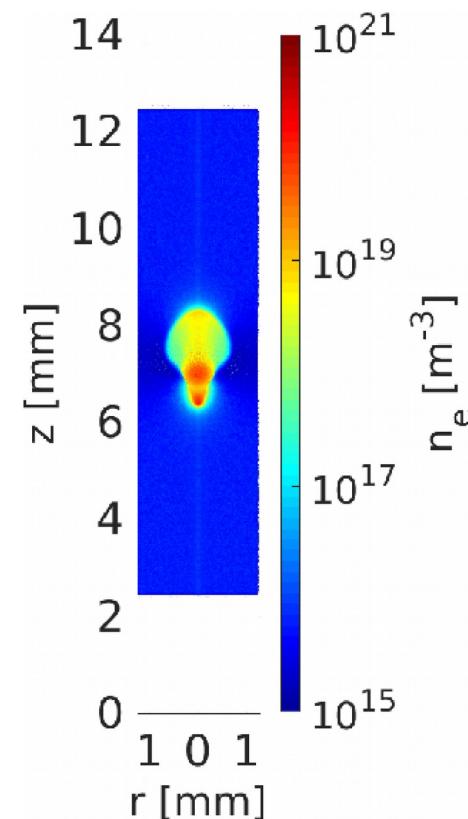
In Kasting mixture after $t=2.85$ ns for different ambient fields:

$$\begin{aligned}E_{\text{amb}} &= 3.6 \text{ MV m}^{-1} \\&= 141.3 \text{ Td}\end{aligned}$$



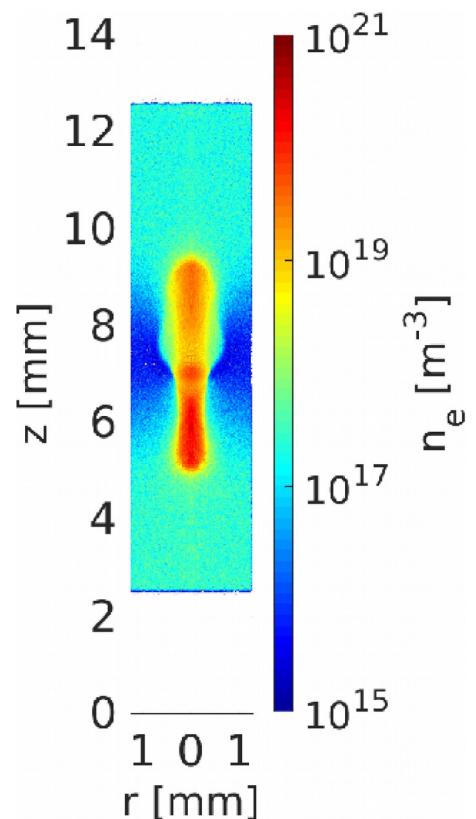
electron avalanches on both
fronts, positive front might
turn to streamer

$$\begin{aligned}E_{\text{amb}} &= 4.0 \text{ MV m}^{-1} \\&= 157.0 \text{ Td}\end{aligned}$$



negative front: electron
avalanche, but no streamer
positive front: avalanche-to-
streamer transition

$$\begin{aligned}E_{\text{amb}} &= 5.0 \text{ MV m}^{-1} \\&= 196.3 \text{ Td}\end{aligned}$$

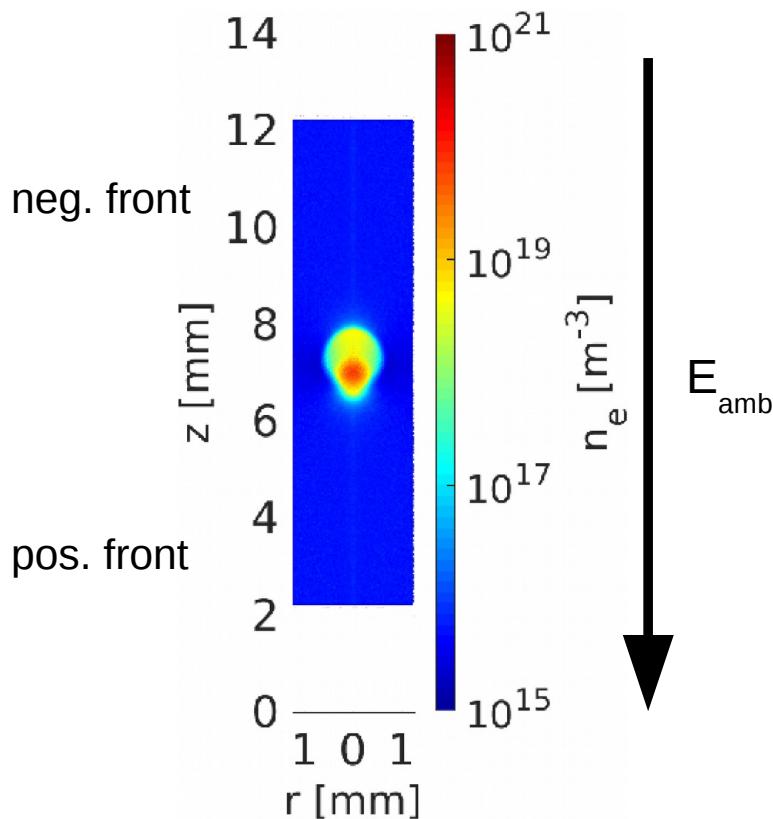


full evolution of
positive and negative
streamer

Electron densities II

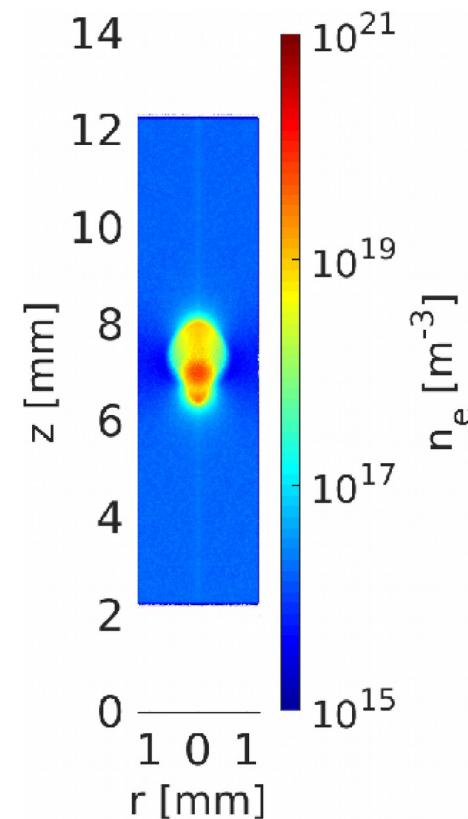
In Urey-Miller mixture after $t=1.81$ ns for different ambient fields:

$$\begin{aligned}E_{\text{amb}} &= 2.5 \text{ MV m}^{-1} \\&= 98.2 \text{ Td}\end{aligned}$$



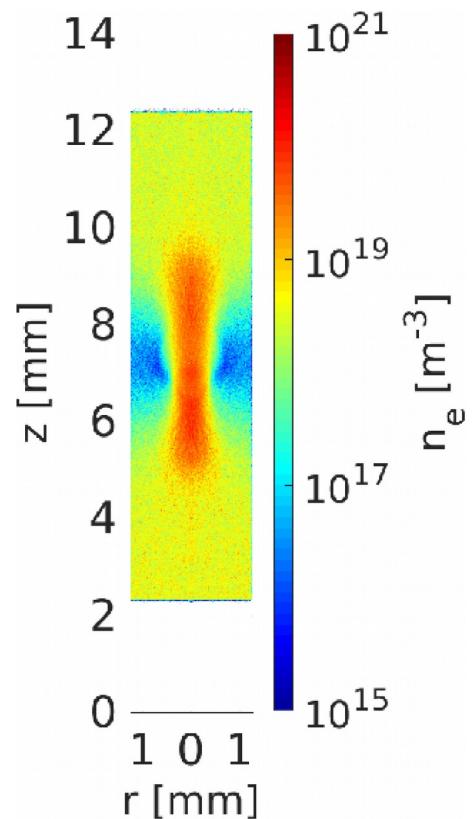
electron avalanches on both
fronts

$$\begin{aligned}E_{\text{amb}} &= 2.9 \text{ MV m}^{-1} \\&= 113.9 \text{ Td}\end{aligned}$$



avalanche-to-streamer
transition for both polarities

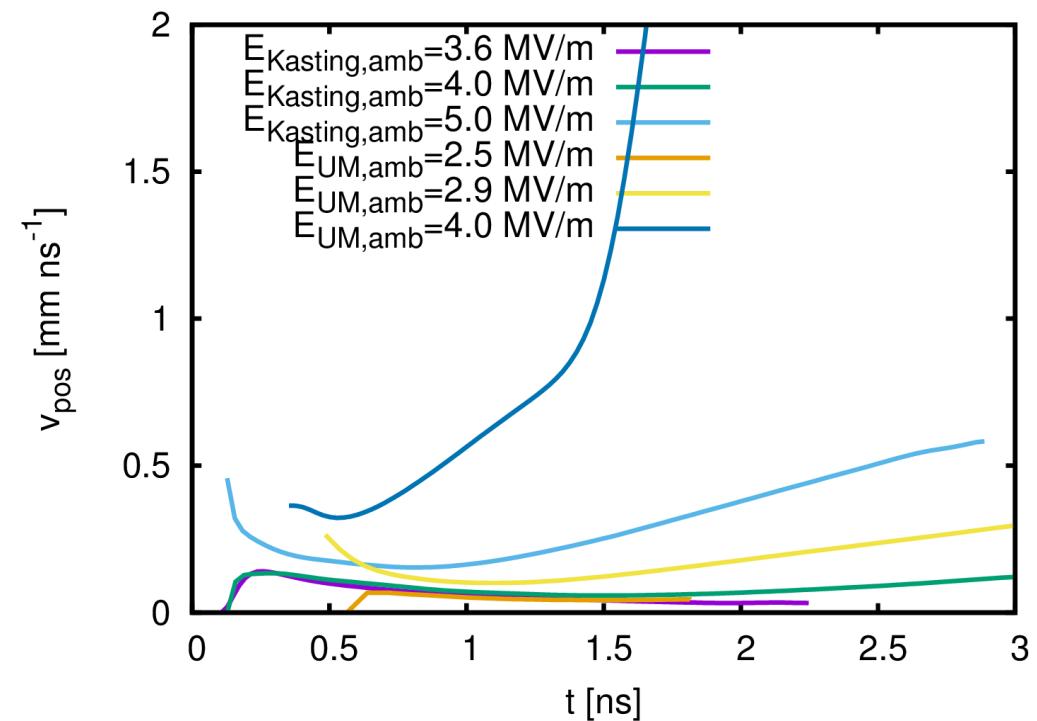
$$\begin{aligned}E_{\text{amb}} &= 4.0 \text{ MV m}^{-1} \\&= 157.0 \text{ Td}\end{aligned}$$



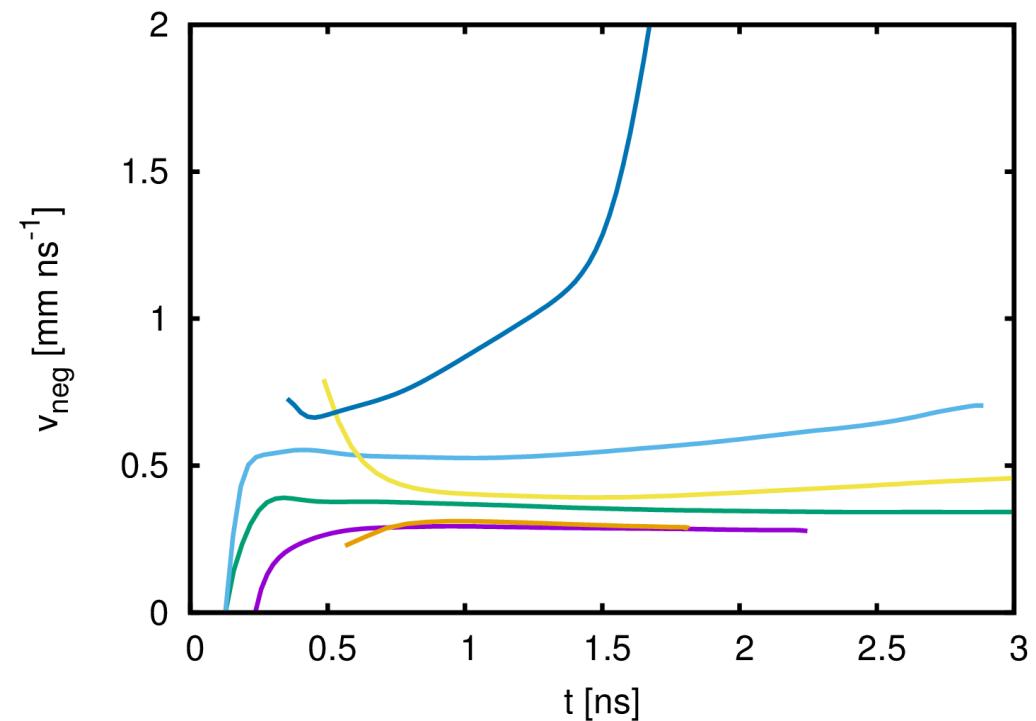
full evolution of
positive and negative
streamer

Front velocities

positive front:



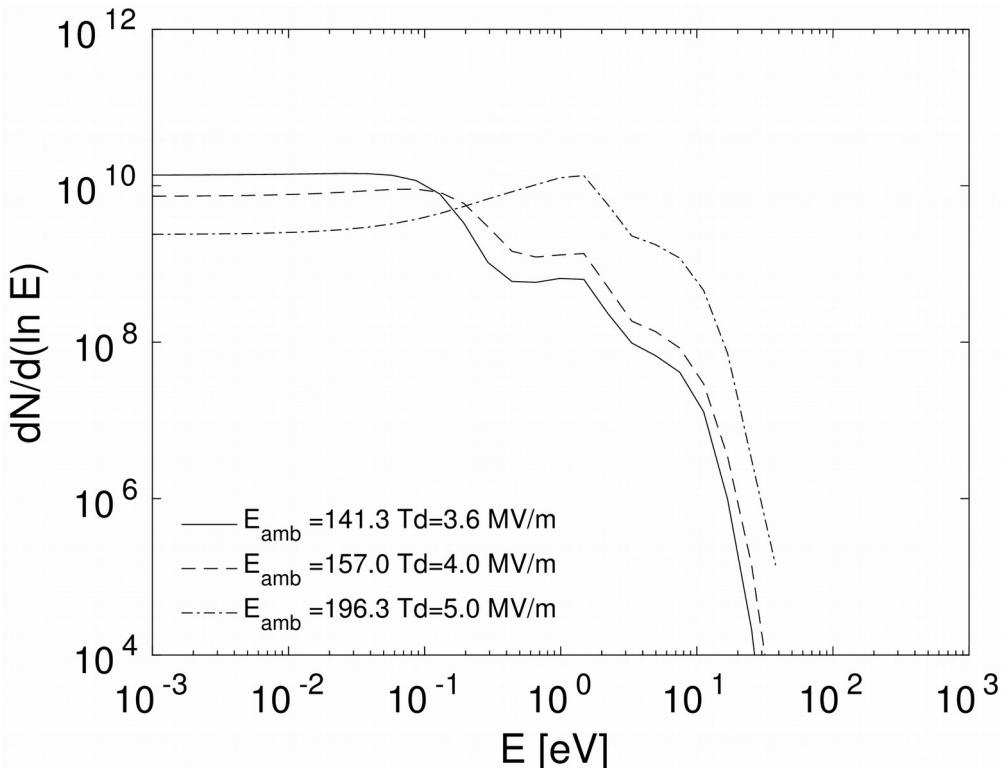
negative front:



- in Kasting mixture: $v < 0.6 \text{ mm ns}^{-1}$
- velocities in Urey-Miller (in average) larger (although smaller electric fields)
- for comparison: $v_{\text{Modern Earth}} \approx 1-10 \text{ mm ns}^{-1}$

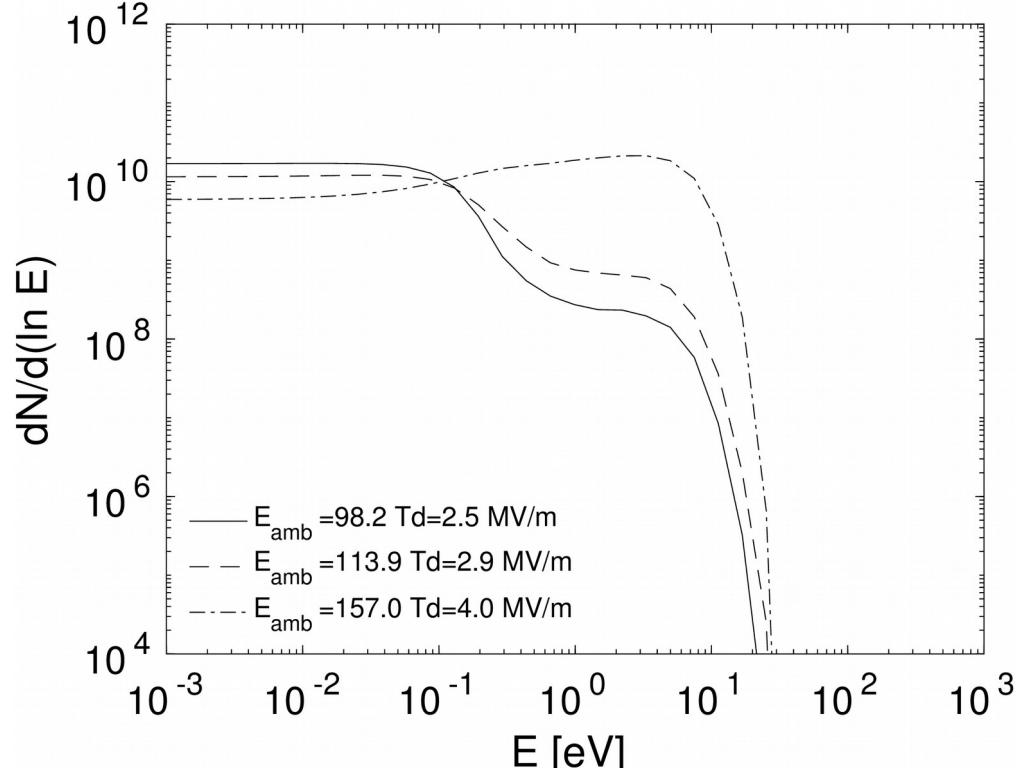
Electron energies

Kasting, t=2.85 ns



- for 141.3 Td and 157.0 Td $E_{\text{max}} < 30 \text{ eV}$
- for 196.3 Td $E_{\text{max}} \approx 40 \text{ eV}$

Urey-Miller, t=1.81 ns



- for all considered cases $E_{\text{max}} < 30 \text{ eV}$

=> similar electron energies
 => mean ionization energy in Kasting mixture larger than in Urey-Miller mixture
 => discharge inception/evolution less efficient in Kasting mixture

Inception fields

(field when avalanche-to-streamer transition occurs)

Gas mixture	$E_{inc, pos} \text{ [MV m}^{-1}\text{]}/[\text{Td}]$	$E_{inc, neg} \text{ [MV m}^{-1}\text{]}/[\text{Td}]$
Kasting $\text{N}_2:\text{CO}_2:\text{H}_2\text{O}:\text{H}_2:\text{CO}$	3.7/145.3	4.6/180.6
Urey-Miller $\text{H}_2\text{O}:\text{CH}_4:\text{NH}_3:\text{H}_2$	2.7/106.0	2.7/106.0

(for comparison, on Modern Earth: $\approx 3.2/125.6$)

Conclusions and outlook

- Streamer discharge inception in Kasting mixture requires higher fields than in Urey-Miller mixture
- Front velocities in Kasting mixtures smaller than in Urey-Miller mixture
- Comparable energies (<40 eV)

=> Discharges (lightning) inception might be more difficult in the oxidizing Kasting mixture

(does not say anything about local reducing environments)

- Future project: include plasma chemistry of Primordial Earth into streamer code