

EGU22-8919

<https://doi.org/10.5194/egusphere-egu22-8919>

EGU General Assembly 2022

© Author(s) 2022. This work is distributed under the Creative Commons Attribution 4.0 License.



## Modelling the propagation of slow magnetoacoustic waves in a multi-stranded coronal loop

**Krishna Prasad Sayamanthula** and Tom Van Doorselaere

Centre for mathematical Plasma Astrophysics, Department of Mathematics, KU Leuven, Leuven, Belgium  
([krishna.prasad@kuleuven.be](mailto:krishna.prasad@kuleuven.be))

The cross-field thermal structure of a coronal loop is one of the critical parameters useful to distinguish the major heating theories. In a recent study we are able to isolate two thermal components of a coronal loop using observations of propagating slow magnetoacoustic waves in two different temperature channels. In order to properly interpret these observations and identify the actual cross-field thermal structure, we develop multiple three-dimensional magnetohydrodynamic models of coronal loop. In each of these models slow magnetoacoustic waves are driven by perturbing the plasma at one end and the corresponding multi-wavelength propagation characteristics are studied by applying forward modelling techniques. We compare the wave propagation properties between different models including a monolithic and a multi-stranded model with those from observations to draw some important inferences which will be discussed in this talk.