

# SATELLITE AND GROUND-BASED MAGNETIC FIELD OBSERVATIONS RELATED TO VOLCANIC ERUPTIONS

K. Schwingenschuh<sup>1</sup>, W. Magnes<sup>1</sup>, X. Shen<sup>2</sup>, J. Wang<sup>3</sup>, B. Cheng<sup>7</sup>, A. Pollinger<sup>1</sup>, Chr. Hagen<sup>1</sup>, R. Lammegger<sup>5</sup>, M. Ellmeier<sup>5</sup>, Chr. Schirnberger<sup>1</sup>, H.-U. Eichelberger<sup>1</sup>, B. Mandl<sup>1</sup>, M. Y. Boudjada<sup>1</sup>, B. P. Besser<sup>1</sup>, A. A. Rozhnoi<sup>4</sup>, T. Zhang<sup>1</sup>, M. Delva<sup>1</sup>, I. Jernej<sup>1</sup>, Ö. Aydogar<sup>1</sup>, and R. Leonhardt<sup>6</sup>

<sup>1</sup>Space Research Institute, Austrian Academy of Sciences, Graz, Austria, <sup>2</sup>Institute of Earthquake Admin., Beijing, China, <sup>3</sup>Center for Space Science and Applied Research, CAS, Beijing, China, <sup>4</sup>Institute of the Earth Physics, RAS, Moscow, Russia, <sup>5</sup>Institute of Experimental Physics, Graz Univ. of Technology, Graz, Austria, <sup>6</sup>Central Institution for Meteorology and Geodynamics (ZAMG), Vienna, Austria, <sup>7</sup>National Space Science Center, CAS, Beijing, China

In this study we investigate volcanic eruption phenomena related to ionospheric disturbances, e.g. Heki (2006) used total electron content (TEC) measurements for this task. In particular, a model is developed how discharge phenomena (e.g. Houghton et al., 2013) can produce magnetic field variations at SWARM and CSES satellite orbits, i.e. altitudes of ~500 km in the F-region. Several coupling mechanism between lithosphere, atmosphere, and ionosphere are discussed by Simões et al. (2012). Experimental evidence is based on magnetic field observations aboard CSES mission in the time frame July 2018 to July 2019. The theoretical considerations include the source mechanism, propagation path, and the signal strength at low earth orbit satellite altitude.

## CSES HPM MAGNETIC FIELD MEASUREMENTS

- China Seismo Electromagnetic Satellite (CSES) launched Feb. 2<sup>nd</sup>, 2018; altitude 507 km, inclination 97.4 deg, Sun synchronous polar orbit, 3 axes stabilized, 8 payloads, among them the High precision magnetometer (HPM) with two Fluxgates and scalar CDSM (Shen et al., 2018), data latitude range -65 deg to +65 deg
- Main objective of the mission is the study of ionospheric response (electromagnetic and plasma phenomena) related to natural hazards, e.g. in particular earthquakes (precursors) and volcanic eruptions (this study)
- Coupled Dark State Magnetometer (CDSM), scalar magnetometer based on two-photon spectroscopy, instrument noise level ~50 pT/sqrt(Hz) @ 1 Hz, accuracy 0.19 nT ( $1\sigma$ ), 1 Hz data product, 5 m nadir pointing magnetometer boom, total magnetic field 1000 - 100000 nT. (Pollinger et al., 2018) (Zhou et al., 2018)

## B-FIELD VARIATIONS AT CSES ALTITUDE

- **SOURCES:** Microfracture crustal electrification (Molchanov and Hayakawa, 1998); volcanic plume charging (Houghton et al., 2013); displacement of magnetized solid material, for the point source  $\mathbf{B} = \mu_0/4\pi (3\mathbf{M}\mathbf{r}/r^5 - \mathbf{M}/r^3)$ , linear source (prop  $1/r^2$ ) or planar source (prop  $1/r$ ); and atmospheric waves (acoustic- and gravity modes)
- **PROPAGATION & COUPLING:** As discussed by Simões et al. (2012), in particular ionospheric fluctuations (electromagnetic, density) triggered by volcanic acoustic waves, see Schirnberger et al. (2019), and for seismic investigations Kherani et al. (2009)

## VOLCANIC ERUPTIONS AND ACTIVE PERIODS

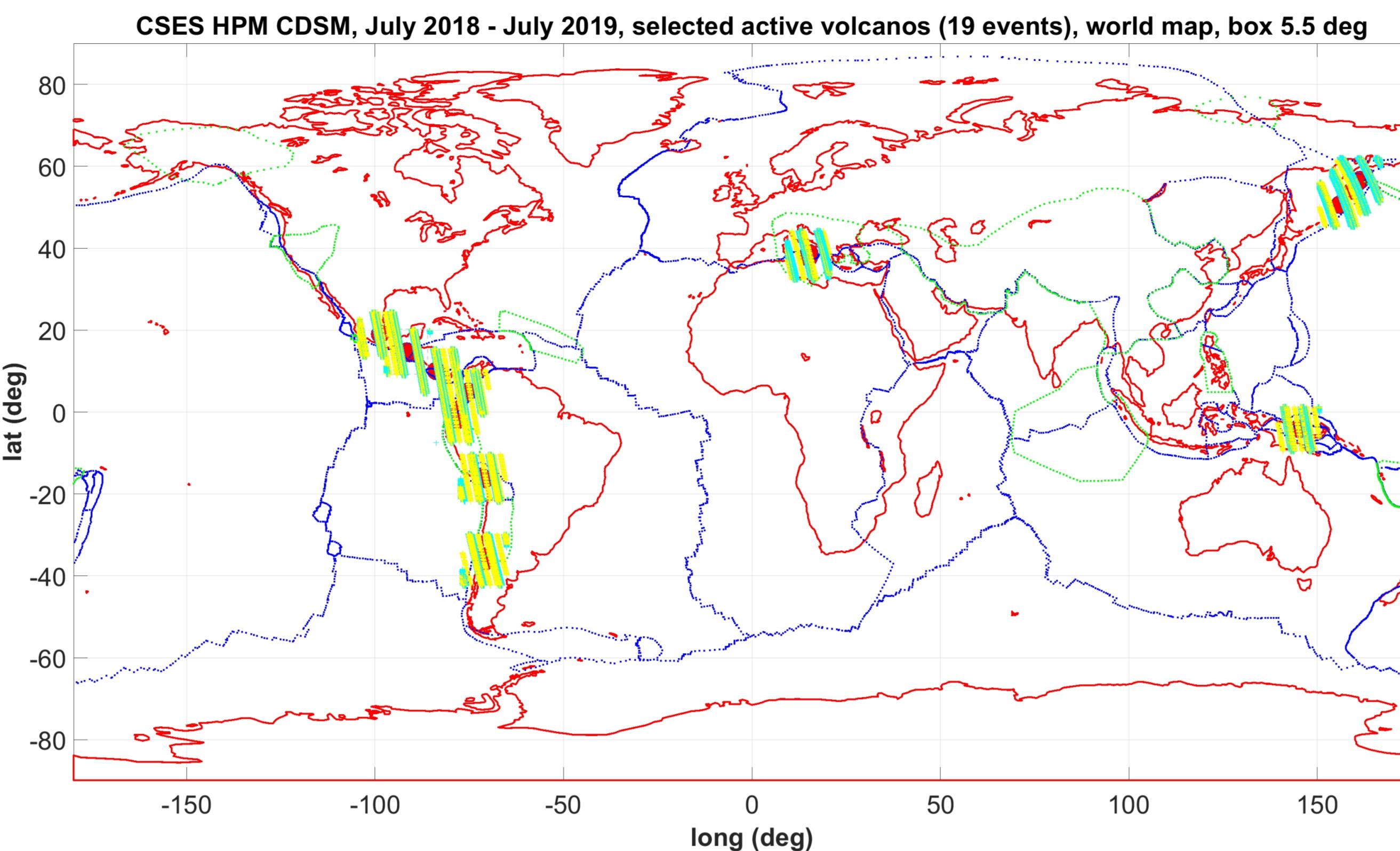


Figure 1: Selected 19 volcanoes with projected CSES paths in a spatial box with 5.5 deg, night-time magnetic field CDSM data shown in yellow/cyan color for major pre/post eruption periods.

CSES HPM magnetic field variations at ionospheric altitude 507 km: volcano list, time span July 2018 – July 2019			
No.	Lat. (deg)	Long. (deg)	Volcano, Country
1	19.02	-98.62	Popocatépetl, Mexico
2	56.65	161.36	Shiveluch, Russia
3	-0.08	-77.66	Reventador, Ecuador
4	-35.24	-70.57	Planchón-Peteroa, Argentina/Chile
5	-15.78	-71.85	Sabancaya, Peru
6	50.68	156.02	Ebeko, Russia
7	-4.08	145.04	Manam, Papua New Guinea
8	14.47	-90.88	Fuego, Guatemala
9	55.98	160.59	Bezymianny, Russia
10	56.06	160.64	Klyuchevskoi, Russia
11	10.03	-83.77	Turrialba, Costa Rica
12	-36.86	-71.38	Nevados de Chillán, Chile
13	4.90	-75.32	Nevado del Ruiz, Colombia
14	38.79	15.21	Stromboli, Italy
15	37.75	14.99	Etna, Italy
16	-2.00	-78.34	Sangay, Ecuador
17	-16.36	-70.90	Ubinas, Peru
18	54.05	159.44	Karymsky, Russia
19	14.76	-91.55	Santa María, Guatemala

Table 1: Selected volcanoes with eruptions and unrest periods in the time frame July 2018 – July 2019 (Database GVP), annotations for CSES/HPM data. The color coding in the first column indicates nearby volcanoes, i.e. the spatial mapping at CSES altitude (~ 507 km) is difficult to resolve because of nearly same data segments.

Parallel ground based magnetic field observations (magnetometer networks) nearby volcanoes can be means to characterize eruptions and unrest periods, similar as for earthquakes (Pratt et al., 2008), but future studies have to demonstrate this in detail.

## REFERENCES

- Heki, K., *Explosion energy of the 2004 eruption of the Asama Volcano, central Japan, inferred from ionospheric disturbances*, GRL, 33, L14303, 2006. doi:10.1029/2006GL026249
- Houghton, I. M. P., K. L. Aplin, and K. A. Nicoll, *Triboelectric Charging of Volcanic Ash from the 2011 Grímsvötn Eruption*, PRL, 111, 118501, 2013. doi:10.1103/PhysRevLett.111.118501
- Simões, F., R. Pfaff, J.-J. Berthelier, J. Klenzing, *A Review of Low Frequency Electromagnetic Wave Phenomena Related to Tropospheric-Ionospheric Coupling Mechanisms*, SSR, 168:551–593, 2012. doi:10.1007/s11214-011-9854-0
- Shen, X., Zhang X., Yuan S., et al., *The state-of-the-art of the China Seismo-Electromagnetic Satellite mission*, Sci China Tech Sci, 61:634–642, 2018. <https://doi.org/10.1007/s11431-018-9242-0>
- Pollinger, A., et al., *Coupled Dark State Magnetometer for the China Seismo-Electromagnetic Satellite*, Measurement Science and Technology, 29, 9, 2018. <https://doi.org/10.1088/1361-6501/aacde4>
- Zhou, B., et al., *Magnetic field data processing methods of the China Seismo-Electromagnetic Satellite*, Earth Planet. Phys., 2(6), 455–461, 2018. <http://doi.org/10.26464/epp2018043>
- Molchanov, O. and Hayakawa, M.: *On the generation mechanism of ULF seismogenic electromagnetic emissions*, Phys. of the Earth and Planet. Int., 105, 201-210, 1998
- Pratt, G., et al., *Multi-point ground-based ULF magnetic field observations in Europe during seismic active periods in 2004 and 2005*, NHESS, 8, 501–507, 2008. <https://doi.org/10.5194/nhess-8-501-2008>
- Schirnberger, C., et al., *Magnetic field investigations of volcanic and seismic phenomena with the China Seismo-Electromagnetic Satellite (CSES)*, Magnetometer workshop, Stubenberg am See, 2019.
- Kherani, E.A., et al., *Response of the ionosphere to the seismic triggered acoustic waves: Electron density and electromagnetic fluctuations*, Geophysical Journal International, 1–13, 176, 1, 2009, <https://doi.org/10.1111/j.1365-246X.2008.03818.x>
- Database: Global Volcanism Program (GVP), Smithsonian Institution, <https://volcano.si.edu/> as of May 2020

## SUMMARY

Satellite B-field meas. can be a tool to observe volcanic eruptions



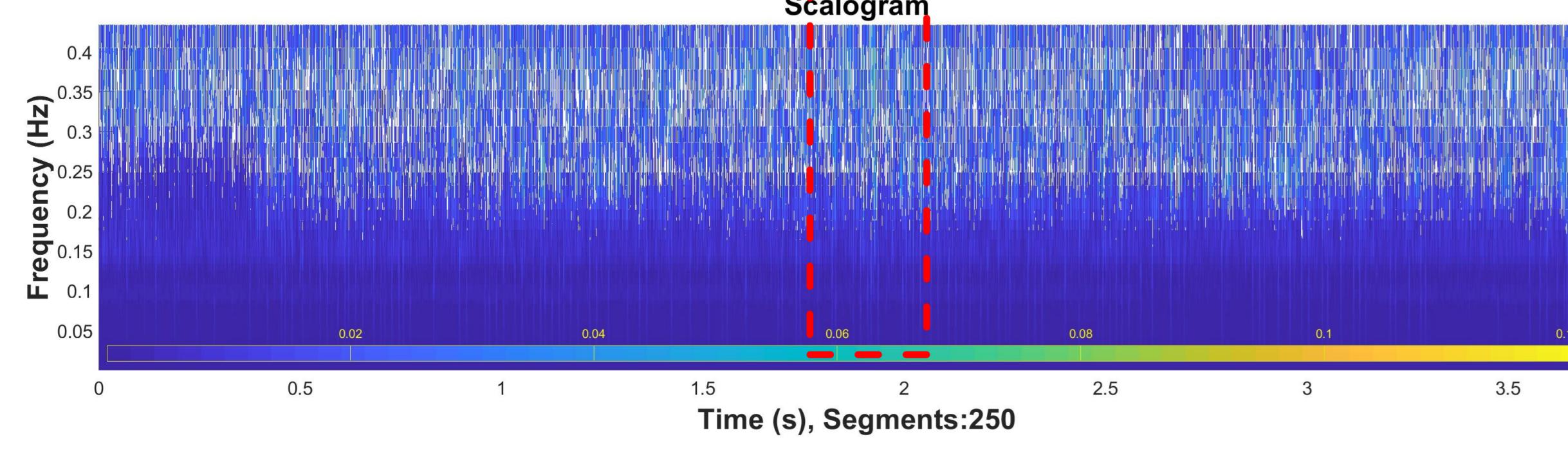
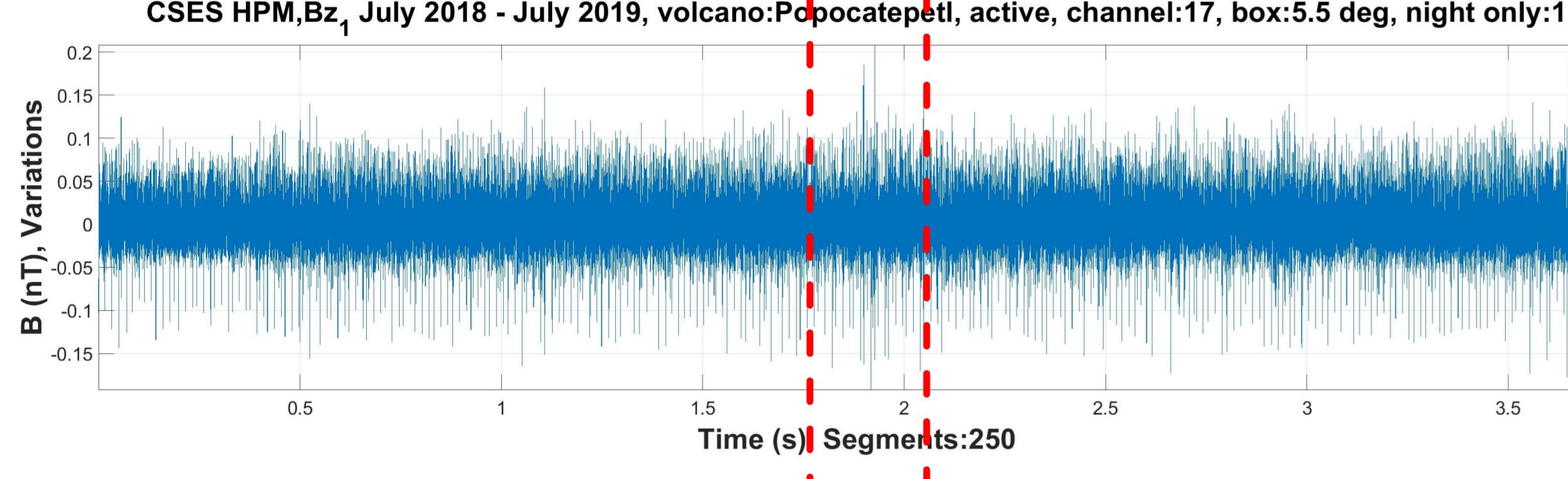
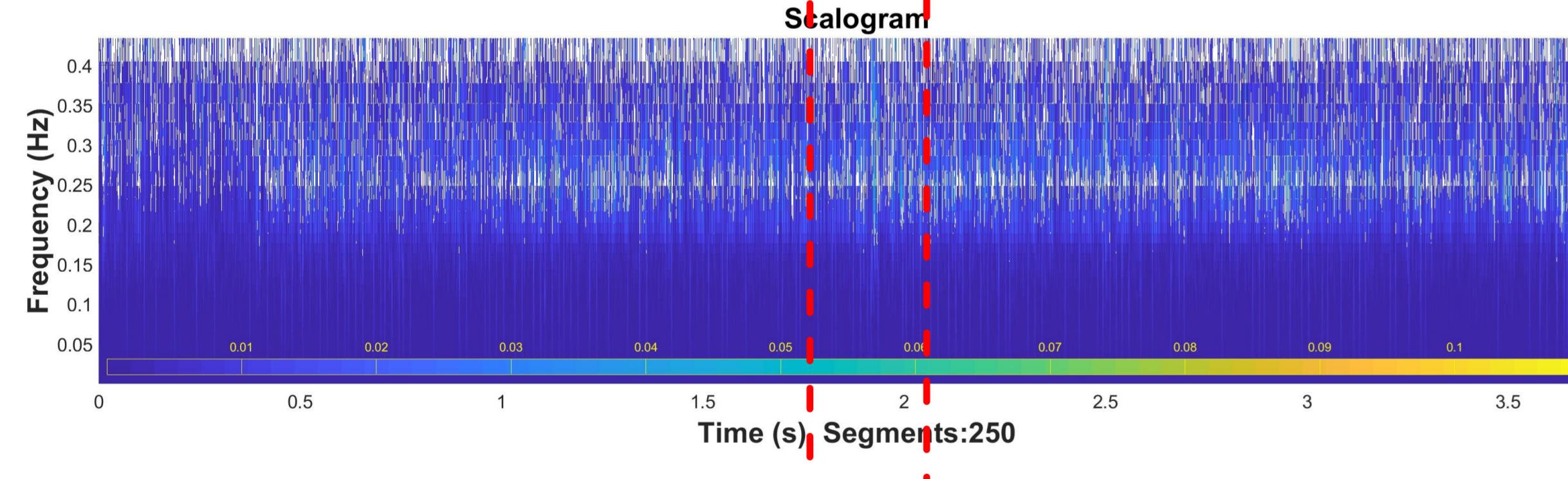
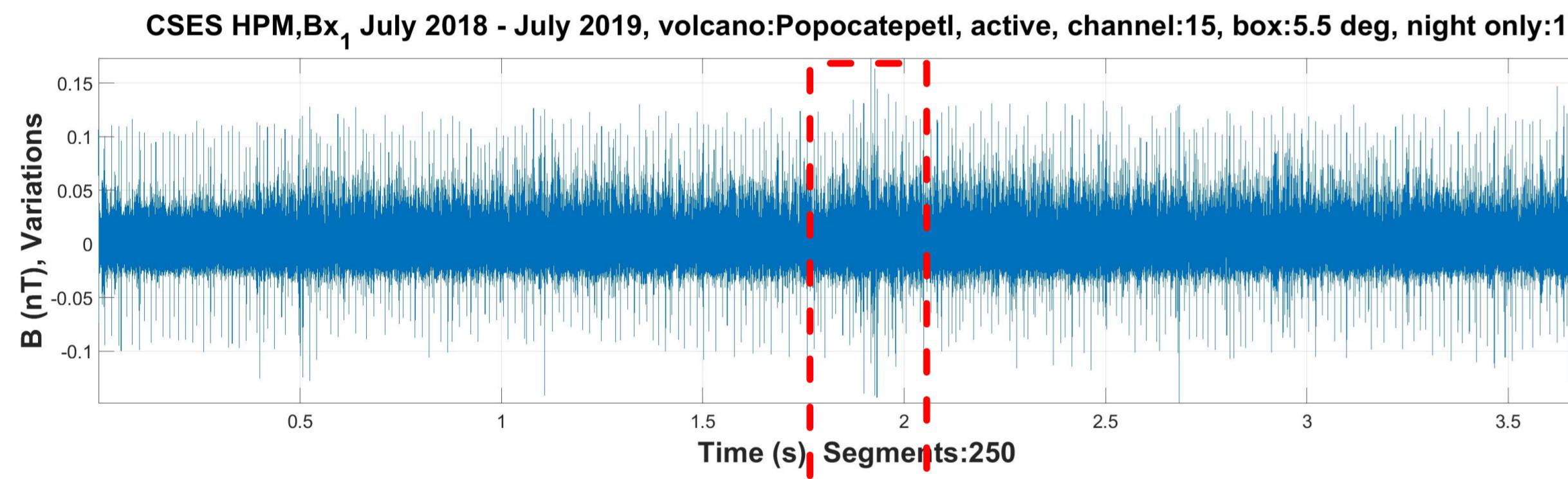
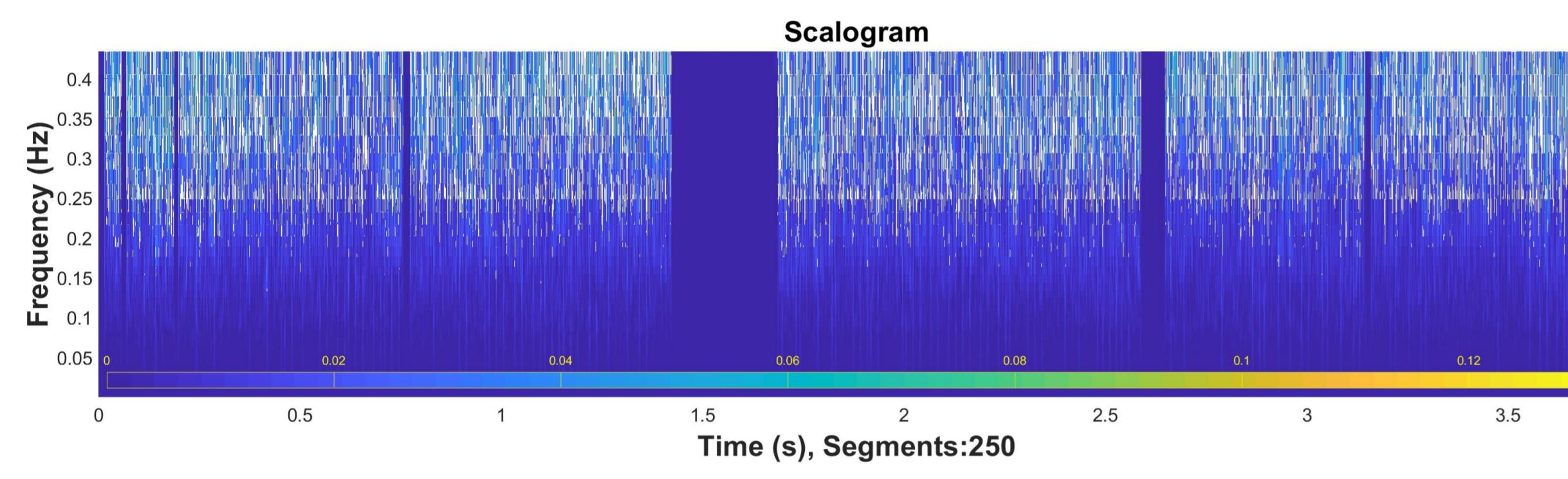
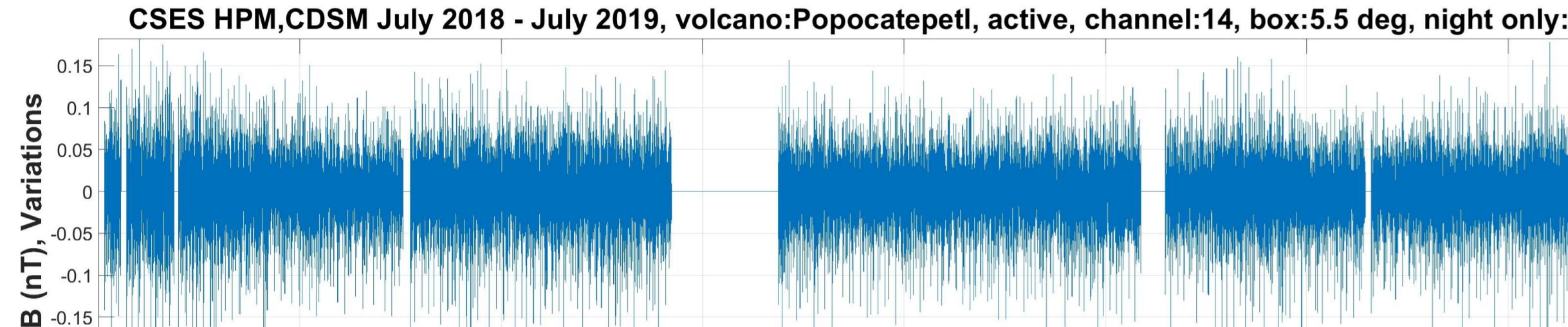
# SATELLITE AND GROUND-BASED MAGNETIC FIELD OBSERVATIONS RELATED TO VOLCANIC ERUPTIONS

K. Schwingenschuh<sup>1</sup>, W. Magnes<sup>1</sup>, X. Shen<sup>2</sup>, J. Wang<sup>3</sup>, B. Cheng<sup>7</sup>, A. Pollinger<sup>1</sup>, Chr. Hagen<sup>1</sup>, R. Lammegger<sup>5</sup>, M. Ellmeier<sup>5</sup>, Christoph Schirniger<sup>1</sup>, H.-U. Eichelberger<sup>1</sup>, B. Mandl<sup>1</sup>, M. Y. Boudjada<sup>1</sup>, B. P. Besser<sup>1</sup>, A. A. Rozhnoi<sup>4</sup>, T. Zhang<sup>1</sup>, M. Delva<sup>1</sup>, I. Jernej<sup>1</sup>, Ö. Aydogar<sup>1</sup>, and R. Leonhardt<sup>6</sup>

<sup>1</sup>Space Research Institute, Austrian Academy of Sciences, Graz, Austria, <sup>2</sup>Institute of Earthquake Admin., Beijing, China, <sup>3</sup>Center for Space Science and Applied Research, CAS, Beijing, China, <sup>4</sup>Institute of the Earth Physics, RAS, Moscow, Russia, <sup>5</sup>Institute of Experimental Physics, Graz Univ. of Technology, Graz, Austria, <sup>6</sup>Central Institution for Meteorology and Geodynamics (ZAMG), Vienna, Austria, <sup>7</sup>National Space Science Center, CAS, Beijing, China

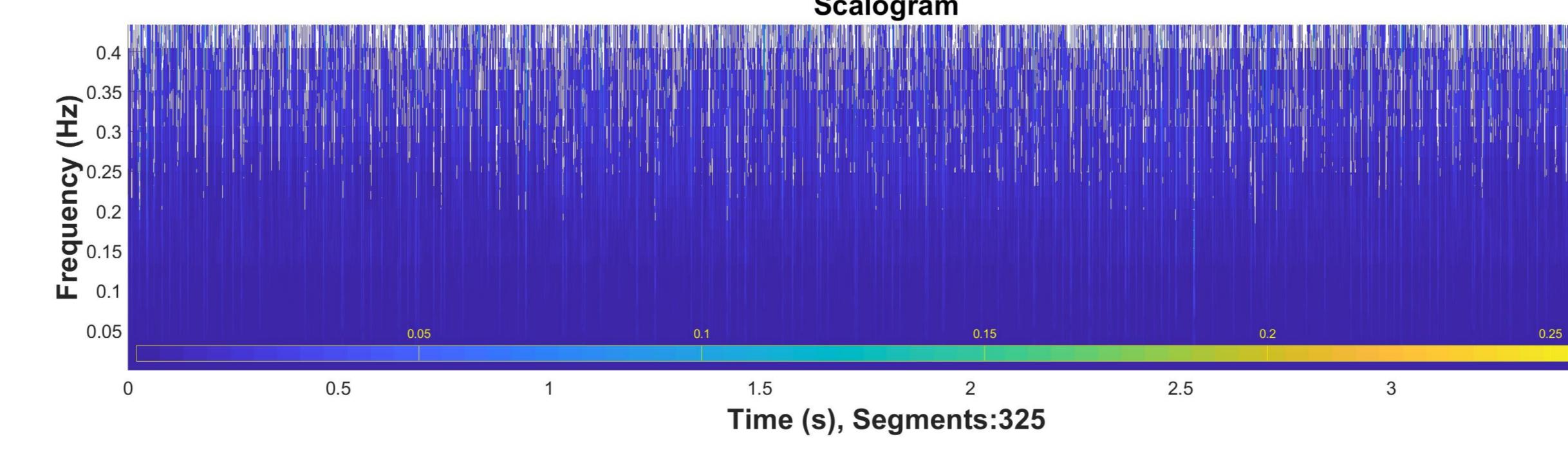
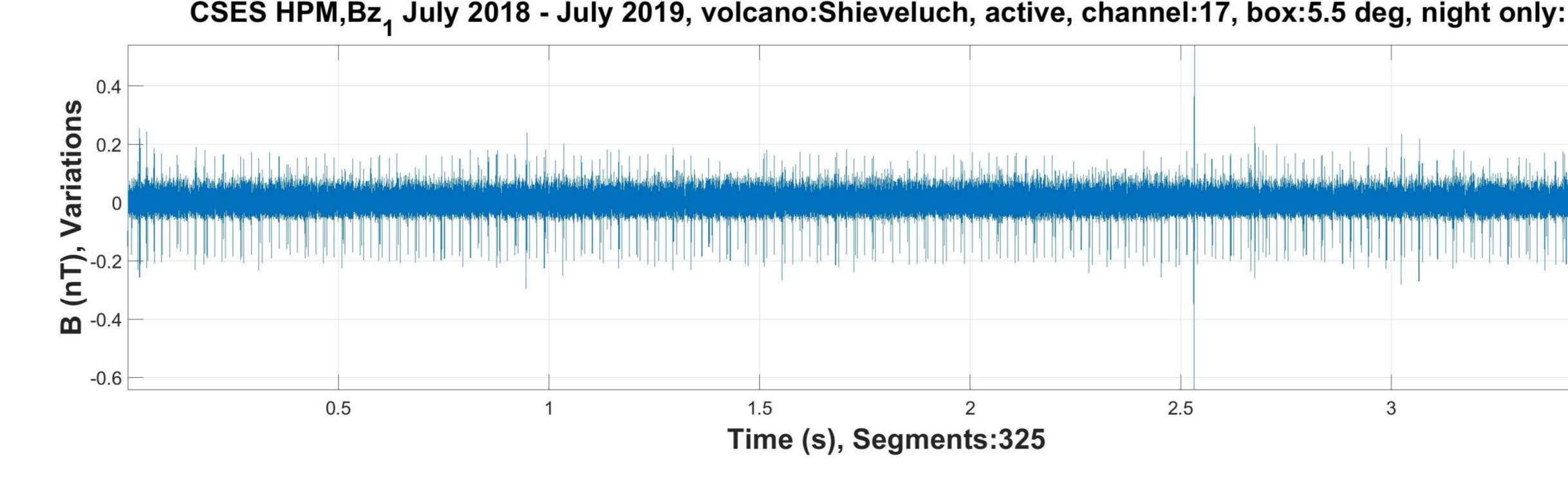
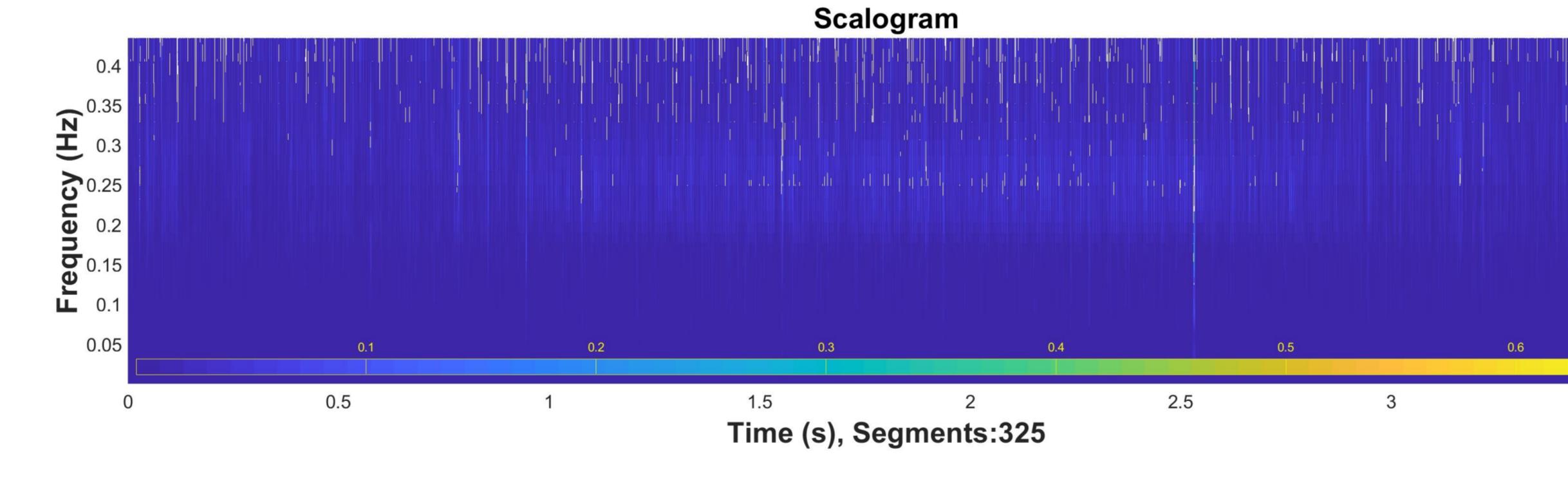
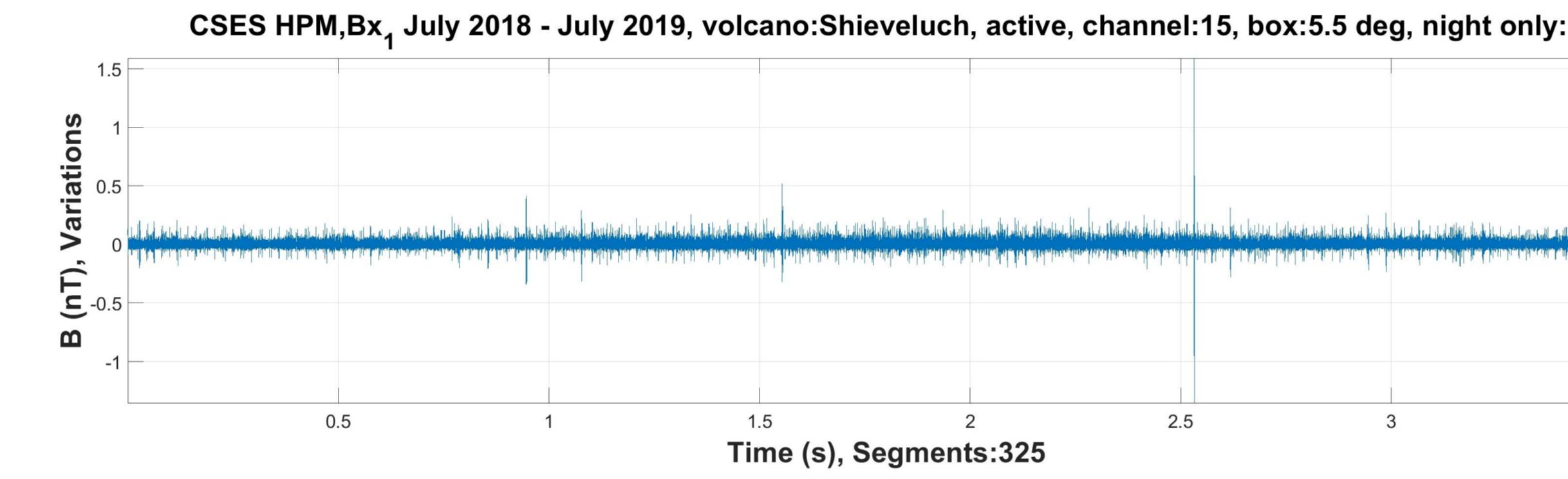
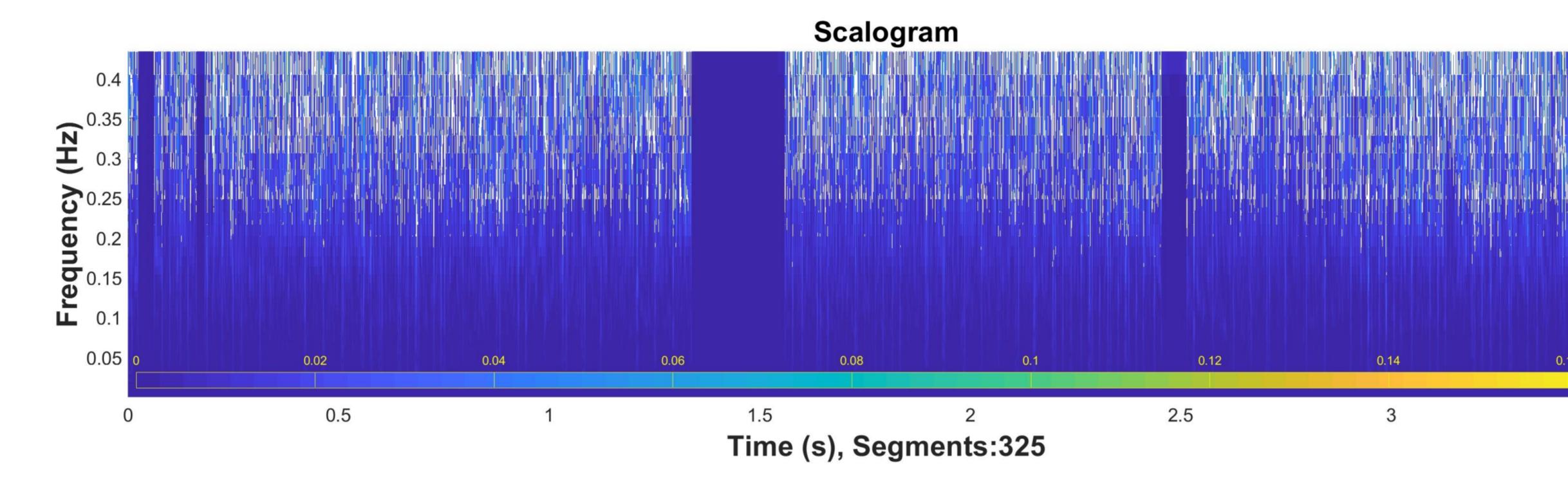
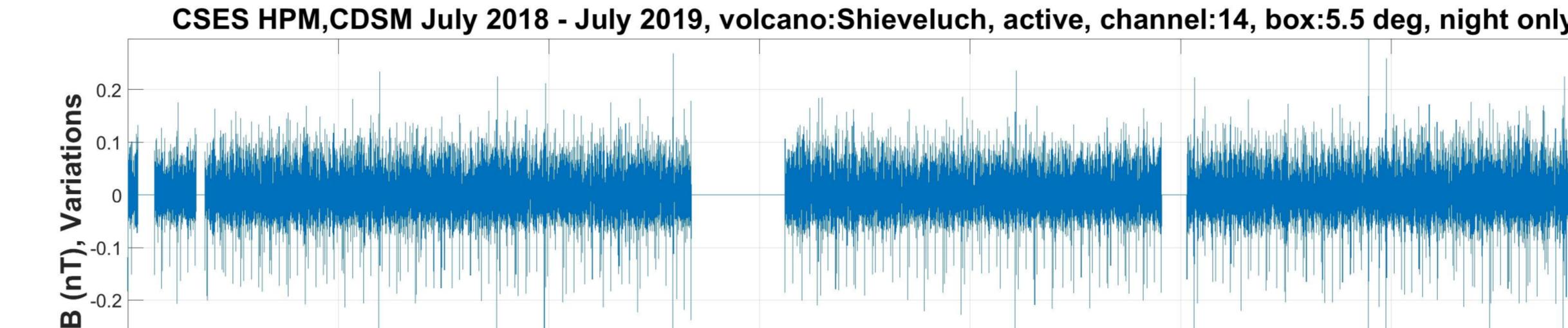
## POPOCATÉPETL

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



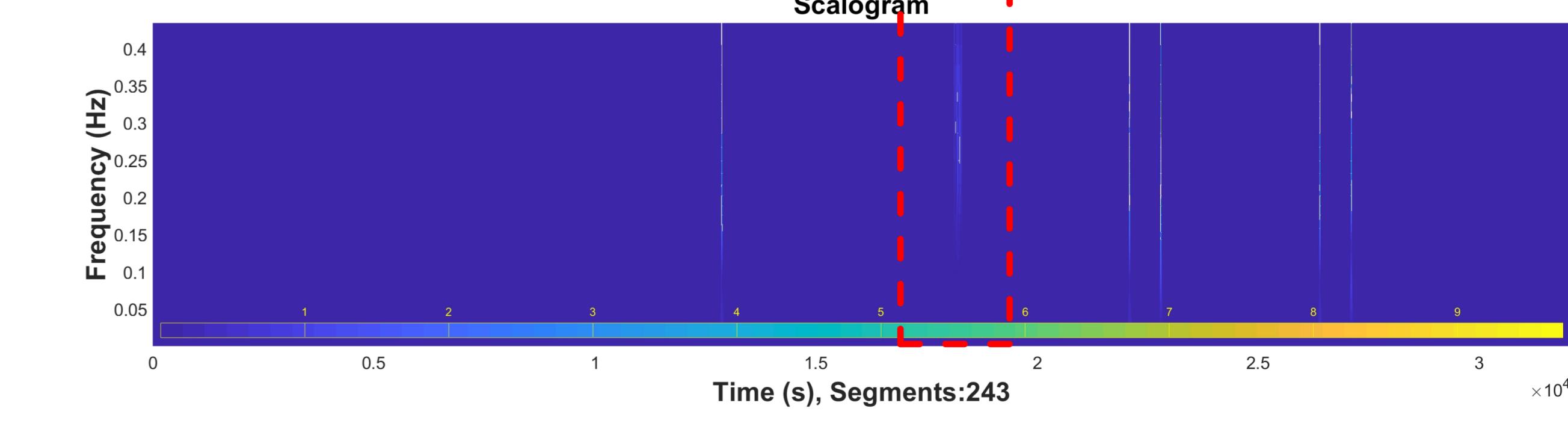
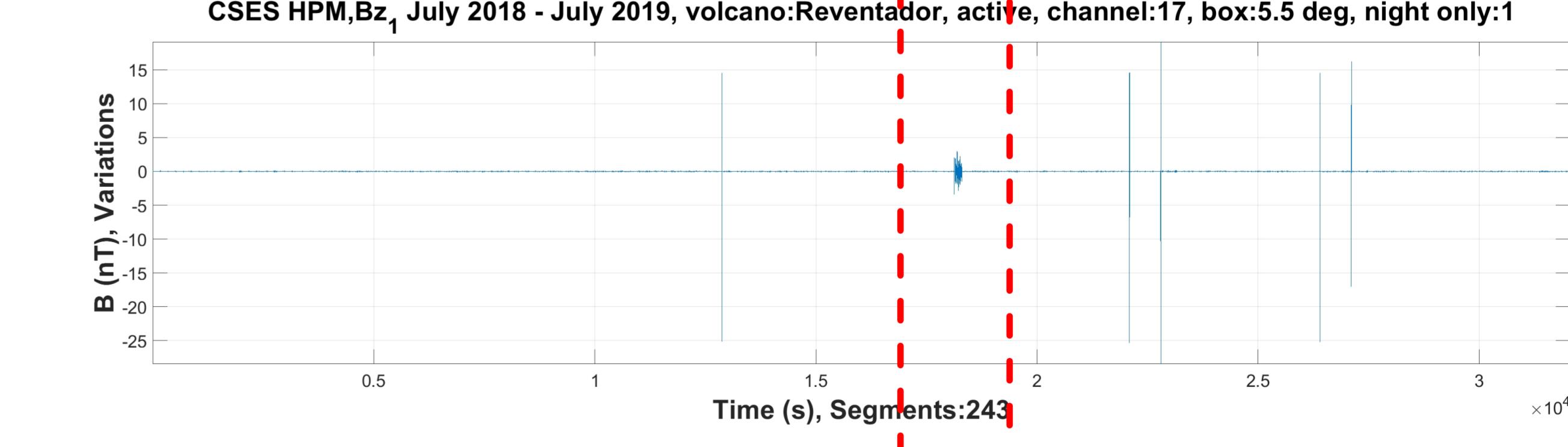
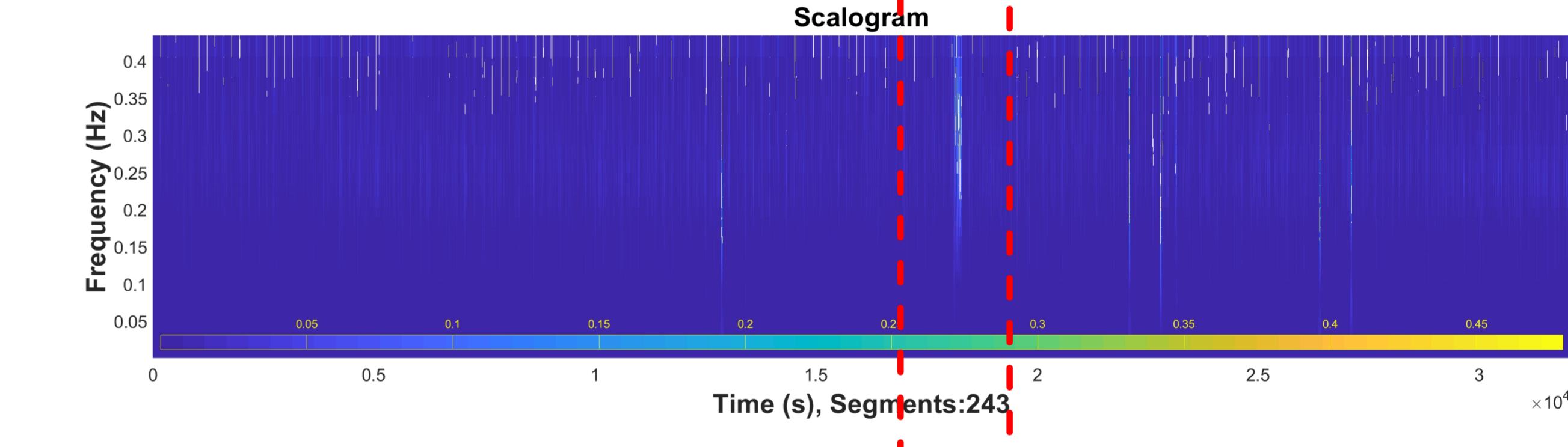
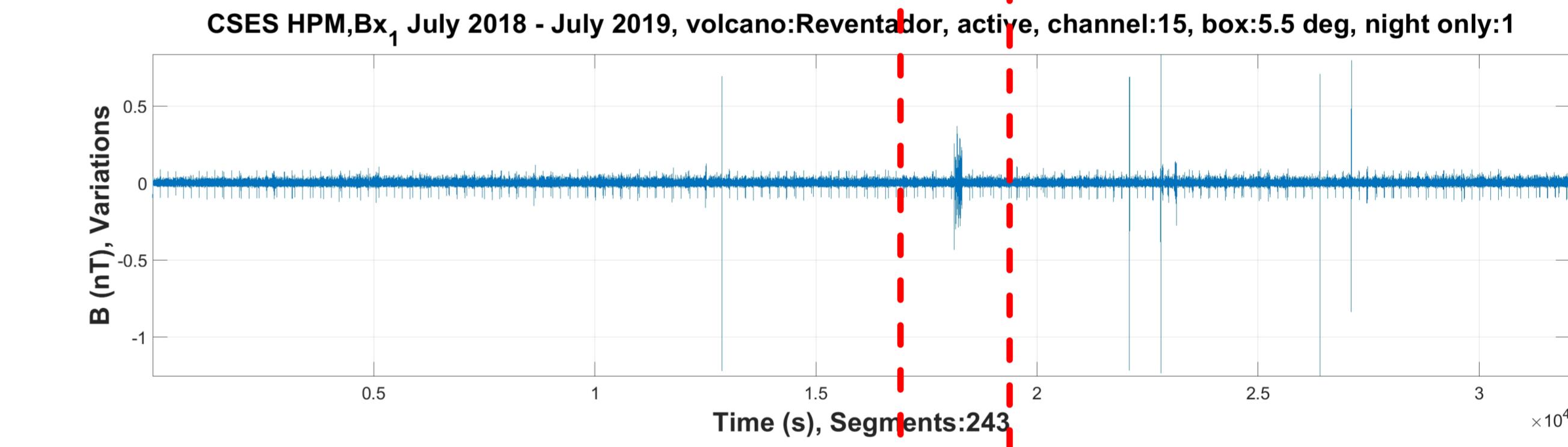
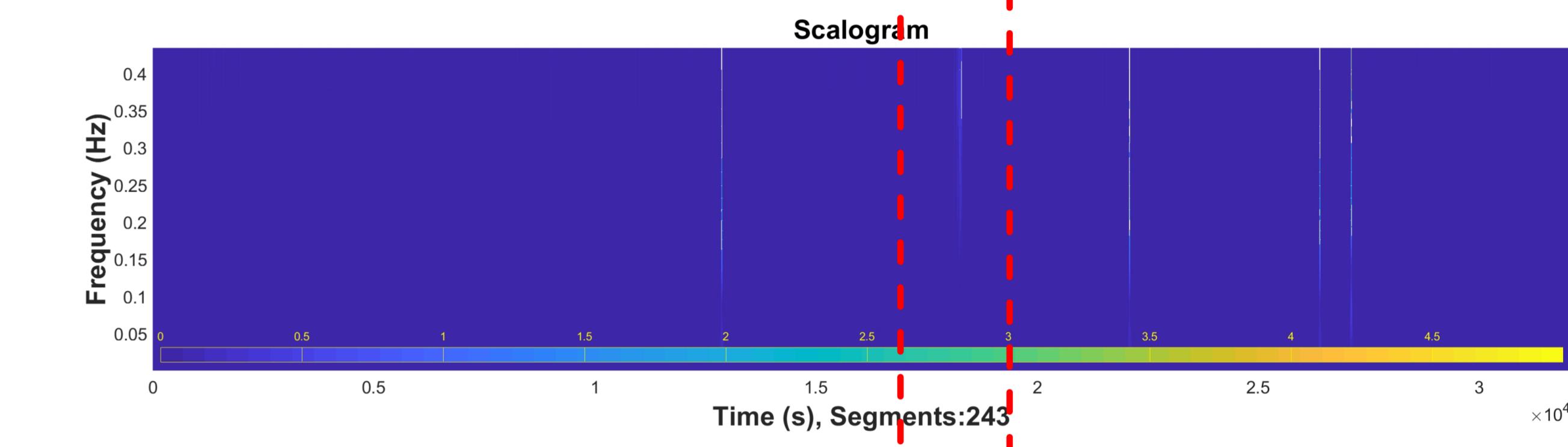
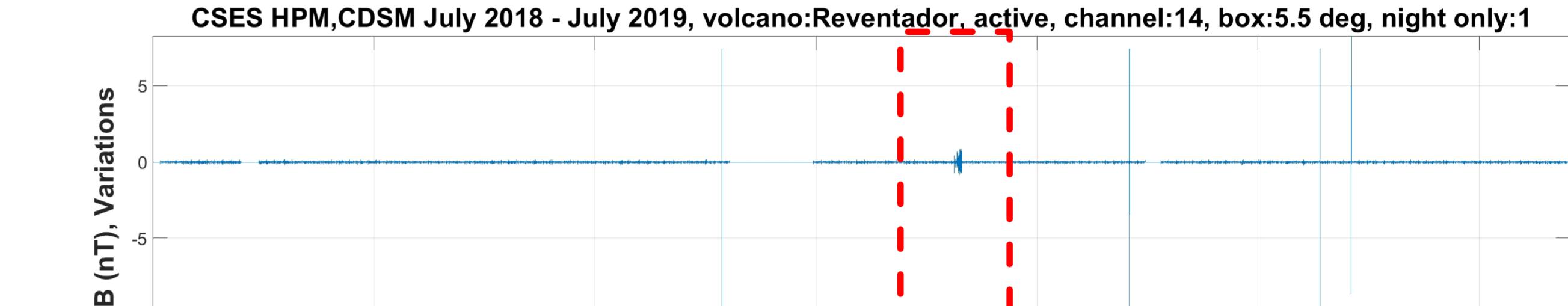
## SHIVELUCH

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



## REVENTADOR

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



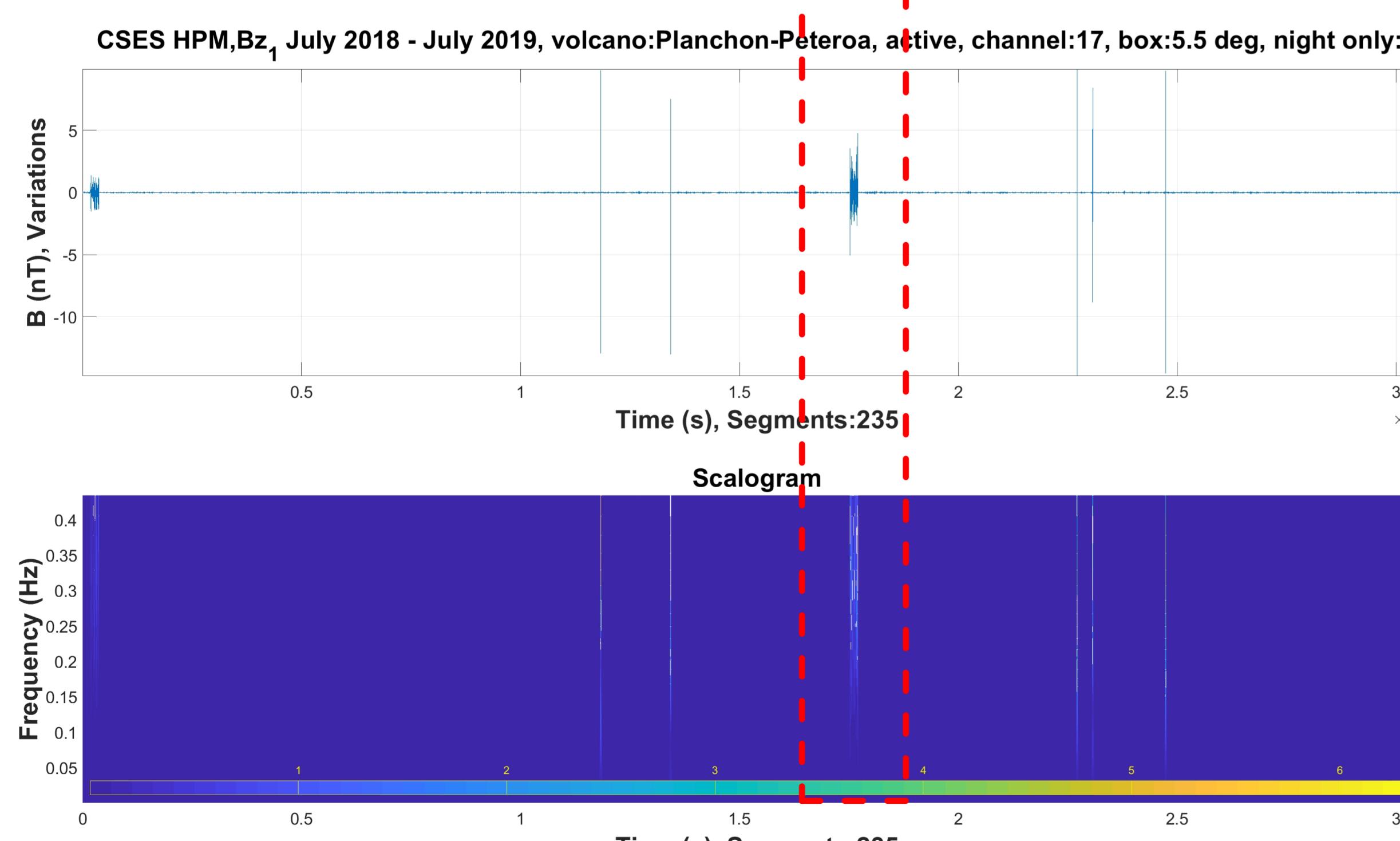
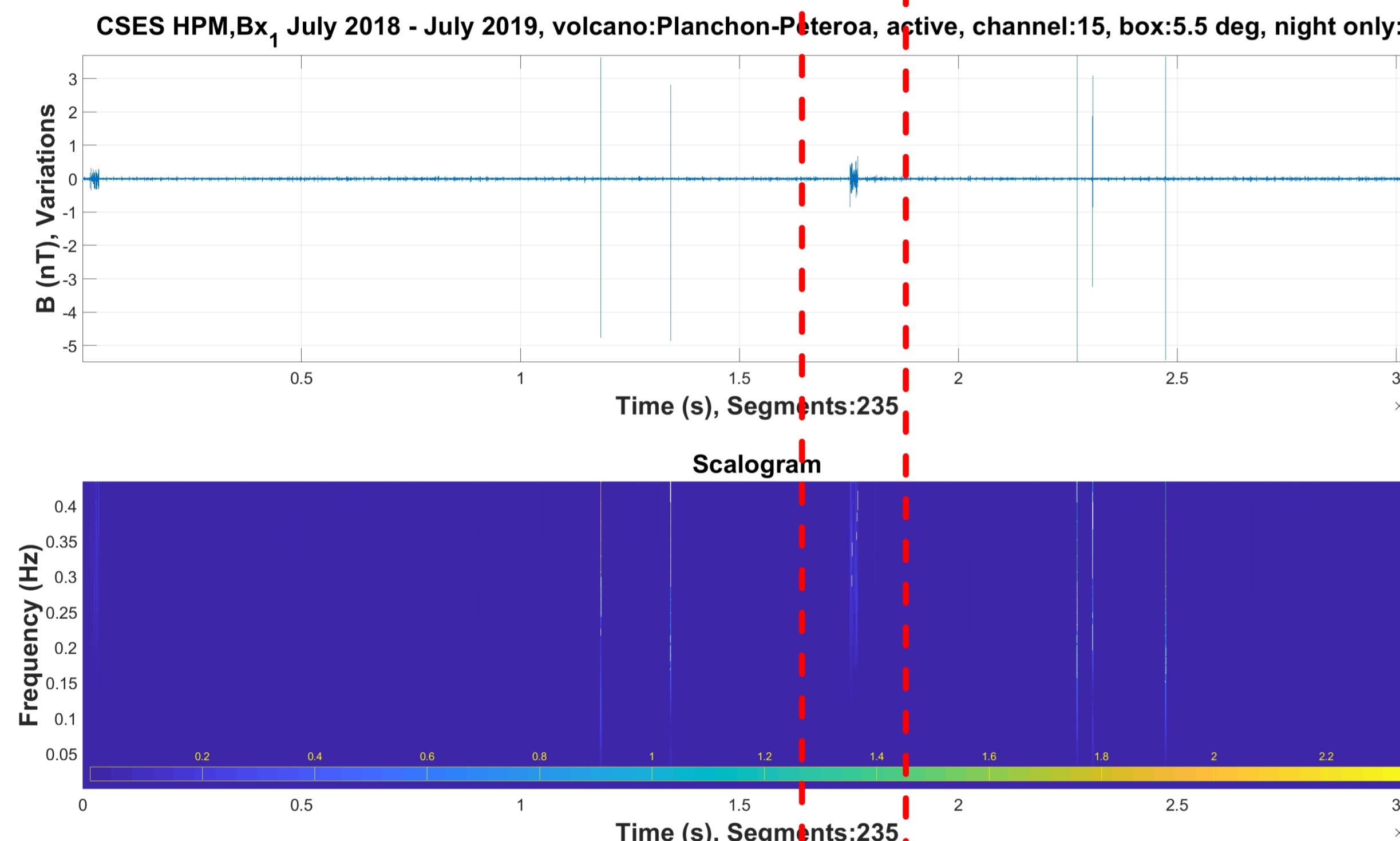
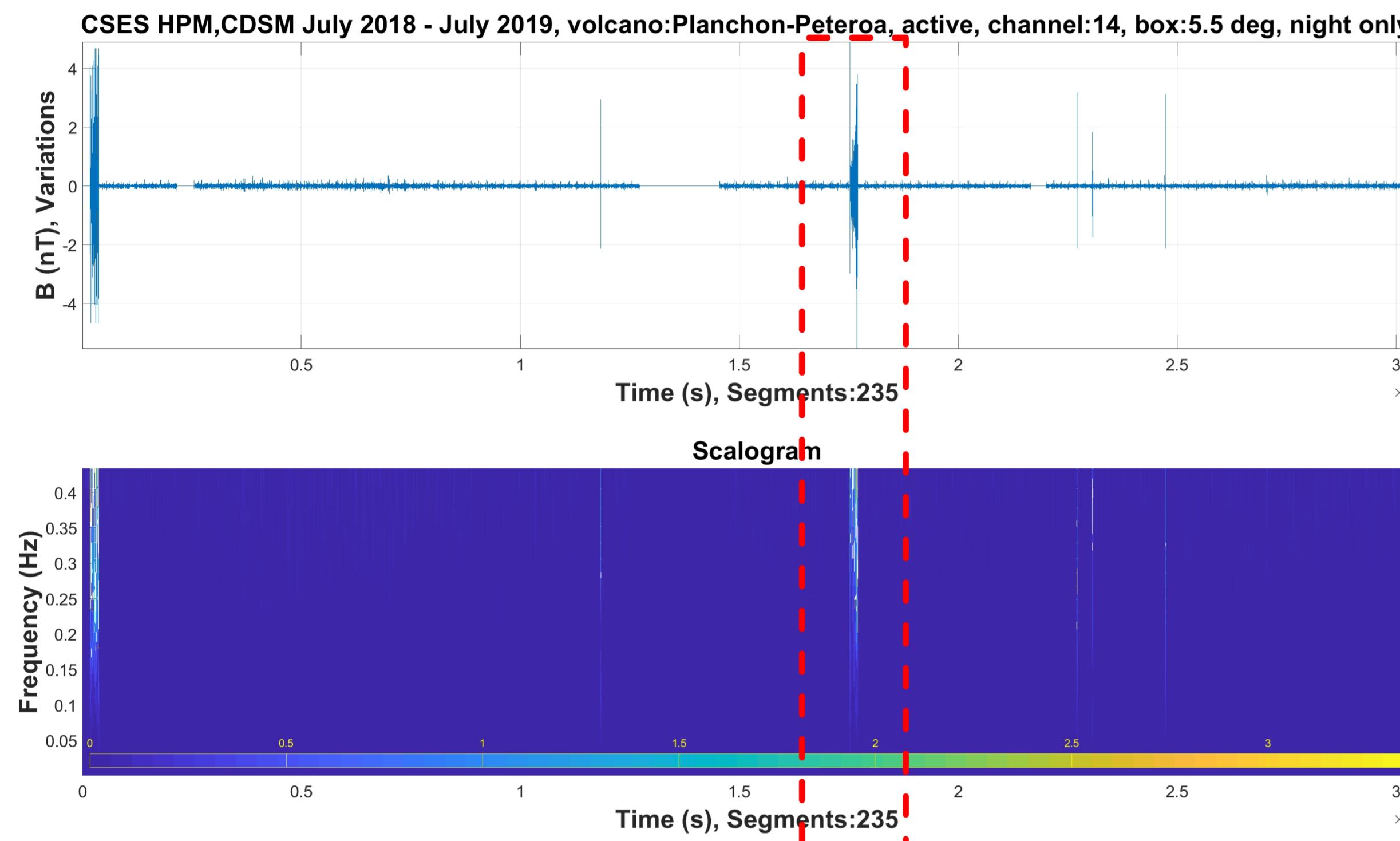
# SATELLITE AND GROUND-BASED MAGNETIC FIELD OBSERVATIONS RELATED TO VOLCANIC ERUPTIONS

K. Schwingenschuh<sup>1</sup>, W. Magnes<sup>1</sup>, X. Shen<sup>2</sup>, J. Wang<sup>3</sup>, B. Cheng<sup>7</sup>, A. Pollinger<sup>1</sup>, Chr. Hagen<sup>1</sup>, R. Lammegeier<sup>5</sup>, M. Ellmeier<sup>5</sup>, Christoph Schirniger<sup>1</sup>, H.-U. Eichelberger<sup>1</sup>, B. Mandl<sup>1</sup>, M. Y. Boudjada<sup>1</sup>, B. P. Besser<sup>1</sup>, A. A. Rozhnoi<sup>4</sup>, T. Zhang<sup>1</sup>, M. Delva<sup>1</sup>, I. Jernej<sup>1</sup>, Ö. Aydogar<sup>1</sup>, and R. Leonhardt<sup>6</sup>

<sup>1</sup>Space Research Institute, Austrian Academy of Sciences, Graz, Austria, <sup>2</sup>Institute of Earthquake Admin., Beijing, China, <sup>3</sup>Center for Space Science and Applied Research, CAS, Beijing, China, <sup>4</sup>Institute of the Earth Physics, RAS, Moscow, Russia, <sup>5</sup>Institute of Experimental Physics, Graz Univ. of Technology, Graz, Austria, <sup>6</sup>Central Institution for Meteorology and Geodynamics (ZAMG), Vienna, Austria, <sup>7</sup>National Space Science Center, CAS, Beijing, China

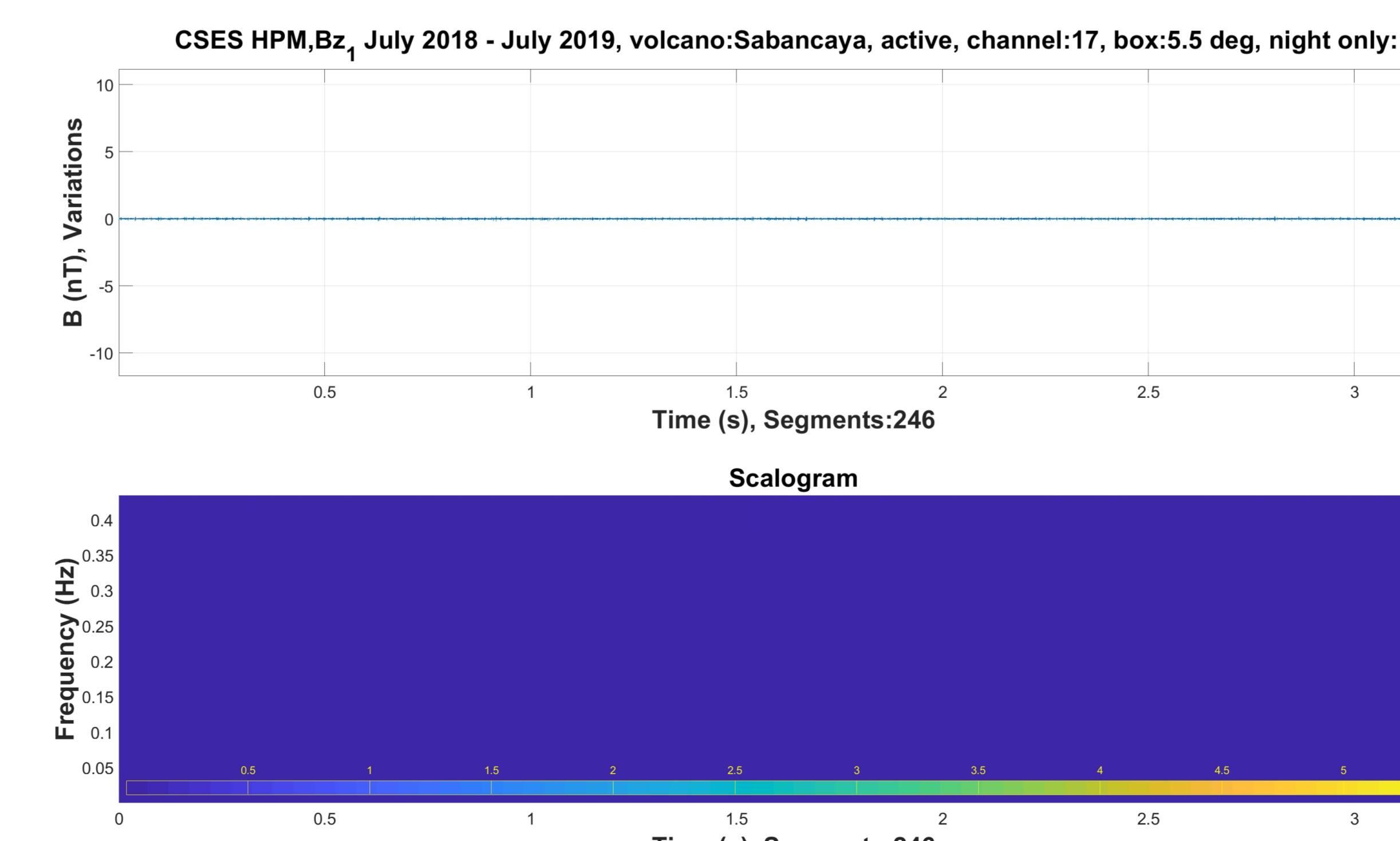
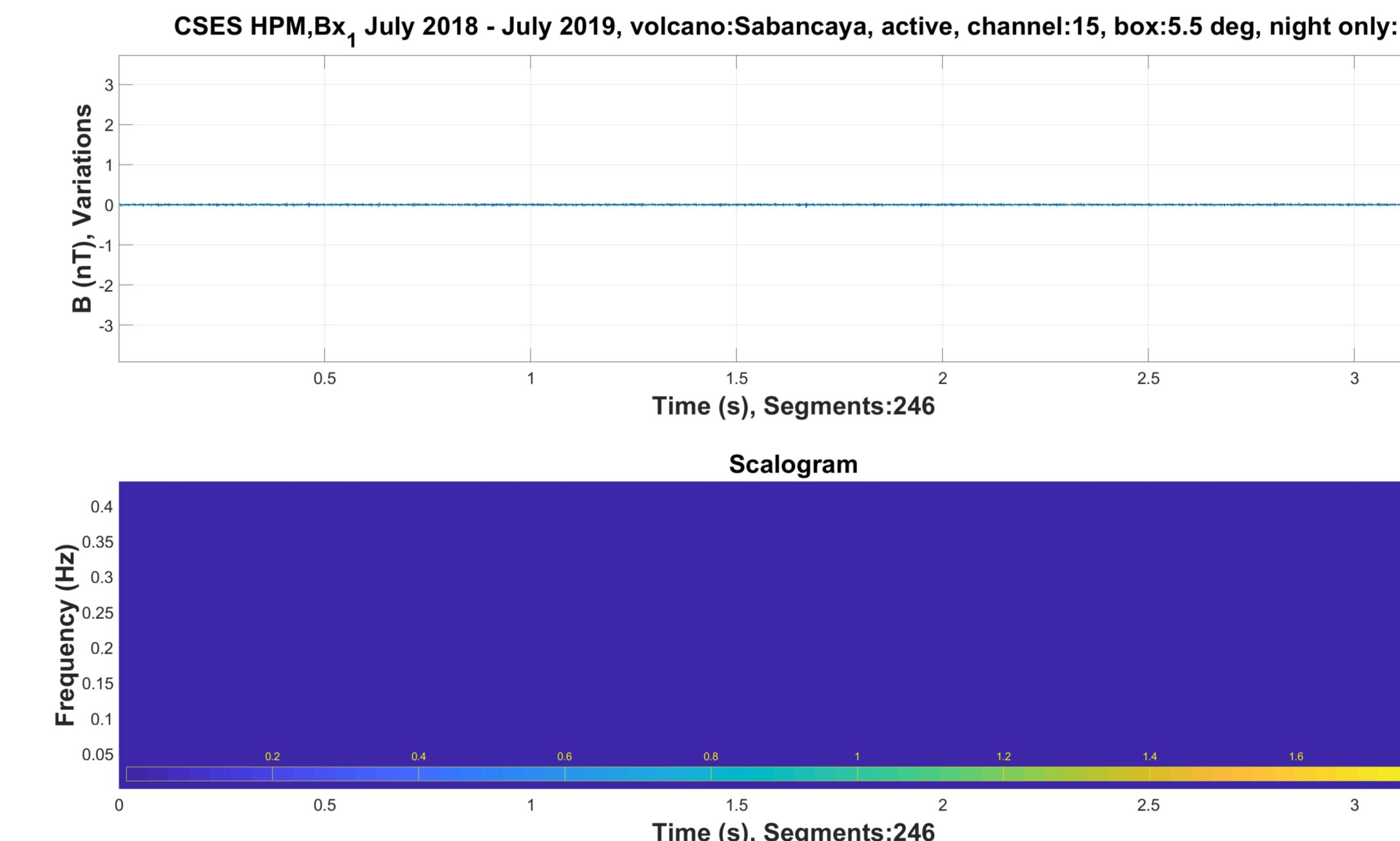
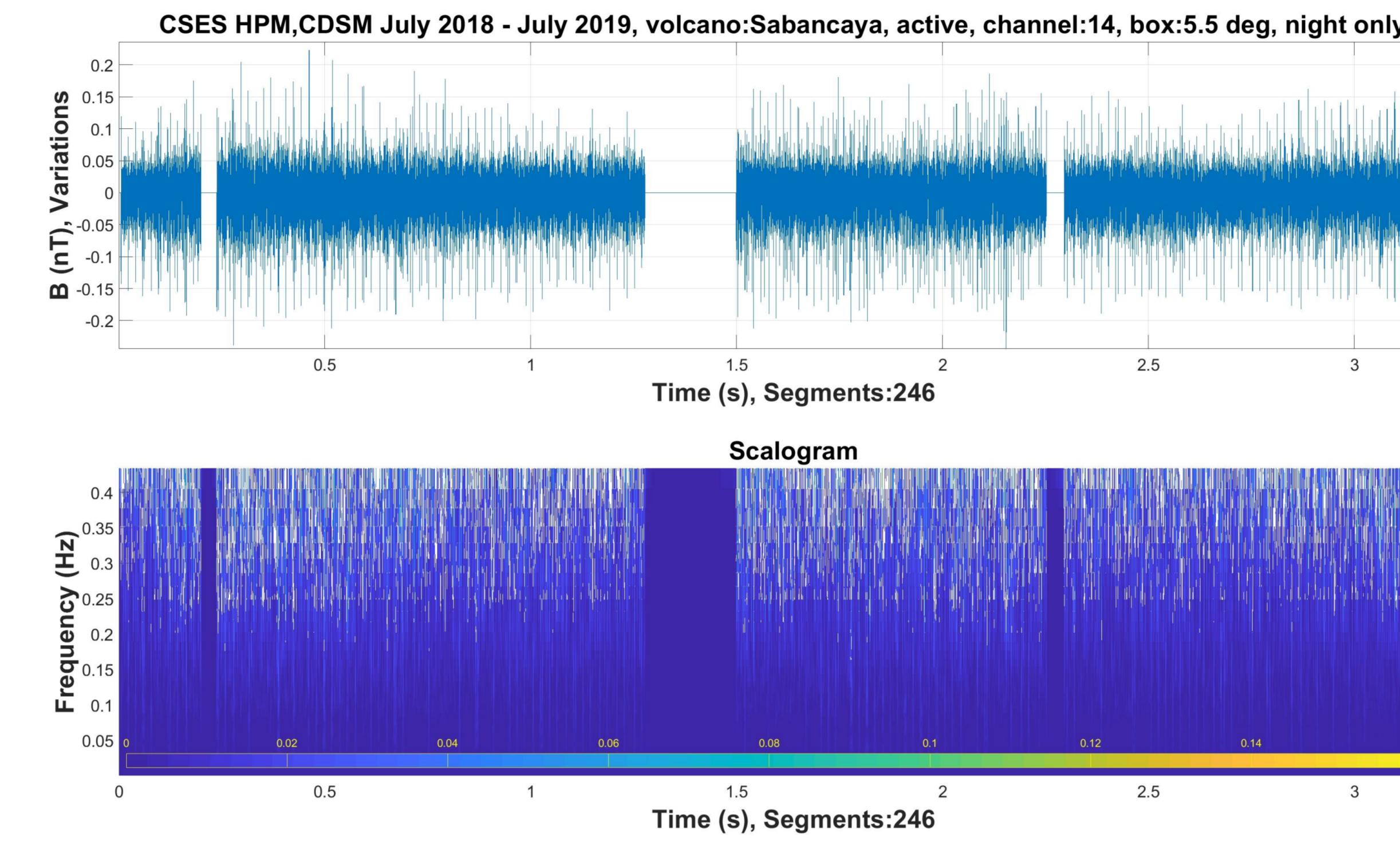
## PLANCHÓN-PETEROA

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



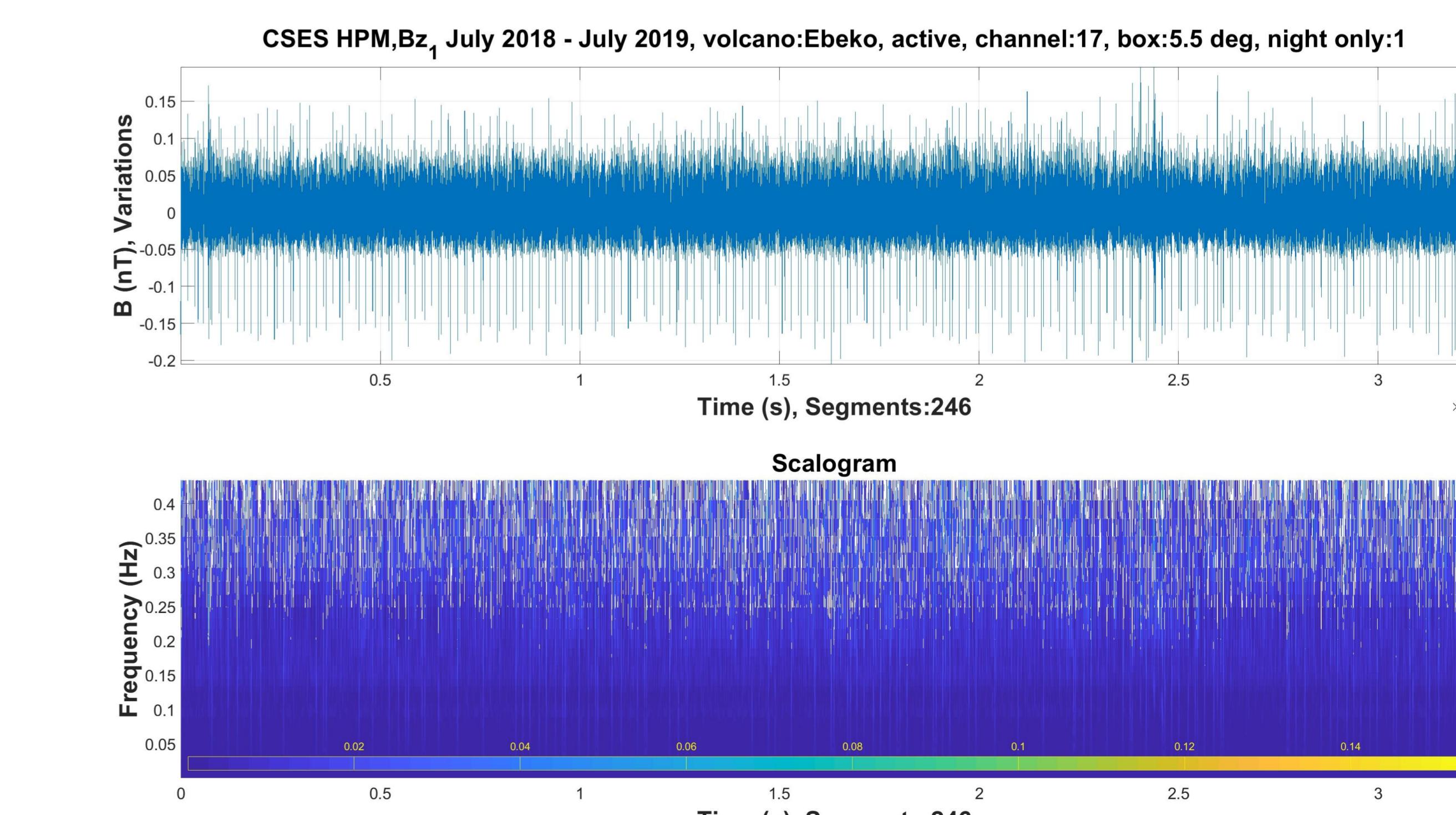
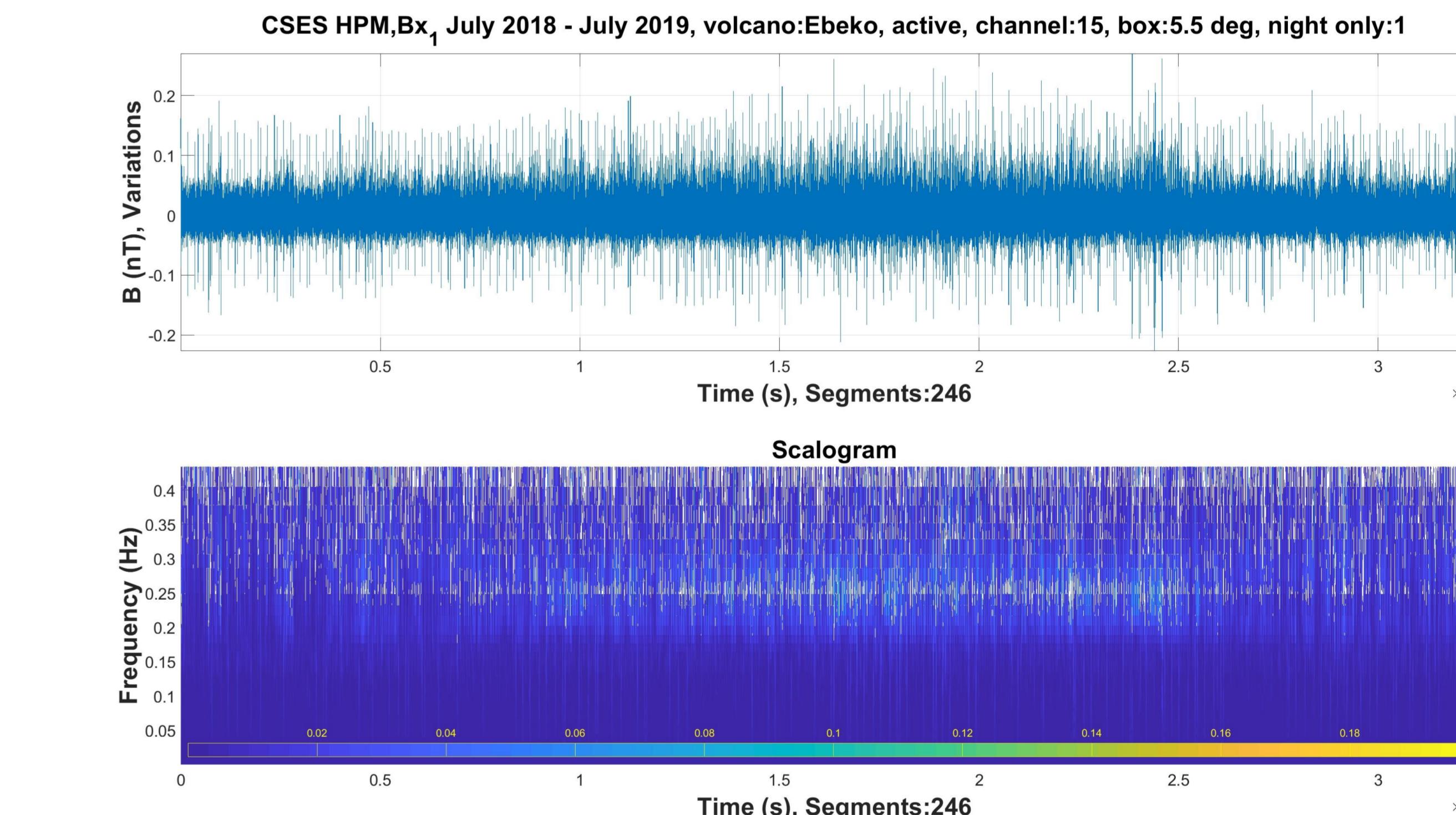
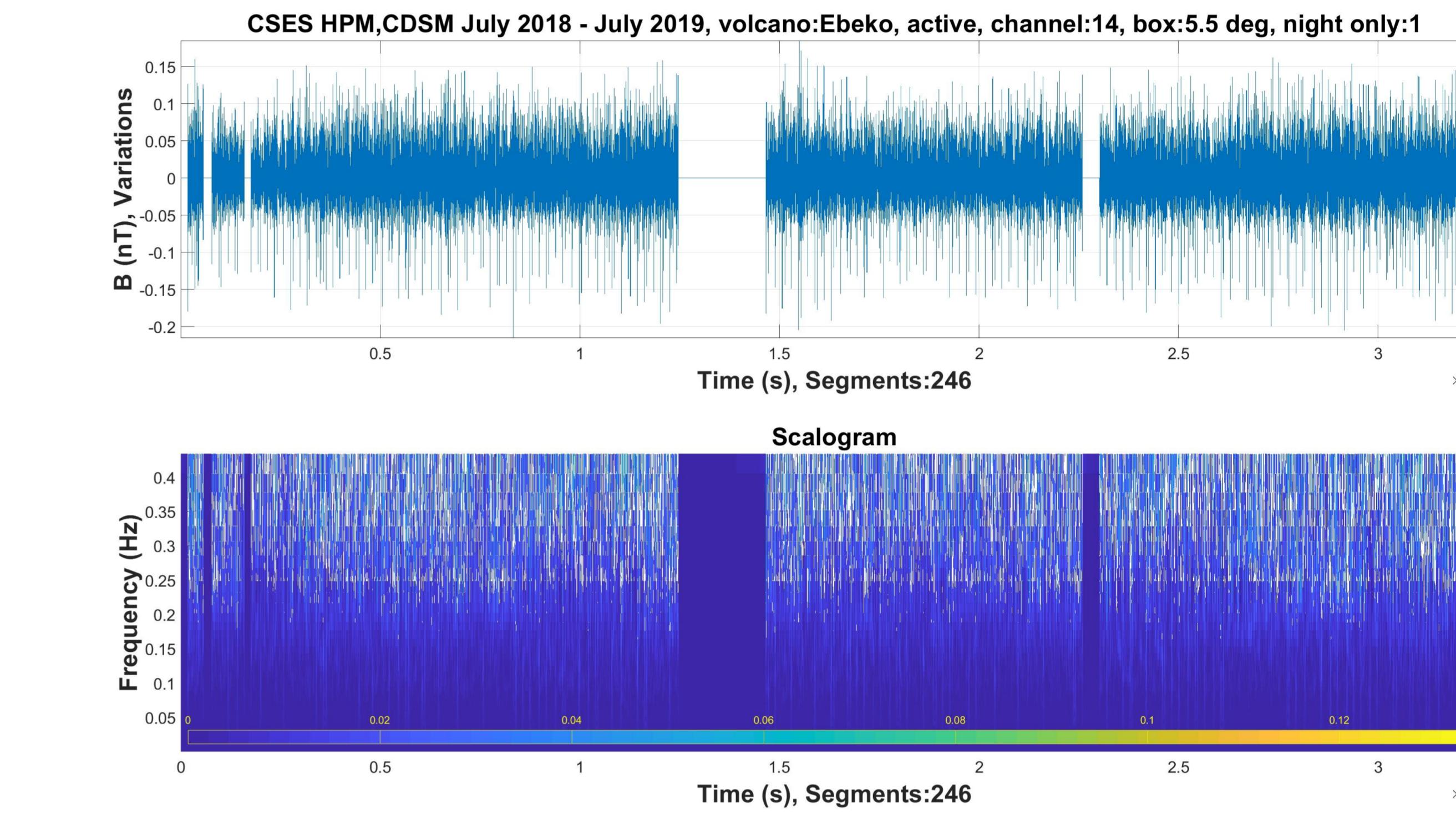
## SABANCAYA

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



## EBEKO

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



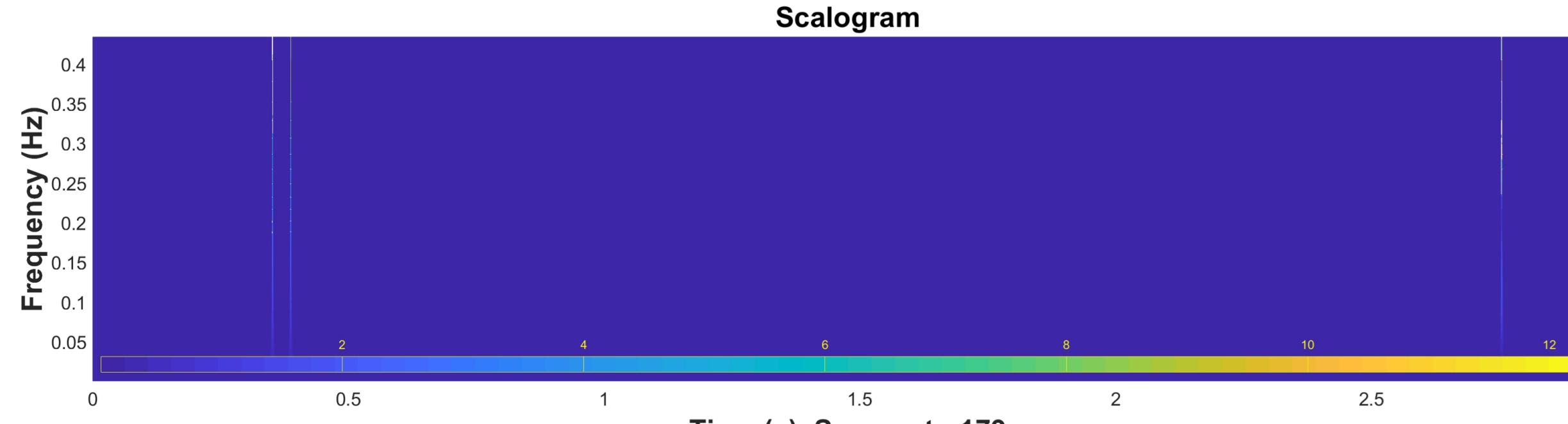
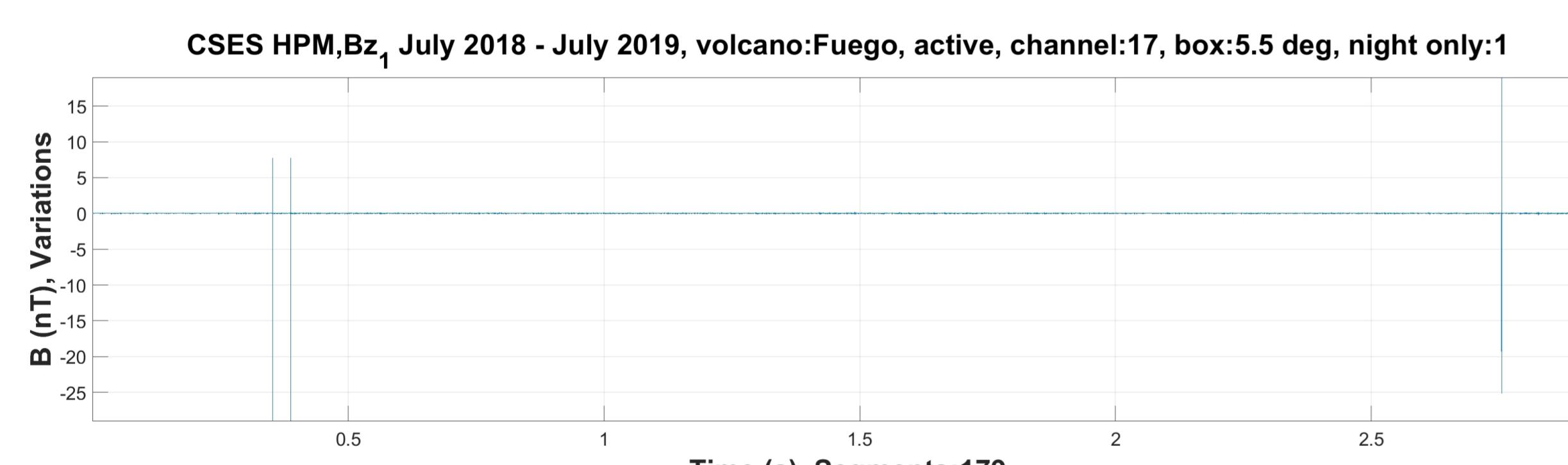
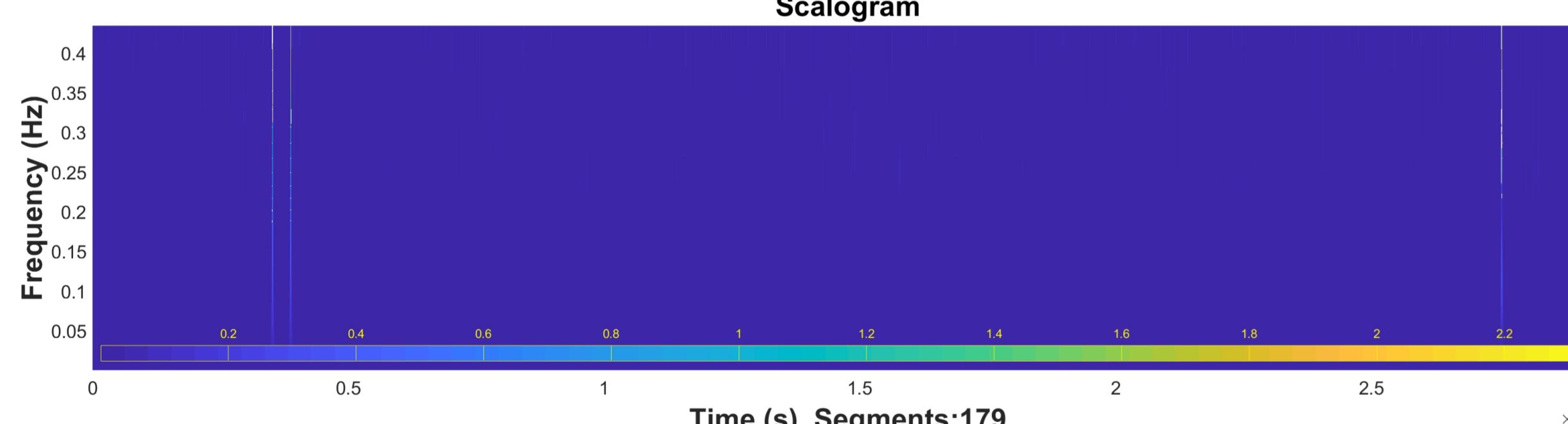
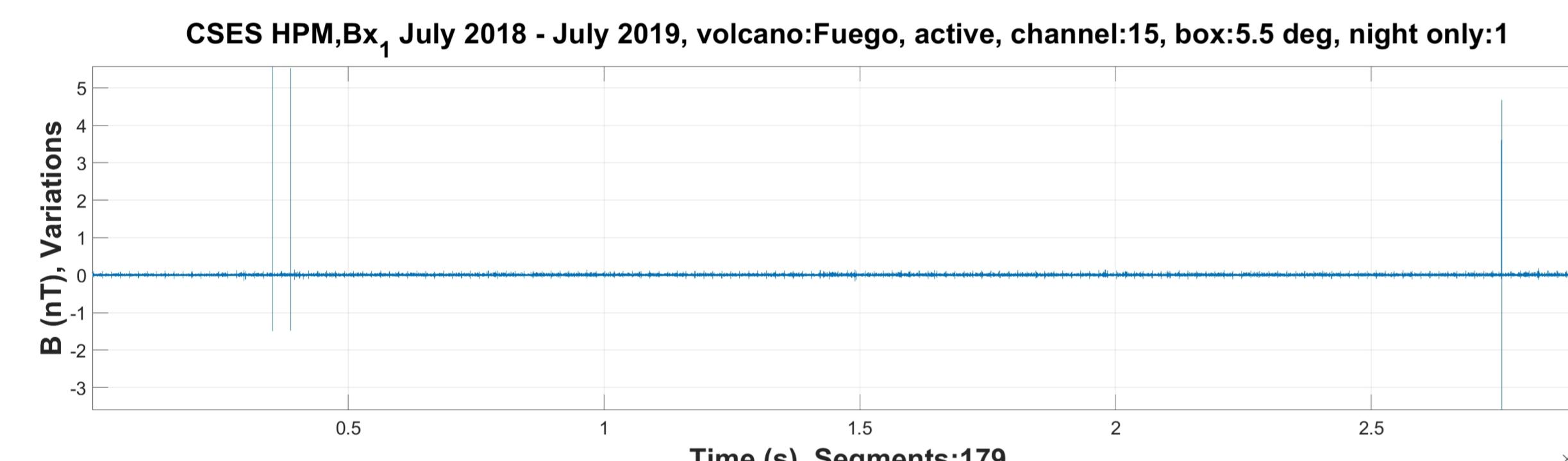
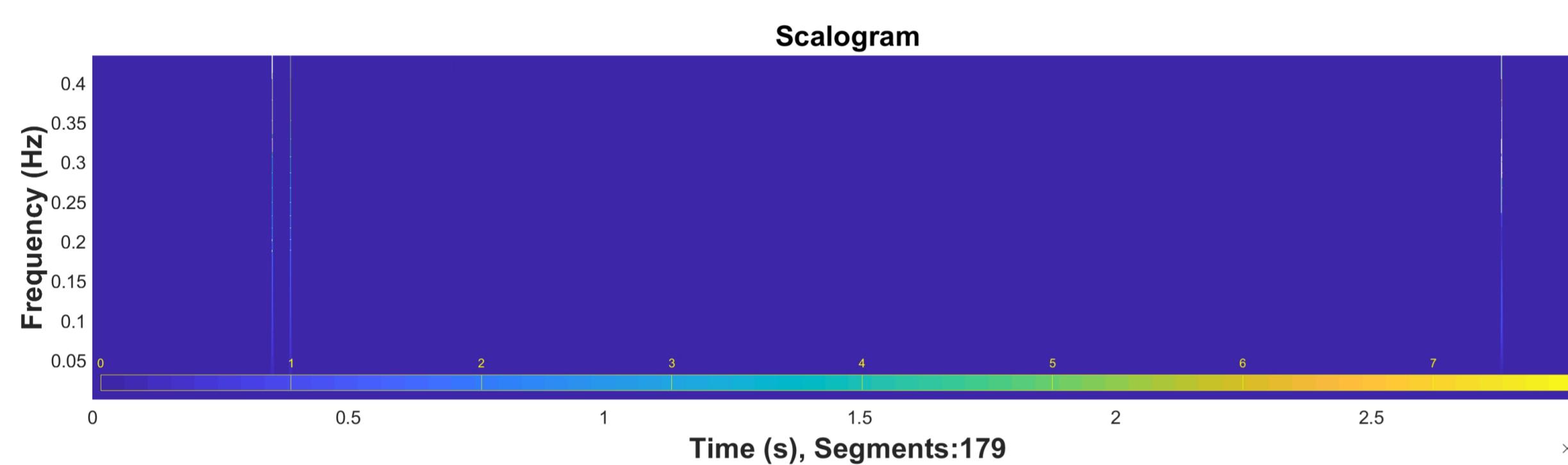
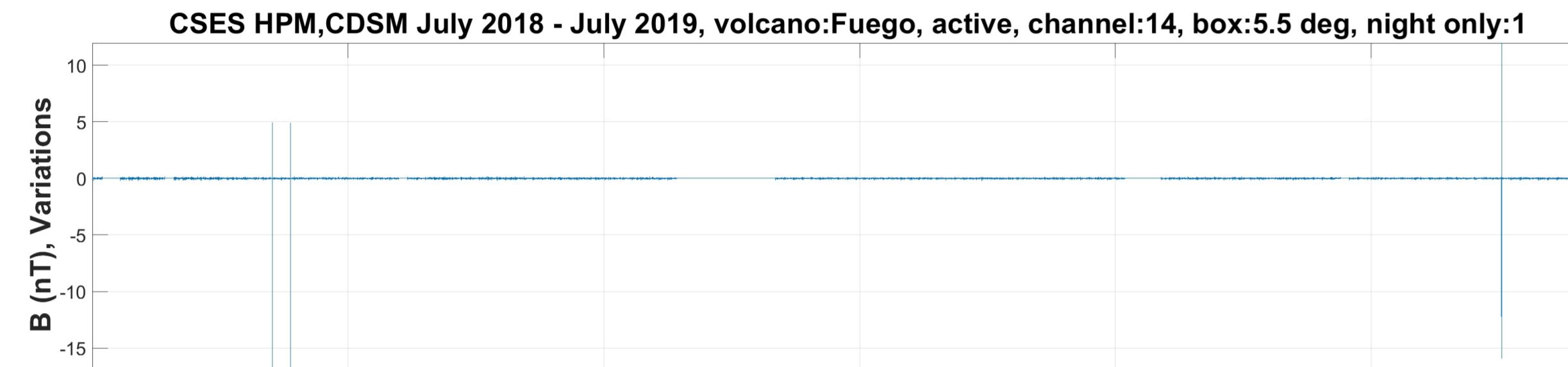
# SATELLITE AND GROUND-BASED MAGNETIC FIELD OBSERVATIONS RELATED TO VOLCANIC ERUPTIONS

K. Schwingenschuh<sup>1</sup>, W. Magnes<sup>1</sup>, X. Shen<sup>2</sup>, J. Wang<sup>3</sup>, B. Cheng<sup>7</sup>, A. Pollinger<sup>1</sup>, Chr. Hagen<sup>1</sup>, R. Lammegeger<sup>5</sup>, M. Ellmeier<sup>5</sup>, Christoph Schirniger<sup>1</sup>, H.-U. Eichelberger<sup>1</sup>, B. Mandl<sup>1</sup>, M. Y. Boudjada<sup>1</sup>, B. P. Besser<sup>1</sup>, A. A. Rozhnoi<sup>4</sup>, T. Zhang<sup>1</sup>, M. Delva<sup>1</sup>, I. Jernej<sup>1</sup>, Ö. Aydogar<sup>1</sup>, and R. Leonhardt<sup>6</sup>

<sup>1</sup>Space Research Institute, Austrian Academy of Sciences, Graz, Austria, <sup>2</sup>Institute of Earthquake Admin., Beijing, China, <sup>3</sup>Center for Space Science and Applied Research, CAS, Beijing, China, <sup>4</sup>Institute of the Earth Physics, RAS, Moscow, Russia, <sup>5</sup>Institute of Experimental Physics, Graz Univ. of Technology, Graz, Austria, <sup>6</sup>Central Institution for Meteorology and Geodynamics (ZAMG), Vienna, Austria, <sup>7</sup>National Space Science Center, CAS, Beijing, China

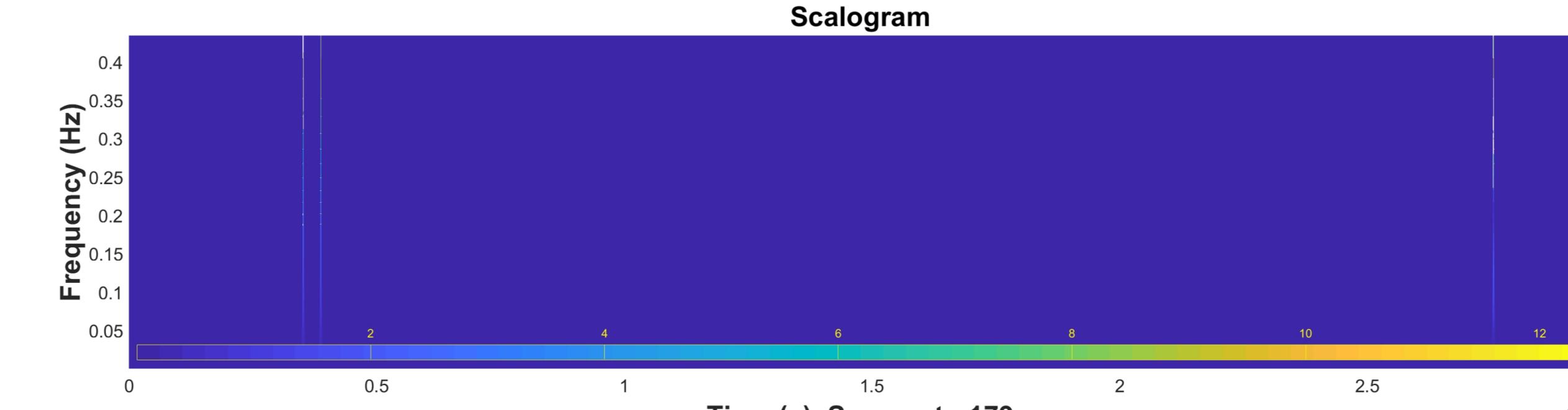
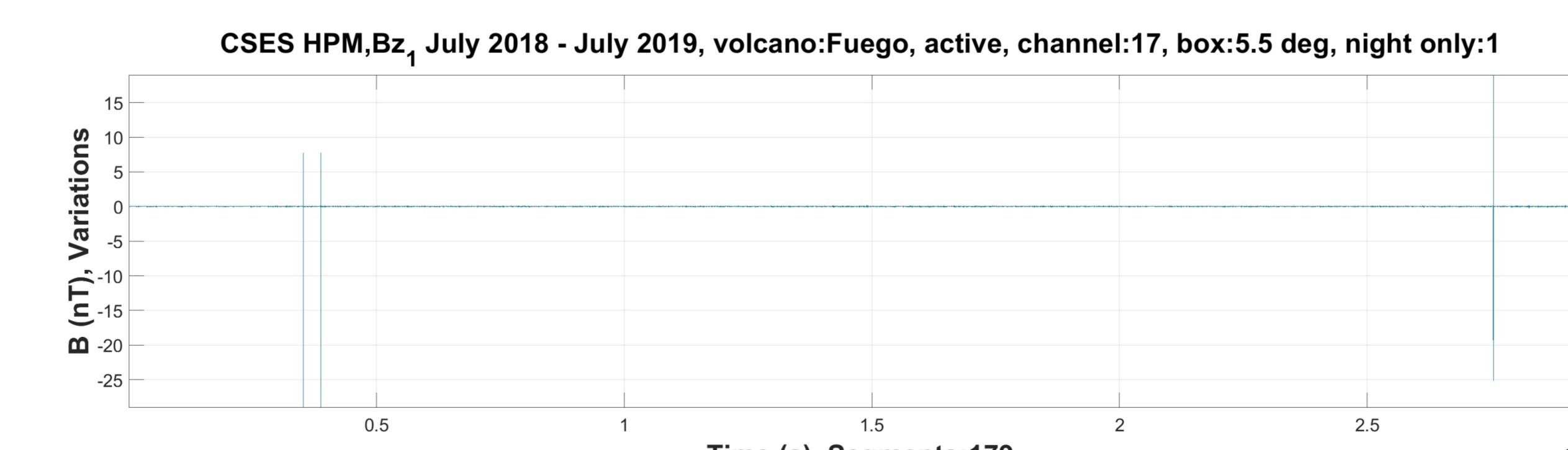
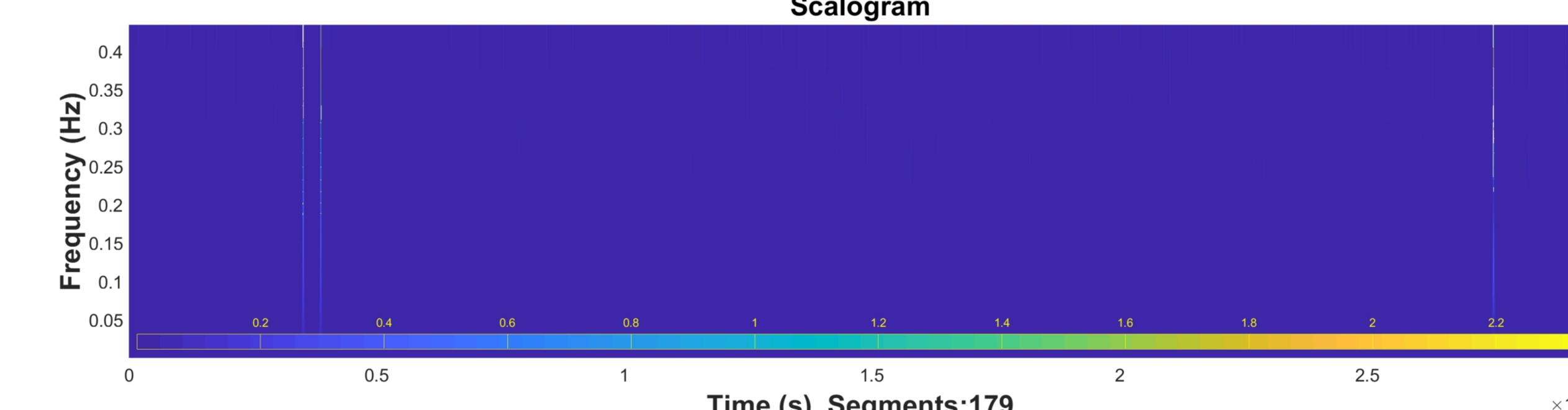
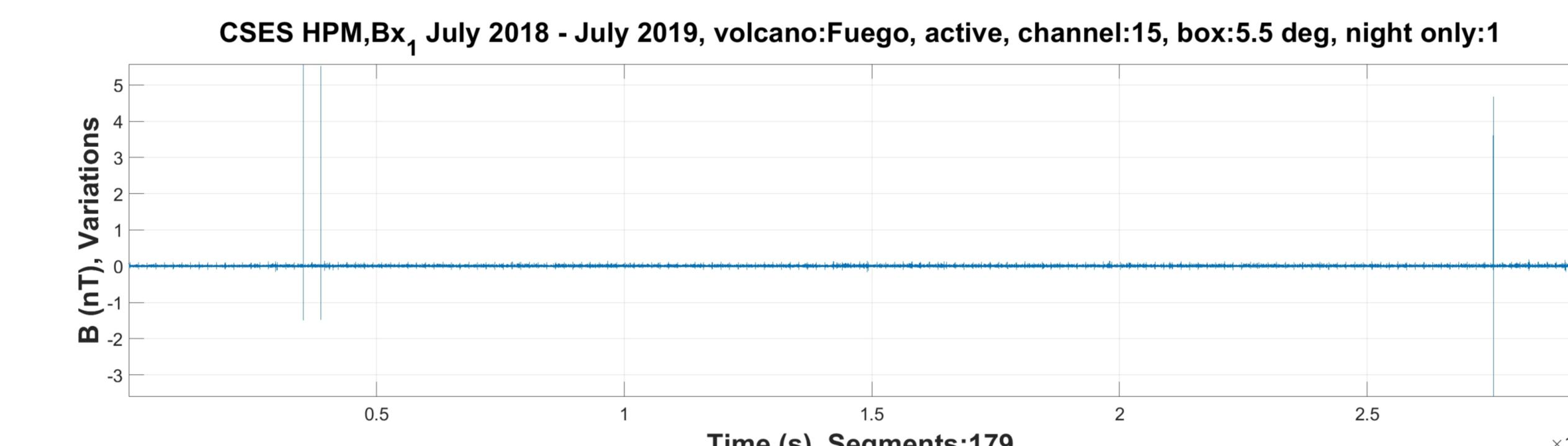
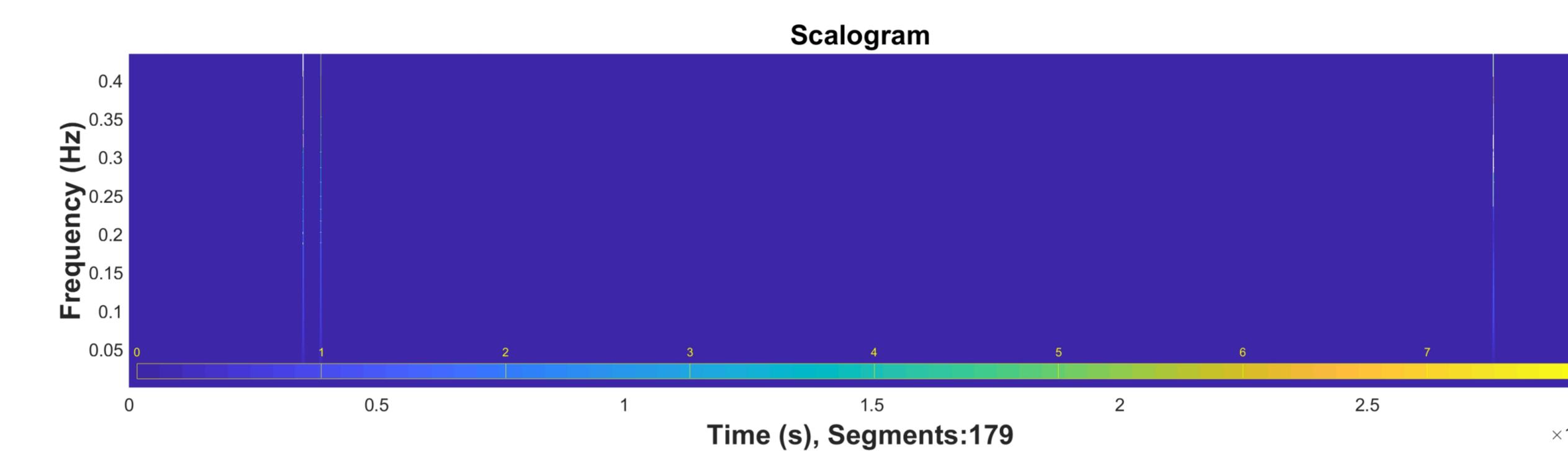
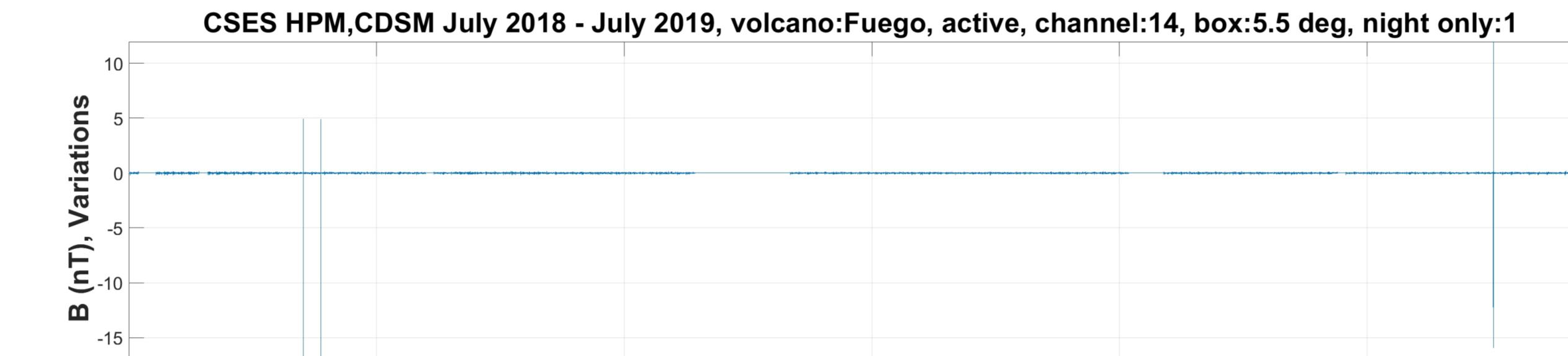
## MANAM

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



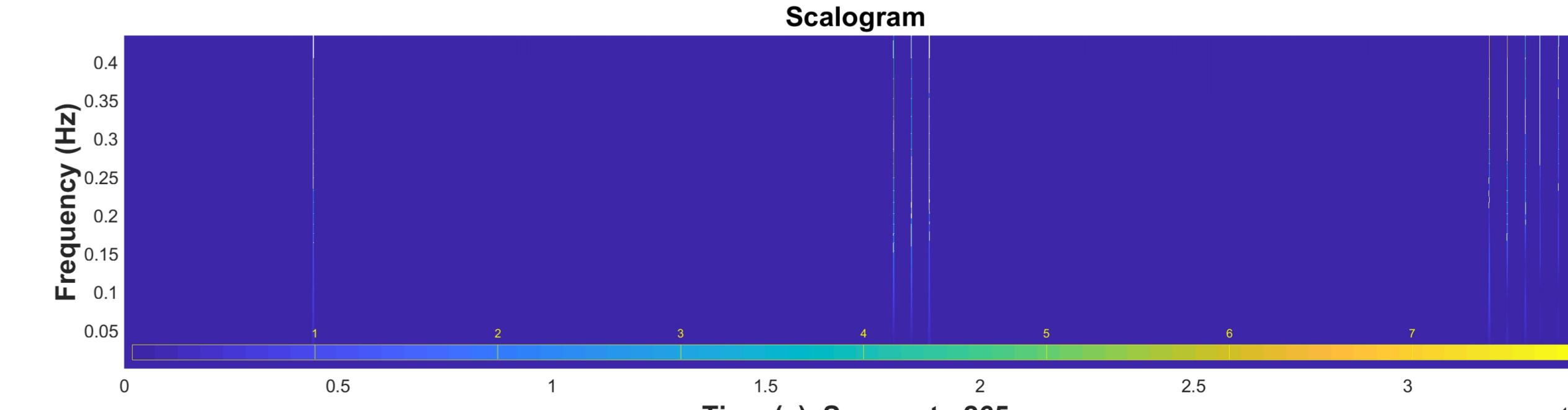
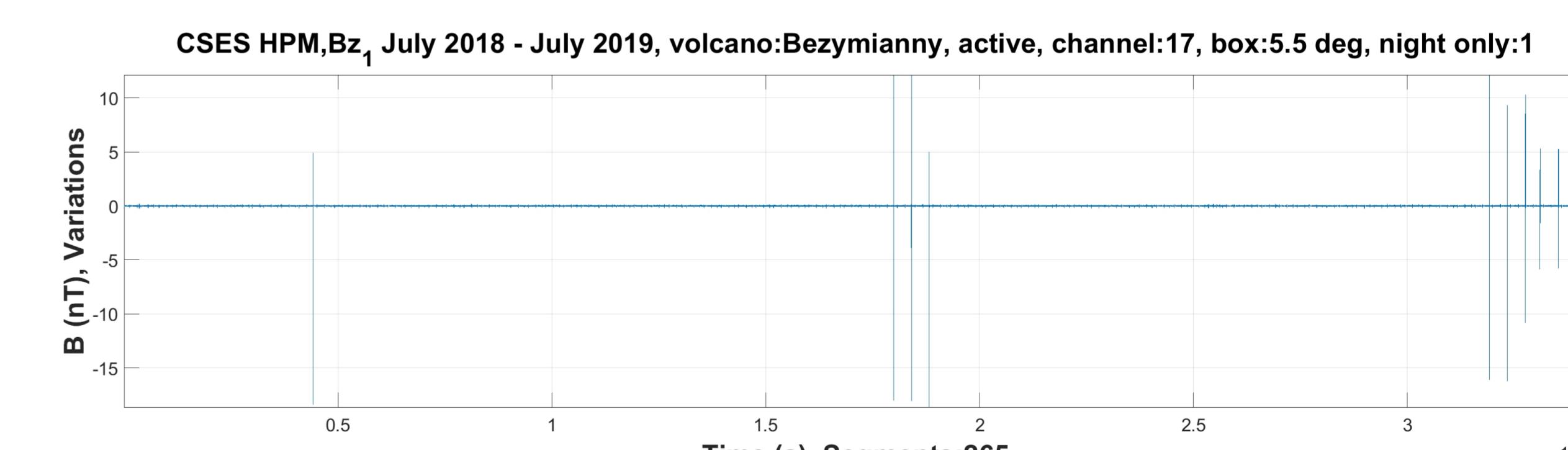
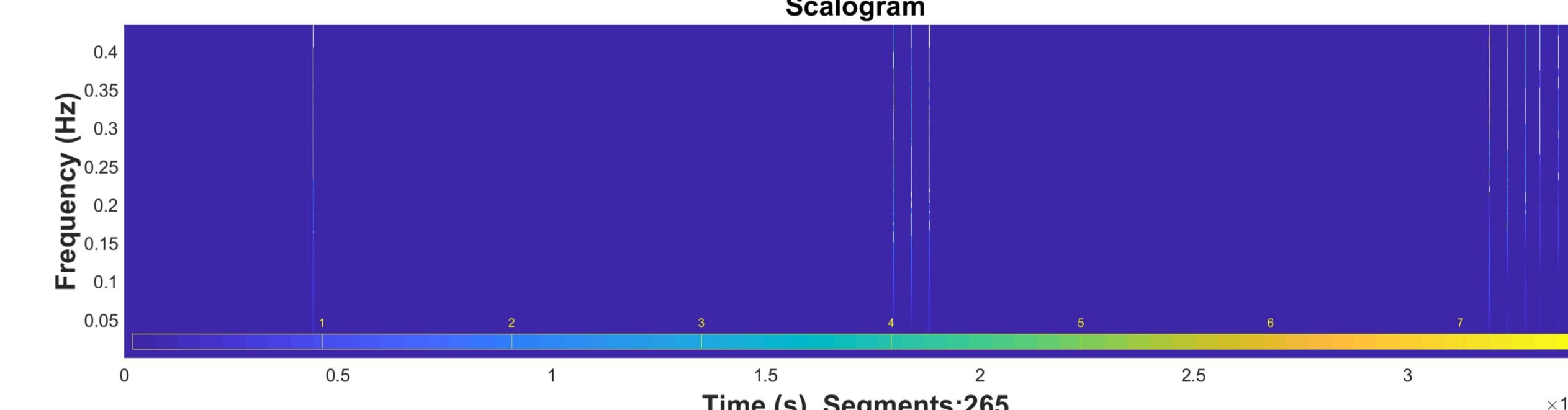
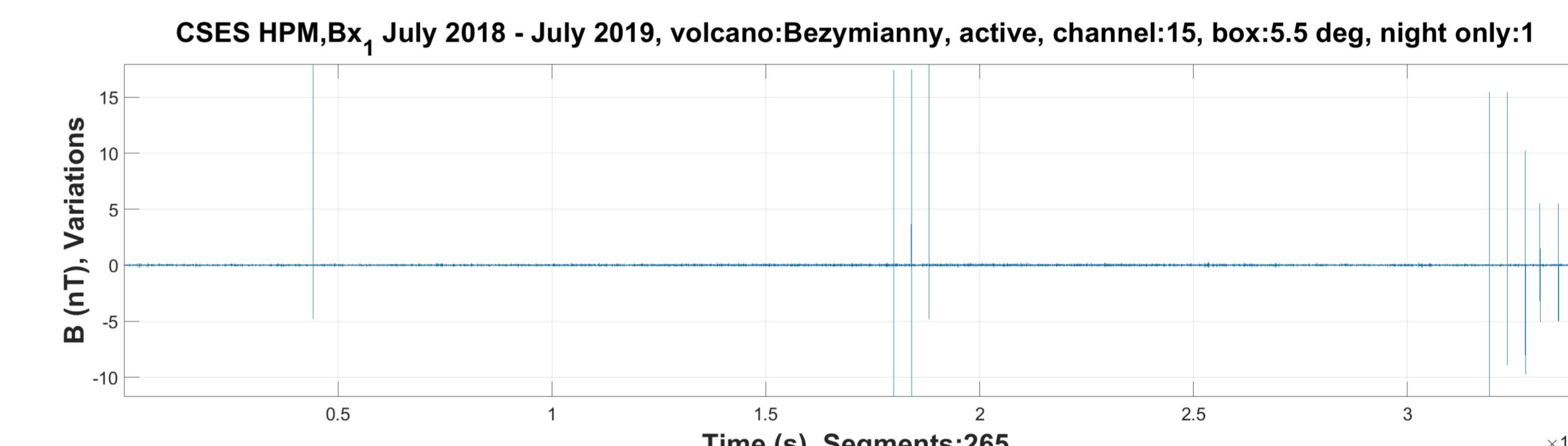
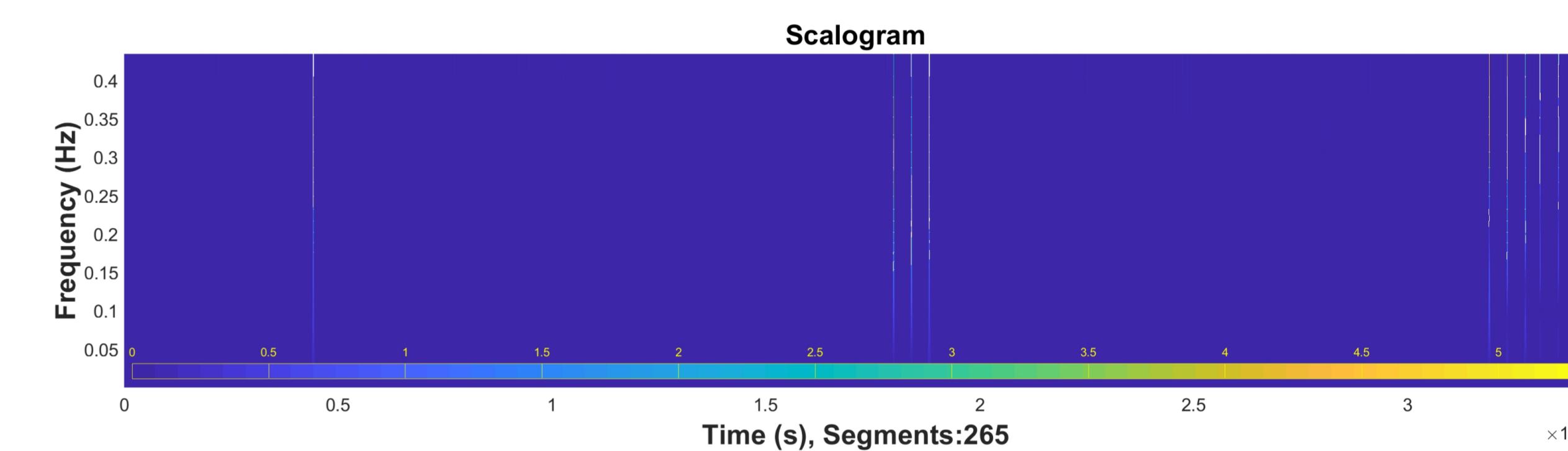
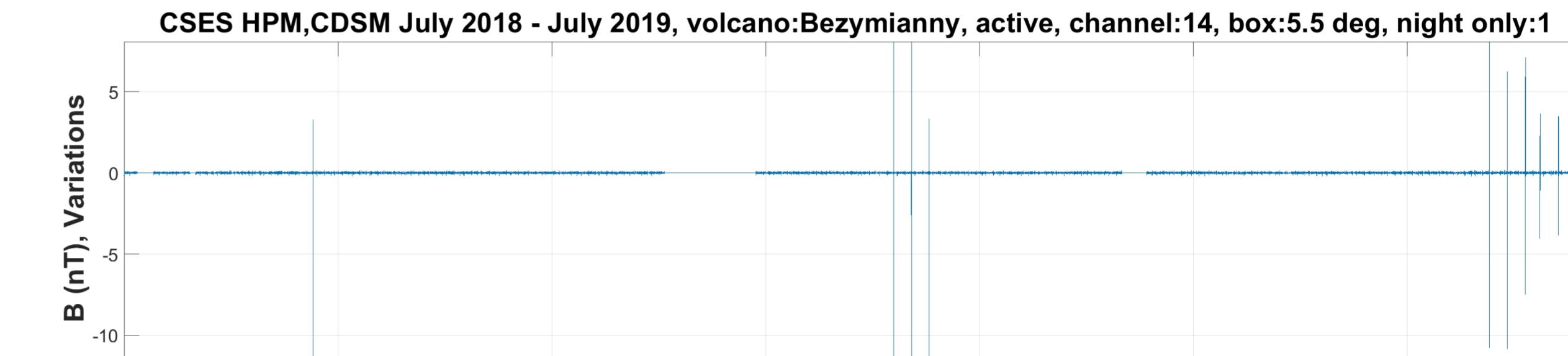
## FUEGO

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



## BEZYMIANNY

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



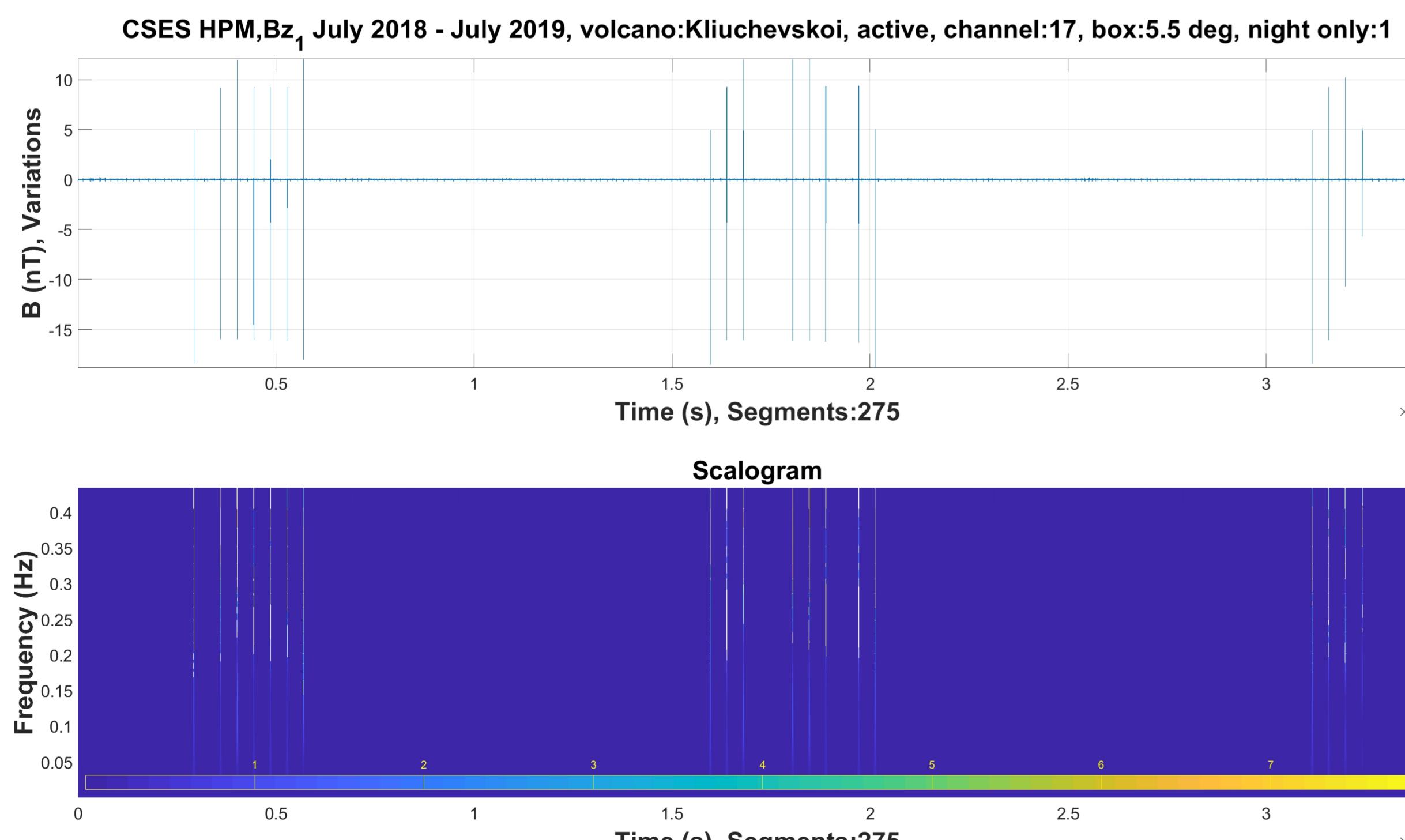
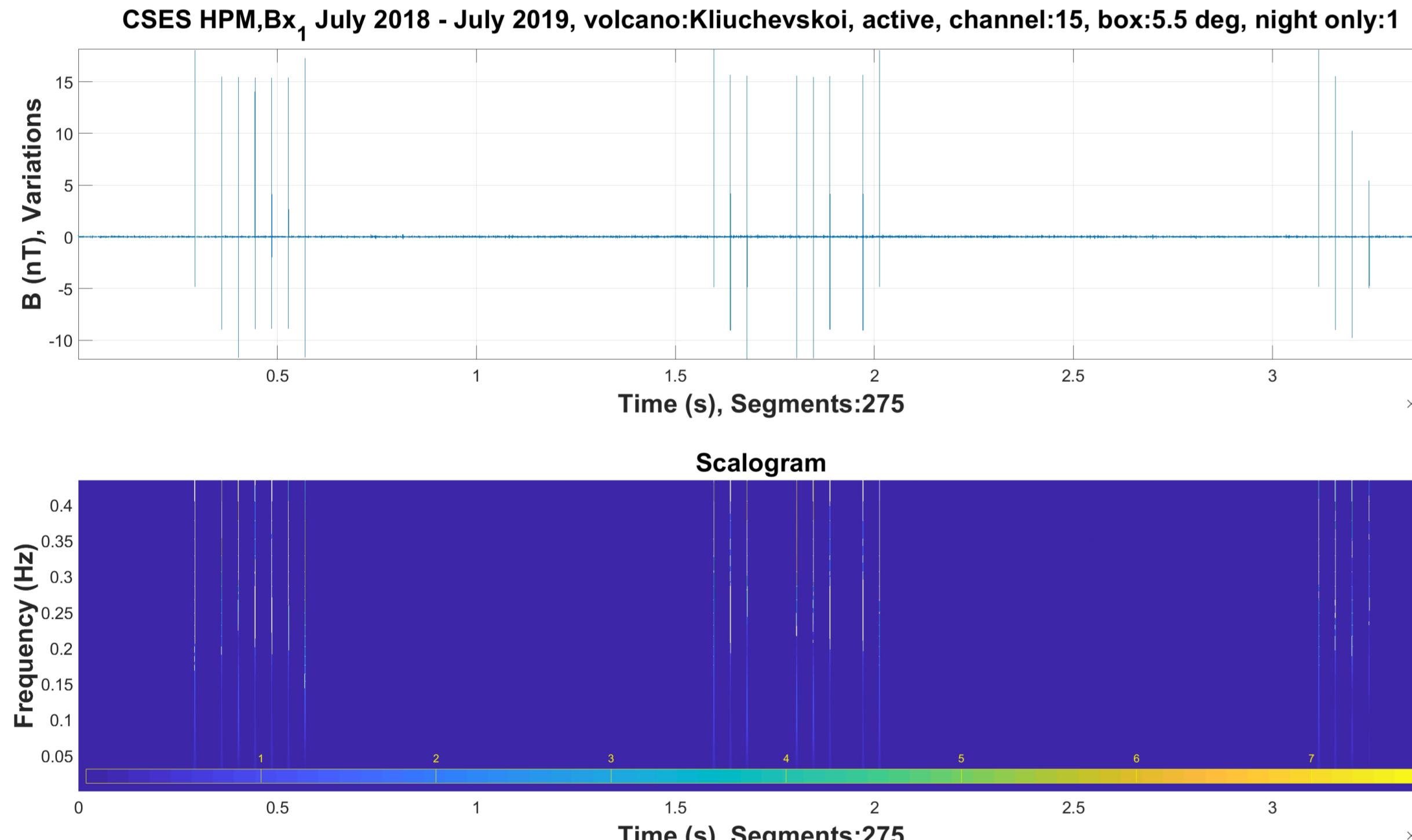
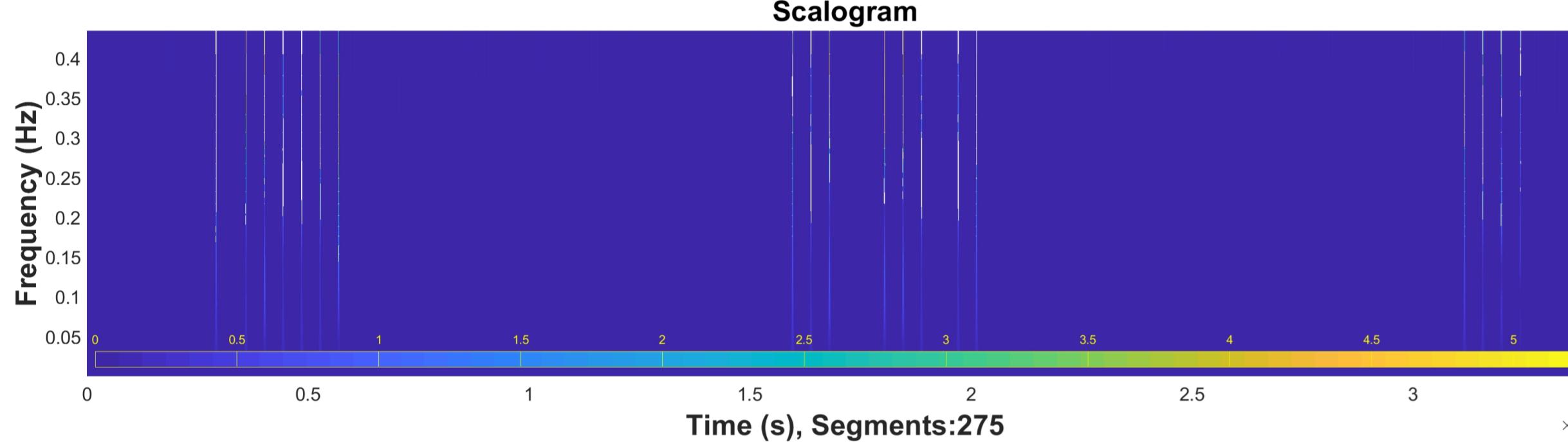
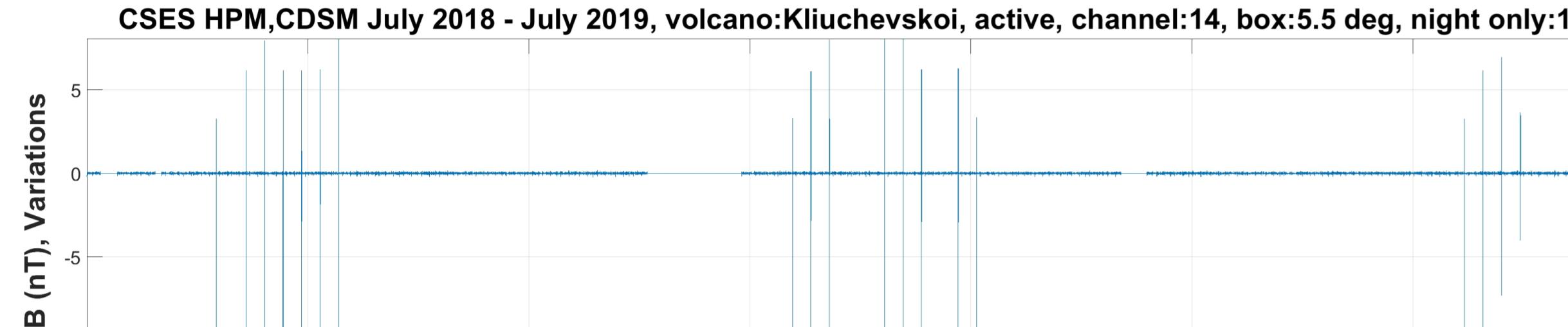
# SATELLITE AND GROUND-BASED MAGNETIC FIELD OBSERVATIONS RELATED TO VOLCANIC ERUPTIONS

K. Schwingenschuh<sup>1</sup>, W. Magnes<sup>1</sup>, X. Shen<sup>2</sup>, J. Wang<sup>3</sup>, B. Cheng<sup>7</sup>, A. Pollinger<sup>1</sup>, Chr. Hagen<sup>1</sup>, R. Lammegeier<sup>5</sup>, M. Ellmeier<sup>5</sup>, Christoph Schirniger<sup>1</sup>, H.-U. Eichelberger<sup>1</sup>, B. Mandl<sup>1</sup>, M. Y. Boudjada<sup>1</sup>, B. P. Besser<sup>1</sup>, A. A. Rozhnoi<sup>4</sup>, T. Zhang<sup>1</sup>, M. Delva<sup>1</sup>, I. Jernej<sup>1</sup>, Ö. Aydogar<sup>1</sup>, and R. Leonhardt<sup>6</sup>

<sup>1</sup>Space Research Institute, Austrian Academy of Sciences, Graz, Austria, <sup>2</sup>Institute of Earthquake Admin., Beijing, China, <sup>3</sup>Center for Space Science and Applied Research, CAS, Beijing, China, <sup>4</sup>Institute of the Earth Physics, RAS, Moscow, Russia, <sup>5</sup>Institute of Experimental Physics, Graz Univ. of Technology, Graz, Austria, <sup>6</sup>Central Institution for Meteorology and Geodynamics (ZAMG), Vienna, Austria, <sup>7</sup>National Space Science Center, CAS, Beijing, China

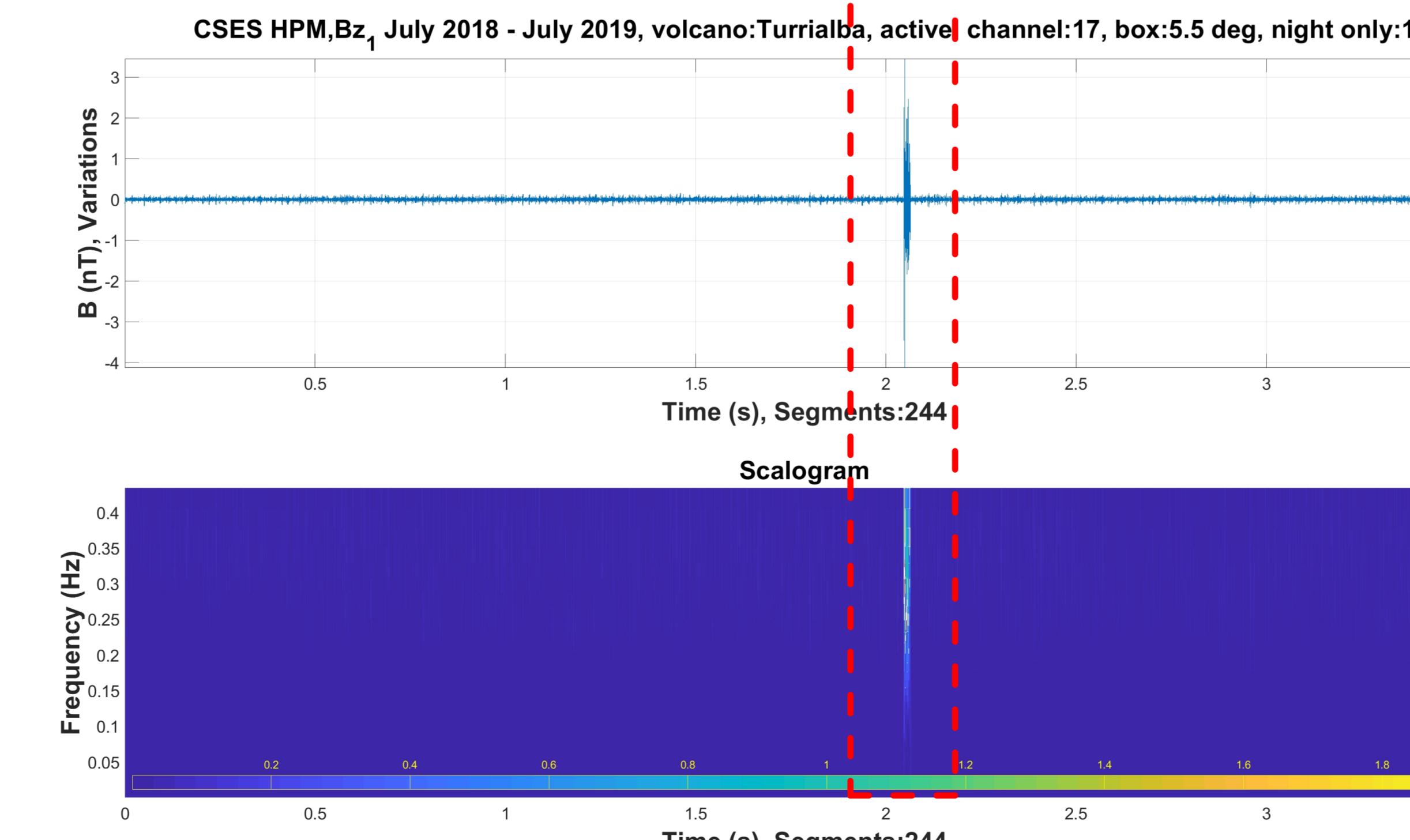
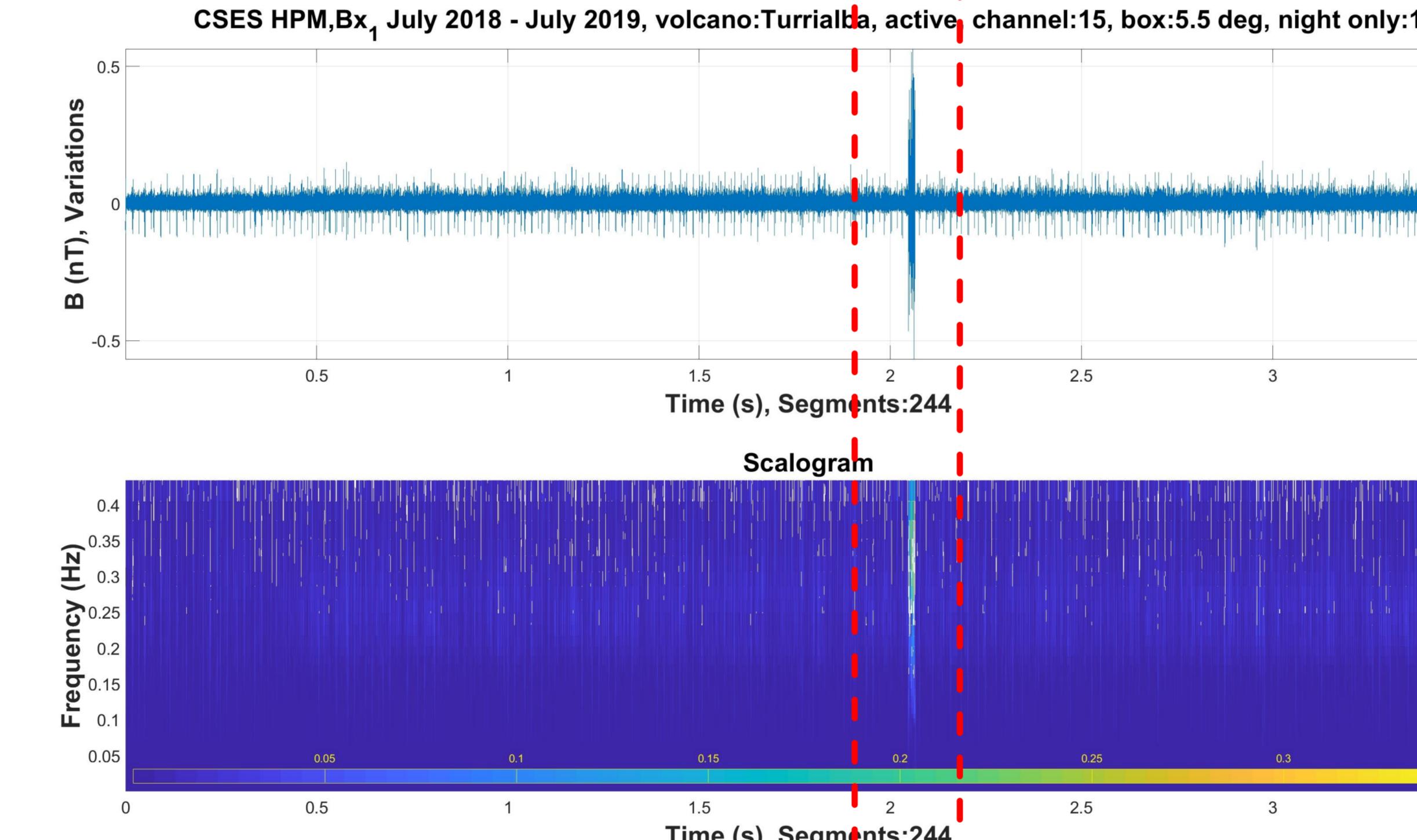
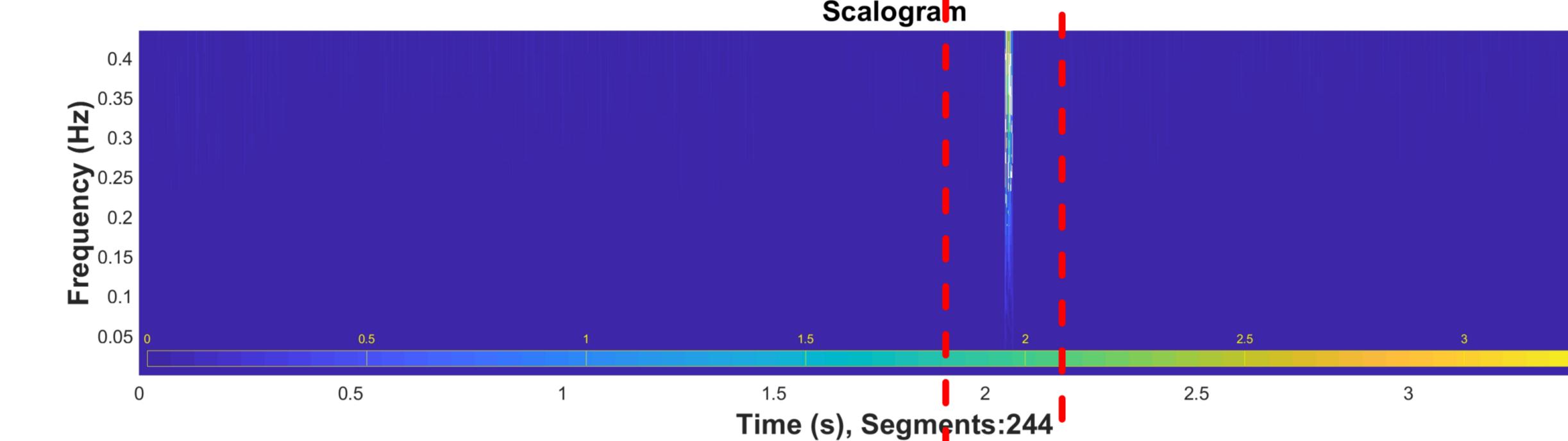
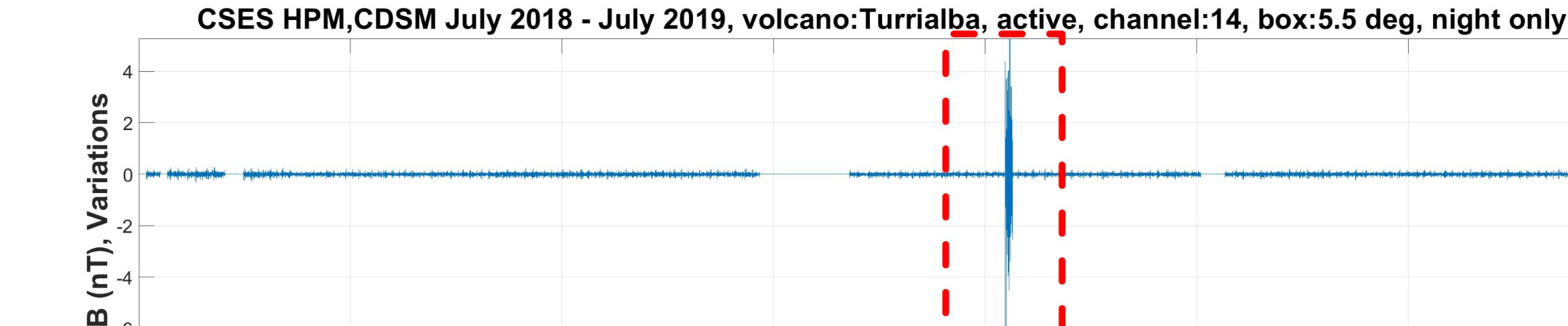
## KLYUCHEVSKOI

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



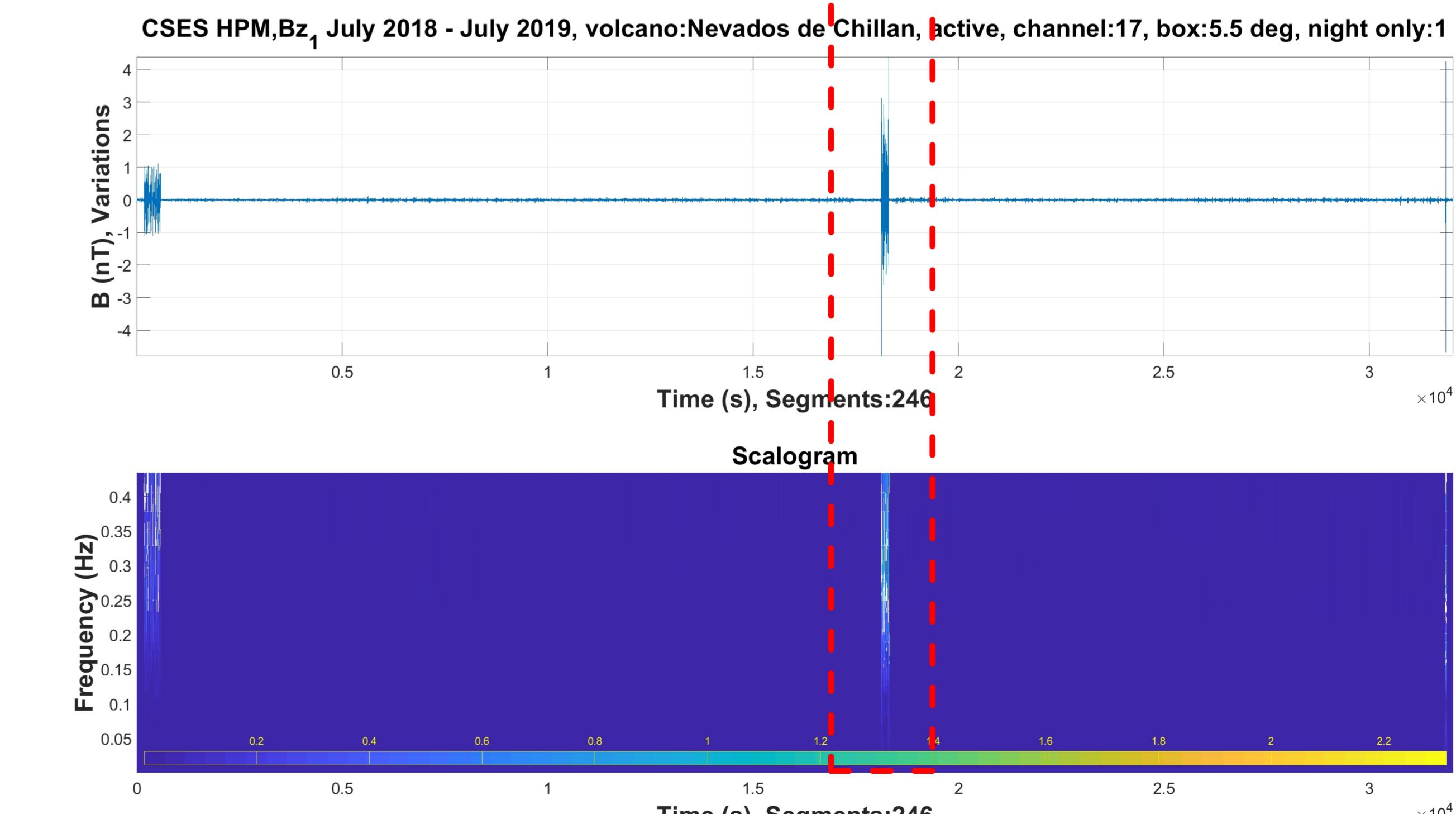
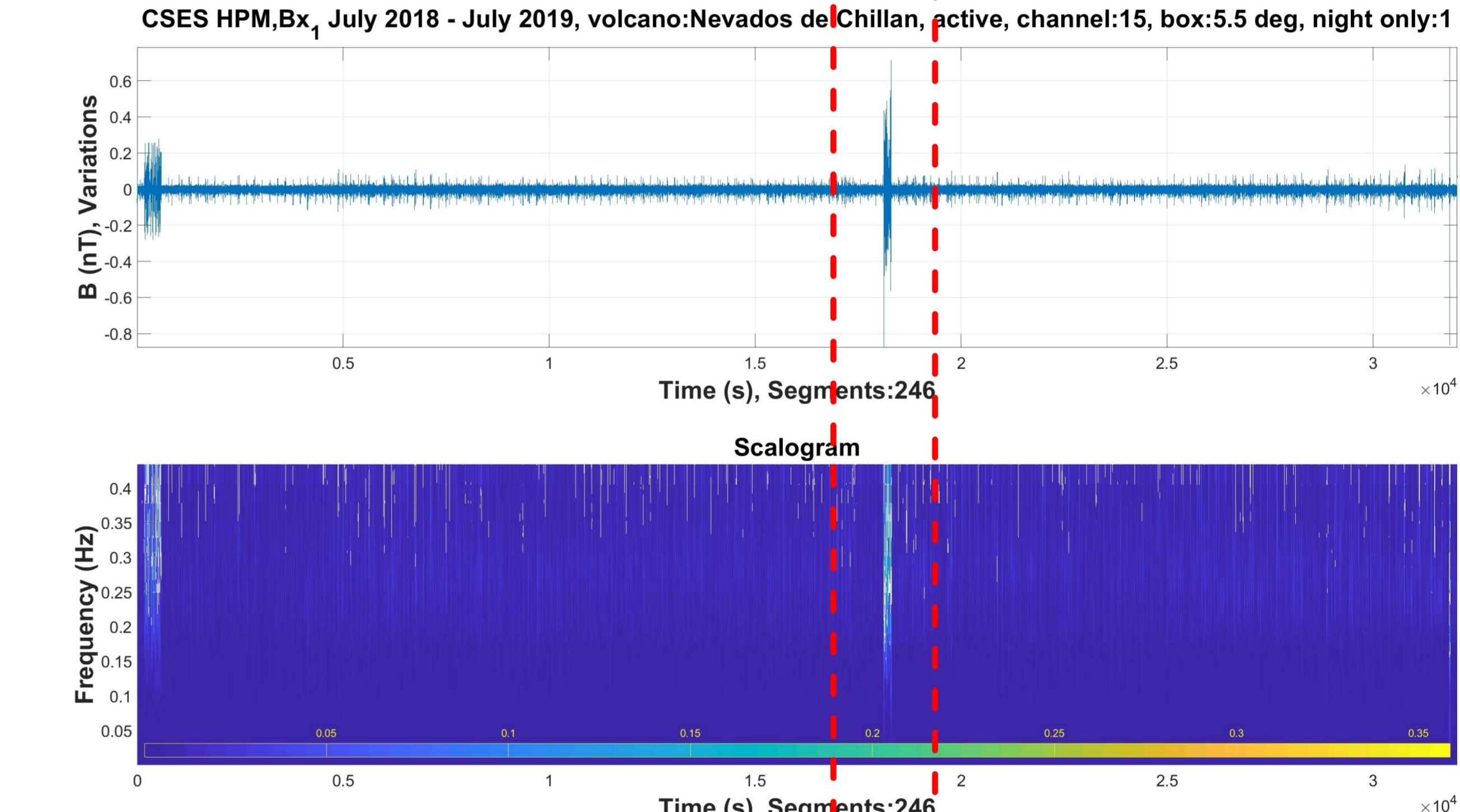
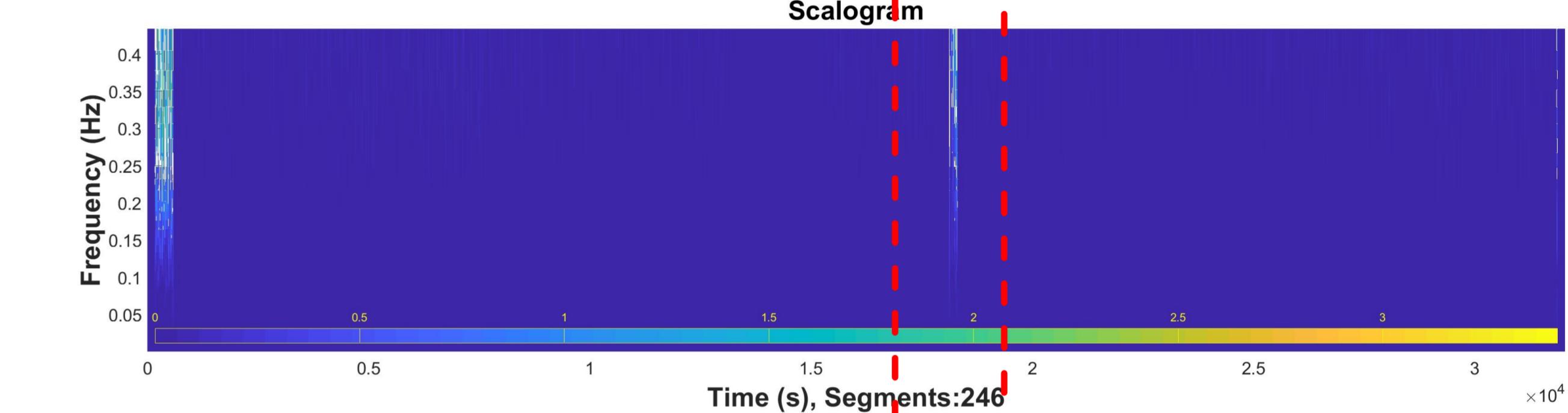
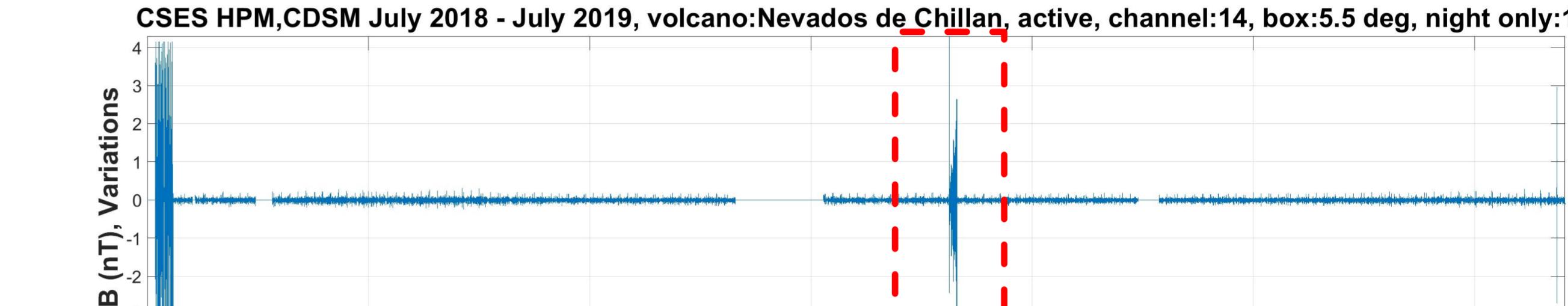
## TURRIALBA

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



## NEVADOS DE CHILLÁN

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



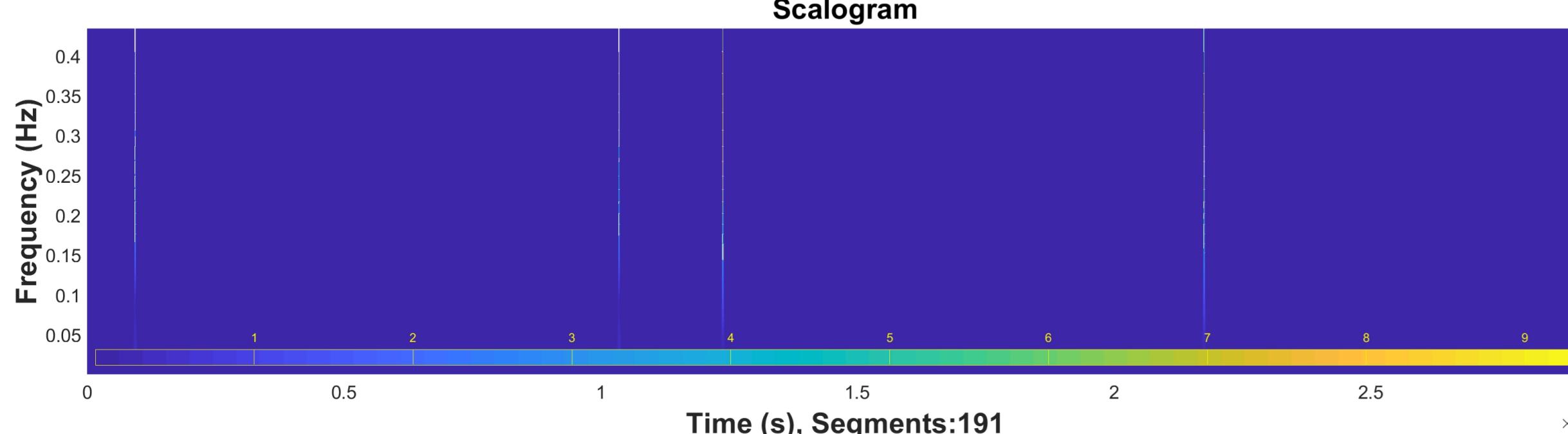
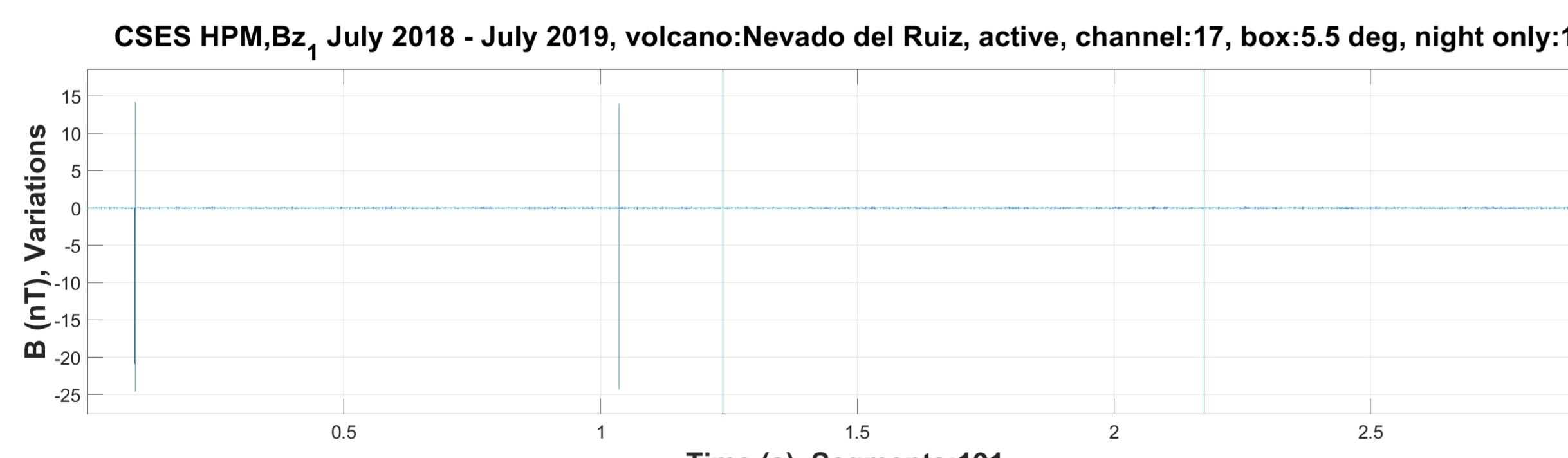
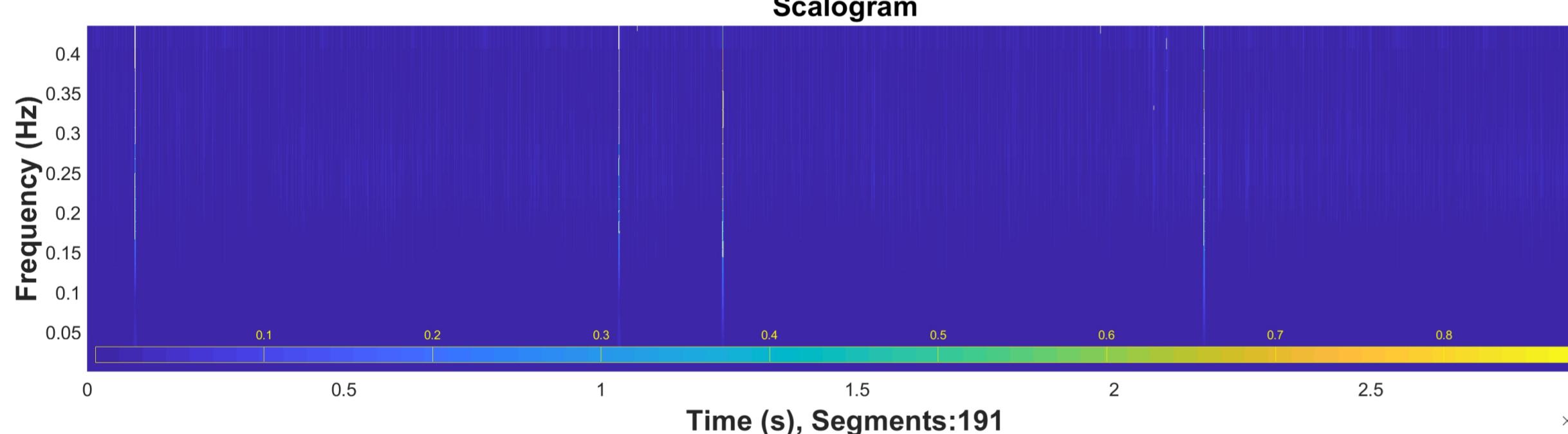
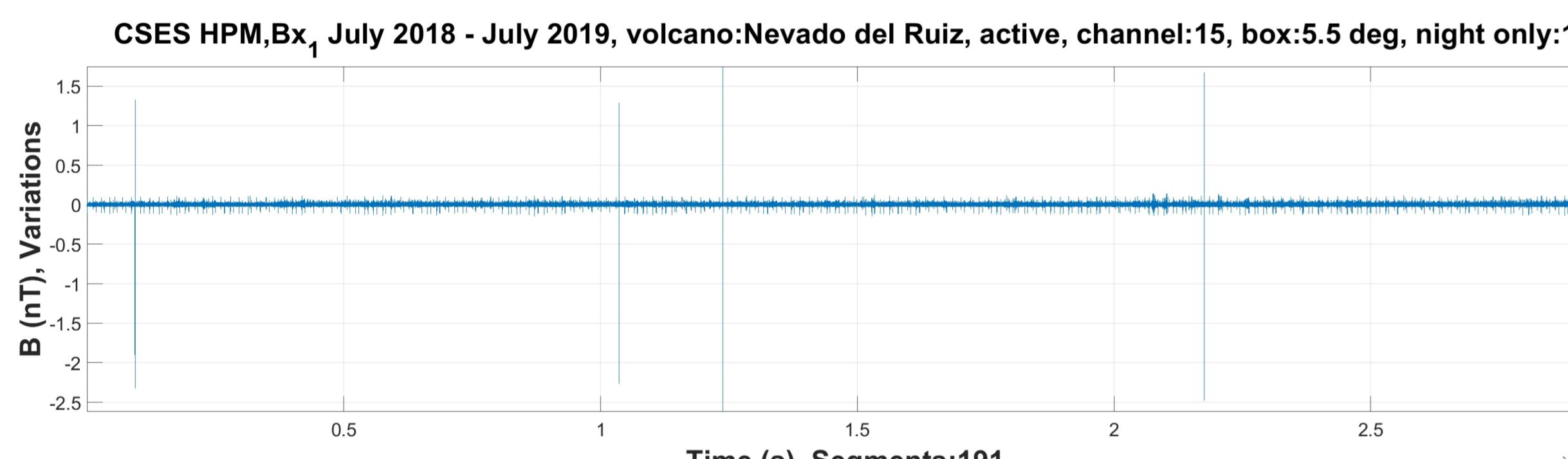
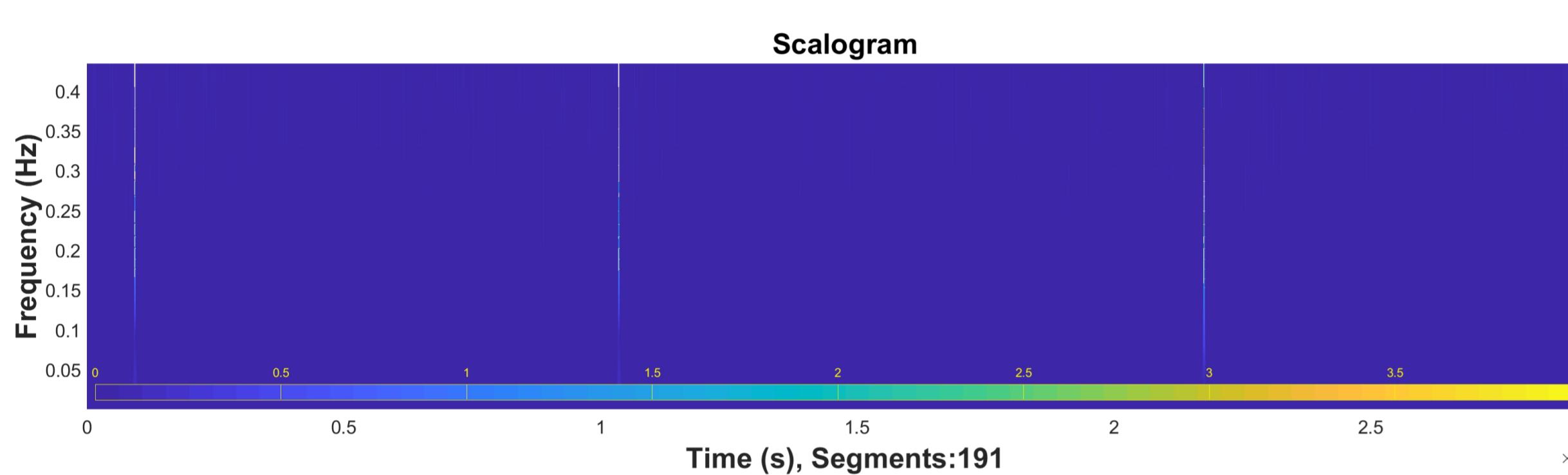
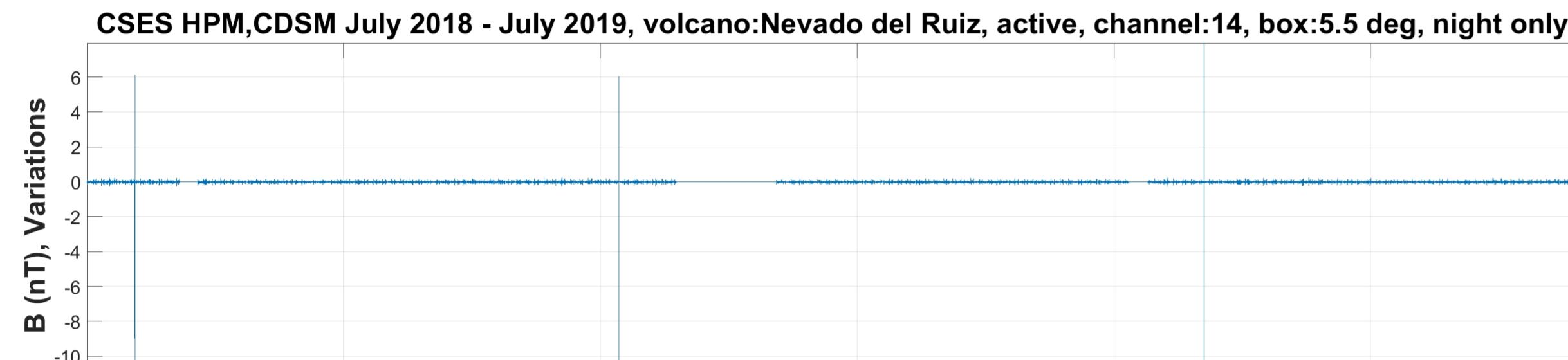
# SATELLITE AND GROUND-BASED MAGNETIC FIELD OBSERVATIONS RELATED TO VOLCANIC ERUPTIONS

K. Schwingenschuh<sup>1</sup>, W. Magnes<sup>1</sup>, X. Shen<sup>2</sup>, J. Wang<sup>3</sup>, B. Cheng<sup>7</sup>, A. Pollinger<sup>1</sup>, Chr. Hagen<sup>1</sup>, R. Lammegeier<sup>5</sup>, M. Ellmeier<sup>5</sup>, Christoph Schirniger<sup>1</sup>, H.-U. Eichelberger<sup>1</sup>, B. Mandl<sup>1</sup>, M. Y. Boudjada<sup>1</sup>, B. P. Besser<sup>1</sup>, A. A. Rozhnoi<sup>4</sup>, T. Zhang<sup>1</sup>, M. Delva<sup>1</sup>, I. Jernej<sup>1</sup>, Ö. Aydogar<sup>1</sup>, and R. Leonhardt<sup>6</sup>

<sup>1</sup>Space Research Institute, Austrian Academy of Sciences, Graz, Austria, <sup>2</sup>Institute of Earthquake Admin., Beijing, China, <sup>3</sup>Center for Space Science and Applied Research, CAS, Beijing, China, <sup>4</sup>Institute of the Earth Physics, RAS, Moscow, Russia, <sup>5</sup>Institute of Experimental Physics, Graz Univ. of Technology, Graz, Austria, <sup>6</sup>Central Institution for Meteorology and Geodynamics (ZAMG), Vienna, Austria, <sup>7</sup>National Space Science Center, CAS, Beijing, China

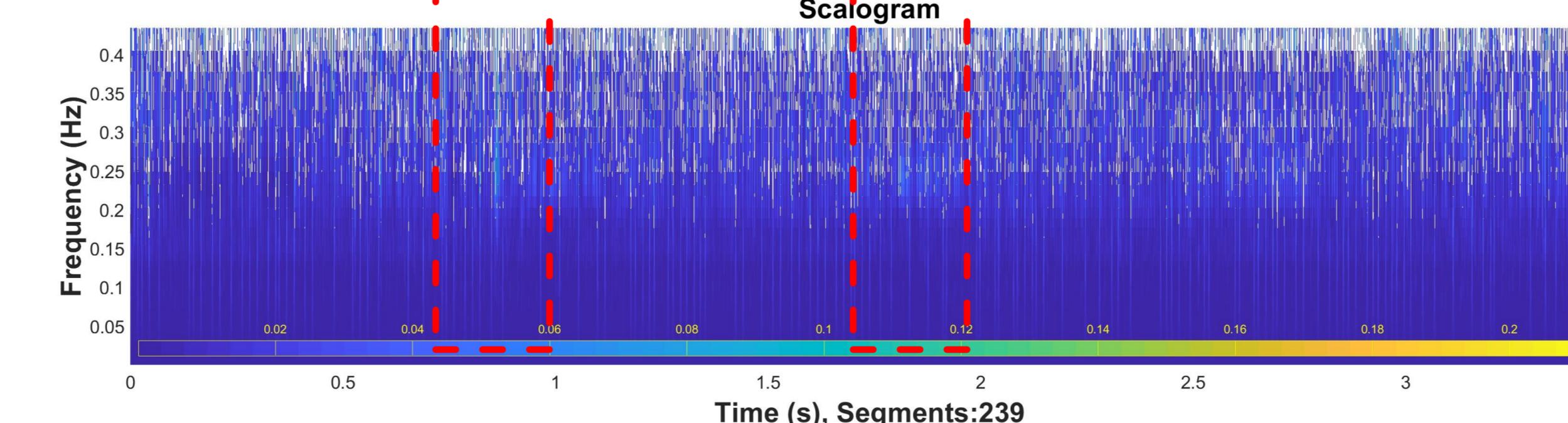
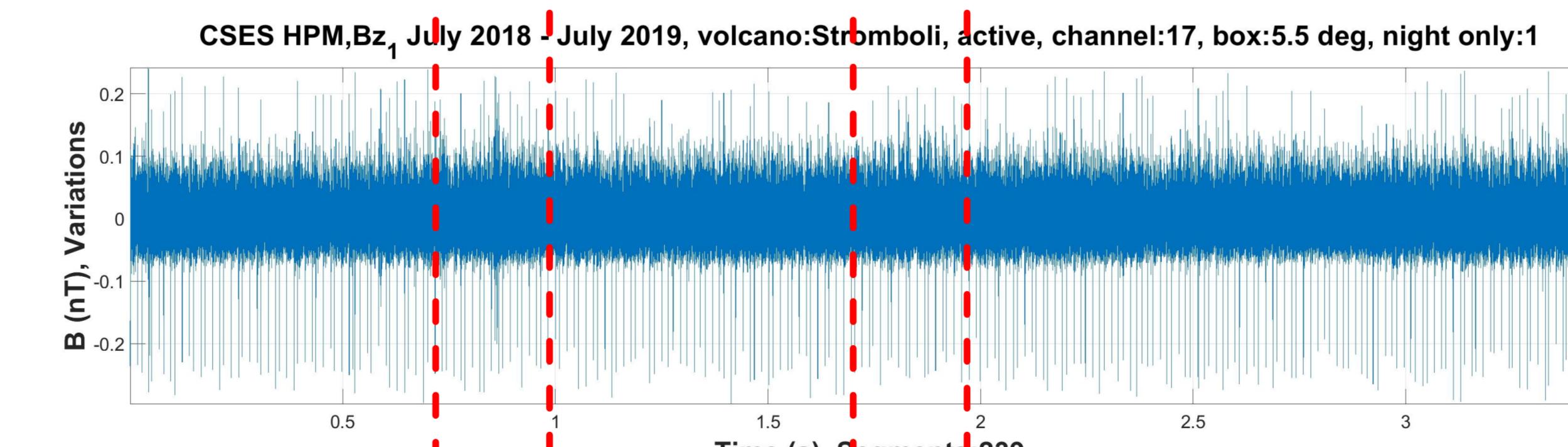
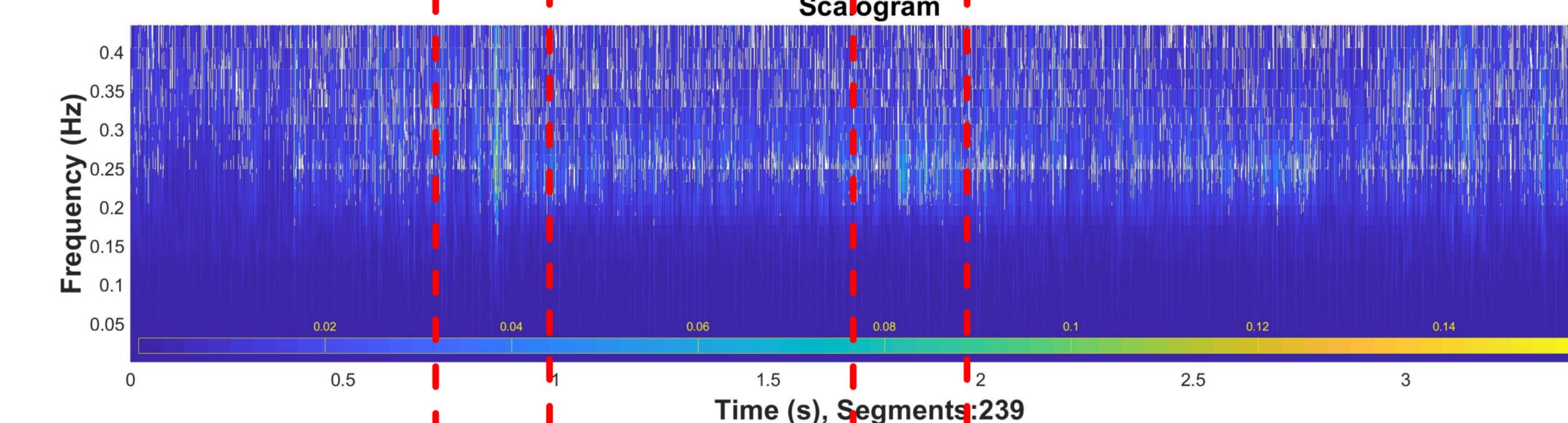
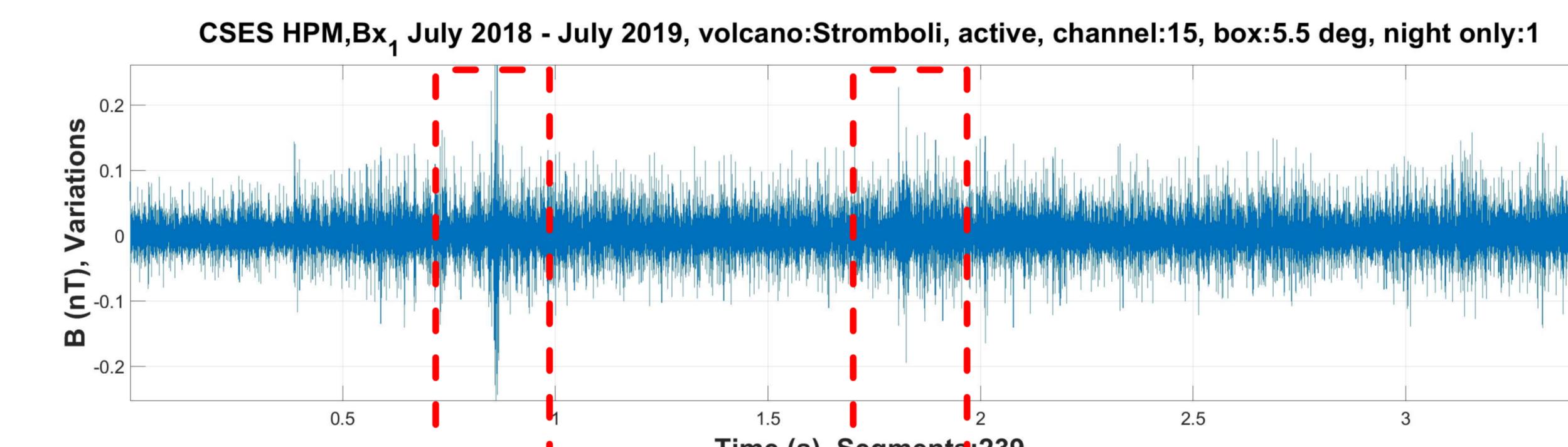
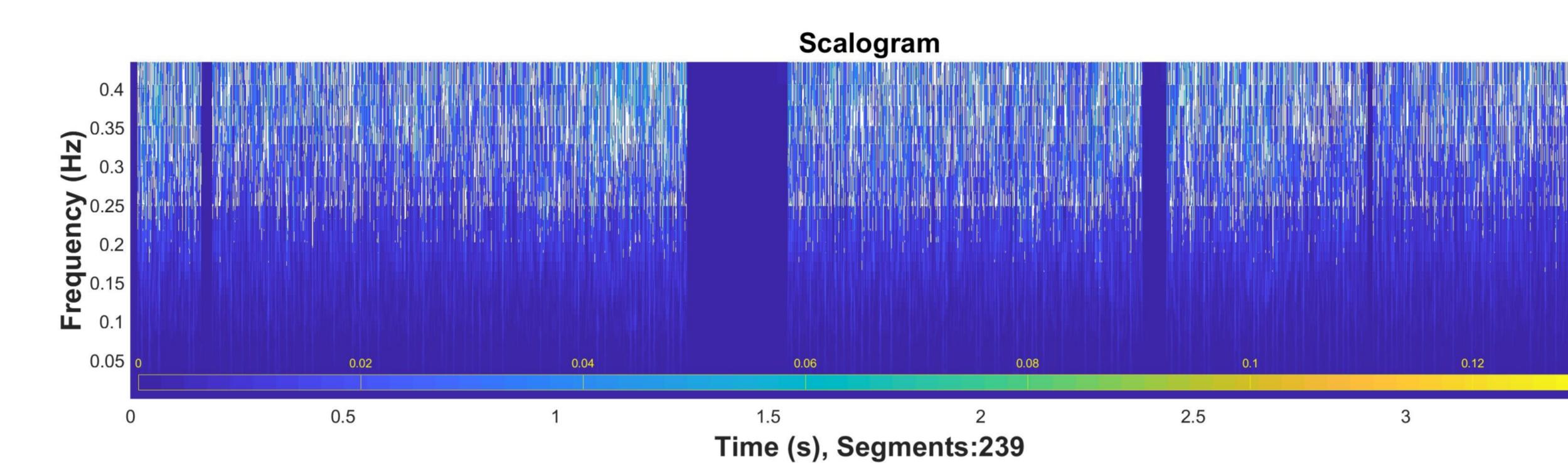
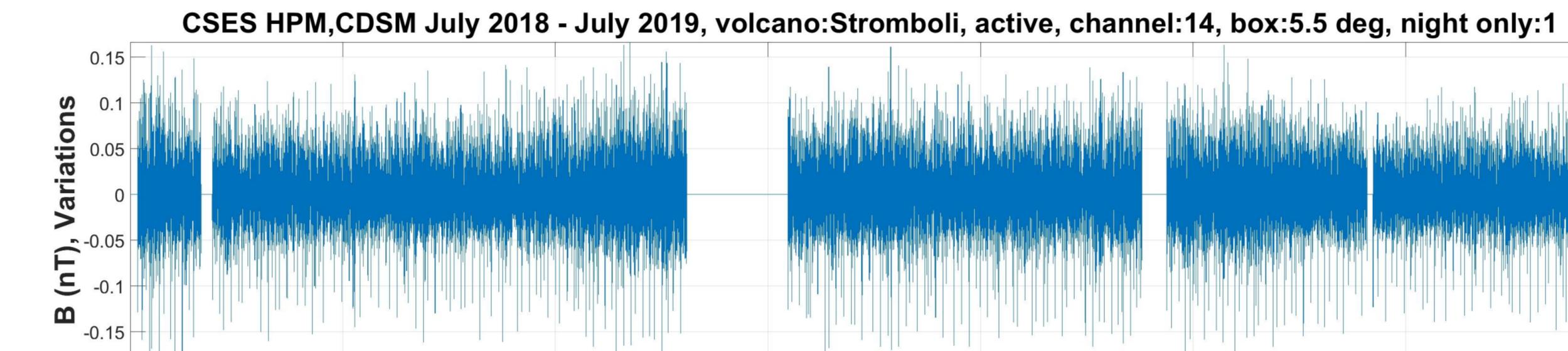
## NEVADO DEL RUIZ

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



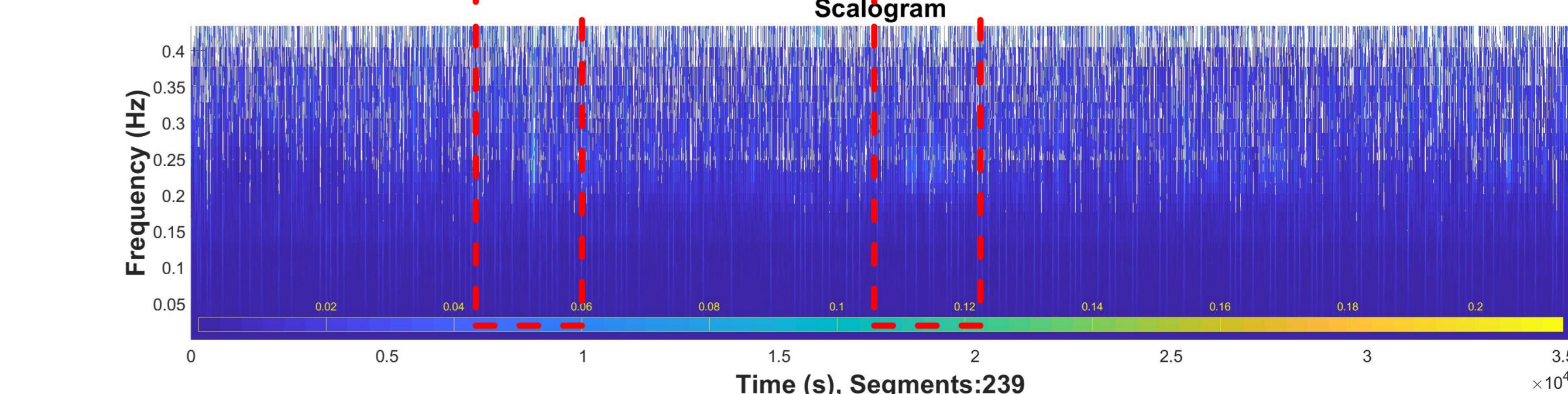
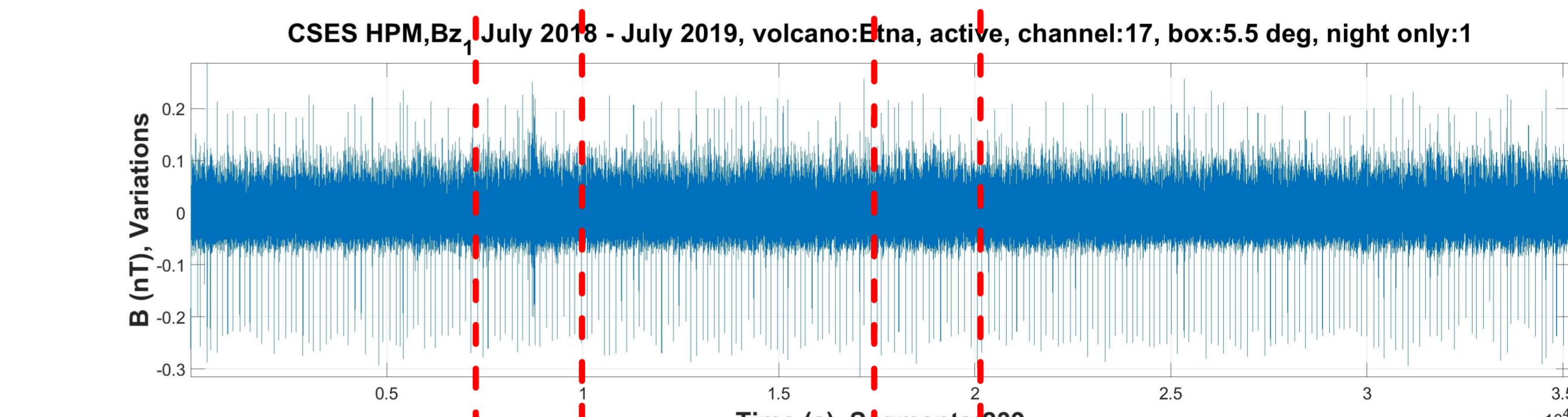
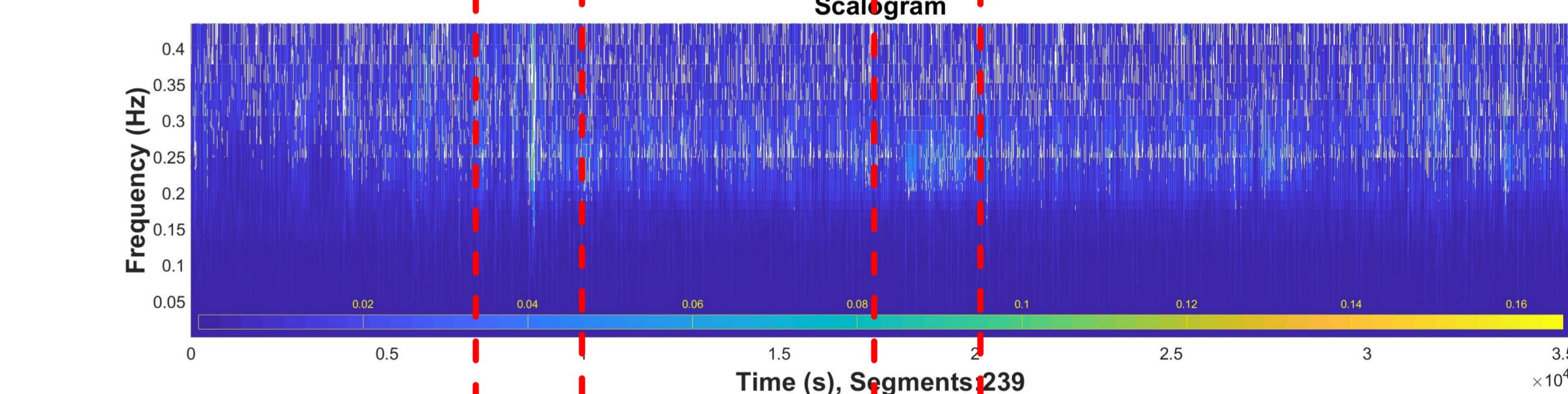
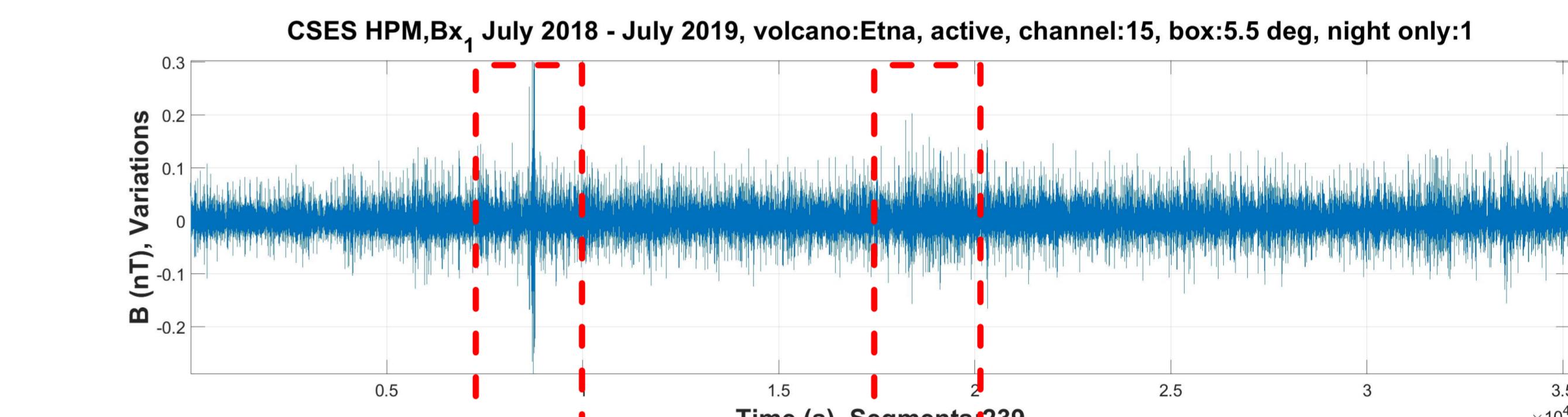
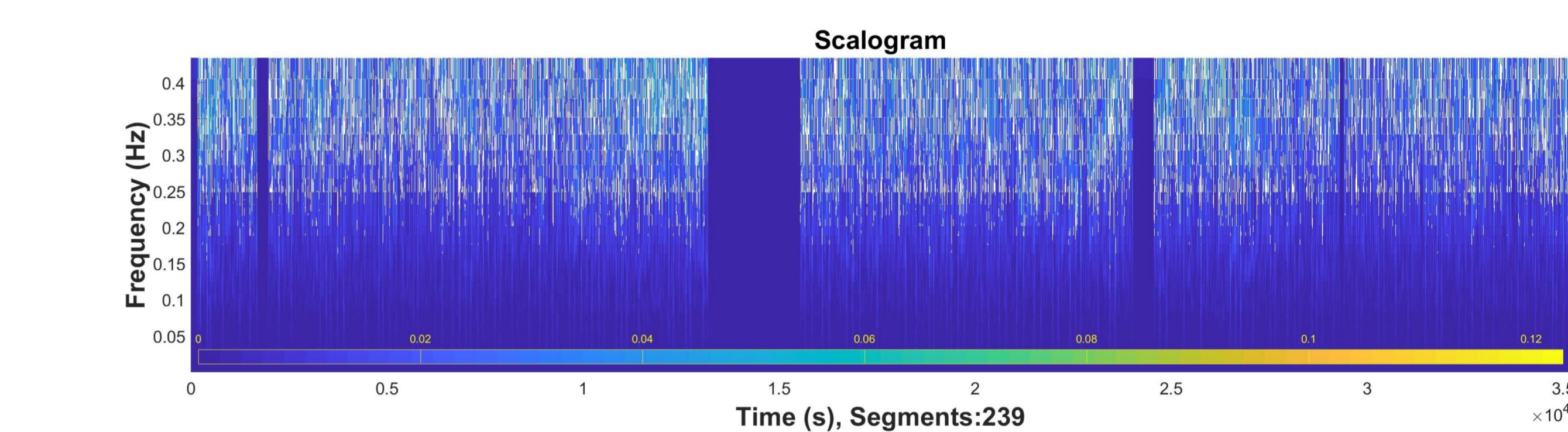
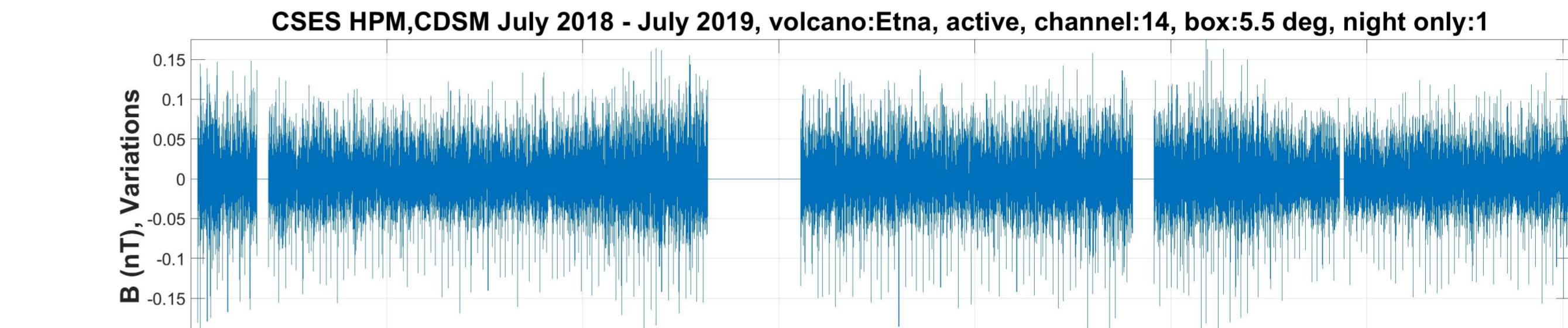
## STROMBOLI

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



## ETNA

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



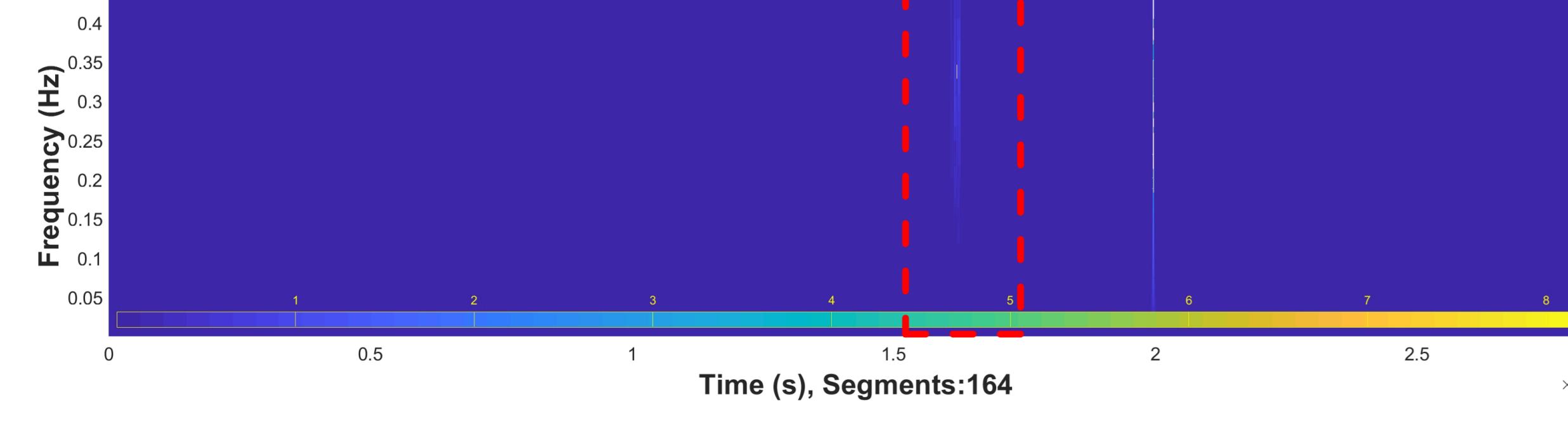
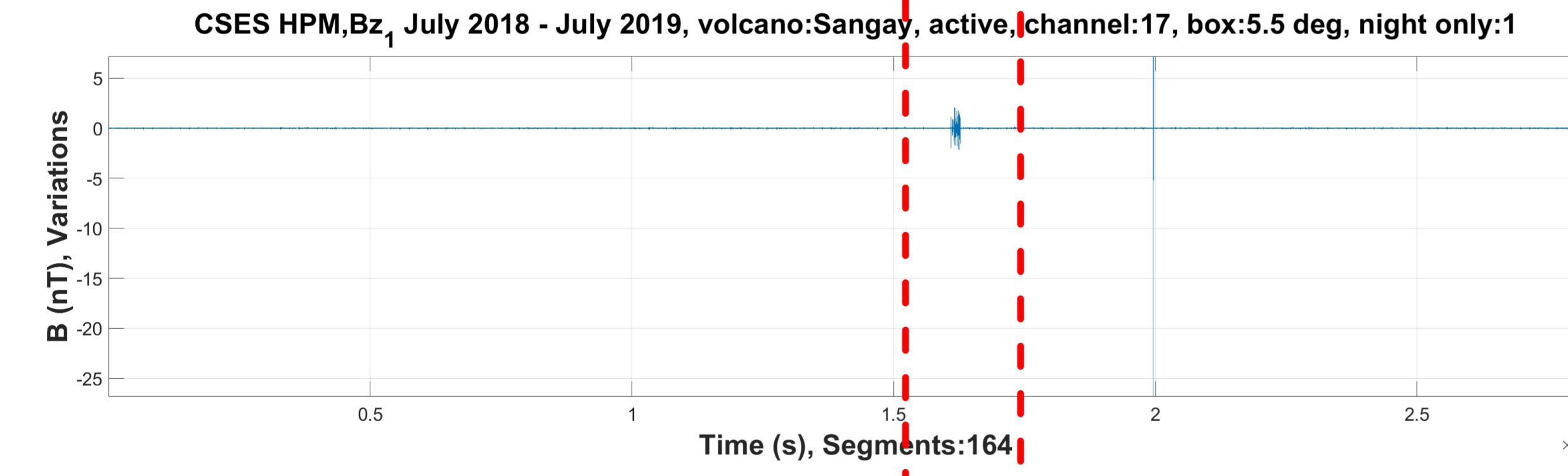
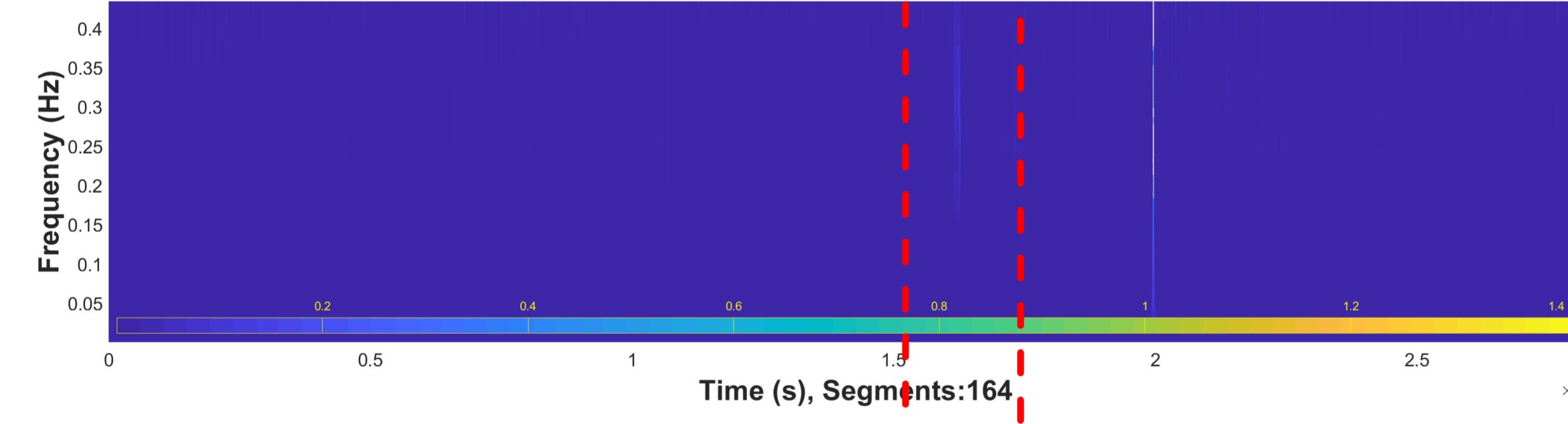
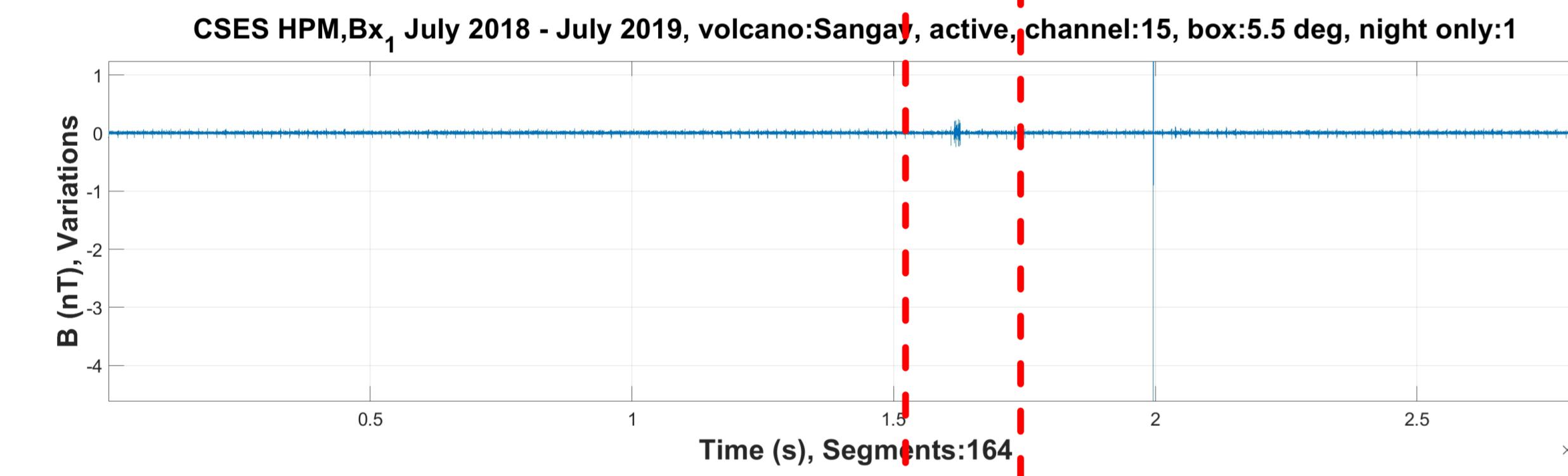
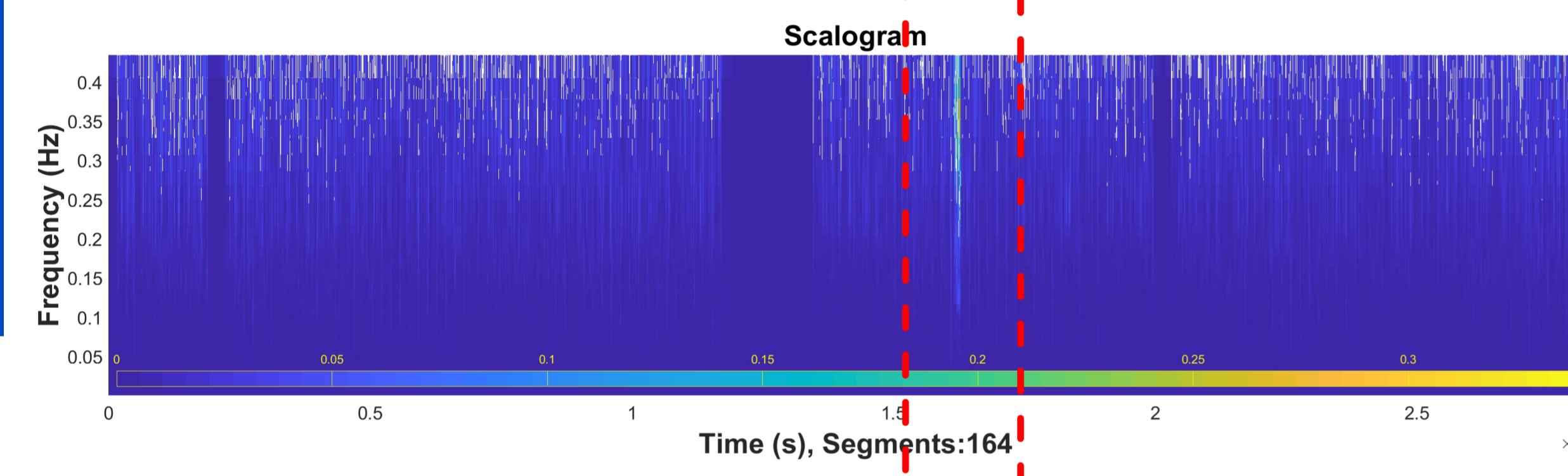
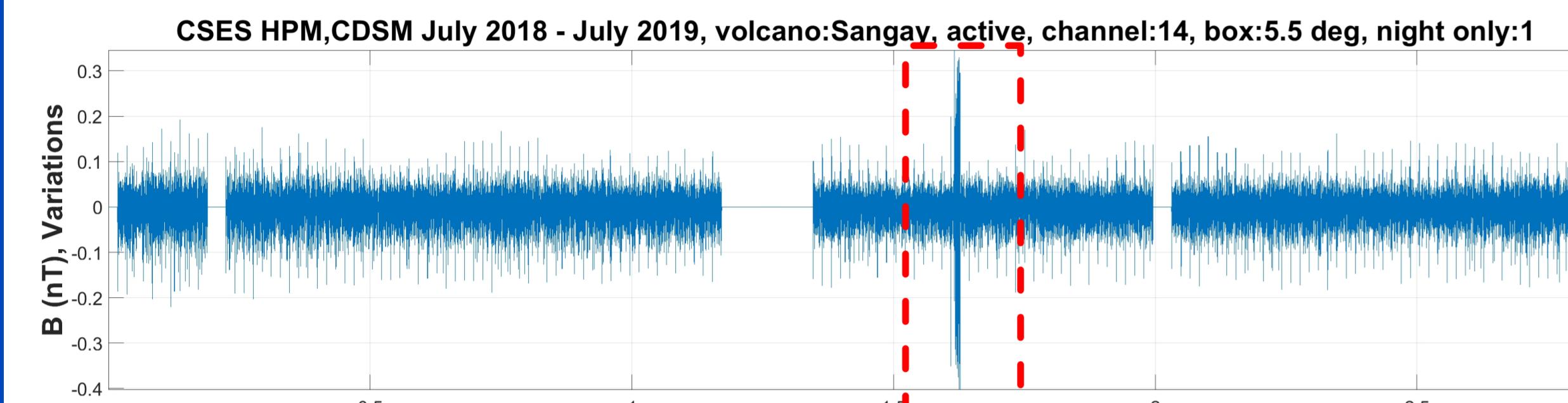
# SATELLITE AND GROUND-BASED MAGNETIC FIELD OBSERVATIONS RELATED TO VOLCANIC ERUPTIONS

K. Schwingenschuh<sup>1</sup>, W. Magnes<sup>1</sup>, X. Shen<sup>2</sup>, J. Wang<sup>3</sup>, B. Cheng<sup>7</sup>, A. Pollinger<sup>1</sup>, Chr. Hagen<sup>1</sup>, R. Lammegger<sup>5</sup>, M. Ellmeier<sup>5</sup>, Christoph Schirniger<sup>1</sup>, H.-U. Eichelberger<sup>1</sup>, B. Mandl<sup>1</sup>, M. Y. Boudjada<sup>1</sup>, B. P. Besser<sup>1</sup>, A. A. Rozhnoi<sup>4</sup>, T. Zhang<sup>1</sup>, M. Delva<sup>1</sup>, I. Jernej<sup>1</sup>, Ö. Aydogar<sup>1</sup>, and R. Leonhardt<sup>6</sup>

<sup>1</sup>Space Research Institute, Austrian Academy of Sciences, Graz, Austria, <sup>2</sup>Institute of Earthquake Admin., Beijing, China, <sup>3</sup>Center for Space Science and Applied Research, CAS, Beijing, China, <sup>4</sup>Institute of the Earth Physics, RAS, Moscow, Russia, <sup>5</sup>Institute of Experimental Physics, Graz Univ. of Technology, Graz, Austria, <sup>6</sup>Central Institution for Meteorology and Geodynamics (ZAMG), Vienna, Austria, <sup>7</sup>National Space Science Center, CAS, Beijing, China

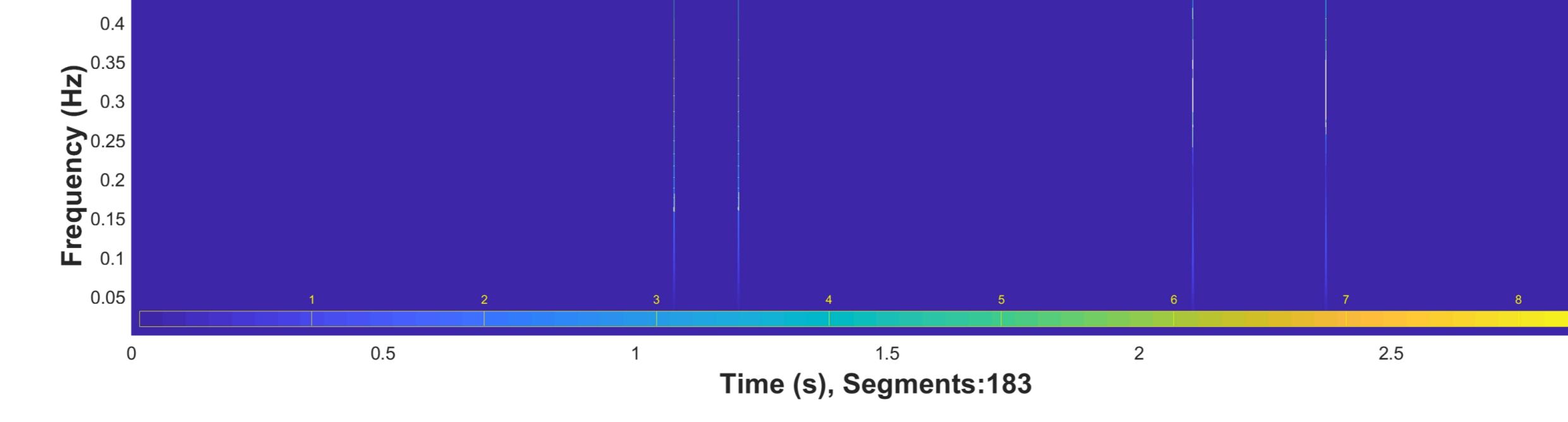
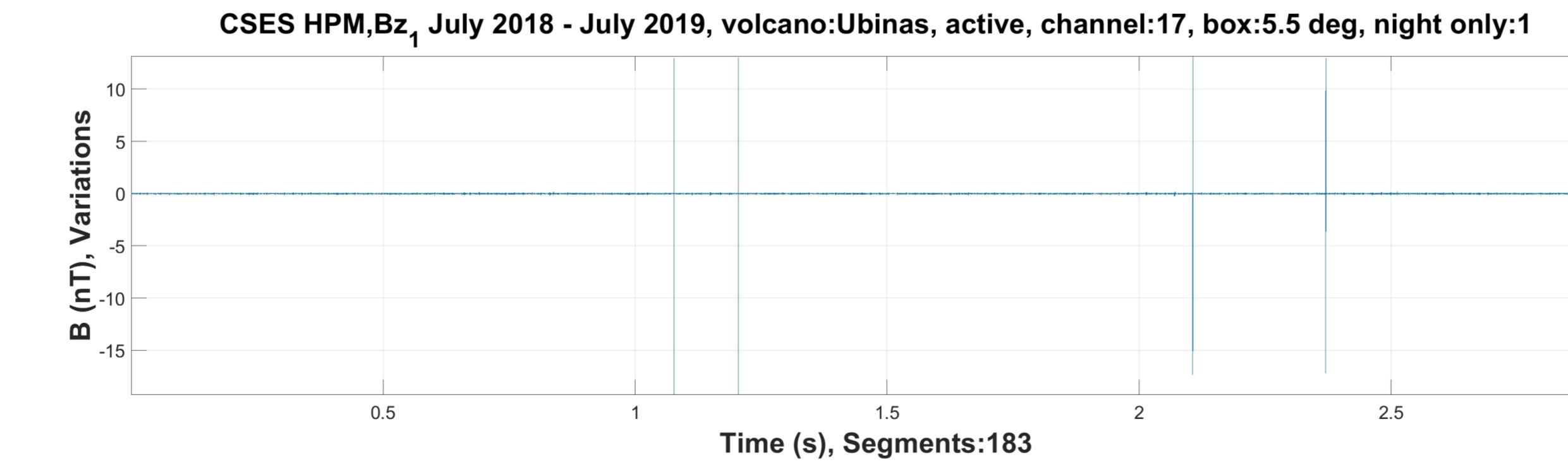
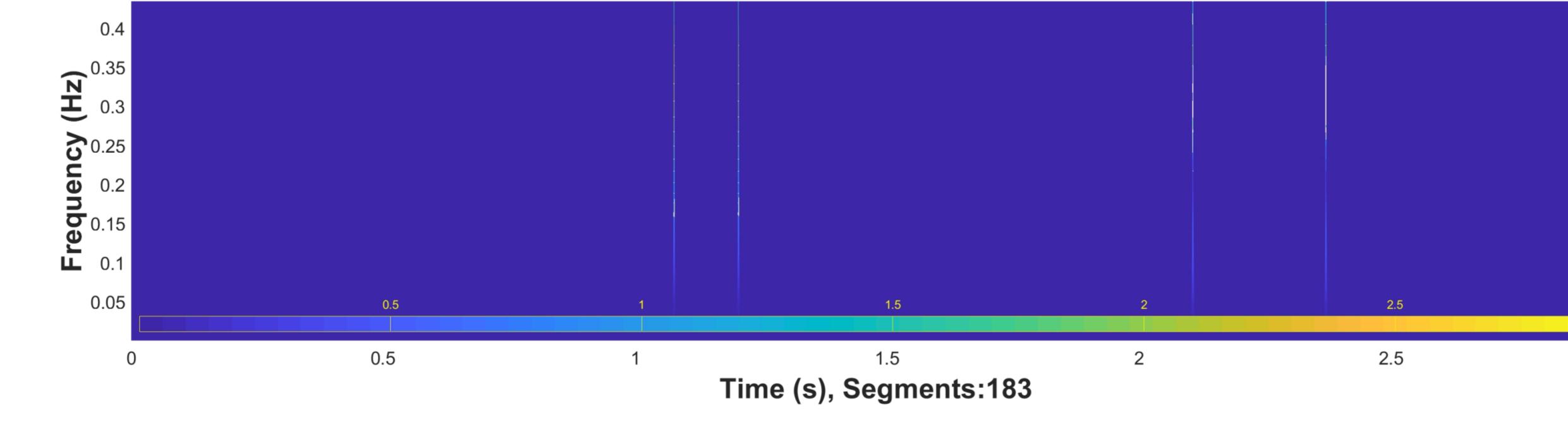
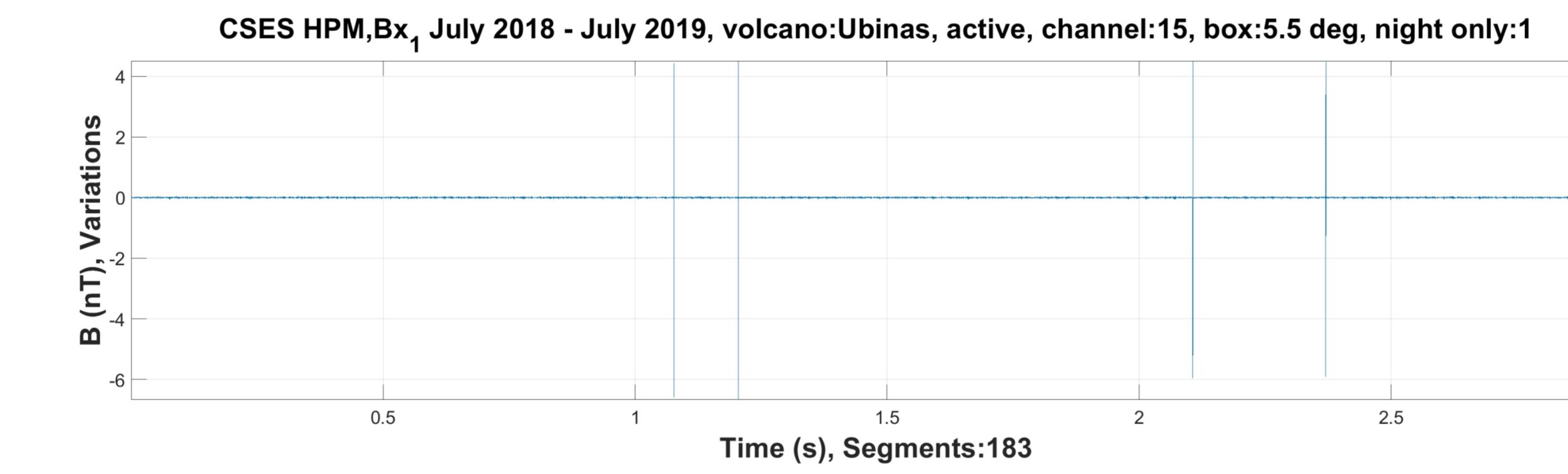
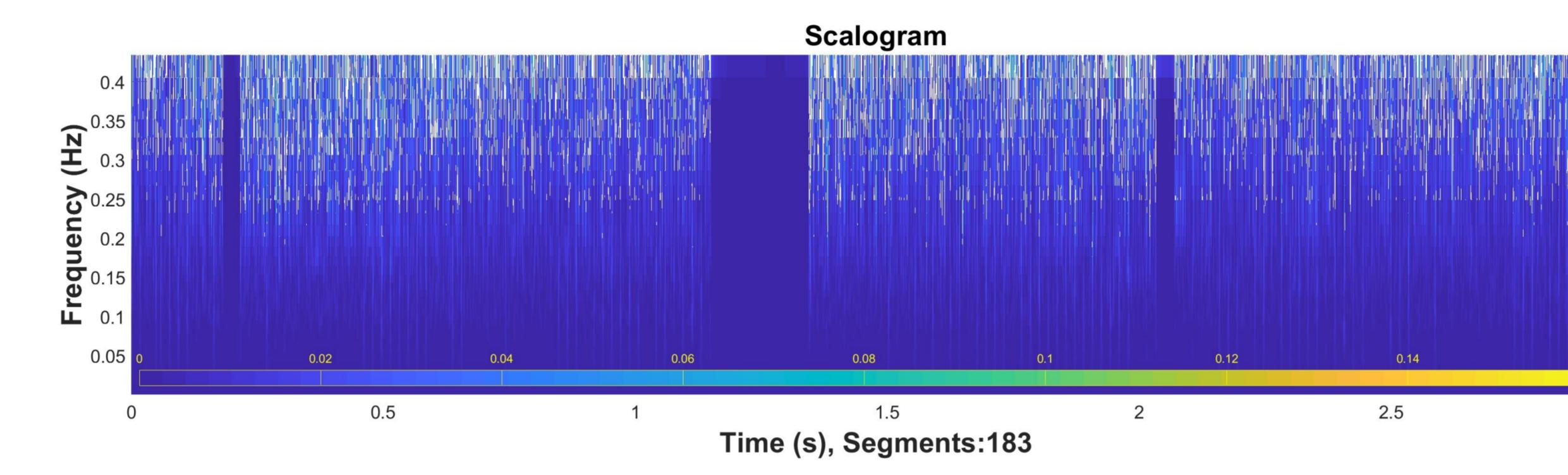
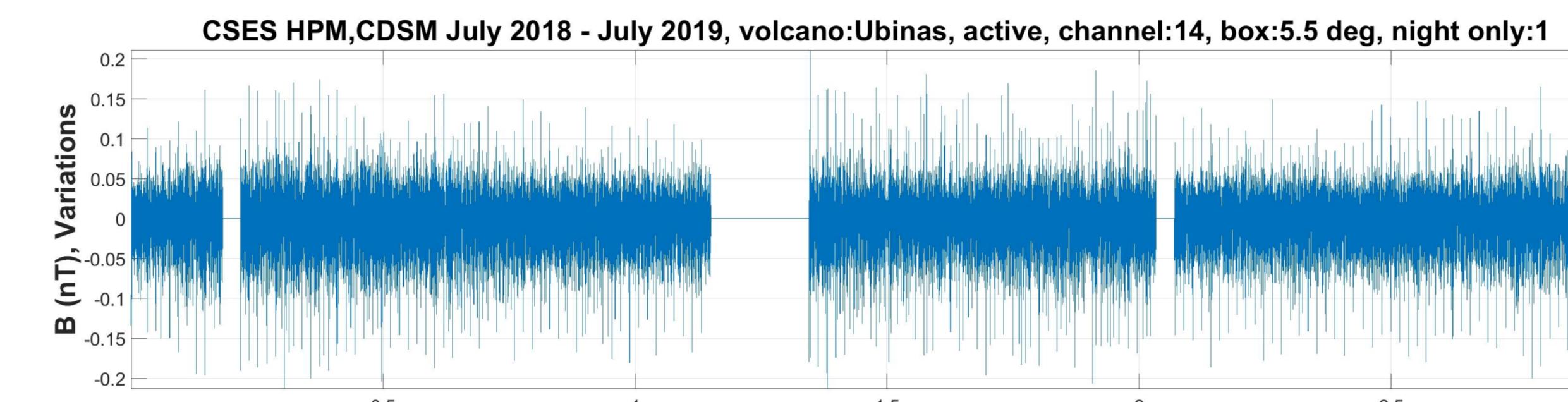
## SANGAY

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



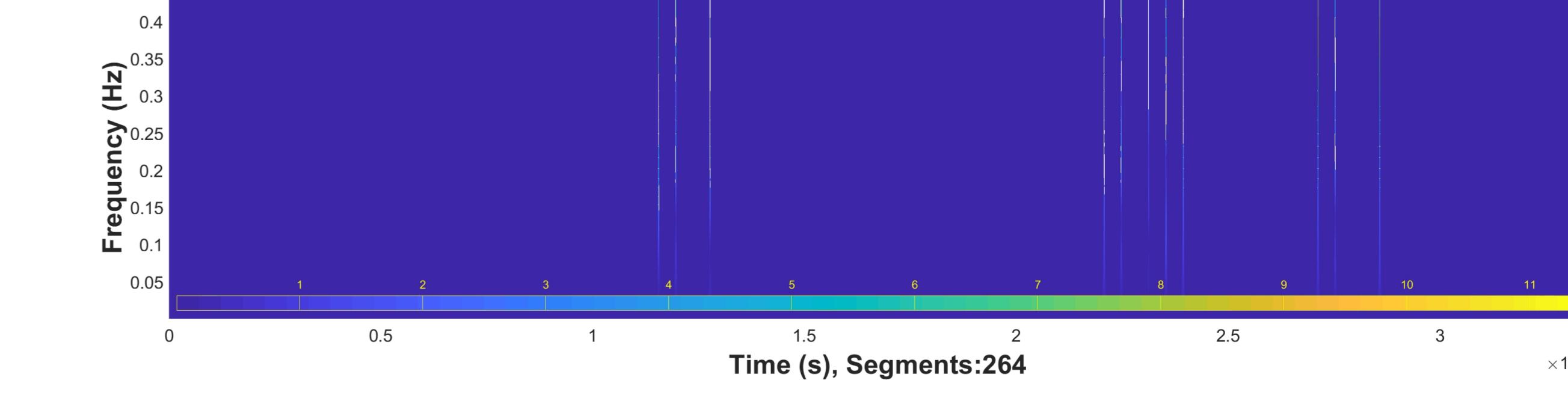
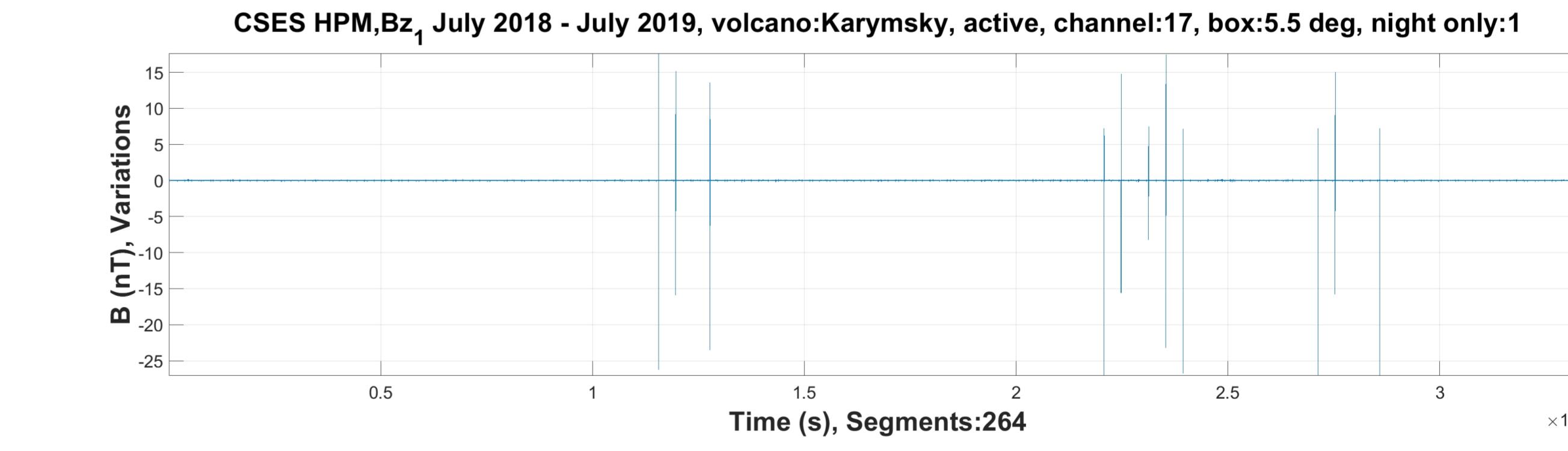
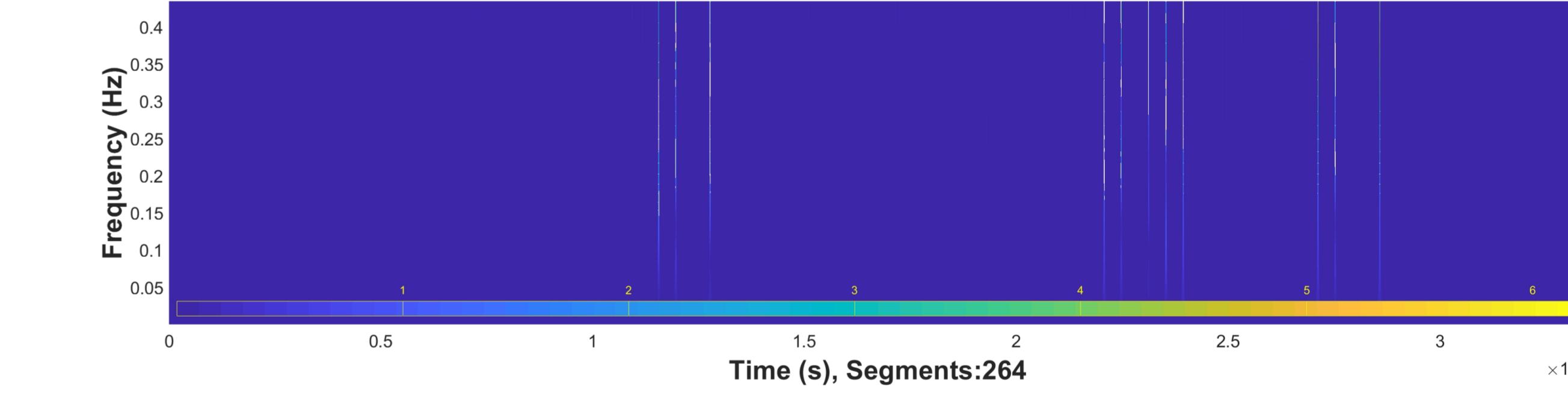
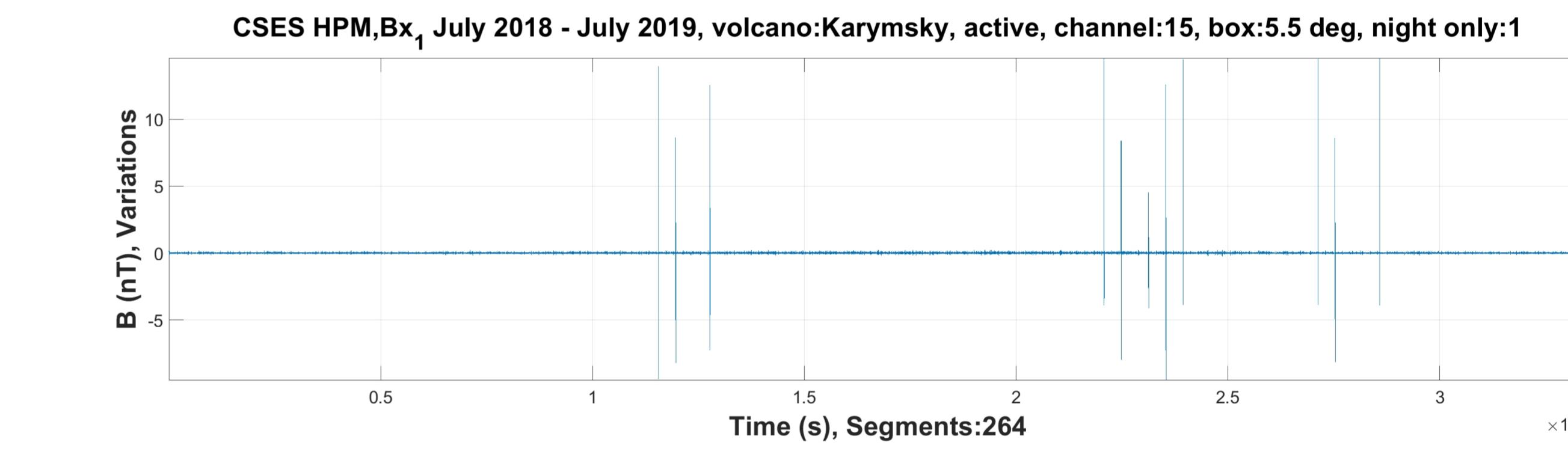
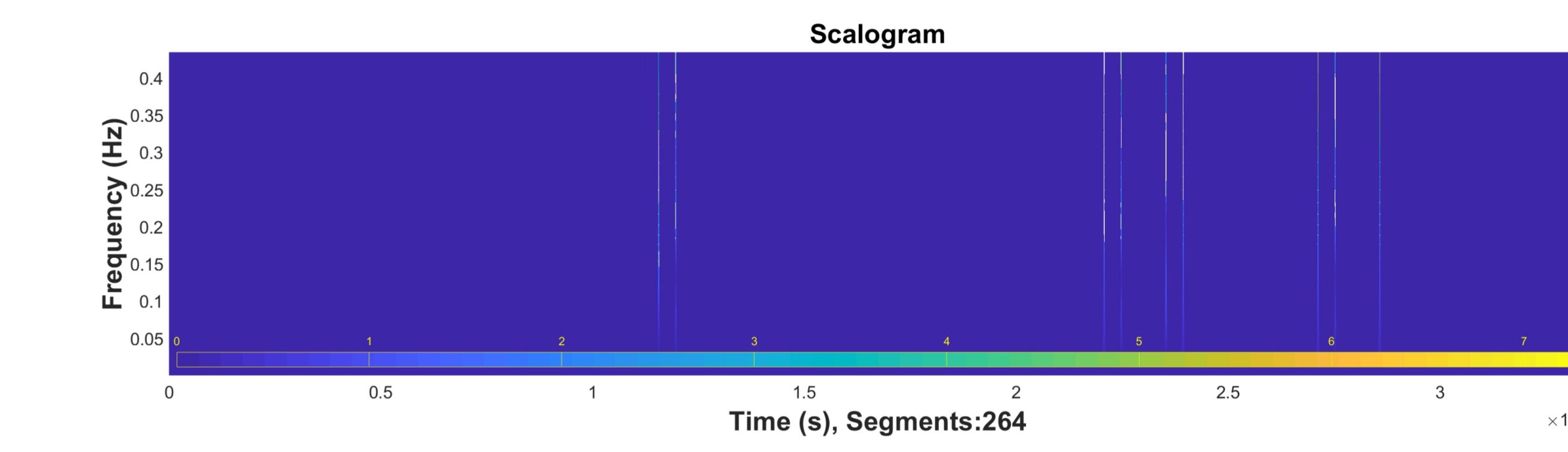
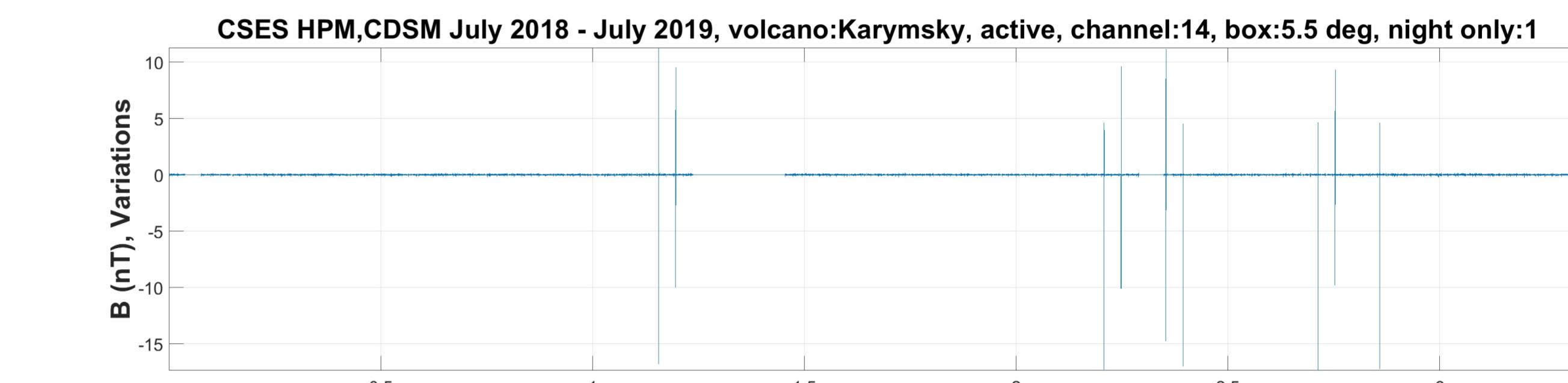
## UBINAS

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



## KARYMSKY

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz



# SATELLITE AND GROUND-BASED MAGNETIC FIELD OBSERVATIONS RELATED TO VOLCANIC ERUPTIONS

K. Schwingenschuh<sup>1</sup>, W. Magnes<sup>1</sup>, X. Shen<sup>2</sup>, J. Wang<sup>3</sup>, B. Cheng<sup>7</sup>, A. Pollinger<sup>1</sup>, Chr. Hagen<sup>1</sup>, R. Lammegger<sup>5</sup>, M. Ellmeier<sup>5</sup>, Christoph Schirniger<sup>1</sup>, H.-U. Eichelberger<sup>1</sup>, B. Mandl<sup>1</sup>, M. Y. Boudjada<sup>1</sup>, B. P. Besser<sup>1</sup>, A. A. Rozhnoi<sup>4</sup>, T. Zhang<sup>1</sup>, M. Delva<sup>1</sup>, I. Jernej<sup>1</sup>, Ö. Aydogar<sup>1</sup>, and R. Leonhardt<sup>6</sup>

<sup>1</sup>Space Research Institute, Austrian Academy of Sciences, Graz, Austria, <sup>2</sup>Institute of Earthquake Admin., Beijing, China, <sup>3</sup>Center for Space Science and Applied Research, CAS, Beijing, China, <sup>4</sup>Institute of the Earth Physics, RAS, Moscow, Russia, <sup>5</sup>Institute of Experimental Physics, Graz Univ. of Technology, Graz, Austria, <sup>6</sup>Central Institution for Meteorology and Geodynamics (ZAMG), Vienna, Austria, <sup>7</sup>National Space Science Center, CAS, Beijing, China

## SANTA MARÍA

Time series and Scalogram // Top: CDSM – Center: Bx – Bottom: Bz

