

Saturn's Other Ring Current

F. J. Crary (1)
(1) Laboratory for Atmospheric and Space Physics,
University of Colorado, United States
(frank.crary@lasp.colorado.edu)

Abstract

Saturn's main rings orbit the planet within an atmosphere and ionosphere of water, oxygen and hydrogen, produced by meteoritic impacts on and ultraviolet photodesorption of the ring particles [Johnson et al., 2006; Luhmann et al., 2006; Tseng et al., 2010]. The neutral atmosphere itself has only been tentatively detected through ultraviolet fluorescents of OH [Hall et al., 1996] while the ionosphere was observed in situ by the Cassini spacecraft shortly after orbital insertion [Coates et al., 2005; Tokar et al. 2005, Waite et al. 2005].

Although the plasma flow velocity of this ionosphere is not well-constrained, but the close association with the rings suggests that its speed would be coupled to the keplarian velocity of the rings themselves. As a result, the motion of the plasma through Saturn's magnetic field would produce an induced voltage, oriented away from the planet outside synchronous orbit and towards the planet inside synchronous orbit. Such a potential could result in currents flowing across the ring plane and closing along magnetic field lines and through Saturn's ionosphere at latitudes between 36° and 48°. Cassini observations of whistler-mode plasma wave emissions [Xin et al., 2006] centered on synchronous orbit (1.76 Rs, mapping to 41° latitude) have been interpreted as a product of field-aligned electron beams associated with such a current.

This presentation will investigate the magnitude of these currents and the resulting Joule heating of the ionosphere. An important constraint is that no auroral ultraviolet emissions have been observed at the relevant latitudes. In contrast, Joule heating could affect infrared emissions from H₃⁺. Variations in H₃⁺ emission associated with Saturn's rings have been reported by O'Donoghue et al., 2013, and interpreted

as a result of ring "rain", i.e. precipitating water group species from the rings which alter ionospheric chemistry and H₃⁺ densities. As noted by O'Donoghue et al., this interpretation may be complicated by Joule heating by the ring-related currents.

References

- Coates, A.J., McAndrews, H.J., Rymer, A.M., Young, D.T., Crary, F.J., Maurice, S., Johnson, R.E., 2005. Plasma electrons above Saturn's main rings: CAPS observations. *Geophys. Res. Lett.* 32, L14S09.
- Johnson, R.E., Luhmann, J.G., Tokar, R.L., Bouhram, M., Berthelier, J.J., Sittler, E.C., Cooper, J.F., Hill, T.W., Crary, F.J., Young, D.T., 2006. Production, ionization and redistribution of Saturn's O₂ ring atmosphere. *Icarus* 180, 393–402.
- Luhmann, J.G., Johnson, R.E., Tokar, R.L., Ledvina, S.A., Cravens, T.E., 2006. A model of the ionosphere of Saturn's rings and its implications. *Icarus* 181, 465–474.
- O'Donoghue, J., T. S. Stallard, H. Melin, G. H. Jones, S. W. H. Cowley, S. Miller, K. H. Baines and J. S. D. Blake, 2013. The domination of Saturn's low-latitude ionosphere by ring 'rain', *Nature* 496, 193-195.
- Tokar, R.L., and 12 colleagues, 2005. Cassini observations of the thermal plasma in the vicinity of Saturn's main ring and the F and G rings. *Geophys. Res. Lett.* 32, L14S04.
- Tseng, W.-L., W.-H. Ip, R.E. Johnson, T.A. Cassidy, M.K. Elrod, 2010. The structure and time variability of the ring atmosphere and ionosphere. *Icarus* 206, 382-389.
- Waite Jr., J.H., and 10 colleagues, 2005. Oxygen ions observed near Saturn's A ring. *Science* 307, 1260–1262.
- Xin, L., D. A. Gurnett, O. Santolik, W. S. Kurth, and G. B. Hospodarsky, 2006. Whistler-mode auroral hiss emissions observed near Saturn's B ring. *J. Geophys. Res.* 111, A06214, doi:10.1029/2005JA011432