

Coriolis Drift in Saturn's Inner Magnetosphere

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Abstract

The theory of adiabatic particle motion in electromagnetic fields was developed well before the discovery of the terrestrial and Jovian radiation belts [1]. However, it has been intensively detailed and widely used to study the motion of trapped particles in magnetospheric systems since the 1960s [2,3,4]. In rapidly rotating magnetospheres, the diffusion equations describing the motion in the guiding center approximation were first revised in the late 1960s, then in the 1970s and 1980s [5,6,7,8]. It was important to take into account the perturbations caused by the *Coriolis* and *Centrifugal* forces on the trapped particle motion and correct the guiding center equations.

In this paper, we propose to examine the role of the *Coriolis* drift on the outward plasma transport in Saturn's magnetosphere. We will first present our theoretical approach and the main results obtained with our diffusion theory model when considering the interactions with neutrals and icy moons environment. Then, we will describe the implementation of inward and outward transports into our transport code. Simulations of outward transport will provide us the opportunity to discuss the origin of outward plasma flows, as well as how they balance the inward transport. We will discuss the results of comparisons between model and *Cassini* CAPS-ELS data sets in order to refine our understanding of the physical processes that drive the electron plasma populations within Saturn's magnetosphere.

References

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