

Solar wind interaction with Venus upper atmosphere: A review of VEX/ASPERA-4 findings and future prospect

Yoshifumi Futaana and Stas Barabash,
 Swedish Institute of Space Physics, Box 812, Kiruna SE-98128, Sweden. (futaana@irf.se)

Abstract

Venus has intrigued planetary scientists since the start of the space-era in 1960s. While Venus is nicknamed "Earth's twin", Venus has huge differences from the Earth. The space environment is significantly different, mainly because of the absence of an intrinsic magnetic field. The Venusian magnetosphere is formed by electric current in the ionosphere induced by the interaction with the solar wind (therefore, we call it an induced magnetosphere). Various signatures have been measured by Venera and Mariner series in the 60s and 70s, and basic signatures of the space environment near Venus were revealed. Pioneer Venus Orbiter was launched in 1978 providing a huge dataset of the Venus environment, and our knowledges about the ionosphere and the far-tail were enhanced. In 2006, ESA's first Venus probe, Venus Express (VEX), was inserted into orbit. VEX made unique measurements in the near-Venus tail for >8 years until VEX consumed all its fuels. VEX was equipped with plasma instrument package, Analyser of Space Plasma and Energetic Atoms (ASPERA-4). ASPERA-4 has provided a huge amount of discoveries about the near-Venus space environment. In this presentation, we will review the findings of ASPERA-4.

1. VEX/ASPERA-4 instrument

In early 2006, Venus Express (VEX) was inserted into a highly-elliptical Venusian orbit. VEX conducted science operation for 8.5 years until its fuel was fully consumed. Throughout the mission, the plasma instrument package, ASPERA-4, was operated, and these instruments provided a large amount of datasets.

The ASPERA-4 instrument comprises five sensors: three energetic neutral atom (ENA) sensors, an electron spectrometer (ELS) and an ion mass spectrometer (IMA) [1]. IMA provides ion measurements in the energy range 0.01-30 keV/q for the main ion components H^+ , H_2^+ , He^+ , O^+ , and heavier ions. ELS provides electron measurements in

the energy range 10 eV–15 keV. The FoV is narrow in elevation but with 360° coverage in the azimuth direction, so that the coverage is quite limited.

2. VEX/ASPERA-4 findings

Together with the magnetometer (MAG [2]), onboard, ASPERA-4 has provided a huge amount of findings about the near-Venus space environment. Figure 1 depicts an overview illustration of VEX findings [3].

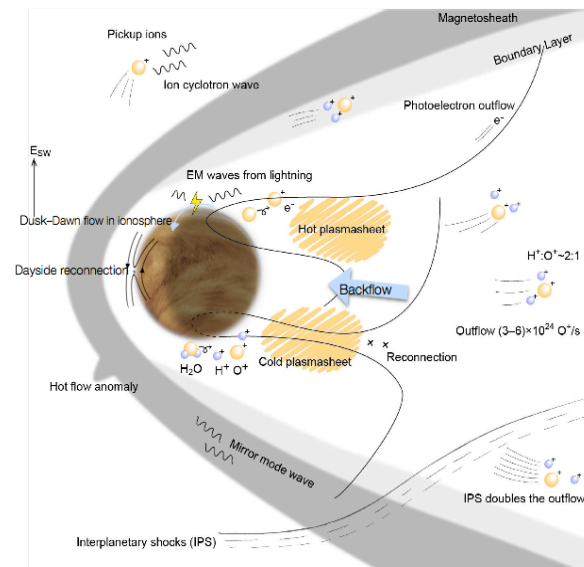


Figure 1: New finding in the near-Venus space environment by VEX/ASPERA-4 together with MAG. The figure is adopted from [3], where a comprehensive review of VEX findings is found. Additional findings after the publication have been added.

The key findings include

- The loss of ionospheric ions to space with their composition rate of H^+ and O^+ as 2 to 1 was identified [4], while the ratio is highly controlled by the solar activity [5].

- The outflowing O^+ flux is enhanced during the interplanetary shock passage [6]
- Significant fractions of outflowing H^+ and O^+ are returned back to Venus inside the plasmashet, particularly when the solar activity is high [5,7,8]
- The structure and dynamics of the Venusian induced magnetosphere respond to the prevailing solar wind conditions with various time scales.
- Magnetic reconnections are in operation inside the Venusian plasmashet [9]
- The plasma flow pattern of the polar ionosphere is different from that of the equatorial ionosphere [10]

We review what has been found by VEX in a context of the plasma environment. We further will formulate the key questions to increase our understandings about the Venus–solar wind interaction for future missions; and more in general, to understanding other Solar System bodies and exoplanets.

3. Future prospect

Venus flybys by ongoing and future missions are planned. BepiColombo (ESA and JAXA) launched in 2018 will flyby Venus in 2020 and 21, and scientific operations during the flyby are planned. Solar Orbiter, launch planned in 2020, also considers flybys with science operations. It is yet unclear for science operations during Venus flybys by JUICE (ESA), but it would be discussed. Parker Solar Probe (NASA) has conducted first Venus flyby in October 2018, and six further flybys are planned. The ASPERA-4 data will be excellent reference data for interpreting the obtained flyby data.

Acknowledgements

The effort of the ASPERA-4 development is highly international, involving 15 research groups from Europe, USA and Japan. The Swedish contribution to the ASPERA-4 experiment was supported by funding from the Swedish National Space Agency (SNSA).

References

- [1] Barabash et al., The Analyser of Space Plasmas and Energetic Atoms (ASPERA-4) for the Venus Express mission, *Planet. Space Sci.*, 55(12), 1772–1792, doi:10.1016/j.pss.2007.01.014, 2007.
- [2] Zhang et al., Magnetic field investigation of the Venus plasma environment: Expected new results from Venus Express, *Planet. Space Sci.*, 54(13–14), 1336–1343, doi:10.1016/j.pss.2006.04.018, 2006.
- [3] Futaana et al., Solar wind interaction and impact on the venus atmosphere, *Space Sci. Rev.*, 212(3), 1453–1509, doi:10.1007/s11214-017-0362-8, 2017.
- [4] Barabash et al., The loss of ions from Venus through the plasma wake, *Nature*, 450(7170), 650–653, doi:10.1038/nature06434, 2007.
- [5] Persson et al., H^+/O^+ escape rate ratio in the Venus magnetotail and its dependence on the solar cycle, *Geophysical Research Letters*, 0 (0), doi:10.1029/2018GL079454, 2018.
- [6] Edberg et al., Atmospheric erosion of Venus during stormy space weather, *J. Geophys. Res.*, 116(A9), doi:10.1029/2011JA016749, 2011.
- [7] Kollmann et al., Properties of planetward ion flows in venus’ magnetotail, *Icarus*, 274, 73–82, doi:10.1016/j.icarus.2016.02.053, 2016.
- [8] Masunaga et al., Effects of the solar wind and the solar EUV flux on O^+ escape rates from Venus, *Icarus*, 321, 379–387, doi: 10.1016/j.icarus.2018.11.017, 2019.
- [9] Zhang et al., Magnetic reconnection in the near venusian magnetotail, *Science*, 336(6081), 567–570, doi:10.1126/science.1217013, 2012.
- [10] Persson et al., Heavy ion flows in the upper ionosphere of the venusian north pole, *J. Geophys. Res.*, in Press.