

# Preliminary retrievals of CO<sub>2</sub> column densities using the first data of TGO/NOMAD

**Arianna Piccialli (1)**, Ann Carine Vandaele (1), Ian R. Thomas (1), Séverine Robert (1), Shohei Aoki (1,2,3), Loïc Trompet (1), Justin T. Erwin (1), Valérie Wilquet (1), Arnaud Mahieux (1), Frank Daerden (1), Lori Neary (1), Sébastien Viscardy (1), Bojan Ristic (1), Ozgur Karatekin (4), Michael D. Smith (5), Giuseppe Sindoni (6), Fabrizio Oliva (6), Sophie Bauduin (7), Paulina Wolkenberg (6), José Juan López-Moreno (8), Giancarlo Bellucci (6), Manish R. Patel (9) and the NOMAD team

(1) Royal Belgian Institute for Space Aeronomy, Belgium, (2) Fonds National de la Recherche Scientifique, Belgium, (3) Tohoku University, Japan, (4) Royal Observatory of Belgium, (5) NASA Goddard Space Flight Center, USA, (6) INAF, Istituto di Astrofisica e Planetologia Spaziali, Italy, (7) Université libre de Bruxelles, Belgium, (8) Instituto de Astrofísica de Andalucía, Spain, (9) Open University, UK. (email: [arianna.piccialli@aeronomie.be](mailto:arianna.piccialli@aeronomie.be), Twitter: [@apic79](#))

## Abstract

The NOMAD (Nadir and Occultation for MArs Discovery) – operating on board the ExoMars 2016 Trace Gas Orbiter mission – started to acquire the first scientific measurements on 21 April 2018.

Here, we will present first retrievals of CO<sub>2</sub> column density and surface pressure measured by the NOMAD LNO channel in the nadir mode.

## 1. The NOMAD LNO channel

NOMAD is a spectrometer operating in 3 channels: 1) a solar occultation channel (SO) operating in the infrared (2.3-4.3 μm); 2) a second infrared channel LNO (2.3-3.8 μm) capable of doing nadir, as well as solar occultation and limb; and 3) an ultraviolet/visible channel UVIS (200-650 nm) that can work in the three observation modes [1,2]. The LNO infrared channel has a high spectral resolution ( $\lambda/d\lambda \sim 10,000$ ) provided by an echelle grating in combination with an Acousto-Optical Tunable Filter which allows the selection of spectral windows (diffraction orders).

In nadir mode, LNO has an instantaneous footprint of 0.5x17 km<sup>2</sup>, therefore it is very well suited to measure the horizontal and local time distribution of total column density of several species such as CO<sub>2</sub>, CO, H<sub>2</sub>O, and isotopic ratio.

## 2. CO<sub>2</sub> column densities and surface pressure

Since carbon dioxide constitutes 95% of the Martian atmosphere, CO<sub>2</sub> column densities can be reasonably used as a proxy for surface pressure [3].

We plan to analyze the data measured in diffraction orders 149, 167, 168, 169 for the CO<sub>2</sub> retrievals (Table 1). Orders 167-169 are particularly interesting since they contain water vapor as well, which may be retrieved simultaneously with CO<sub>2</sub> (See [4]). We will use the line-by-line radiative transfer code ASIMUT-ALVL developed at IASB-BIRA [5] to retrieve CO<sub>2</sub> column density and surface pressure. These retrieved quantities will then be compared with values predicted by the 3D GEM-Mars v4 Global Circulation Model (GCM) [6].

**Table 1:** Wavenumber ranges of the diffraction orders for the CO<sub>2</sub> analysis by the NOMAD LNO channel.

Diffraction order	LNO wavenumber limits [cm <sup>-1</sup> ]
149	3349.24-3375.99
167	3753.84-3783.83
168	3776.32-3806.49
169	3798.80-3829.15

## Acknowledgements

The NOMAD experiment is led by the Royal Belgian Institute for Space Aeronomy (BIRA-IASB), assisted by Co-PI teams from Spain (IAA-CSIC), Italy (INAF-IAPS), and the United Kingdom (Open University). This project acknowledges funding by the Belgian Science Policy Office (BELSPO), with the financial and contractual coordination by the ESA Prodex Office (PEA 4000103401, 4000121493), by MICIIN through Plan Nacional (AYA2009-08190 and AYA2012-39691), as well as by UK Space Agency through grant ST/P000886/1 and Italian Space Agency through grant 2018-2-HH.0. The research was performed as part of the “Excellence of Science” project “Evolution and Tracers of Habitability on Mars and the Earth” (FNRS 30442502. This research was supported by the FNRS CRAMIC project under grant number T.0171.16 and by the BrainBe SCOOP project.

## References

- [1] Vandaele, A.C., Neefs, E., Drummond, R. et al.: Science objectives and performances of NOMAD, a spectrometer suite for the Exo-Mars TGO mission, *Planetary and Space Science*, Vol. 119, pp. 233–249, 2015.
- [2] Neefs, E., Vandaele, A.C., Drummond, R. et al.: NOMAD spectrometer on the ExoMars trace gas orbiter mission: part 1 – design, manufacturing and testing of the infrared channels, *Applied Optics*, Vol. 54 (28), pp. 8494-8520, 2015.
- [3] Forget, F. Spiga, A.; Dolla, B.; Vinatier, S.; Melchiorri, R.; Drossart, P.; Gendrin, A.; Bibring, J-P; Langevin, Y.; Gondet, B. Remote sensing of surface pressure on Mars with the Mars Express/OMEGA spectrometer: 1. Retrieval method. *JGR*, 2007, 112, doi: 10.1029/2006JE002871
- [4] Aoki, S., EPSC 2018.
- [5] Vandaele, A.C., M. De Mazière, R. Drummond, A. Mahieux, E. Neefs, V. Wilquet, O. Korablev, A. Fedorova, D. Belyaev, F. Montmessin, and J.L. Bertaux, Composition of the Venus mesosphere measured by SOIR on board Venus Express. *JGR*, 2008. 113 doi:10.1029/2008JE003140.
- [6] Neary, L., and F. Daerden (2018), The GEM-Mars general circulation model for Mars: Description and evaluation, *Icarus*, 300, 458–476, doi:10.1016/j.icarus.2017.09.028

## The NOMAD Team

*Scientific team:* Vandaele, Ann Carine; Lopez Moreno, Jose Juan; Bellucci, Giancarlo; Patel, Manish; Allen, Mark; Alonso-Rodrigo, Gustavo; Altieri, Francesca; Aoki, Shohei; Bauduin, Sophie; Bolsée, David; F. Giacomo Carrozzo, Clancy, Todd; Cloutis, Edward; Daerden, Frank; D'Aversa, Emiliano; Depiesse, Cédric; Erwin, Justin; Fedorova, Anna; Formisano, Vittorio; Funke, Bernd; Fussen, Didier; Garcia-Comas, Maia; Geminale, Anna; Gérard, Jean-Claude; Gillotay, Didier; Giuranna, Marco; Gonzalez-Galindo, Francisco; Hewson, Will; Homes, James; Ignatiev, Nicolai; Kaminski, Jacek; Karatekin, Ozgur; Kasaba, Yasumasa; Lanciano, Orietta; Lefèvre, Franck; Lewis, Stephen; López-Puertas, Manuel; López-Valverde, Miguel; Mahieux, Arnaud; Mason, Jon; Mc Connell, Jack; Mumma, Mike; Nakagawa, Hiromu; Neary, Lori; Neefs, Eddy; Novak, R.; Oliva, Fabrizio; Piccialli, Arianna; Renotte, Etienne; Robert, Severine; Sindoni, Giuseppe; Smith, Mike; Stiepen, Arnaud; Thomas, Ian; Trokhimovskiy, Alexander; Vander Auwera, Jean; Villanueva, Geronimo; Viscardi, Sébastien; Whiteway, Jim; Willame, Yannick; Wilquet, Valérie; Wolff, Michael; Wolkenberg, Paulina – *Tech team:* Alonso-Rodrigo, Gustavo; Aparicio del Moral, Beatriz; Barzin, Pascal; Beeckman, Bram; BenMoussa, Ali; Berkenbosch, Sophie; Biondi, David; Bonnewijn, Sabrina; Candini, Gian Paolo; Clairquin, Roland; Cubas, Javier; Giordanengo, Boris; Gissot, Samuel; Gomez, Alejandro; Hathi, Brijen; Jeronimo Zafra, Jose; Leese, Mark; Maes, Jeroen; Mazy, Emmanuel; Mazzoli, Alexandra; Meseguer, Jose; Morales, Rafael; Orban, Anne; Pastor-Morales, M; Perez-grande, Isabel; Queirolo, Claudio; Ristic, Bojan; Rodriguez Gomez, Julio; Saggin, Bortolino; Samain, Valérie; Sanz Andres, Angel; Sanz, Rosario; Simar, Juan-Felipe; Thibert, Tanguy