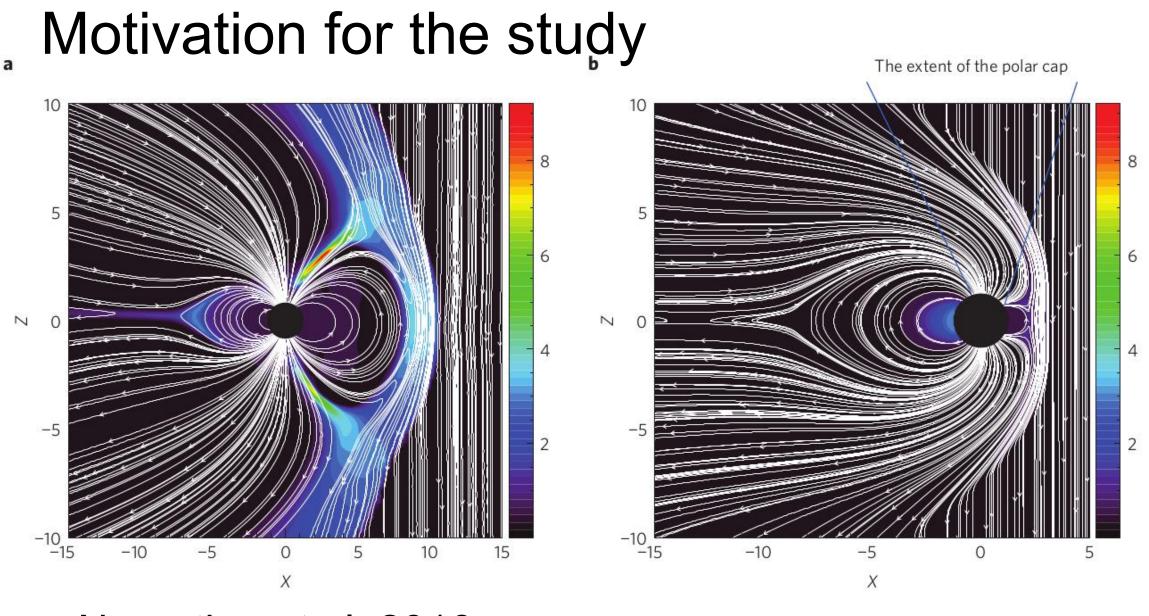
## Evolution of the polar outflow from the Earth's cusps

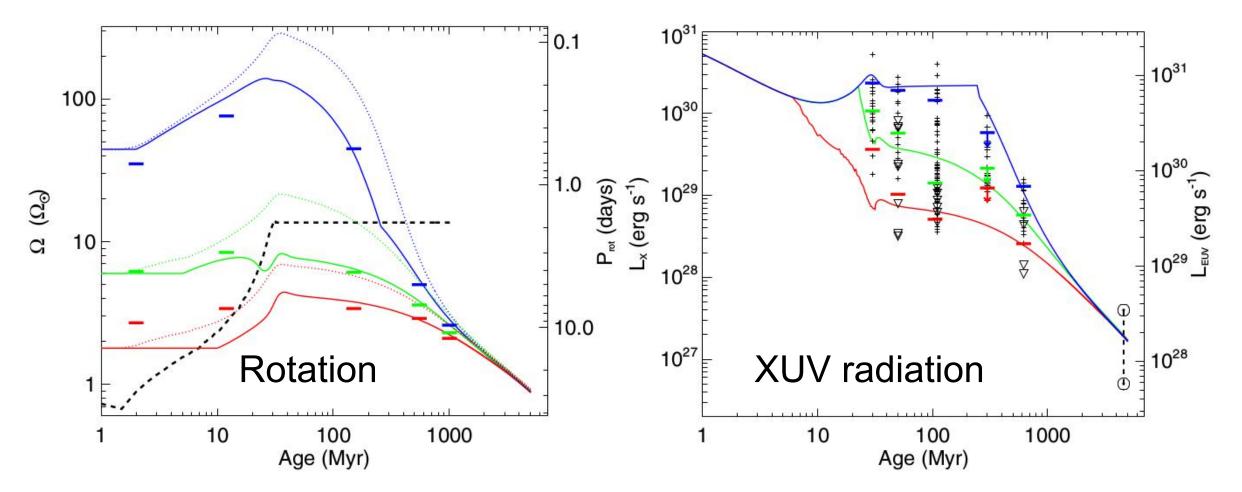
K.G. Kislyakova<sup>1</sup>, C.P. Johnstone<sup>1</sup>, M. Scherf<sup>2</sup>, H. Lammer<sup>2</sup>, M. Holmström<sup>3</sup>, M. Khodachenko<sup>2</sup>, M. Güdel<sup>1</sup>
<sup>1</sup>University of Vienna, Deparment of Astrophysics, Vienna, Austria
<sup>2</sup>Space Research Institute, Austrial Academy of Sciences, Graz, Austria
<sup>3</sup>Swedish Institute of Space Physics, Kiruna, Sweden



Airapetian et al. 2016



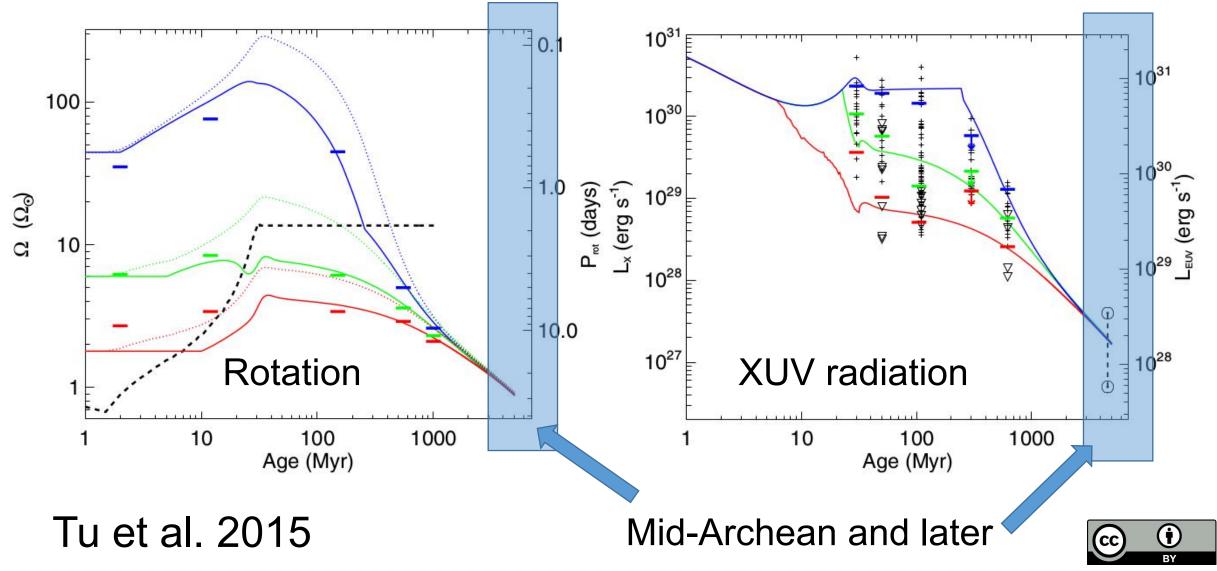
# Evolution of the Sun and its XUV (X-ray + EUV) output



Tu et al. 2015

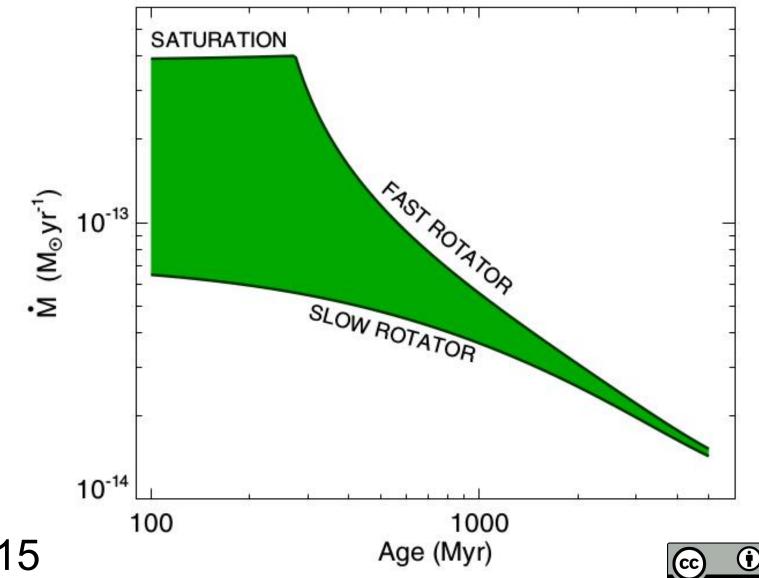


# Evolution of the Sun and its XUV (X-ray + EUV) output



### Evolution of the solar wind

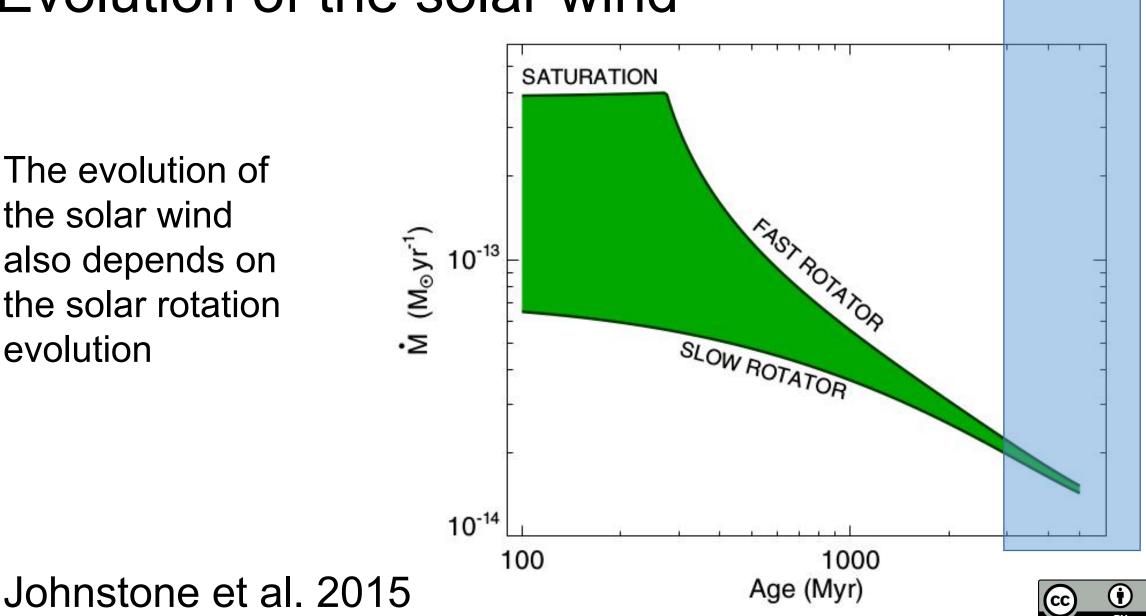
The evolution of the solar wind also depends on the solar rotation evolution



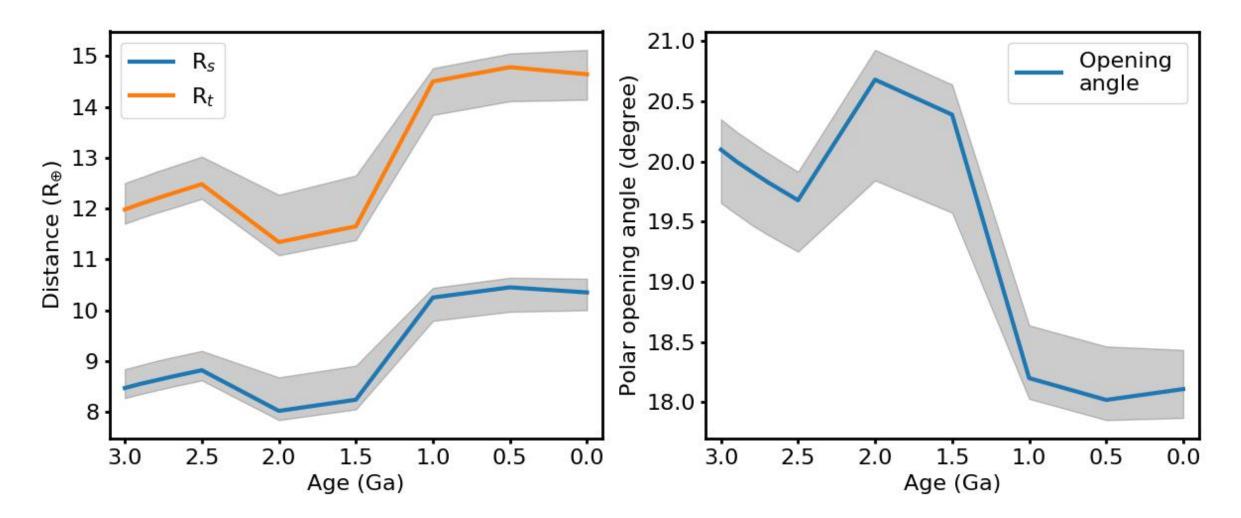
Johnstone et al. 2015

### Evolution of the solar wind

The evolution of the solar wind also depends on the solar rotation evolution



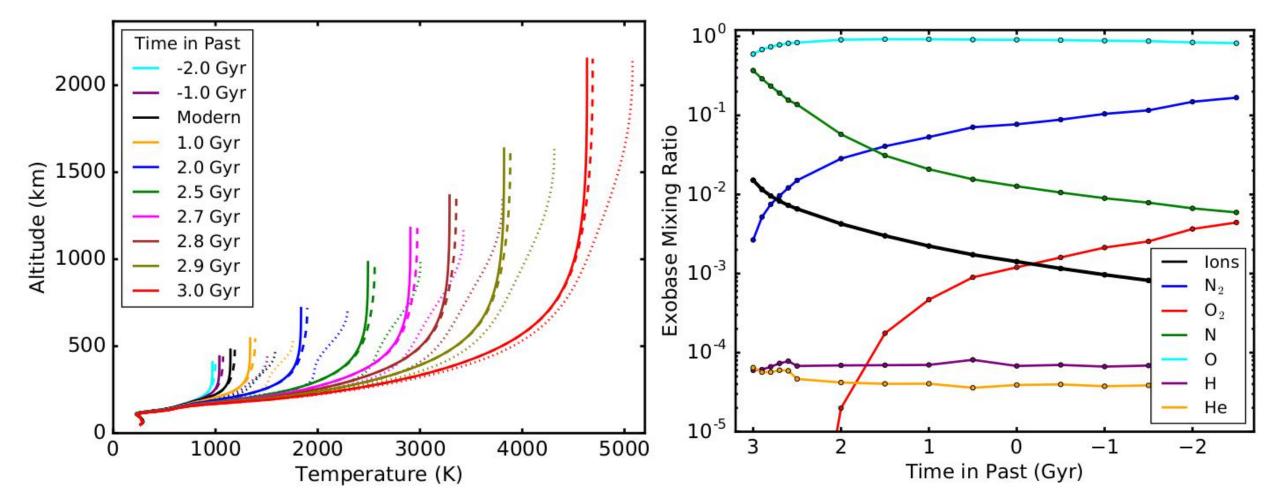
#### Evolution of the Earth's dynamo



Magnetic moment evolution following Biggin et al. 2015



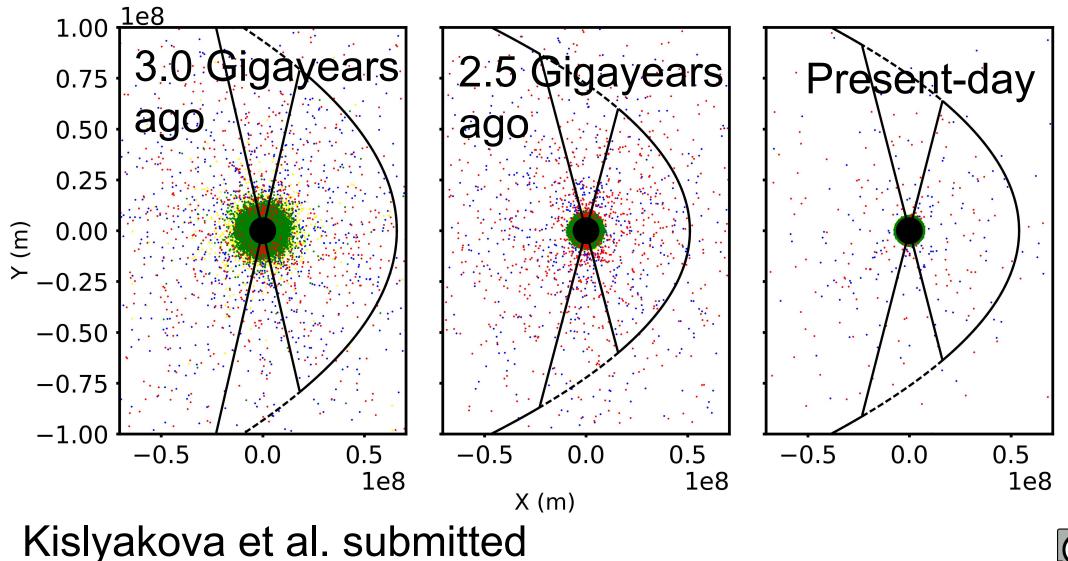
#### Evolution of the Earth's atmosphere



Johnstone et al. 2018



#### Upper atmosphere structure

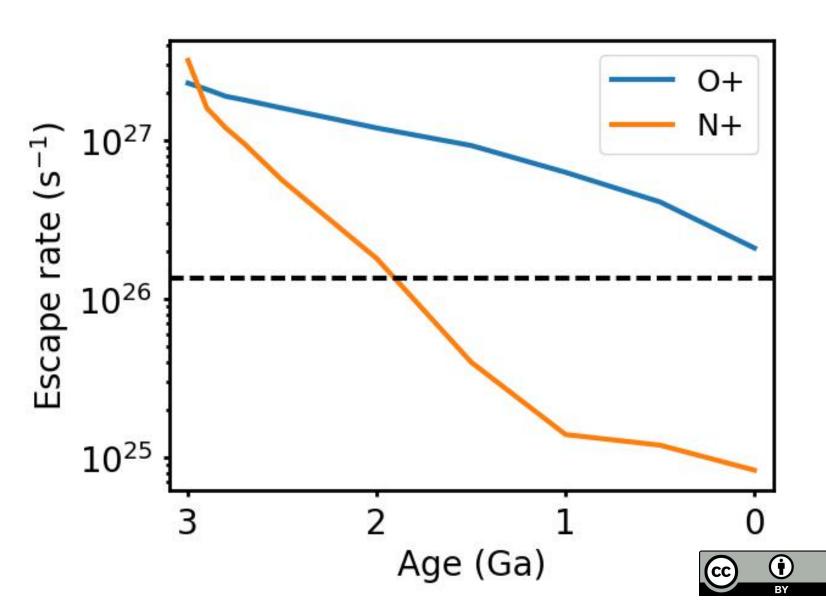




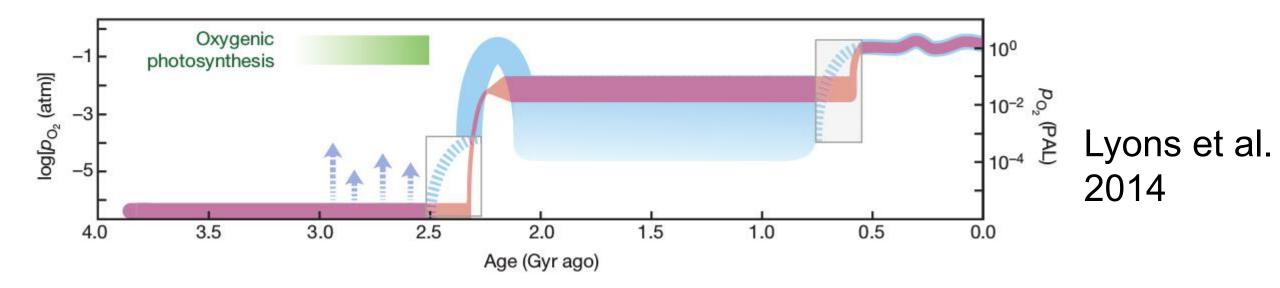
#### Evolution of the polar outflow

The outflow is assumed to be source limited

We assume the modern-day composition of the atmosphere



# The Great Oxydation Event (GOE) and the polar outflow



Simulation 2.5 Ga with modern composition: O<sup>+</sup> escape =  $1.6 \times 10^{27} \text{ s}^{-1}$ , N<sup>+</sup> escape =  $5.6 \times 10^{26} \text{ s}^{-1}$ 

Simulation 2.5 Ga with 1% oxygen: O<sup>+</sup> escape =  $2.0 \times 10^{26}$  s<sup>-1</sup>, N<sup>+</sup> escape =  $2.9 \times 10^{27}$  s<sup>-1</sup>

### Conclusions

- Polar outflow of the nitrogen ions three gigayears ago increases by two orders of magnitude compared to its present value.
- Polar outflow of oxygen ions from the Earth's open field line regions varies greatly depending on the oxygen mixing ratio
- Polar outflow escape is governed primarily by the evolution of the solar short-wavelength radiation and the atmosphere's composition
- $\rightarrow$  Kislyakova et al., JGR, under review

