

The Ballistic Spreading of Debris Clouds over a Planetary Ring

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Abstract

In CASSINI images of Saturn's rings, taken close to the planet's equinox, elongated features were detected in the A and C rings [Tiscareno et al., 2009]. These features stand out bright in front of the darker ring and they are slightly inclined with respect to the orbital direction, with a cant angle consistent with the direction of Keplerian shear. They were interpreted [Tiscareno et al., 2009] as clouds of debris, spreading over the rings, released in a catastrophic impact of an external projectile on a ring particle. Such a bombardment of planetary rings has long been suspected [Cook and Franklin, 1970, Durisen et al., 1989, Cuzzi and Durisen, 1990, Cuzzi and Estrada, 1998, Cuk et al., 2000, Chambers et al., 2008], with consequences for origin, evolution, structure, and composition of the rings.

I develop a theoretical model for the evolution of a debris cloud, spreading from the location of the impact. For simplicity the details of the disruption are not addressed. I start with an isotropic model, assuming that a 30 km/s hypervelocity projectile, perhaps centimeter to decimeter sized, hits and destroys a ring particle (perhaps meter to tens of meters in size), leading to a cloud of debris spreading uniformly from the point of impact, with power law distributed speeds of ejection. This conceptually simple model allows us to study basic properties of the evolution of the cloud. Effects of anisotropy, arising from the direction of the projectile and the effect of momentum conservation, can be incorporated without principal difficulty in an improved model.

Other applications of the model could be the material vertically expelled from the ring plane in parts of the outer B ring edge (see e.g. Fig 5 (b) and (c) from [Spitale and Porco, 2010]), or the release of small particles in the perturbed regions of propellers [Sremčević et al., 2007].

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