

# SMART-1 & recent missions: results from combining data

B.H. Foing , (1) ESA/ ESTEC, Postbus 299 , 2200 AG Noordwijk, NL ([Bernard.Foing@esa.int](mailto:Bernard.Foing@esa.int))

## Abstract

We highlight some results from combined data analysis using SMART-1 archive with other recent lunar missions. We discuss in particular impact craters, volcanic, photometry and studies of ILEWG sites.. SMART-1 demonstrated the use of Solar Electric Propulsion for deep space, tested new technologies for spacecraft and instruments miniaturisation, and provided an opportunity for science [1-12] until impact on 3 September 2006. To date 75 refereed papers and more than 325 conference or technical papers have been published based on SMART-1 (see ADS & SMART-1 website [sci.esa.int/smart-1](http://sci.esa.int/smart-1) or [www.esa.int/smart-1](http://www.esa.int/smart-1)). The SMART-1 data are accessible on the ESA Planetary Science Archive PSA [13] <http://www.rssd.esa.int/psa>

## 1. Recent SMART-1 archive results

These include: multi-angular photometry of Mare and specific regions to diagnose the regolith roughness and constrain models of light reflection and scattering [14] and comparison to laboratory granular photometric studies [15]; the lunar North and South polar illumination was mapped and monitored over the entire year, permitting to identify “SMART-1 peaks of quasi-eternal light” and study their topography [16, 17]; SMART-1 was used for radio occultation experiments [18], positioning reduction of SMART-1, Chang'E1 and 2 VLBI tracking data [19]; the X-Ray Solar Monitor studied the Sun as a flare star with GOES and RHESSI [20,21]; SMART-1 SIR and Chandrayaan-1 HySI were used to study the composition of the central peak of craters [22]; the SMART-1 impact observed from Earth was modeled using laboratory experiments [23].. The South Pole Aitken Basin was mapped and studied combining data from Clementine, SMART-1, and other missions [24]. The SMART-1 archive observations have been used

to support Kaguya, Chandrayaan-1, Chang'E 1, the US LRO and to characterise potential sites for lunar science and exploration.

## 2. Acknowledgements

We thank SMART-1 team: G.Racca and S1 Project team, O.Camino and S1 spacecraft operations team, D.Frew (STOC), D.Koschny, B.Grieger, M.Almeida, J.Volp, D.Heather, H.Metselaar, S.Martinez and STWT members. including J.-L.Josset, S.Beauvivre, M.Grande, J.Huovelin, H.U.Keller, U.Mall, A.Nathues, A.Malkki, G.Noci, P.McMannamon, Z.Sodnik, B.Kellett, P.Pinet, S.Chevrel, P.Cerroni, M.C.de Sanctis, M.A.Barucci, S.Erard, D.Despan, K.Muinonen, V.Shevchenko, Y.Shkuratov, C.Veillet, P.Ehrenfreund, M.Ellouzi, S.Peters, A.Borst, F.Bexkens, G.Davies, W.van Westrenen, E.Martellato for their contribution.

## 3. References

- [1] Foing etal (2001) EMP85-523;[2] Racca et al (2002) EMP85-379; [3] Racca et al. (2002) PSS50-1323; [4] Grande et al. (2003) PSS51-427; [5] Dunkin et al. (2003) PSS51-435; [6] Huovelin et al. (2002) PSS50-1345; [7] Shkuratov et al (2003) JGRE108-E4-1; [8] Foing et al (2003) ASR31-2323;[9] Grande et al (2007) PSS55-494; [10] Pinet et al (2005) PSS53-1309; [11] Josset etal (2006) ASR37-14; [12] Foing et al (2006) ASR37-6; [13] Heather et al, EPSC-DPS 2011-873 [14] Muinonen et al (2011) A&A531-150; [15] Souchon et al EPSC-DPS2011-928, [16] Grieger (2010) cosp38-417; [17] Bussey et al (2011) LPICO-1621-5; [18] Pluchino et al MSAItS 16-152; [19] Qiao (2011) AcASn 52, 539; [20] Vaananen et al (2009) SolarPhys260-479; [21] Alha et al (2012) NIMPA 664, 358; [22] Bhattacharya et al EPSC-DPS2011-1842; [23] Burchell et al (2010) Icarus207-28[24] Borst, Foing et al (2012) PSS 68, 76; SMART-1 site: [www.esa.int/smart-1](http://www.esa.int/smart-1).,

