

Does the innermost occurrence distribution measure tidal dissipation, reveal a flow of giant planets, or both?

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Abstract

The occurrence distribution of the shortest period giant exoplanets as found by Kepler show a drop-off that is a remarkable match to the drop-off expected by taking migration due to tides in the star [9], [6]. We present a comparison that can show the level of tidal dissipation (friction) as a function of the distribution of the ages of the star and planet system, with known dependencies on basic star and planet parameters. Use of this relation enables constraints to be put on the value of the tidal dissipation, constraints that will be improved as the distribution of the ages are determined. For the giant planets, this leads to an unexpectedly low value of tidal dissipation. This over-abundance of short period giant planets may be due to a continuing resupply of longer period giant planets migrating into a shorter period pileup, disrupting the presence of smaller planets along the way. Perhaps the occurrence distribution of close Neptune sized planets will better measure the tidal friction, while the distribution of Jupiter sized planets reveals that giant planets are more likely to complete a gradual migration into the star.

1. Introducing Infall versus Q'_*

The "hot Jupiter" planets, those giant planets with the shortest periods, have been shown to mostly be tidally migrating to their destruction by merger with their star, [4], [3]. Kepler planet (candidate) occurrence distributions from [2] show a drop-off in occurrence at short periods, which have a power law index of 13/3, which [9] show is consistent with tidal migration due to tides on the star, using the equations of [3]. The homogeneous Kepler statistics are sufficient to allow us to compare calculated drop-offs as a function of tidal dissipation (friction) Q'_* to the actual innermost occurrence distributions. We show in figure 1 and figure 2 that these distributions give different apparent values of Q'_* for large and medium planets, which could be explained by proposing a larger inflow of larger planets.

We find that the tidal migration increases so rapidly

with smaller semi-major axis that the initial distribution matters little, so for the initial distribution we simply extend inward the power law that [2] find for the occurrence distribution beyond the drop-off. Though it has long been expected that the process of planet formation would leave a drop off of planets at the shortest period, we find that the ongoing migration erases the inner-most distribution left by planet formation.

2. Further Work

We are currently preparing models where the shift in the drop-off point can be interpreted in relation of tidal dissipation factor Q'_* versus realistic stellar age distribution for realistic star and planet distributions. Here we use single representative ages and star parameters for the planet masses.

3. Summary and Conclusions

We compare what the tidal dissipation Q'_* in the star appears to be when looking at Jupiter and Neptune mass planets. We find a discrepancy that could be better explained by a flow of planets such as proposed by [7]. This flow could be producing the "pileup" of hot Jupiters. If there is a higher rate of ongoing inward migration of Jupiter mass than of Neptune mass planets, this could explain why we find "too many" hot Jupiters shortly before they merge with the star, without requiring an especially low amount of tidal friction for giant planets. This discrepancy in the apparent value of Q'_* is the result of the pileup of giant planets. Compared to Neptune mass planets, there is a higher number of short period Jupiter mass planets but a lower number of longer period Jupiter mass planets. Kepler is now finding that the general trend of more smaller radii planets also holds in the short period range as more smaller planets are detected, further emphasizing the anomalous nature of the large number of Jupiter radii short period planets. That the number of short period massive planets is anomalously large is made more clear by how the pattern of more planets at smaller

radii in now being observed to continue into the super-earth size range at all measurable periods. Another sign of ongoing large planet migration is how the short period large planets are anti-correlated with the presence of other planets further out, but there is no such correlation for smaller short period planets. This suggests that inwardly scattered large planets make their way to the star, disrupting the orbits of smaller planets along the way, but that inwardly scattered midsize planets do not make it to the star. The resulting transient events created by the merger of planets with stars, as suggested by [8], [5], and [1], will provide an important means of quantifying planet migration.

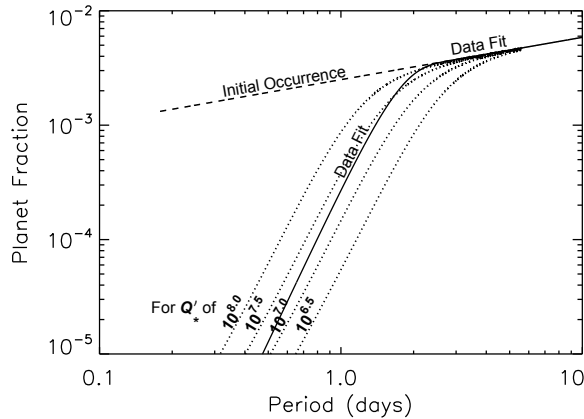


Figure 1: Comparison of a range of four calculated occurrence distribution drop-offs (dotted lines) with the fit to Kepler planet candidate data (solid line) for large planets, defined as 8 to 32 earth radii or one Jupiter mass. found by [2] giant radii or mass planets. Calculated infall for four values of stellar tidal dissipation $Q'_* = (10^{8.0}, 10^{7.5}, 10^{7.0}, 10^{6.5})$, is based on [3] using solar mass stars taken to be at an age of 4.5 Giga-years, and using the power law of [2] without the inward dropoff as an initial planet fraction (dashed line).

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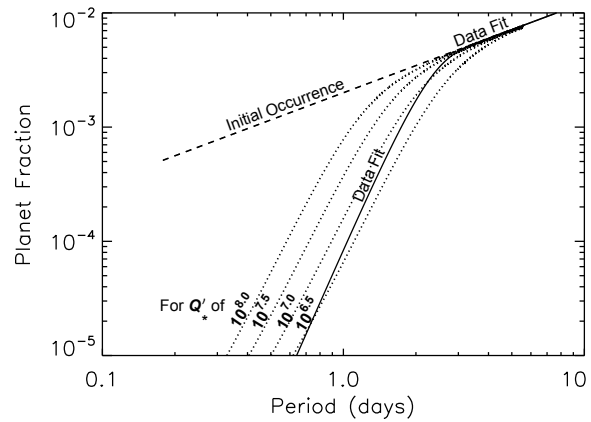


Figure 2: Similar comparison as figure 1 for “medium planets”. The fit to Kepler data for 4 to 8 earth radii is shown with calculated infall for 15 earth mass planets and solar mass stars, at an age of 4.5 Giga-years.

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