

# Hydrodynamic Simulations of Asymmetric Propeller Structures in the Saturnian Ring System

Michael Seiler, Martin Seiß, Holger Hoffmann and Frank Spahn  
Institute of Physics and Astronomy, University of Potsdam, Germany (miseiler@uni-potsdam.de)

## Abstract

Small sub-kilometer sized objects (called moonlets) embedded in the dense rings of Saturn cause density structures due to their gravitational interaction with the surrounding ring material which resemble a propeller, giving the structure its name in this way. The prediction of the existence of propeller structures within Saturn's rings [1, 2] led to their detection [3, 4, 5]. The recurrent observation of the large outer A ring propellers in Cassini ISS images allowed the reconstruction of their orbits. This analysis yielded that the observed propellers are deviating considerably from their expected Keplerian orbit [6]. The offset motion of the largest propeller structure called *Blériot* can be astonishingly well composed by a three-mode harmonic fit [7]. The origin of this offset motion still is on debate. Two hypotheses are in discussion: Whether the moonlet is perturbed resonantly by one or several of the large outer moons [7] or whether it is stochastically migrating [8, 9, 10]. Independent which of these hypotheses is finally correct, the changed orbital motion of the moonlet is effecting the shape of its created propeller structure. Here, we perform hydrodynamic simulations to study the changes of the propeller structure due to a disk-embedded moonlet which is librating around its mean orbital position. We present results showing how the induced propeller structure changes because of the libration of the moonlet and if these changes might be visible in Cassini images. Further, we apply our results to the propeller structures *Blériot* and *Santos Dumont*.

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